# Determination of mass hierarchy with reactor neutrino experiment

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arXiv: 1210.8141

 $\sin^2 2 heta_{13} \sim 0.1$  is observed.



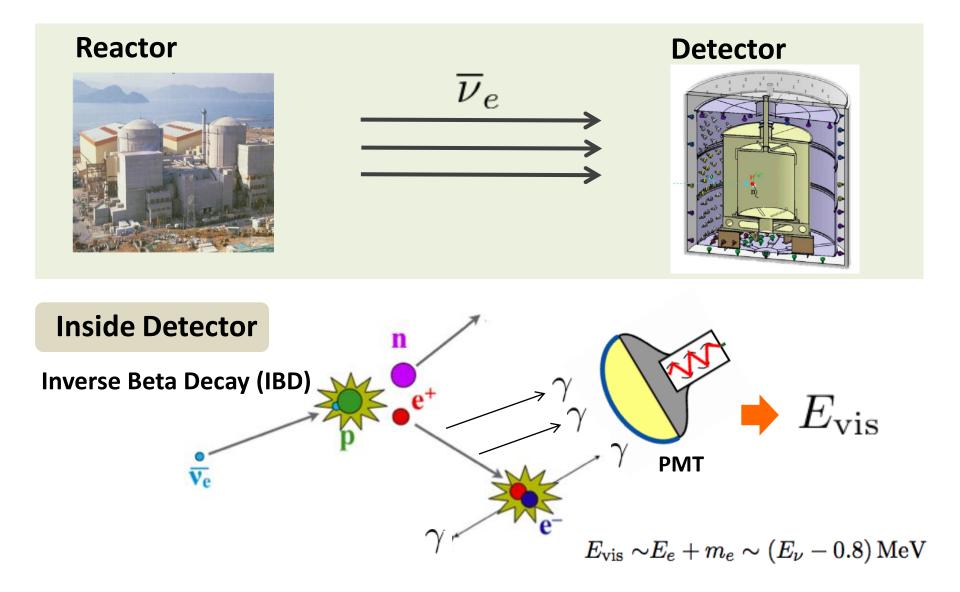
### *Reactor Experiments* can determine the *Neutrino Mass Hierarchy* .



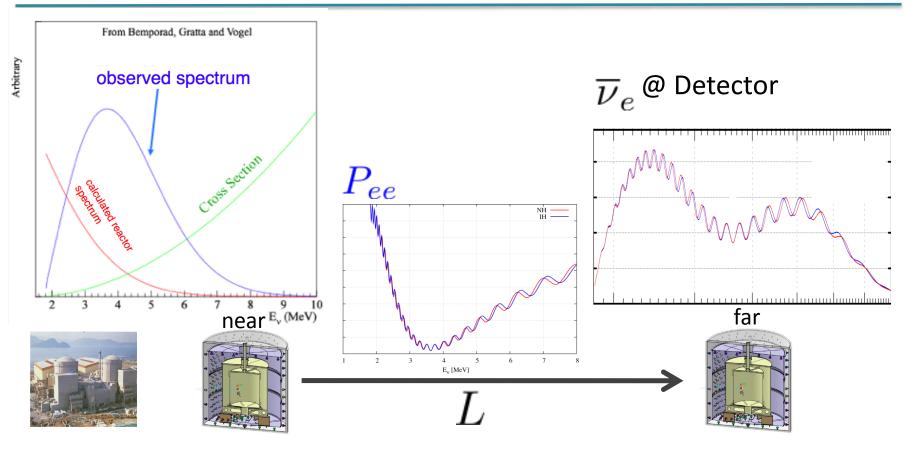
#### The sensitivity is studied

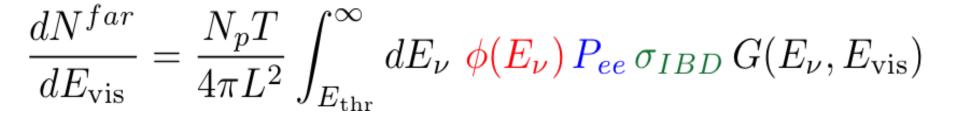
of future Medium Baseline Reactor Experiments for determining the Neutrino Mass Hierarchy

