

# Constraints on new physics from long baseline neutrino oscillation experiments

ニュートリノで探る “non-standard” physics

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based on

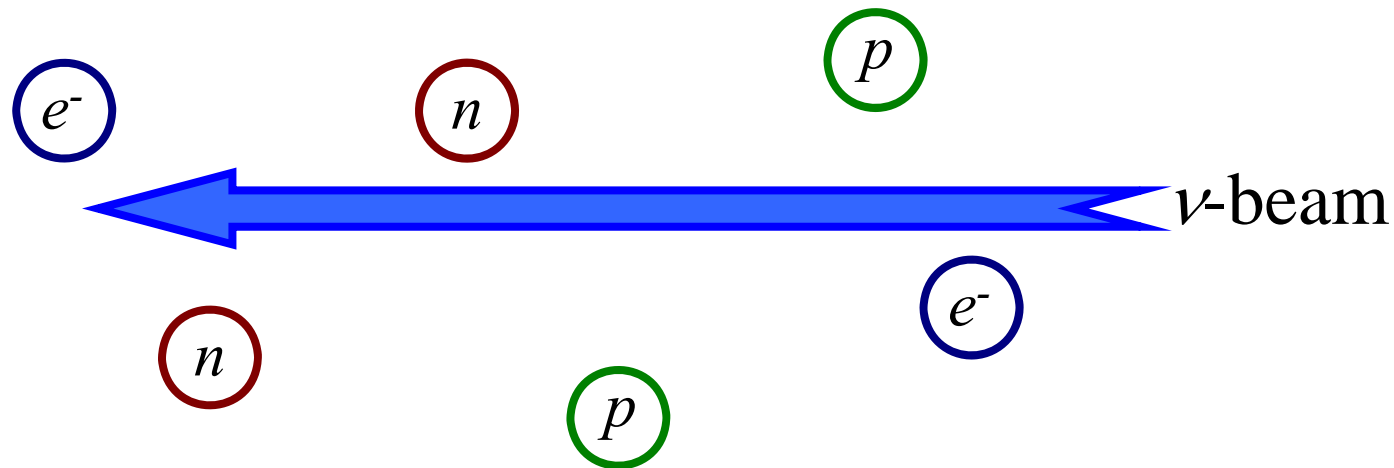
[hep-ph/0602115](#)

[hep-ph/0603268](#)

[hep-ph/0702xxx](#)

## Motivation

1. Can interactions between neutrinos and matter due to New Physics have any effect on neutrino oscillations ?

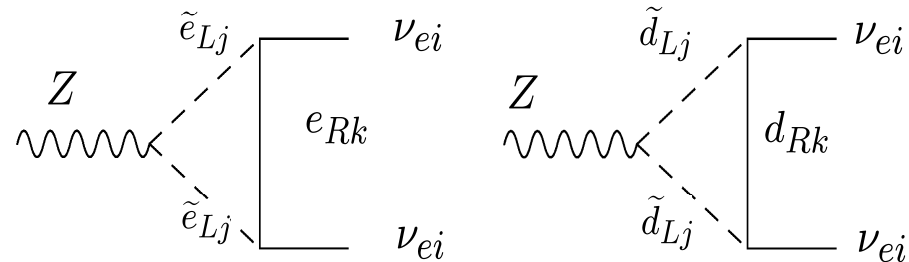


2. Can such effects be measured experimentally and be used to constrain New Physics ?

# New Physics effect

- New Physics breaks the NC interaction

- radiative corrections  
e.g. MSSM with  
R-parity violation



- Direct exchange of the New Particle

- Gauged  $B - \alpha L_e - \beta L_\mu - \gamma L_\tau$
- $Z'$  in topcolor assisted Technicolor
- Leptoquarks that couple different generations
- MSSM with R-Parity violating couplings
- *etc etc, Please consider the "fantastic model".*

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# Contents

- Motivation
- Can be measured ? (FNAL-to-HK)
- New Physics
  - $Z'$  gauge boson
  - Leptoquarks
  - $R$ -parity violation, SUSY
  - Extended Higgs Model
  - .....as you wish.....
- Summary

# The effective Hamiltonian

**U:MNS matrix**

$$H_{eff}^2 = U \begin{bmatrix} 0 & 0 & 0 \\ 0 & \delta m_{21}^2 & 0 \\ 0 & 0 & \delta m_{31}^2 \end{bmatrix} U^\dagger + \underbrace{\begin{bmatrix} a & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{bmatrix}}_{CC} + \underbrace{\begin{bmatrix} b_e & 0 & 0 \\ 0 & b_\mu & 0 \\ 0 & 0 & b_\tau \end{bmatrix}}_{NP}$$

derived from universality violation in NC and/or New Physics

$$= U \begin{bmatrix} 0 & 0 & 0 \\ 0 & \delta m_{21}^2 & 0 \\ 0 & 0 & \delta m_{31}^2 \end{bmatrix} U^\dagger$$

$$= \tilde{U} \begin{bmatrix} \lambda_1 & 0 & 0 \\ 0 & \lambda_2 & 0 \\ 0 & 0 & \lambda_3 \end{bmatrix} \tilde{U}^\dagger$$

central value of CHARM  
 $\xi = 0.025$

$\xi$  : universality violation in NC, New Physics

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 $\tilde{U}$  : effective mixing matrix  
 $\lambda_1, \lambda_2, \lambda_3$  : effective mass-squares

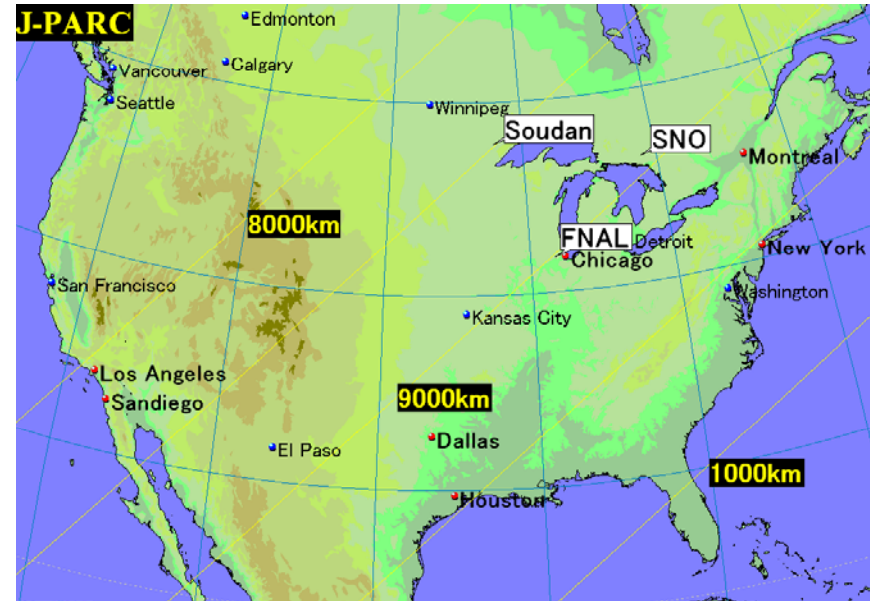
$$a \equiv 2EV_{CC} = 2\sqrt{2}G_F N_e E = 7.56 \times 10^{-5} (\text{eV})^2 \left( \frac{n_e}{g/\text{cm}^3} \right) \left( \frac{E}{g/\text{GeV}} \right)$$

$$b \equiv \underbrace{2EV_{NC}}_{\text{loop correction}} = -\sqrt{2}G_F N_n E \approx -\frac{1}{2}a \quad (N_e = N_n = N_p)$$

loop correction

# Can be measured ?

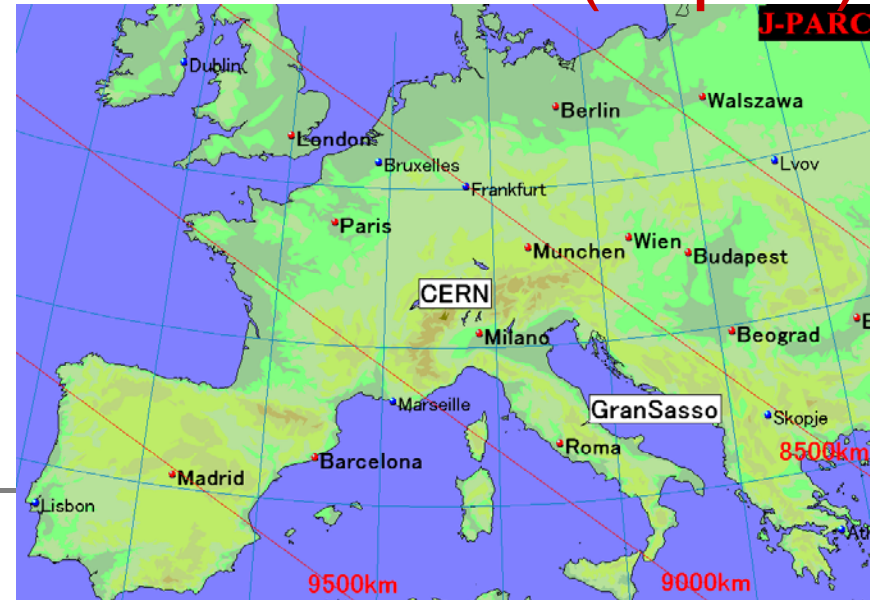
small value( $\xi$ )  
large matter effect  
 $\Leftrightarrow$  high energy  
 $\Leftrightarrow$  long base-line



to Japan



from J-PARC (Japan)



# Fermilab $\xrightarrow{\nu_\mu}$ HK [ L~9,120 km ]

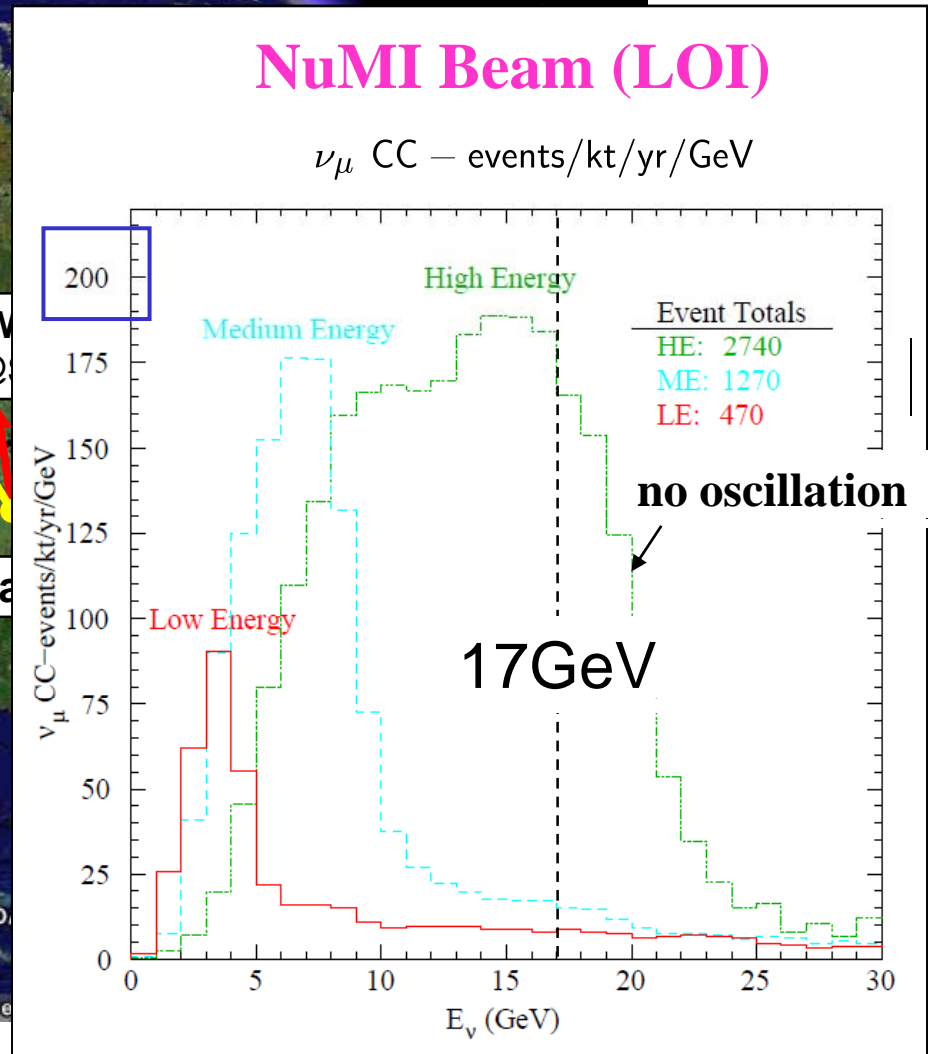
$$N_{\text{HK}} = N_{\text{Soudan}} \times V_{\text{HK}} \times \left( \frac{L_{\text{Soudan}}}{L_{\text{HK}}} \right)^2$$