## **Reactor neutrino experiment**



# $\overline{\nu}_e$ Energy Distribution

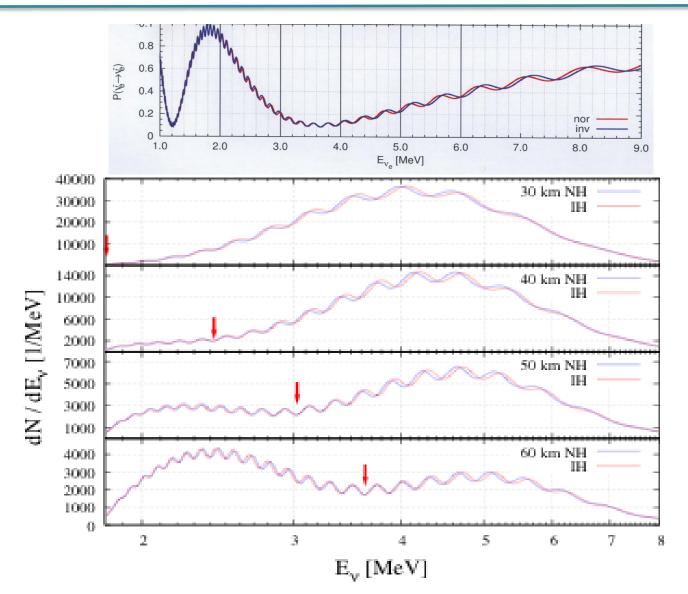




# How to distinguish Mass Hierarchy ?

$$\begin{split} P_{ee} &= \left| \sum_{i=1}^{3} U_{ei} \exp\left(-i\frac{m_{i}^{2}}{2E_{i}}\right) U_{ei}^{*} \right|^{2} \\ &= 1 - \cos^{4} \theta_{13} \sin^{2} 2\theta_{12} \sin^{2} (\Delta_{21}) \qquad \Delta_{ij} \equiv \frac{\Delta m_{ij}^{2}L}{4E_{\nu}}, \quad (\Delta m_{ij}^{2} \equiv m_{i}^{2} - m_{j}^{2}), \\ &- \sin^{2} 2\theta_{13} \sin^{2} (|\Delta_{31}|) \\ &- \sin^{2} \theta_{12} \sin^{2} 2\theta_{13} \sin^{2} (\Delta_{21}) \cos (2|\Delta_{31}|) \\ &= \frac{\sin^{2} \theta_{12}}{2} \sin^{2} 2\theta_{13} \sin (2\Delta_{21}) \sin (2|\Delta_{31}|) \\ &\qquad \\ &\text{Mass Hierarchy difference} \\ &\text{Max: } 2\Delta_{21} = \frac{\pi}{2} (2n - 1) \sim 36 \text{ km}, \cdots (E_{\nu} \sim 4 \text{MeV}) \\ &\text{vanish: } 2\Delta_{21} = n\pi \sim 72 \text{ km}, \cdots (E_{\nu} \sim 4 \text{MeV}) \end{split}$$

# MH difference in $\overline{\nu}_e$ spectrum



- Can determine MH independently from CP phase and matter effects.
- •Need only smaller detector than other experiments (e.g., LBL, atmospheric).
- •Free neutrino source with adjustable baseline length

# **Analysis methods for MH**

# determination

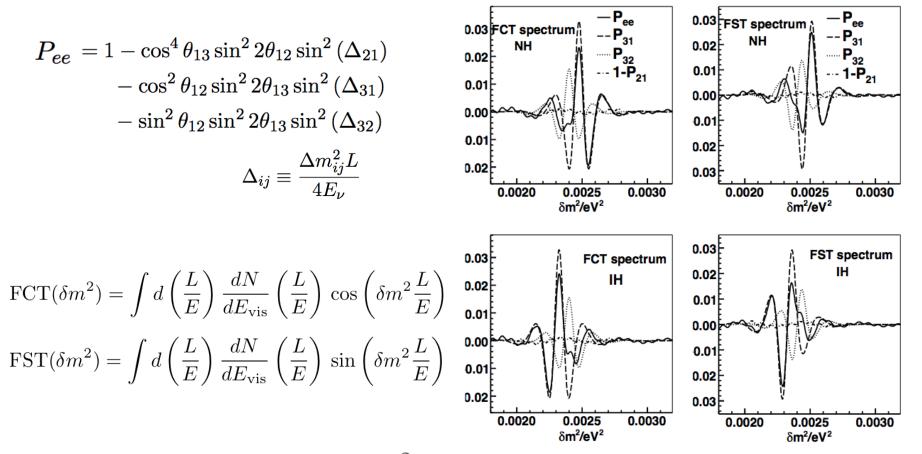
# Fourier analysis

hep-ph/0612022 J.G. Learned et al. arXiv: 0807.3203 L. Zhan et al. arXiv: 0901.2976 L. Zhan et al. arXiv: 1208.1551 X. Qian et al. arXiv: 1208.1991 E. Ciuffoli et al.



hep-ph/030601 S. Choubey et.al. arXiv: 0810.2580 M. Batygov et.al. arXiv: 1011.1646 P. Ghoshal et.al.

# **Fourier Analysis**



- has sensitivity for small  $\sin^2 2\theta_{13} \le 0.005$ .
- Don't need accurate knowledge of  $|\Delta m^2_{31(32)}|$  or  $\overline{
  u}_e$  flux.



$$\chi^{2} = \sum_{i=1}^{\text{nbin}} \left( \frac{N_{i}^{\text{fit}} - N_{i}^{\text{data}}}{\sqrt{N_{i}^{\text{data}}}} \right)^{2} + \underbrace{\sum_{i=1}^{\text{nparam}} \left( \frac{X_{i} - X_{i}^{\text{input}}}{\delta X_{i}} \right)^{2}}_{\text{Penalty term}},$$

Data is fitted with the theoretical prediction  $N_i^{\text{fit}}$ , assuming NH or IH.  $N_i^{\text{fit}} = \int dE_{\text{vis}} \frac{N_p T}{4\pi L^2} \int_{E_{\text{thr}}}^{\infty} dE_{\nu} \phi(E_{\nu}) P_{ee} \sigma_{IBD} G(E_{\nu}, E_{\text{vis}})$  $\chi^2_{\min}(NH) \chi^2_{\min}(IH)$ 

Fitting parameters are  $\theta_{12}, \theta_{13}, \Delta m_{21}^2, |\Delta m_{31}^2|, f_{sys}$ 

Y	$\sin^2 2 heta_{12}$	$\sin^2 2 heta_{13}$	$\Delta m^2_{21}\mathrm{eV}^2$	$ \Delta m^2_{31} \mathrm{eV}^2$	$f_{ m sys}$
$Y^{ ext{input}}$	0.857	0.089	$7.50  imes 10^{-5}$	$2.32  imes 10^{-3}$	1
$\delta Y$	0.024	0.005	$0.20  imes 10^{-5}$	$0.1  imes 10^{-3}$	0.03