$$\approx 200 \times 1000(\text{kt}) \times \left( \frac{732}{9120} \right)^2$$

$$\approx 1300(\text{event/Mt/yr}) \quad (@17\text{GeV})$$

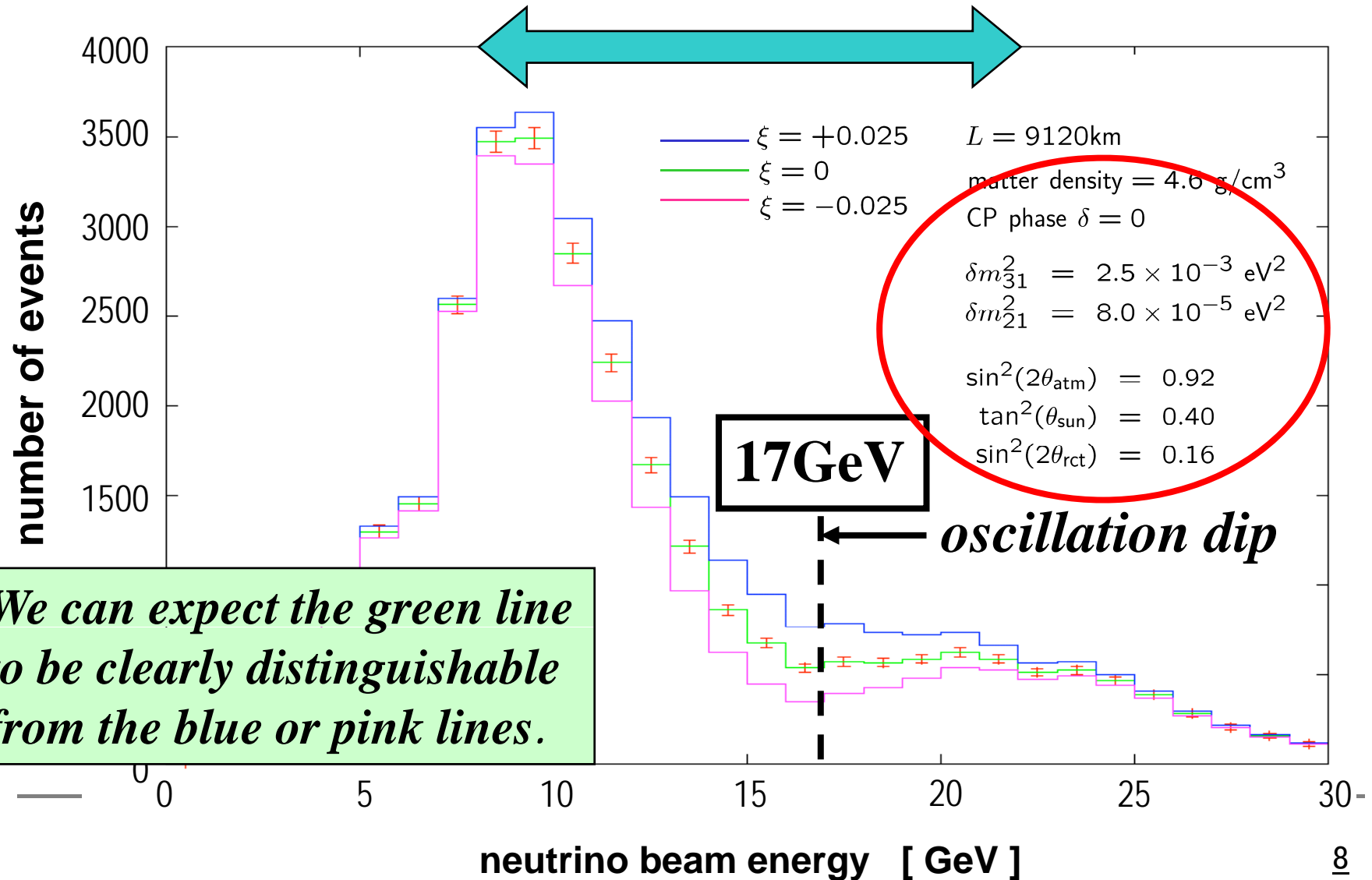
*We can expect  
a non-  
negligible event  
rate*

**even when oscillation  
is taken into account**



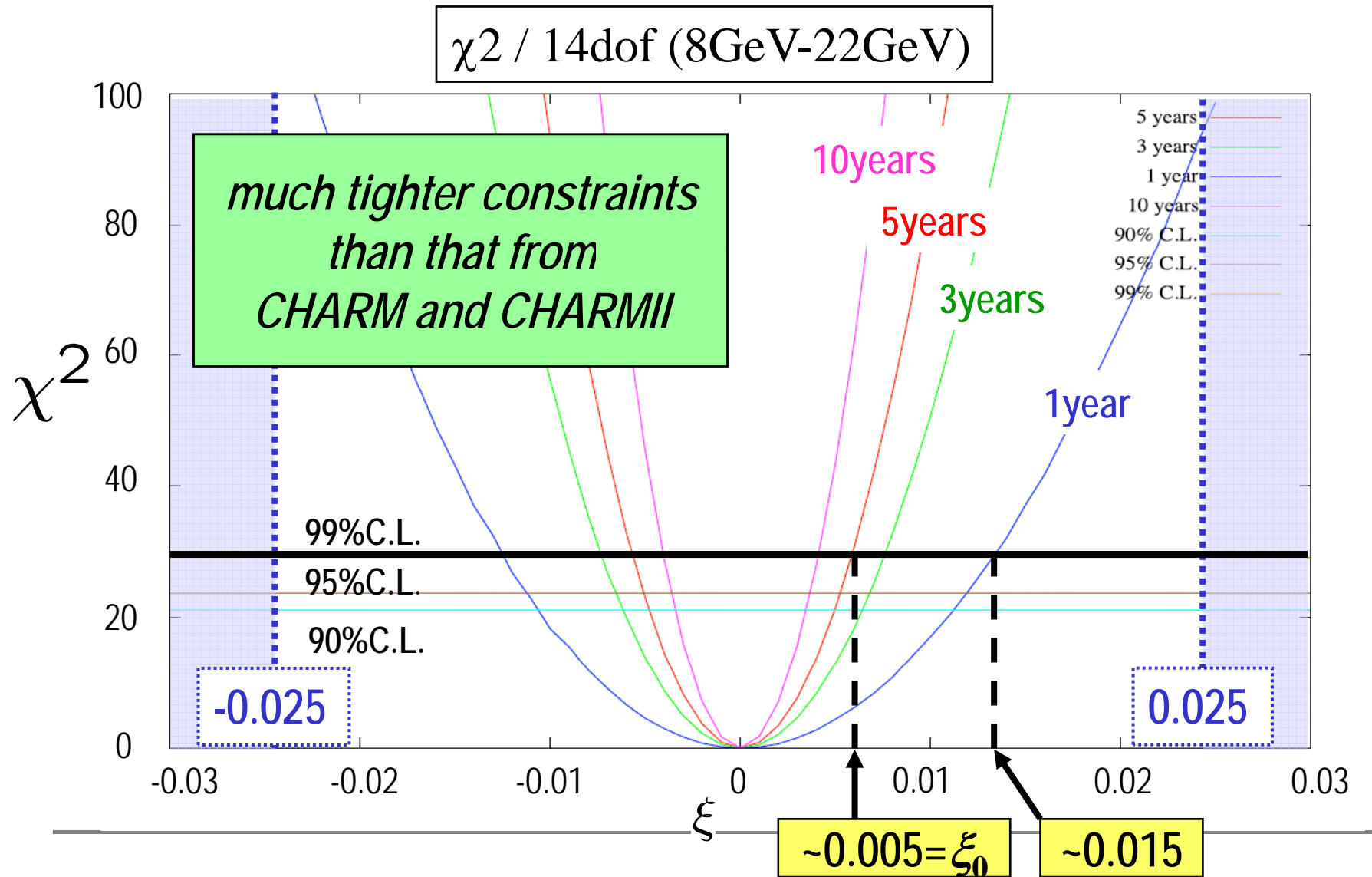
# The expected number of $\nu_\mu \rightarrow \nu_\mu$ events @ HK

*after 5 years of data taking (1Mton)*





# The constraints on universality violation $\xi$



# New Physics

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FNAL-to-HK

$$\xi > \xi_0 = 0.005$$



constraints on new physics

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# Constraints on the New Physics

- $Z'$  gauge boson
  - gauged  $B-(\alpha L_e + \beta L_\mu + \gamma L_\tau)$  with  $(\alpha + \beta + \gamma = 3)$
  - Topcolor Assisted Technicolor
- Leptoquarks
  - Scalar leptoquark
  - Vector leptoquark
- $R$ -parity violation
- Extended Higgs Model

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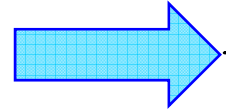
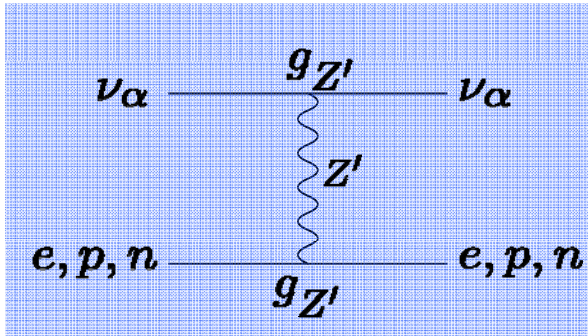
## Z' boson

$$SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)_X$$

# Z' boson

1. gauged  $B-(\alpha L_e + \beta L_\mu + \gamma L_\tau)$  with  $(\alpha + \beta + \gamma = 3)$

$$L_{Z'} = g_{Z'} \left[ \sum_q (\bar{q} \gamma^\mu q) - \alpha (\bar{l}_e \gamma^\mu l_e) - \beta (\bar{l}_\mu \gamma^\mu l_\mu) - \gamma (\bar{l}_\tau \gamma^\mu l_\tau) \right] Z'_\mu$$



$$\begin{cases} V_{\nu_e} = \alpha [\alpha N_e - (N_p + N_n)] \frac{g_{Z'}^2}{M_{Z'}^2} \approx \alpha(\alpha - 2)N \frac{g_{Z'}^2}{M_{Z'}^2} \\ V_{\nu_\mu} = \beta [\alpha N_e - (N_p + N_n)] \frac{g_{Z'}^2}{M_{Z'}^2} \approx \beta(\alpha - 2)N \frac{g_{Z'}^2}{M_{Z'}^2} \\ V_{\nu_\tau} = \gamma [\alpha N_e - (N_p + N_n)] \frac{g_{Z'}^2}{M_{Z'}^2} \approx \gamma(\alpha - 2)N \frac{g_{Z'}^2}{M_{Z'}^2} \end{cases}$$

$$\xi_{Z'} = \frac{V_{\nu_\mu} - V_{\nu_\tau}}{2V_{NC}}$$

$$= -4(\alpha - 2)(\alpha + 2\beta - 3) \frac{(g_{Z'}/M_{Z'})^2}{(g/M_W)^2}$$

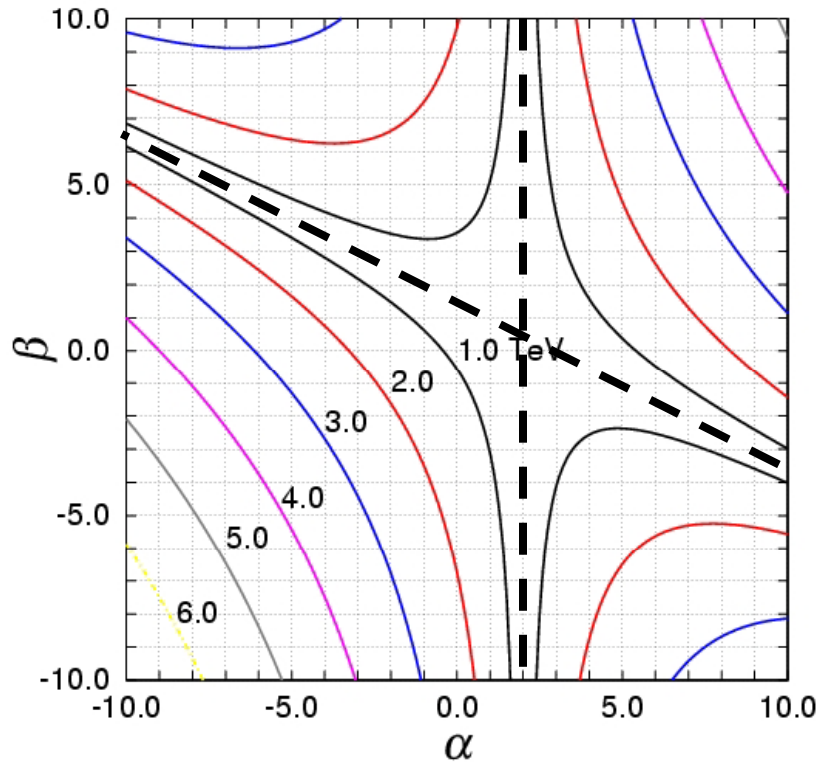
$$|\xi_{Z'}| \leq \xi_0 = 0.005$$

If  $\xi_0$  is larger than 0.005, the effect on the neutrino oscillation gives the constraint on the mass of  $Z'$

$$M_{Z'} \geq g_{Z'} \sqrt{\frac{|\alpha - 2| |\alpha + 2\beta - 3|}{\sqrt{2} G_F \xi_0}}$$

# Constraint on $M_{Z'}$ , (gauged)

$$|\xi_0|=0.005, \alpha+\beta+\gamma=3, g_{Z'}=0.10$$



$$M_{Z'} \geq g_{Z'} \sqrt{\frac{|\alpha - 2| |\alpha + 2\beta - 3|}{\sqrt{2} G_F \xi_0}}$$

**Current Limit**

Improve the limit

$$\alpha = \beta = 0 \quad (\gamma = 3)$$

$$g_{Z'} = 0.65:$$

$$M_{Z'} \geq 860 \text{ GeV} \Rightarrow 5500 \text{ GeV}$$

$$g_{Z'} = 0.35:$$

$$M_{Z'} \geq 370 \text{ GeV} \Rightarrow 3000 \text{ GeV}$$

Not "Improve" the limit

**Current Limit**

$$\beta = \gamma = 1.5 \quad (\alpha = 0)$$

$$g_{Z'} = 0.65:$$

$$M_{Z'} \geq 1100 \text{ GeV} \Rightarrow \text{no limit}$$

$$g_{Z'} = 0.35:$$

$$M_{Z'} \geq 490 \text{ GeV} \Rightarrow \text{no limit}$$

# Z' boson

C.T.Hill, PLB345,483(95)

G.Buchalla,etal, PRD53,5185(96)

W.Loinaz,etal,PRD60,015005(99)

## 2. Topcolor Assisted Technicolor

$$L = g'(\cot \theta \cdot J_{1s}^\mu - \tan \theta \cdot J_{1w}^\mu)Z'_\mu + g'(J_{1s}^\mu + J_{1w}^\mu)B_\mu \quad \theta \ll 1$$

$$J_{1s}^\mu = \frac{1}{6}(\bar{t}_L \gamma^\mu t_L + \bar{b}_L \gamma^\mu b_L) + \frac{2}{3}\bar{t}_R \gamma^\mu t_R - \frac{1}{2}\bar{b}_R \gamma^\mu b_R - \frac{1}{2}(\bar{\tau}_L \gamma^\mu \tau_L + \bar{\nu}_\tau \gamma^\mu \nu_\tau) - \bar{\tau}_R \gamma^\mu \tau_R$$

$$J_{1w}^\mu = [\text{1st. and 2nd. generation}]$$

Current limit  
 $M_{Z'} \sim 200\text{GeV}$

$$V_{\nu_e} \approx 0, \quad V_{\nu_\mu} \approx 0, \quad V_{\nu_\tau} \approx -\frac{N}{8} \frac{g_{Z'}^2}{M_{Z'}^2}$$

$$\xi_{TT} = -\frac{V_{\nu_\tau}}{2V_{NC}} = -\frac{1}{2} \frac{(g'/M_{Z'})^2}{(g/M_W)^2} = -\frac{1}{2} \sin^2 \theta_W \frac{M_Z^2}{M_{Z'}^2} \approx -0.1 \frac{M_Z^2}{M_{Z'}^2}$$

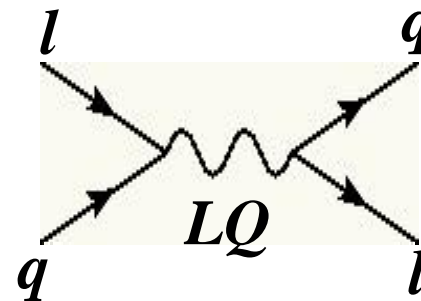
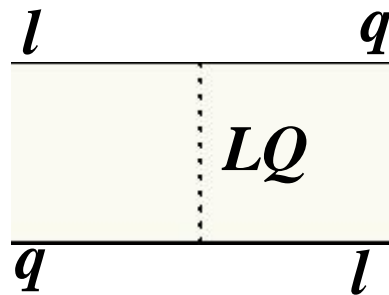
$$M_{Z'} \geq \sqrt{\frac{0.2311 \times (91.19)^2}{2 \times 0.005}} \approx 440\text{GeV} \left( \frac{0.005}{\xi_{TT}} \right)^{1/2}$$

If  $M_{Z'} < 440\text{ GeV} \Leftrightarrow$  The effect can be measured.



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# Leptoquarks





# Leptoquarks

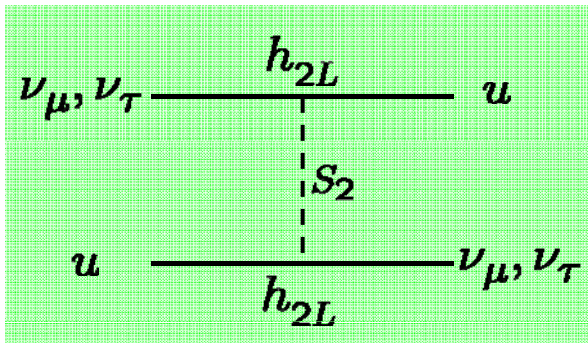
There are many leptoquarks (more than 10)

Fermion number :  $F=3B+L$ ,  $S$  : scalar,  $V$  : vector, subscript:dim.SU(2)

$$F = 2 : S_1, \tilde{S}_1, \vec{S}_3, V_{2\mu}, \tilde{V}_{2\mu} \quad F = 0 : S_2, \tilde{S}_2, V_{1\mu}, \tilde{V}_{1\mu}, \vec{V}_{3\mu}$$

$$L_{F=2} = [g_{1L} \bar{q}_L^C i\tau_2 l_L + g_{1R} \bar{u}_R^C e_R] S_1 + \tilde{g}_{1R} [\bar{d}_R^C e_R] \tilde{S}_1 + g_{3L} [\bar{q}_L^C i\tau_2 \vec{\tau} l_L] \vec{S}_3 \\ + [g_{2L} \bar{d}_R^C \gamma^\mu l_L + g_{2R} \bar{q}_R^C \gamma^\mu e_R] V_{2\mu} + \tilde{g}_{2L} [\bar{u}_R^C \gamma^\mu l_L] \tilde{V}_{2\mu} + h.c.$$

$$L_{F=0} = [h_{2L} \bar{u}_R l_L + h_{2R} \bar{q}_L i\tau_2 e_R] S_2 + \tilde{h}_{2L} [\bar{d}_R l_L] \tilde{S}_2 \\ + [h_{1L} \bar{q}_L \gamma^\mu l_L + h_{1R} \bar{d}_R \gamma^\mu e_R] V_{1\mu} + \tilde{h}_{1R} [\bar{u}_R \gamma^\mu e_R] \tilde{V}_{1\mu} + h_{3L} [\bar{q}_L \vec{\tau} \gamma^\mu l_L] \vec{V}_{3\mu} + h.c.$$

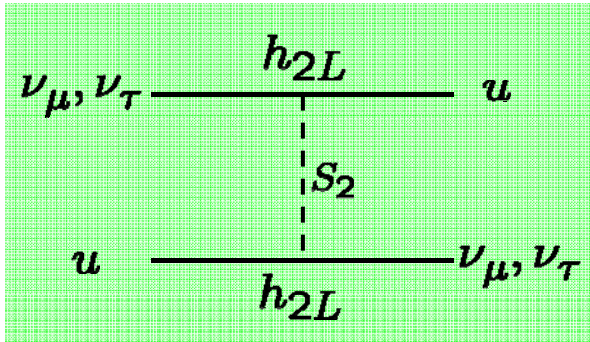


Only  $S_2$  and  $V_3$  cases are shown, today.  
(The others are in our paper)  
There is no special reason.