# $\chi^2$ analysis – MH determination

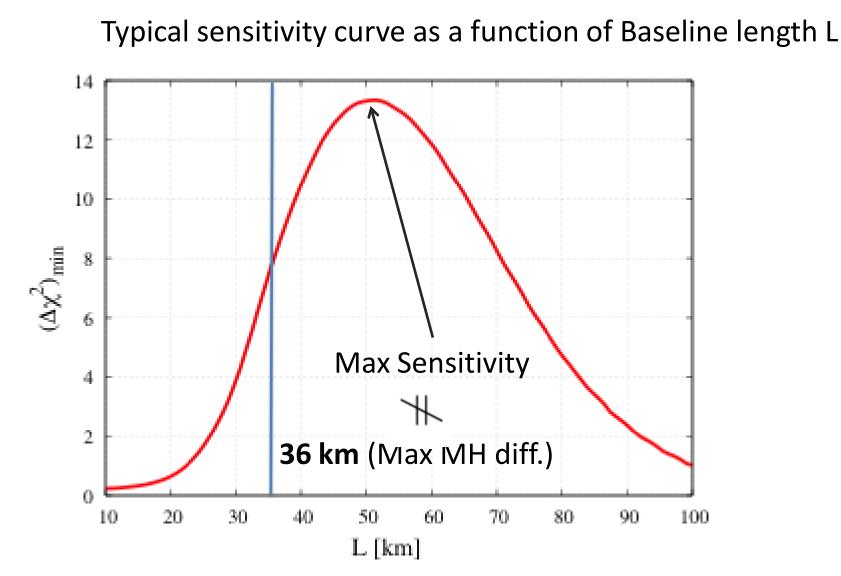
MH is determined as the hierarchy which gives smaller  $\chi^2_{\min}$ . Ex)  $\chi^2_{\min}(IH) > \chi^2_{\min}(NH) \implies$  MH is NH.

#### Sensitivity is estimated by

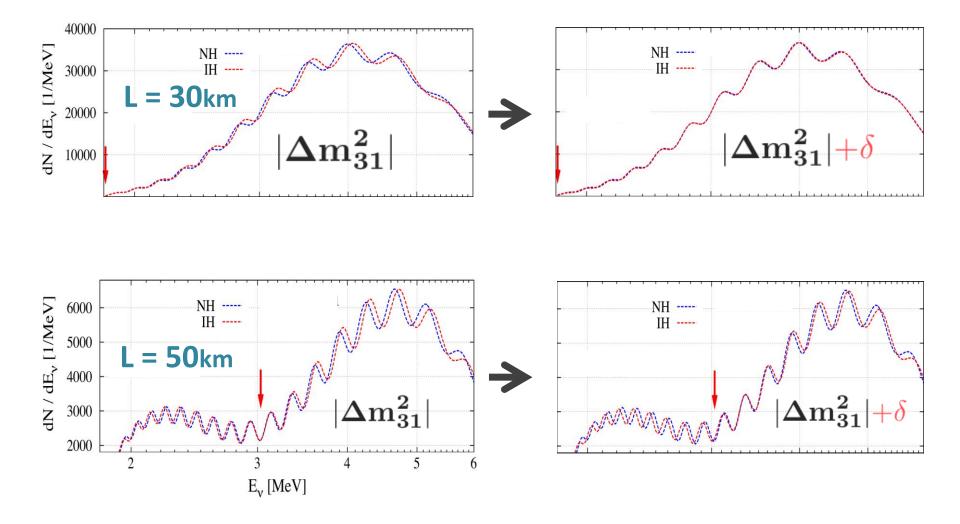
$$(\Delta \chi^2)_{\min} = \chi^2_{\min} (\text{rejected MH}) - \chi^2_{\min} (\text{accepted MH})$$
We claim that MH is determined with  $\sqrt{(\Delta \chi^2)_{\min}} \sigma$  significane.