# Leptoquarks ( $S_2$ )

$$L_{F=0} = \left[ h_{2L}^{ijk} \bar{u}_{iR} l_{jL} + h_{2R}^{ijk} \bar{q}_{iL} i\tau_2 e_{jR} \right] S_{2k} + \tilde{h}_{2L}^{ijk} \bar{d}_{iR} l_{jL} \tilde{S}_{2k} + h.c. \quad (k=1 \dots L)$$

$$= h_{2L}^{12k} (\bar{u}_R v_{\mu L}) S_{2k}^- + h_{2L}^{13k} (\bar{u}_R v_{\tau L}) S_{2k}^- + \tilde{h}_{2L}^{12k} (\bar{d}_R v_{\mu L}) \tilde{S}_{2k}^- + \tilde{h}_{2L}^{13k} (\bar{d}_R v_{\tau L}) \tilde{S}_{2k}^- + h.c.$$



$$V_{\nu_\mu} = \frac{3}{4} N \sum_{k=1}^L \frac{|h_{2L}^{12k}|^2}{M_{S_{2k}}^2}, \quad V_{\nu_\tau} = \frac{3}{4} N \sum_{k=1}^L \frac{|h_{2L}^{13k}|^2}{M_{S_{2k}}^2}$$

$$\xi_{S_2^-} = \frac{V_{\nu_\mu} - V_{\nu_\tau}}{2V_{NC}} = -3 \sum_{k=1}^L \frac{|h_{2L}^{12k}/M_{S_{2k}}|^2 - |h_{2L}^{13k}/M_{S_{2k}}|^2}{(g/M_W)^2}$$

For  $M_{S_2^-} = 100\text{GeV}$  and  $L = 1$

$$\left| |h_{2L}^{12}|^2 - |h_{2L}^{13}|^2 \right| < 1.1 \times 10^{-3}$$

$$\left| |\tilde{h}_{2L}^{12}|^2 - |\tilde{h}_{2L}^{13}|^2 \right| < 1.1 \times 10^{-3}$$

$$(h_{2L}^{12})^2 \leq 1 \quad (\mu N \rightarrow \mu X)$$

$$(\tilde{h}_{2L}^{12})^2 \leq 2 \quad (\mu N \rightarrow \mu X)$$

**Current Limit**

# Leptoquarks ( $V_3$ )

$$\begin{aligned}
 L_{F=0} &= h_{3L}^{ijk} \left[ \bar{q}_{iL} \vec{\tau} \gamma^\mu l_L \right] \vec{V}_{3k,\mu} \quad (k=1 \cdots L) \\
 &= h_{3L}^{12k} \left[ (\bar{u}_L \gamma^\mu \nu_{\mu L}) V_{3k,\mu}^0 + \sqrt{2} (\bar{d}_L \gamma^\mu \nu_{\mu L}) V_{3k,\mu}^- \right] \\
 &\quad + h_{3L}^{13k} \left[ (\bar{u}_L \gamma^\mu \nu_{\tau L}) V_{3k,\mu}^0 + \sqrt{2} (\bar{d}_L \gamma^\mu \nu_{\tau L}) V_{3k,\mu}^- \right] + h.c.
 \end{aligned}$$

$$\begin{aligned}
 V_{\nu_\mu} &= 3N \sum_{k=1}^L |h_{3L}^{12k}|^2 \left[ \frac{1}{2M_{V_{3k}^0}^2} + \frac{1}{M_{V_{3k}^-}^2} \right], \quad V_{\nu_\tau} = 3N \sum_{k=1}^L |h_{3L}^{13k}|^2 \left[ \frac{1}{2M_{V_{3k}^0}^2} + \frac{1}{M_{V_{3k}^-}^2} \right] \\
 \xi_{\vec{V}_3} &= \frac{V_{\nu_\mu} - V_{\nu_\tau}}{2V_{NC}} = -\frac{6}{(g/M_W)^2} \sum_{k=1}^L \left( |h_{3L}^{12k}|^2 - |h_{3L}^{13k}|^2 \right) \left[ \frac{1}{M_{V_{3k}^0}^2} + \frac{2}{M_{V_{3k}^-}^2} \right]
 \end{aligned}$$

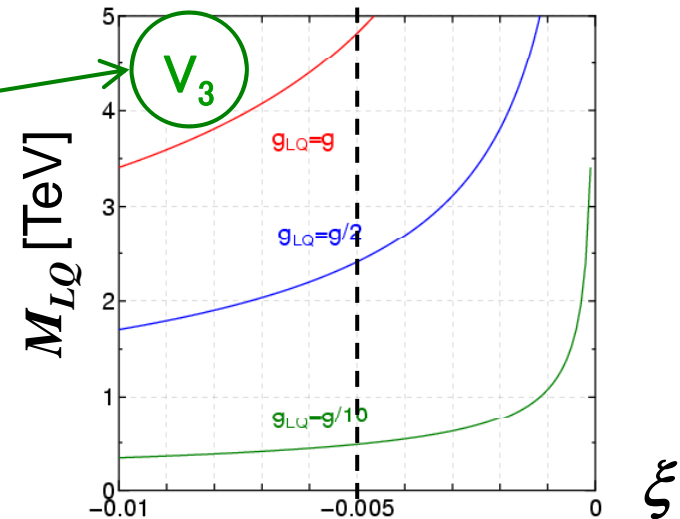
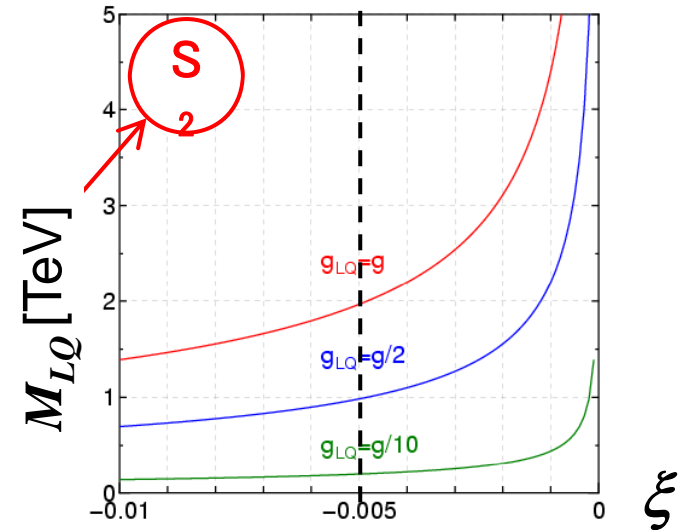
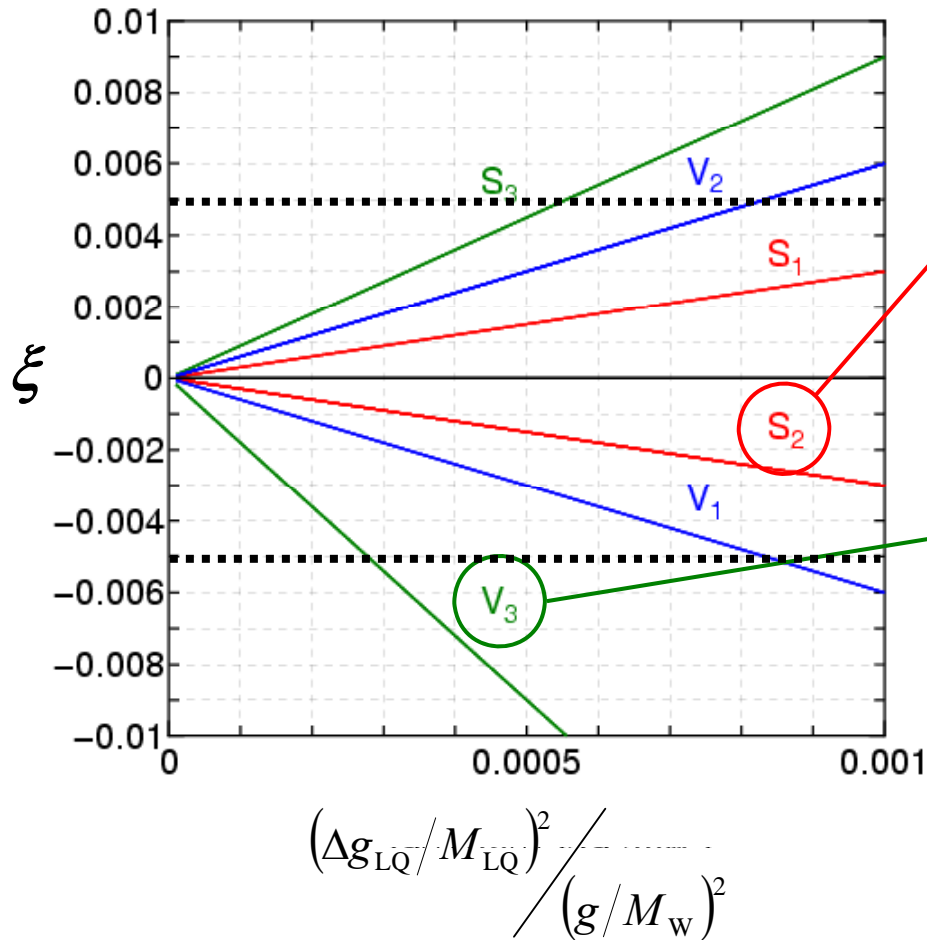
For  $M_{S_2^-} = 100\text{GeV}$  and  $L = 1$

**Current Limit**

$$\left| |h_{3L}^{12}|^2 - |h_{3L}^{13}|^2 \right| < 1.8 \times 10^{-4} \quad (h_{3L}^{12})^2 \leq 0.004 (R_\pi), \quad (h_{3L}^{13})^2 \leq 0.1 (D \rightarrow \mu\nu)$$

# Leptoquarks

S : scalar, V : vector,  
subscript : dim of SU(2) rep.



# SUSY

## R-parity Violation

$$W_R = \frac{1}{2} \lambda_{ijk} \hat{L}_i \hat{L}_j \hat{E}_k + \lambda'_{ijk} \hat{L}_i \hat{Q}_j \hat{D}_k + \frac{1}{2} \lambda''_{ijk} \hat{U}_i \hat{D}_j \hat{D}_k$$

tightly  
constrained  
by lepton univ.

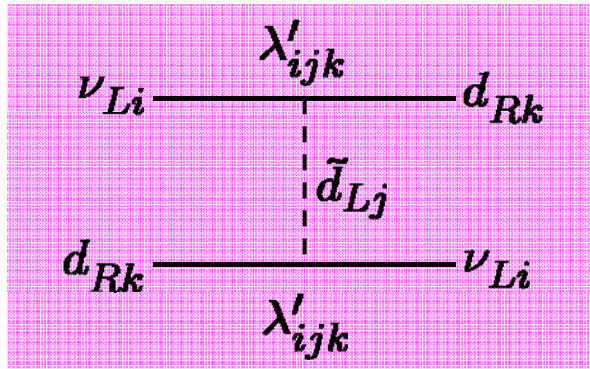
less stringent

irrelevant

# SUSY

R-parity violation

$$W_R = \frac{1}{2} \lambda_{ijk} \hat{L}_i \hat{L}_j \hat{E}_k + \lambda'_{ijk} \hat{L}_i \hat{Q}_j \hat{D}_k + \frac{1}{2} \lambda''_{ijk} \hat{U}_i \hat{D}_j \hat{D}_k$$



Similar to  $S_1$  and  $\tilde{S}_2^-$  type leptoquarks

$$V_{\nu_\mu} = \sum_{j=1}^3 \frac{3N}{4M_{\tilde{d}_j}^2} \left( |\lambda'_{2j1}|^2 - |\lambda'_{21j}|^2 \right)$$

$$V_{\nu_\tau} = \sum_{j=1}^3 \frac{3N}{4M_{\tilde{d}_j}^2} \left( |\lambda'_{3j1}|^2 - |\lambda'_{31j}|^2 \right)$$

$$\xi_{\tilde{d}} = \frac{V_{\nu_\mu} - V_{\nu_\tau}}{2V_{NC}} = -3 \sum_{j=1}^3 \frac{\left( |\lambda'_{2j1}|^2 - |\lambda'_{21j}|^2 \right) - \left( |\lambda'_{3j1}|^2 - |\lambda'_{31j}|^2 \right)}{(g/M_W)^2} \frac{1}{M_{\tilde{d}_j}^2} = -\frac{3\sqrt{2}}{8G_F} \frac{\Delta\lambda_{\tilde{d}}^2}{M_{\tilde{d}}^2}$$

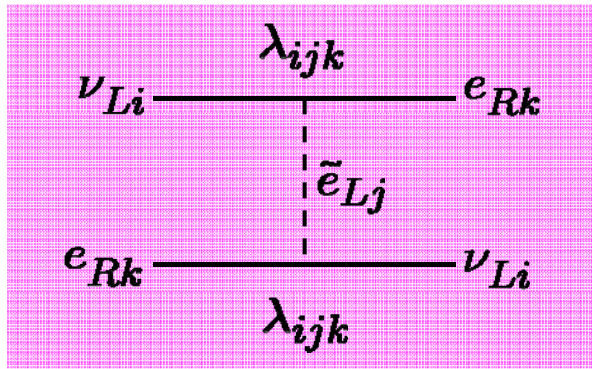
$$M_{\tilde{d}} \geq \sqrt{|\Delta\lambda_{\tilde{d}}^2|} \frac{\sqrt{3}}{2} \sqrt{\frac{1}{\sqrt{2}G_F \xi_{\tilde{d}}}} \approx 3.0[\text{TeV}] \sqrt{|\Delta\lambda_{\tilde{d}}^2|} \left( \frac{0.005}{\xi_{\tilde{d}}} \right)^{1/2}$$

Current Limit  
 $M_d > 300\text{GeV}$

# SUSY

R-parity violation

$$W_R = \frac{1}{2} \lambda_{ijk} \hat{L}_i \hat{L}_j \hat{E}_k + \lambda'_{ijk} \hat{L}_i \hat{Q}_j \hat{D}_k + \frac{1}{2} \lambda''_{ijk} \hat{U}_i \hat{D}_j \hat{D}_k$$



Similar  $S_1$ ,  $S_2^-$  and  $\tilde{S}_2^-$  type leptoquarks

$$V_{\nu_\mu} = \frac{N}{4} \left( \sum_{i=1,3} \frac{|\lambda'_{i21}|^2}{M_{\tilde{e}_i}^2} - \sum_{i=1}^3 \frac{|\lambda'_{21i}|^2}{M_{\tilde{e}_i}^2} \right)$$

$$V_{\nu_\tau} = \frac{N}{4} \left( \sum_{i=1,2} \frac{|\lambda'_{i31}|^2}{M_{\tilde{e}_i}^2} - \sum_{i=1}^3 \frac{|\lambda'_{31i}|^2}{M_{\tilde{e}_i}^2} \right)$$

$$\xi_{\tilde{e}} = \frac{V_{\nu_\mu} - V_{\nu_\tau}}{2V_{NC}} = \frac{\sum_{i=1,3} |\lambda_{i21}|^2 / M_{\tilde{e}_i}^2 + \sum_{i=1,2} |\lambda_{i31}|^2 / M_{\tilde{e}_i}^2 - \sum_{i=1}^3 (|\lambda_{21i}|^2 + |\lambda_{31i}|^2) / M_{\tilde{e}_i}^2}{(g/M_W)^2}$$

$$= -\frac{\sqrt{2}}{8G_F} \frac{\Delta\lambda_{\tilde{e}}^2}{M_{\tilde{e}}^2},$$

$$M_{\tilde{e}} \geq 1.7[\text{TeV}] \sqrt{|\Delta\lambda_{\tilde{e}}^2|} \left( \frac{0.005}{\xi_{\tilde{e}}} \right)^{1/2}$$

Current Limit  
 $\lambda \sim \mathcal{O}(10^{-2})$

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# Extended Higgs

triplet Higgs for neutrino mass



# Extended Higgs Models

K.S.Babu, PLB203, 132 (88)

A.Zee, PLB93, 389 (80)

E.Accomande etal,  
hep-ph/0608079, p497

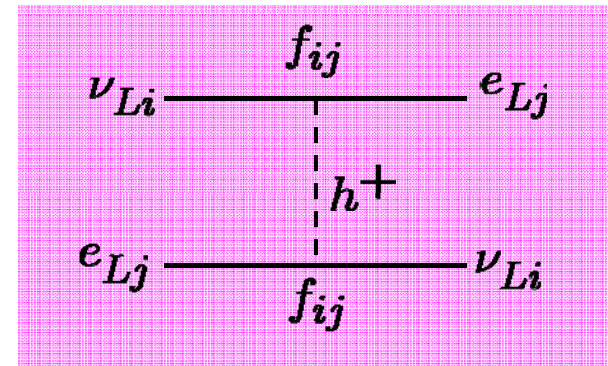
$$L = f_{ij} (\bar{l}_{iL}^C i \tau_2 l_{jL}) h^+ + h.c.$$

$$= -2(f_{e\mu} \bar{\nu}_\mu^C e_L + f_{e\tau} \bar{\nu}_\tau^C e_L) h^+ + h.c. + \dots$$

$f_{ij}$ : Yukawa couplings (anti-symmetric)  
 $h^+$ : Higgs particle with Y=1

$$V_{\nu_\mu} = -N \frac{|f_{e\mu}|^2}{M_{h^+}^2}, \quad V_{\nu_\tau} = -N \frac{|f_{e\tau}|^2}{M_{h^+}^2}$$

$$\xi_{h^+} = \frac{V_{\nu_\mu} - V_{\nu_\tau}}{2V_{NC}} = 4 \frac{(|f_{e\mu}|^2 - |f_{e\tau}|^2) / M_{h^+}^2}{(g/M_W)^2} = \frac{4\sqrt{2}}{8G_F} \frac{\Delta\lambda_{h^+}^2}{M_{h^+}^2}$$



$$\left| \frac{|f_{e\mu}|^2 - |f_{e\tau}|^2}{M_{h^+}^2} \right| = \sqrt{2} G_F \xi_0 \approx 8.2 \times 10^{-8} \quad \longleftrightarrow \quad 3.4 \times 10^{-8} (\tau^- \rightarrow \nu_\tau e^- \bar{\nu}_e)$$

**Current Limit**

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## Summary

- New Physics effects on neutrino oscillation
- FNAL -to- HK
  - $\xi_0 = (V_{\nu\mu} - V_{\nu\tau})/V_{NC} \sim 0.005$
  - (coupling / mass)<sup>2</sup>
- Loop correction to the NC
- Direct interaction
  - $Z'$  gauge boson
  - Leptoquarks
  - $R$ -parity violation
  - Extended Higgs Model

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*Thank you very much  
for your attention*



# topcolor assisted technicolor

## topcolor

*top quark* condensate gives rise to a triplet of NG bosons  $\Leftrightarrow$  **W,Z**.

$$f_\pi^2 = m_t^2 \left( \frac{N_C}{16\pi^2} \ln \frac{\Lambda^2}{\mu^2} \right), \quad O(\mu) \sim O(m_t)$$

**top-pions**

$f_\pi = 174 \text{ GeV} \Leftrightarrow \Lambda \sim 10^{13-14} \text{ GeV} \Leftrightarrow$  large hierarchy  $\Leftrightarrow$  fine tuning of the c.c.

## T.A.T.

$$SU(3)_S \times SU(3)_W \times U(1)_S \times U(1)_W \times SU(2)_L \xrightarrow{\text{TechniC}} SU(3)_C \times U(1)_Y \times SU(2)_L$$

$$\xrightarrow{\text{TopC}} SU(3)_C \times U(1)_{em}$$

$$F_\pi \approx 167 \text{ GeV}, \text{ TechniC} \quad f_\pi \approx 50 \text{ GeV}, \quad \Lambda_{\text{TopC}} \sim \text{TeV}$$

The majority of the **W** and **Z** masses come from the technifermion condensate.

$$\begin{aligned} SU(3)_S \times SU(3)_W &\rightarrow SU(3)_C : \text{colorons and gluons} \\ U(1)_S \times U(1)_W &\rightarrow U(1)_Y : \mathbf{Z}' \quad \text{and} \quad \mathbf{B} \end{aligned}$$

3<sup>rd</sup>. gene.

1<sup>st</sup>. 2<sup>nd</sup>. gene.

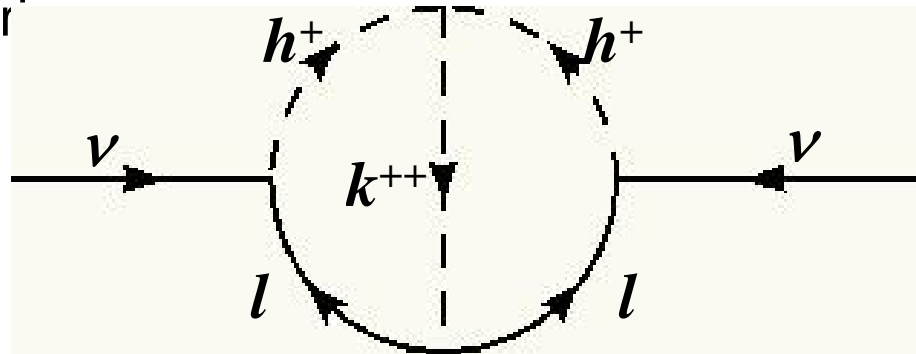
# Extended Higgs Models

$$L = f_{ij} (\bar{l}_{iL}^C i \tau_2 l_{jL}) h^+ + h_{ij} (\bar{l}_{iR}^C l_{jR}) k^{++} - \mu h^+ h^+ k^{--} + h.c.$$

$f_{ij}$  : anti-symmetric,  $h_{ij}$  : symmetric

$h^+$  : Higgs #L = -2

$k^{++}$  : Higgs #L = -2



$$(M_\nu)_{ij} = 8\mu f_{ia} m_a h_{ab}^* m_b f_{bj} \cdot I_{ab}$$

$$I_{ab} \approx \frac{1}{(16\pi^2)^2} \frac{1}{m_h^2} \tilde{I} \left( \frac{m_k^2}{m_h^2} \right)$$

$$\tilde{I}(r) = -\int_0^1 dx \int_0^{1-x} dy \frac{1-y}{x + (r-1)y + y^2} \log \frac{y(1-y)}{x+ry}$$

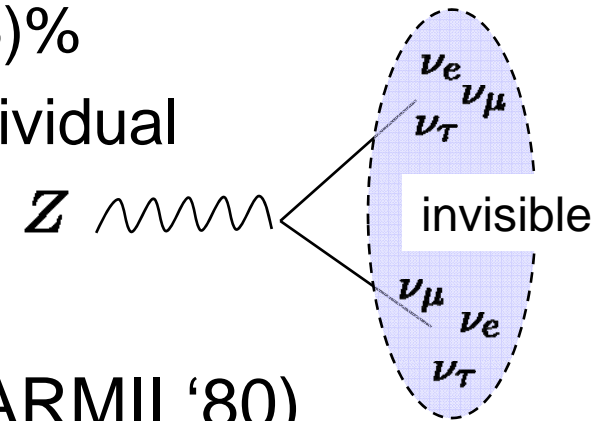
“ペンギンだいたいあぐらむ”って  
知ってる？

こんな格好  
でしょ？



# Experimental Constraints (NC)

- Invisible decay width (LEP/SLC)
  - Invisible Z decay width,  $(20.00 \pm 0.06)\%$
  - No information is available on the individual decay widths into each flavor.