2.46

# **Sensitivity for MH determination**

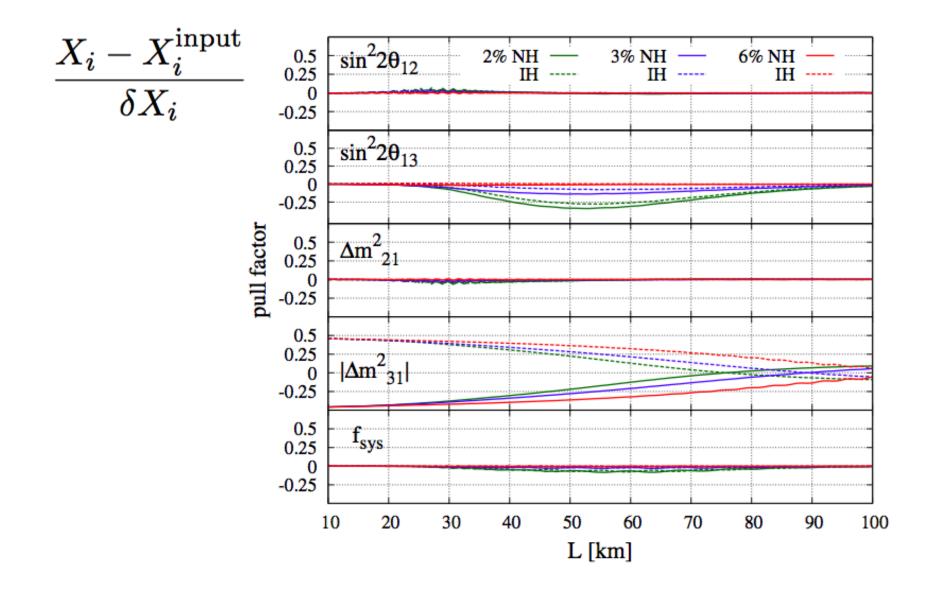


# Effect of $\delta |\Delta m_{31}^2|$

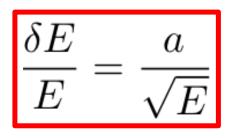


#### Baseline Length L should be long enough.

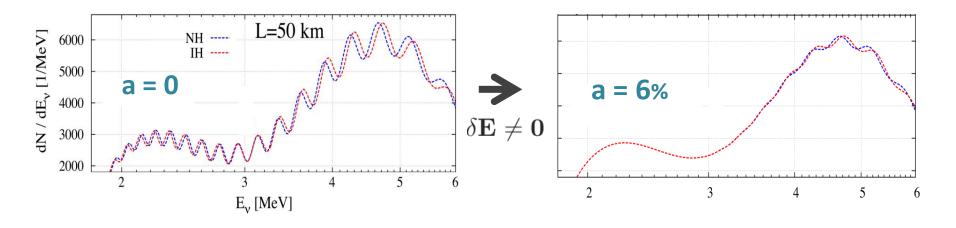
# **Pull Factor**



# **Effect of Energy Resolution**



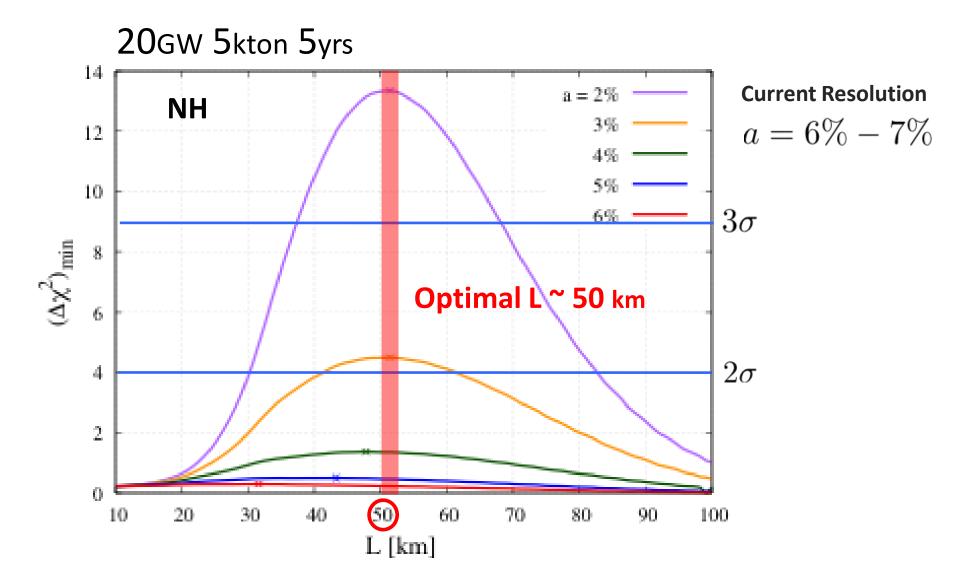
#### Evergy Resolution affects the sensitivity significantly.