- Individual decay width (CHARM/CHARMII '80)
  - Z couplings to neutrinos [PLB180,303('86), PLB320 203('94)]

$$g^{\nu_e} : 0.528 \pm 0.085 \quad g^{\nu_e} : 0.502 \pm 0.017$$

$$g^{\nu_e}/g^{\nu_\mu} = 1.05^{+0.15}_{-0.18} \rightarrow \boxed{0.9-1.2}$$

- The theoretical possibility of universality violation in many models
- The weakness of the current experimental bounds

Z DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )
$e^+ e^-$	( 3.363 $\pm$ 0.004 ) %
$\mu^+ \mu^-$	( 3.366 $\pm$ 0.007 ) %
$\tau^+ \tau^-$	( 3.370 $\pm$ 0.008 ) %
invisible	(20.00 $\pm$ 0.06 ) %
hadrons	(69.91 $\pm$ 0.06 ) %

**CHARM,CHARM II**

$$g^{\nu e} = 0.528 \pm 0.085$$

$$g^{\nu \mu} = 0.502 \pm 0.017$$

$$g^{\nu e}/g^{\nu \mu} = 1.05^{+0.15}_{-0.18}$$

*It would be interesting to analyze the effect of such a violation on neutrino oscillations.*



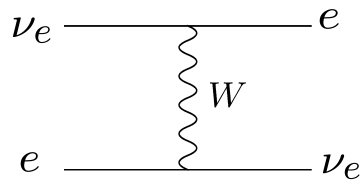
**Yeah!!**



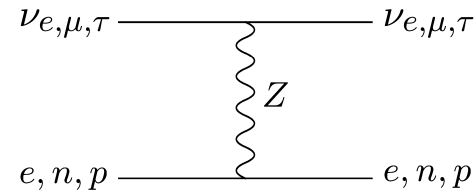
**Yeah!!**

# The effective potentials

## charged current interaction



## neutral current interaction



$$M_{CC} = -\frac{G_F}{\sqrt{2}} [\bar{e}\gamma^\mu(1-\gamma_5)\nu_e] [\bar{\nu}_e\gamma_\mu(1-\gamma_5)e]$$

↓ *Fierz tran.*

$$= +\frac{G_F}{\sqrt{2}} \underbrace{[\bar{e}\gamma^\mu(1-\gamma_5)e]}_{N_e} \underbrace{[\bar{\nu}_e\gamma_\mu(1-\gamma_5)\nu_e]}_2$$

$$M_{NC} = \rho \frac{G_F}{\sqrt{2}} [\bar{\nu}_\alpha\gamma^\mu(1-\gamma_5)\nu_\alpha] \quad (\alpha = e, \mu, \tau)$$

$$\times \left[ \underbrace{g_L^f \bar{\psi}_f \gamma_\mu (1-\gamma_5) \psi_f}_{N_f} + \underbrace{g_R^f \bar{\psi}_f \gamma_\mu (1+\gamma_5) \psi_f}_{N_f} \right]$$

$$g_L^f = I_3^f - Q_f \sin^2 \theta_W, \quad g_R^f = -Q_f \sin^2 \theta_W \quad (f = e, n, p)$$

$$V_{CC} = \sqrt{2} G_F N_e$$

$N_e$ : electron number density in matter

$$g_L^f + g_R^f = I_3^f - 2Q_f \sin^2 \theta_W$$

$$f = e^- : (-1/2) - 2(-1)\sin^2 \theta_W$$

$$p : (+1/2) - 2(+1)\sin^2 \theta_W$$

$$n : (-1/2) - 2(0)\sin^2 \theta_W \sim -\frac{1}{2}$$

$$V_{NC} = \rho \sqrt{2} G_F N_f (g_L^f + g_R^f) \\ = \rho \sqrt{2} G_F N_n (-1/2)$$

$$V_{NC} \approx (-\sqrt{2}/2) G_F N_n \\ \approx (-1/2) V_{CC}$$

$N_n$ : neutron number density

# effective mixing angles (normal)

$$\delta = 0 \text{ case, expanded in powers of } \varepsilon \equiv \sqrt{\frac{\delta m_{21}^2}{|\delta m_{31}^2|}} = 0.15 \sim 0.24$$

$$\text{in the region of } \frac{a}{|\delta m_{31}^2|} \sim O(\varepsilon^{-1}): (E(\text{GeV}) \sim O(10)), \xi \sim O(\varepsilon^2)$$

neutrino

$$\tilde{\theta}_{13} = \theta'_{13} \approx \frac{\pi}{2} - \left( \frac{\delta m_{31}^2}{a} \right) \theta_{13}$$

$$\tilde{\theta}_{12} = \frac{\pi}{2} - \frac{c_{13}}{c'_{13}} \left( \frac{\delta m_{21}^2}{2a} \right) \sin 2\theta_{12} - \xi \frac{a}{2\delta m_{31}^2}$$

$$\tilde{\theta}_{23} = \theta_{23} + \frac{s_\phi}{c'_{13}} \left( \frac{\delta m_{21}^2}{2a} \right) \sin 2\theta_{12}$$

$$\tilde{\delta} = 0$$

$$\left( \phi = \frac{\pi}{2} - \theta_{13} - \left( \frac{\delta m_{31}^2}{a} \right) \theta_{13} \right)$$

anti - neutrino

$$\tilde{\theta}_{13} = \theta'_{13} \approx \left( \frac{\delta m_{31}^2}{a} \right) \theta_{13}$$

$$\tilde{\theta}_{12} = \left( \frac{\delta m_{21}^2}{2a} \right) \sin 2\theta_{12}$$

$$\tilde{\theta}_{23} = \theta_{23} + \xi \frac{a}{2\delta m_{31}^2}$$

$$\tilde{\delta} = 0$$

# effective mixing angles (inverted)

$$\delta = 0 \text{ case, expanded in powers of } \varepsilon \equiv \sqrt{\frac{\delta m_{21}^2}{|\delta m_{31}^2|}} = 0.15 \sim 0.24$$

in the region of  $\frac{a}{|\delta m_{31}^2|} \sim O(\varepsilon^{-1}) : (E(\text{GeV}) \sim O(10)), \xi \sim O(\varepsilon^2)$

neutrino

$$\tilde{\theta}_{13} = \theta'_{13} \approx -\left(\frac{\delta m_{31}^2}{a}\right)\theta_{13}$$

$$\tilde{\theta}_{12} = \frac{\pi}{2} - \left(\frac{\delta m_{21}^2}{2a}\right)\sin 2\theta_{12}$$

$$\tilde{\theta}_{23} = \theta_{23} + \xi \frac{a}{2|\delta m_{31}^2|}$$

$$\tilde{\delta} = 0$$

anti - neutrino

$$\tilde{\theta}_{13} = \theta'_{13} \approx \frac{\pi}{2} + \left(\frac{\delta m_{31}^2}{a}\right)\theta_{13}$$

$$\tilde{\theta}_{12} = \frac{c_{13}}{c'_{13}} \left(\frac{\delta m_{21}^2}{2a}\right)\sin 2\theta_{12} - \xi \frac{a}{2|\delta m_{31}^2|}$$

$$\tilde{\theta}_{23} = \theta_{23} - \frac{s_\phi}{c'_{13}} \left(\frac{\delta m_{21}^2}{2a}\right)\sin 2\theta_{12}$$

$$\tilde{\delta} = 0$$

$$\left(\phi = \frac{\pi}{2} - \theta_{13} - \left(\frac{|\delta m_{31}^2|}{a}\right)\theta_{13}\right)$$

## shift of effective mixing by $\xi$

	<i>neutrino</i>	<i>anti-neutrino</i>
<i>normal hierarchy</i>	$\tilde{\theta}_{12}$ is shifted by $-\frac{a}{2 \delta m_{31}^2 }\xi$	$\tilde{\theta}_{23}$ is shifted by $+\frac{a}{2 \delta m_{31}^2 }\xi$
<i>inverted hierarchy</i>	$\tilde{\theta}_{23}$ is shifted by $+\frac{a}{2 \delta m_{31}^2 }\xi$	$\tilde{\theta}_{12}$ is shifted by $-\frac{a}{2 \delta m_{31}^2 }\xi$

in the region of

$$\frac{a}{|\delta m_{31}^2|} \approx O(\varepsilon^{-1}): (E(\text{GeV}) \sim O(10)), \quad \xi \approx O(\varepsilon^2), \quad \left( \varepsilon \equiv \sqrt{\frac{\delta m_{21}^2}{|\delta m_{31}^2|}} = 0.15 \sim 0.24 \right)$$

# Survival Probability

$$\begin{aligned} \tilde{P}(v_\mu \rightarrow v_\mu) &= 1 - 4|\tilde{U}_{\mu 2}|^2 \left(1 - |\tilde{U}_{\mu 2}|^2\right) \sin^2 \frac{\tilde{\Delta}_{21}}{2} - 4|\tilde{U}_{\mu 3}|^2 \left(1 - |\tilde{U}_{\mu 3}|^2\right) \sin^2 \frac{\tilde{\Delta}_{31}}{2} \\ &\quad + 2|\tilde{U}_{\mu 2}|^2 |\tilde{U}_{\mu 3}|^2 \left(4 \sin^2 \frac{\tilde{\Delta}_{21}}{2} \sin^2 \frac{\tilde{\Delta}_{31}}{2} + \sin \tilde{\Delta}_{21} \sin \tilde{\Delta}_{31}\right) \\ \tilde{U}_{\mu 2} &= \tilde{c}_{12} \tilde{c}_{23} - \tilde{s}_{12} \tilde{s}_{13} \tilde{s}_{23}, \quad \tilde{U}_{\mu 3} = \tilde{c}_{13} \tilde{s}_{23} \end{aligned}$$

neutrino (normal hierarchy)

$$\tilde{U}_{\mu 2} \approx \cos(\tilde{\theta}_{12} + \tilde{\theta}_{23}), \quad \tilde{U}_{\mu 3} \approx 0$$

$$\tilde{P}(v_\mu \rightarrow v_\mu) \approx 1 - \sin^2 \left[ 2 \left( \theta_{23} - \frac{a}{2|\delta m_{31}^2|} \xi \right) \right] \sin^2 \frac{\tilde{\Delta}_{21}}{2}$$

$$\tilde{\Delta}_{21} \approx \delta m_{31}^2 c_{13}^2 - \delta m_{21}^2 c_{12}^2$$

neutrino (inverted hierarchy)

$$\tilde{U}_{\mu 2} \approx 0, \quad \tilde{U}_{\mu 3} \approx \tilde{s}_{23}$$

$$\tilde{P}(v_\mu \rightarrow v_\mu) \approx 1 - \sin^2 \left[ 2 \left( \theta_{23} + \frac{a}{2|\delta m_{31}^2|} \xi \right) \right] \sin^2 \frac{\tilde{\Delta}_{31}}{2}$$

$$\tilde{\Delta}_{31} \approx \delta m_{31}^2 c_{13}^2 - \delta m_{21}^2 c_{12}^2$$

anti - neutrino (normal hierarchy)

$$\tilde{U}_{\mu 2} \approx \tilde{c}_{23}, \quad \tilde{U}_{\mu 3} \approx \tilde{s}_{23}$$

$$\tilde{P}(\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu) \approx 1 - \sin^2 \left[ 2 \left( \theta_{23} + \frac{a}{2|\delta m_{31}^2|} \xi \right) \right] \sin^2 \frac{\tilde{\Delta}_{32}}{2}$$

$$\tilde{\Delta}_{32} \approx \delta m_{31}^2 c_{13}^2 - \delta m_{21}^2 c_{12}^2$$

anti - neutrino (inverted hierarchy)

$$\tilde{U}_{\mu 2} \approx \cos(\tilde{\theta}_{12} + \tilde{\theta}_{23}), \quad \tilde{U}_{\mu 3} \approx 0$$

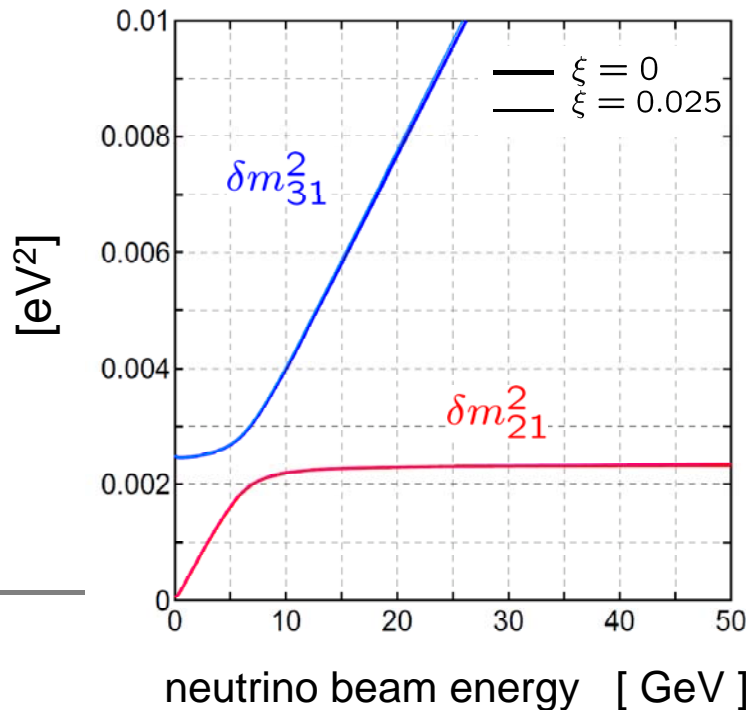
$$\tilde{P}(\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu) \approx 1 - \sin^2 \left[ 2 \left( \theta_{23} - \frac{a}{2|\delta m_{31}^2|} \xi \right) \right] \sin^2 \frac{\tilde{\Delta}_{21}}{2}$$

$$\tilde{\Delta}_{21} \approx -\delta m_{31}^2 c_{13}^2 + \delta m_{21}^2 c_{12}^2$$

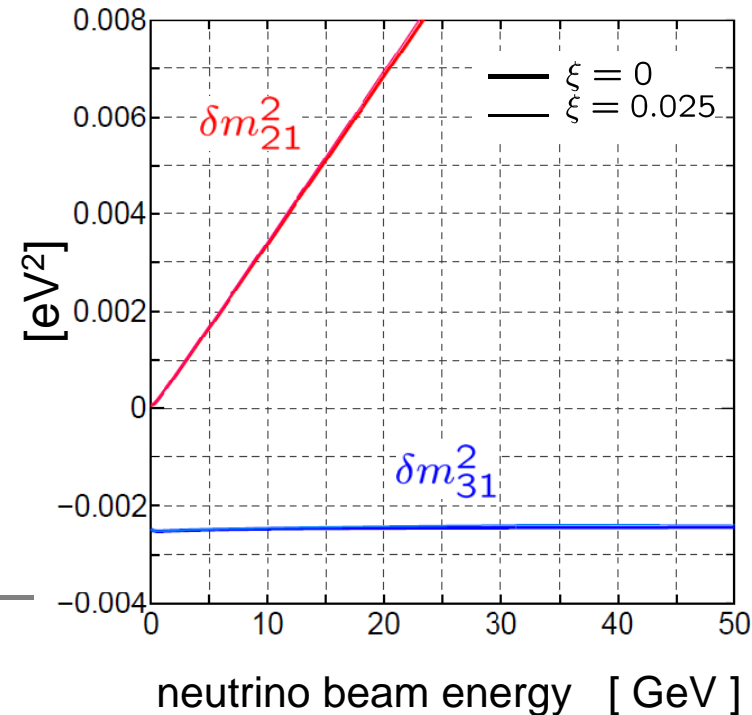
# The energy dependence of the effective mass squared differences

matter density = 4.6 g/cm <sup>3</sup>	$\sin^2(2\theta_{\text{atm}}) = 1.0$	$\delta m_{31}^2 = 2.5 \times 10^{-3} \text{ eV}^2$
CP phase $\delta = 0$	$\tan^2(\theta_{\text{sun}}) = 0.40$	$\delta m_{21}^2 = 8.0 \times 10^{-5} \text{ eV}^2$
	$\sin^2(2\theta_{\text{rc}}) = 0.16$	

*normal hierarchy*



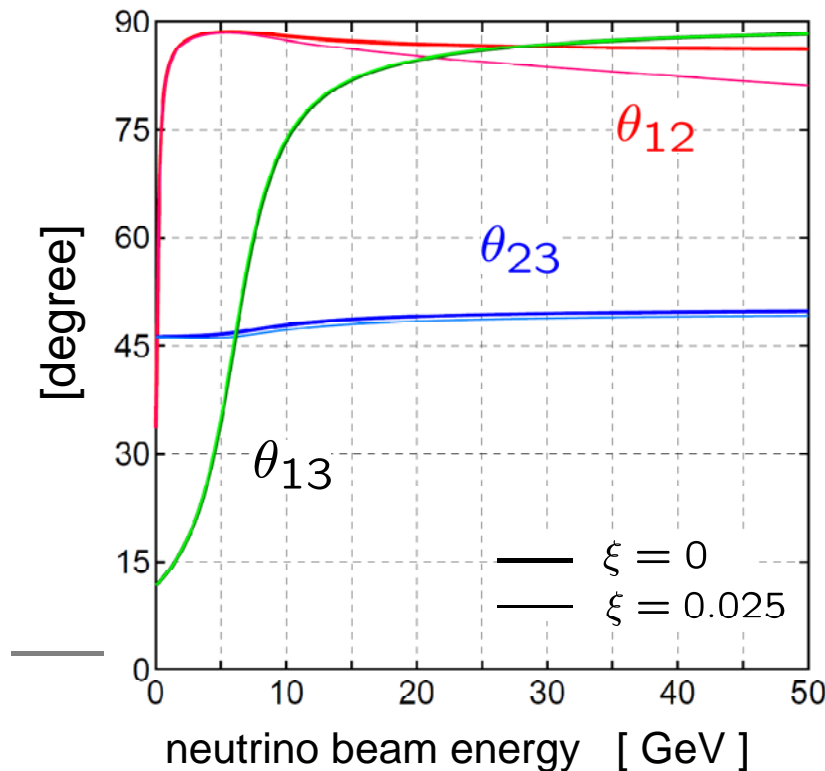
*inverted hierarchy*



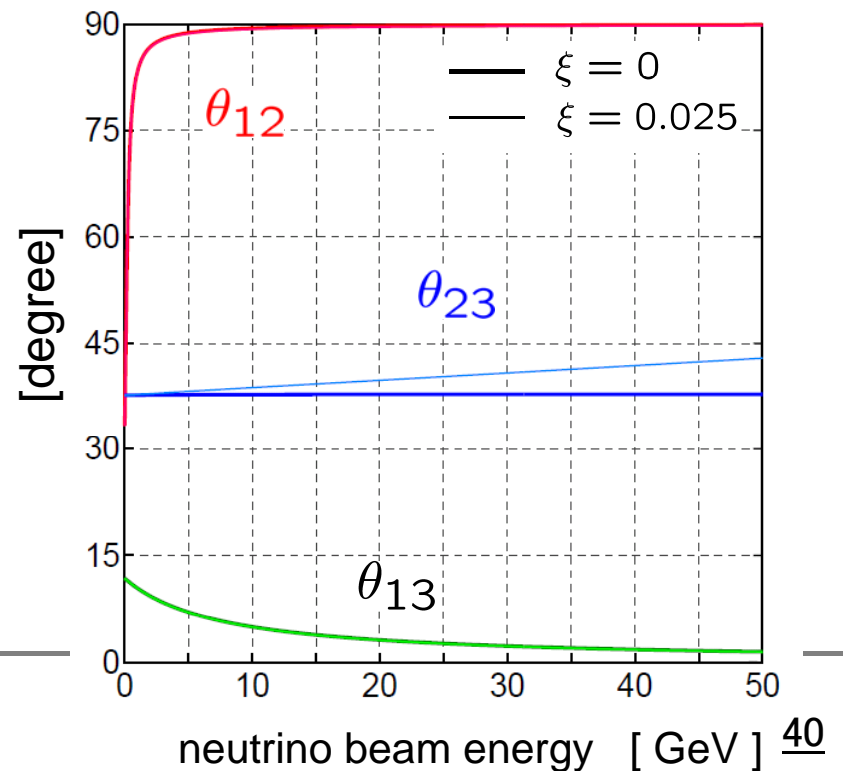
# The energy dependence of the effective mixing angles

matter density = 4.6 g/cm <sup>3</sup>	$\sin^2(2\theta_{\text{atm}}) = 1.0$	$\delta m_{31}^2 = 2.5 \times 10^{-3} \text{ eV}^2$
CP phase $\delta = 0$	$\tan^2(\theta_{\text{sol}}) = 0.40$	$\delta m_{21}^2 = 8.0 \times 10^{-5} \text{ eV}^2$
	$\sin^2(2\theta_{\text{rd}}) = 0.16$	

*normal hierarchy*



*inverted hierarchy*





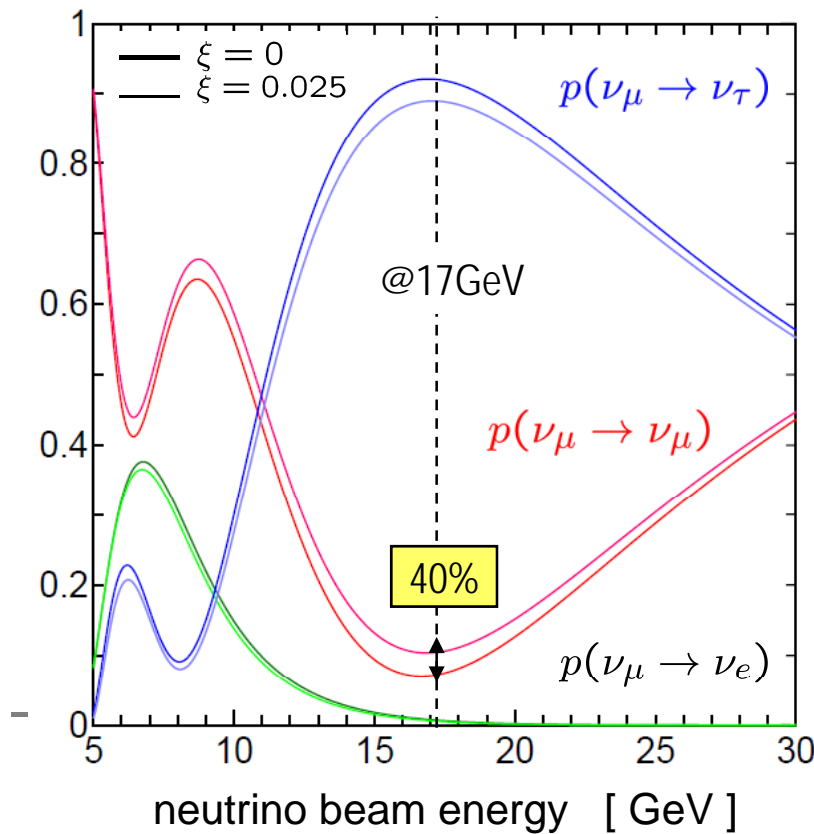
# The oscillation probabilities ( L=9,120km )

$$\sin(2\theta_{\text{atm}}) = \sin \theta_{23} \cos \theta_{13} \approx \sin \theta_{23}$$