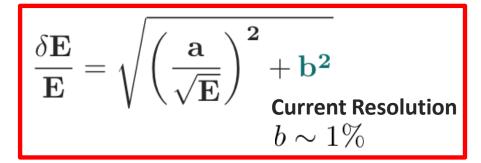
#### How good should the energy resolution be?

**3%**? → Not enough for RENO50 class of detector (5kton).

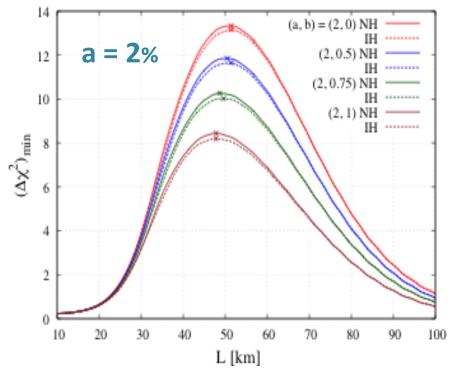
# **Sensitivity for MH determination**



# **Systematic part of Resolution**

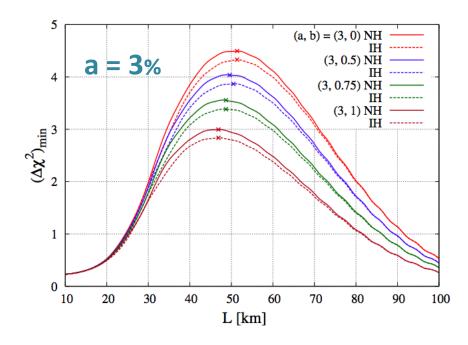


20GW 5kton 5yrs

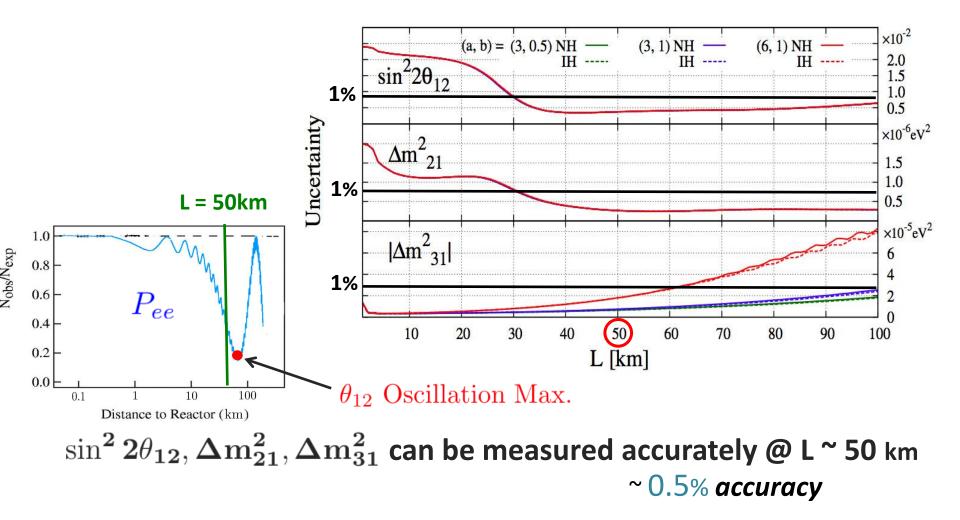


 $b = 0 \longrightarrow b = 1\%$ 

- Sensitivity is reduced by 40%.
- Optimal L is shortened by a few km.