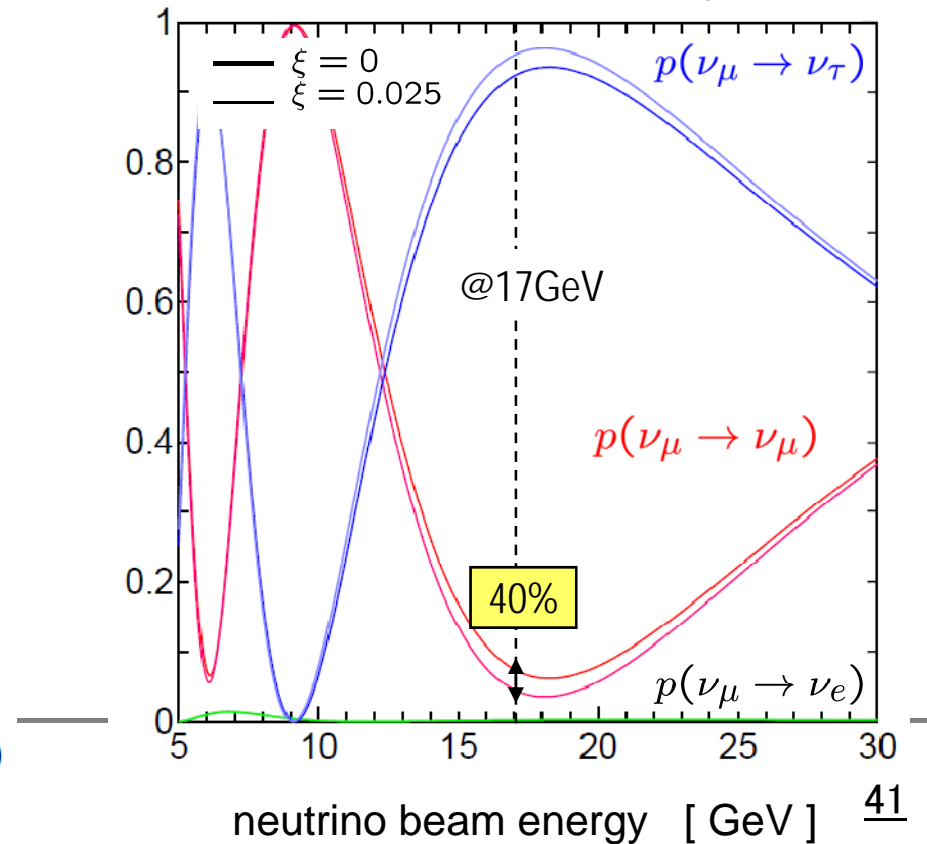
matter density = 4.6 g/cm <sup>3</sup>	$\sin^2(2\theta_{\text{atm}}) = 0.92$	$\delta m_{31}^2 = 2.5 \times 10^{-3} \text{ eV}^2$
CP phase $\delta = 0$	$\tan^2(\theta_{\text{sun}}) = 0.40$	$\delta m_{21}^2 = 8.0 \times 10^{-5} \text{ eV}^2$
	$\sin^2(2\theta_{\text{rx}}) = 0.16$	

$$\theta_{\text{atm}} < \pi/4$$

*normal hierarchy*



*inverted hierarchy*

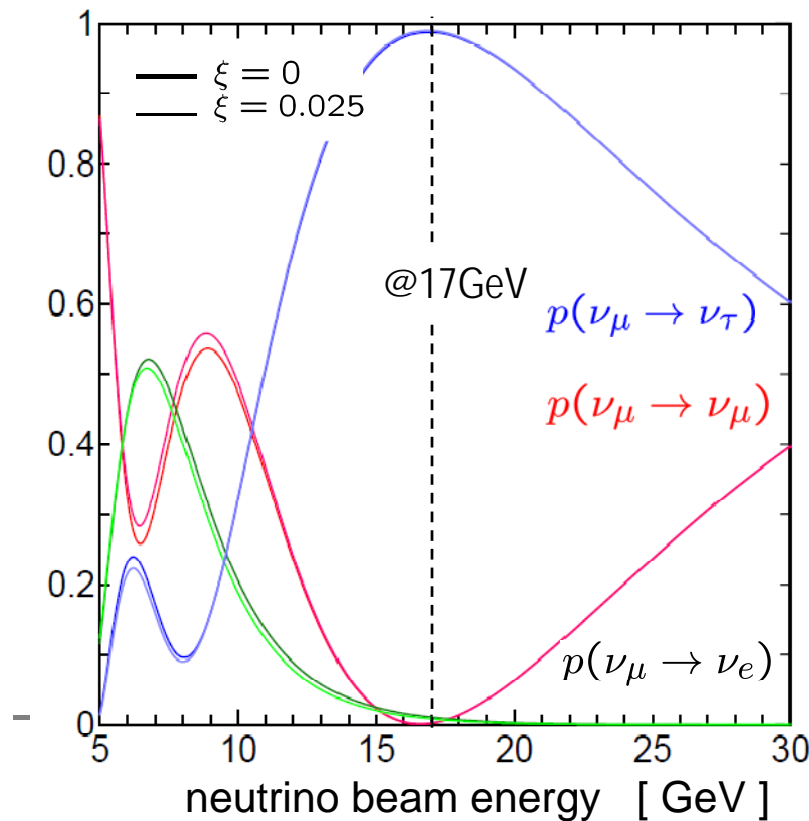


# The oscillation probabilities ( L=9,120km )

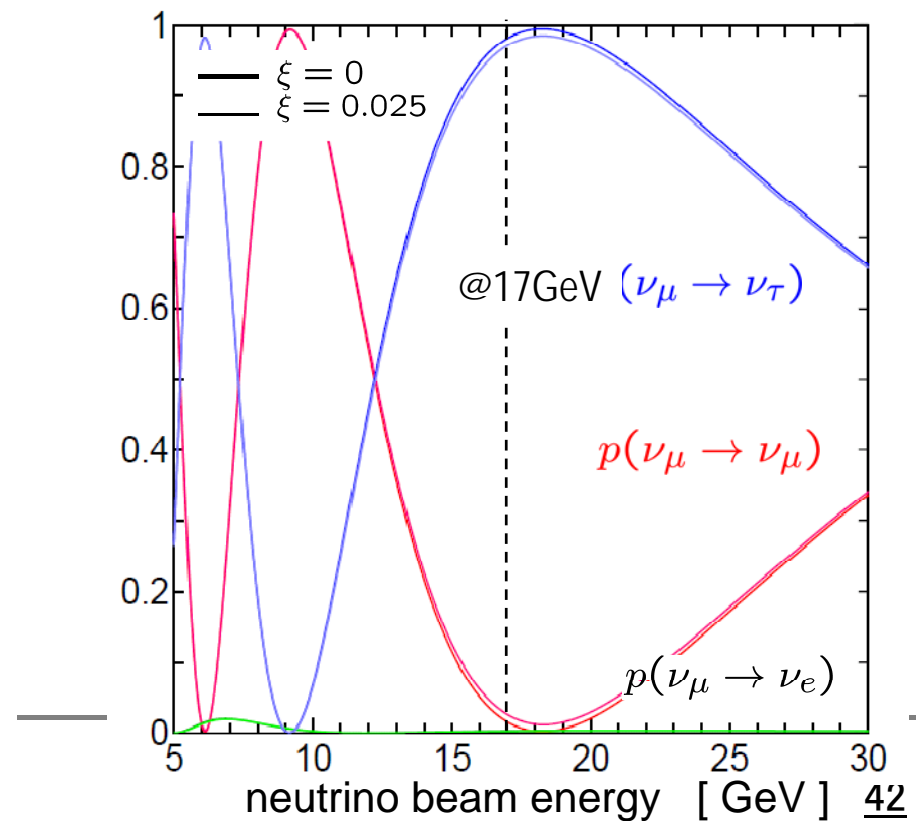
$$\sin(2\theta_{\text{atm}}) = \sin \theta_{23} \cos \theta_{13} \approx \sin \theta_{23}$$

matter density = 4.6 g/cm <sup>3</sup>	$\sin^2(2\theta_{\text{atm}}) = 1$	$\delta m_{31}^2 = 2.5 \times 10^{-3} \text{ eV}^2$
CP phase $\delta = 0$	$\tan^2(\theta_{\text{sun}}) = 0.40$	$\delta m_{21}^2 = 8.0 \times 10^{-5} \text{ eV}^2$
	$\sin^2(2\theta_{\text{rel}}) = 0.16$	

*normal hierarchy*

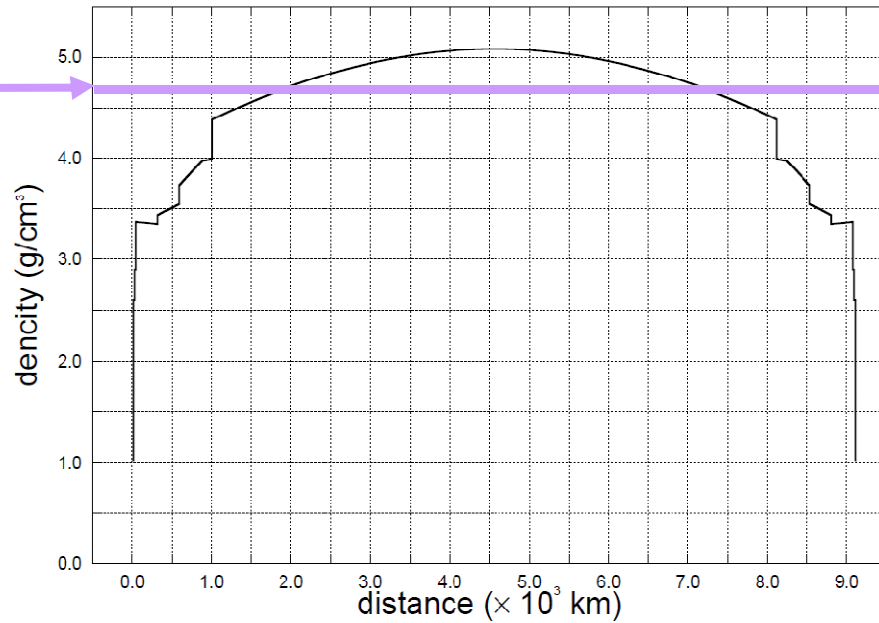


*inverted hierarchy*



# Matter Profile

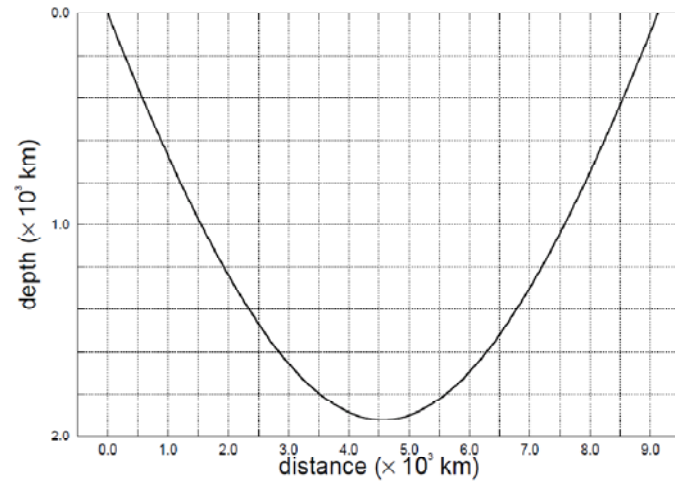
density (FNAL  $\leftrightarrow$  HK)



average matter density

$$\rho = 4.6 \text{ g/cm}^3$$

depth (FNAL  $\leftrightarrow$  HK)



+

density profile