## Parameter measurement @ L ~ 50 km



Parameter measurement is not sensitive to the Energy Resolution.

# **Fluctuation of Sensitivity**

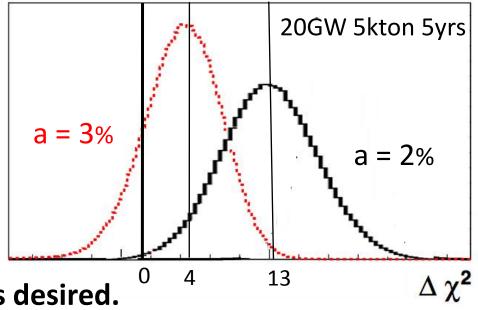
So far, we have discussed on the sensitivity using the "typical data set".

In the real experiments, event number in each bin,  $N_i^{\text{data}}$ , fluctuates from the "typical event number",  $\overline{N}_i$ .

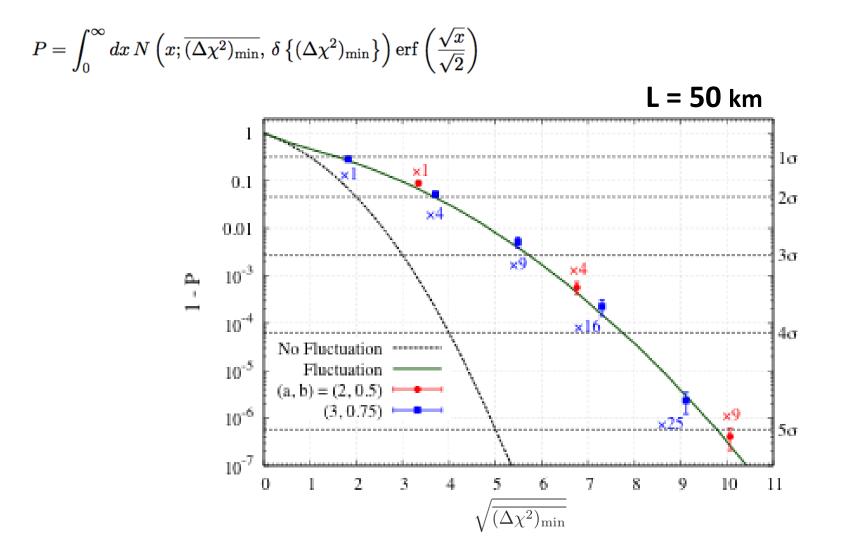
$$N_i^{\text{data}} = \overline{N}_i \pm \sqrt{\overline{N}_i}$$

Then,  $(\Delta \chi^2)_{\rm min}$  follows Gaussian distribution.

$$\delta(\Delta\chi^2)_{\rm min} \sim 2\sqrt{(\Delta\chi^2)_{\rm min}}$$



**2% level** of energy resolution is desired.



# How to achieve 2% energy resolution ?

#### Increase the photo-electron (p.e.)detected by PMT.

~ 200 pe/MeV 📥 ~ 2000 pe/MeV Ex) Increase PMT coverage Increase QE of PMT More transparent LS

## More photons, how?

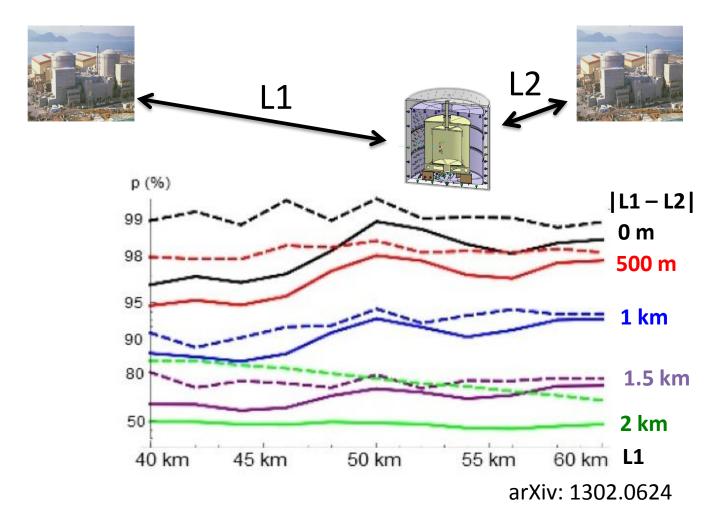
<ul> <li>♦ Highly transparent LS:</li> <li>⇒ Attenuation length/D: 15m/16m → 30m/34m</li> </ul>	×0.9
High light yield LS:	
⇒ KamLAND: 1.5g/l PPO → 5g/l PPO	
Light Yield: 30%→ 45%;	× 1.5
Photocathode coverage :	
⇒ KamLAND: 34% → ~80%	× 2.3
High QE "PMT":	
⇒ 20" SBA PMT QE: 25% → 35%	× 1.4
or New PMT QE: 25% → 40%	× 1.6
Both: 25% → 50%	× 2.0
4.3 - 5.0	→ (3.0 – 2.5)% /√E

**Other contributions: 0.5% constant term & 0.5% neutron recoil uncertainty** 

By Channgen Yang @ NuMass 2013

# **Multi-reactor interference**

Baseline difference should be < 500m.

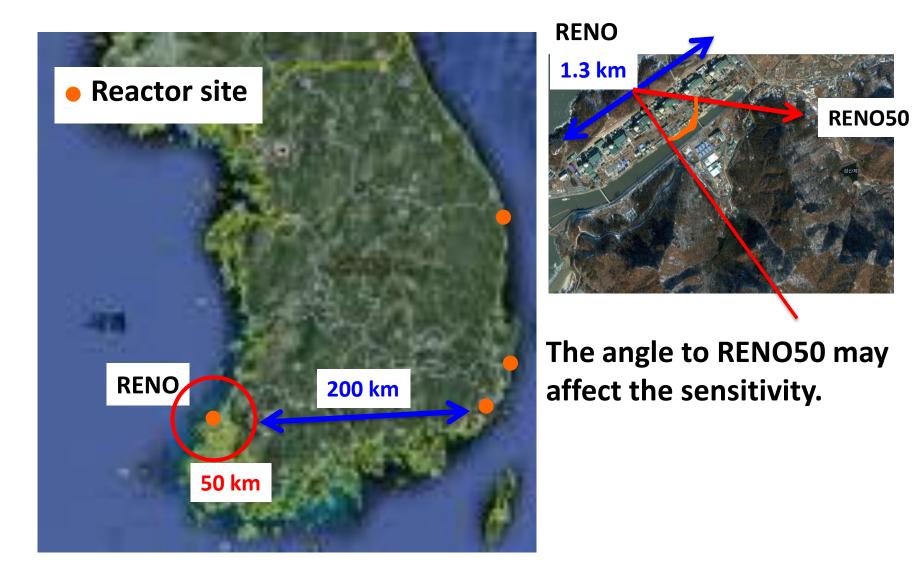


# **Multi-reactor interference**

#### Because of this effect, DayaBayII has changed its location.



# **Multi-reactor interference**



# **Summary**

We have discussed *the sensitivity* of *Medium Baseline Reactor Experiments* for *MH determination*.

### For 20GW 5kton 5yrs exposure,

 $\begin{array}{ll} \textbf{50km} & \text{optimal \textit{Baseline Length}} & \frac{\delta E}{E} = \sqrt{\left(\frac{a}{\sqrt{E}}\right)^2 + b^2} \\ \textbf{a < 3\%} \\ \textbf{b < 1\%} \end{array} \right\} \quad \text{of \textit{Energy Resolution}} \text{ is required.} \end{array}$ 

**0.5%** accuracy is achieved for  $\sin^2 2\theta_{12}, \Delta m_{21}^2, \Delta m_{31}^2$