

June 24th, 2004

@ ICRR

ニュートリノ研究会

K2K実験の最新結果

*M. Yokoyama (Kyoto University)
for K2K collaboration*

Introduction

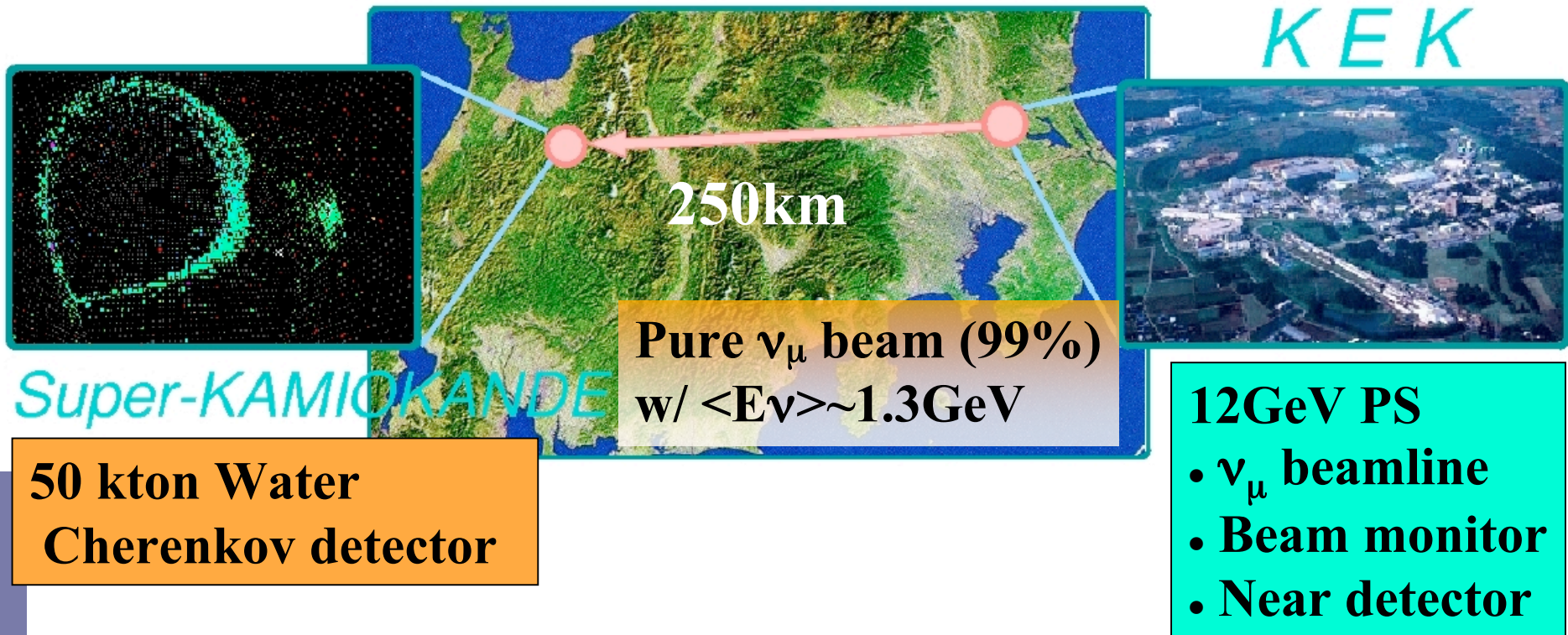


1. Introduction

K2K experiment since 1999

First accelerator-based long baseline (**250km**) neutrino experiment.

Search for ν_μ disappearance and ν_e appearance



Brief history of K2K

- 1995
 - Proposed to study neutrino oscillation for atmospheric neutrinos anomaly.
- 1999
 - Started taking data.
- 2000
 - Detected the less number of neutrinos than the expectation at a distance of 250 km. **Disfavored null oscillation at the 2σ level.**
- 2002
 - Observed indications of neutrino oscillation. **The probability of null oscillation is less than 1%.**
- 2004
 - **This result!**

K2K Collaboration



JAPAN: High Energy Accelerator Research Organization (KEK) / Institute for Cosmic Ray Research (ICRR), Univ. of Tokyo / Kobe University / Kyoto University / Niigata University / Okayama University / Tokyo University of Science / Tohoku University

KOREA: Chonnam National University / Dongshin University / Korea University / Seoul National University

U.S.A.: Boston University / University of California, Irvine / University of Hawaii, Manoa / Massachusetts Institute of Technology / State University of New York at Stony Brook / University of Washington at Seattle

POLAND: Warsaw University / Solton Institute

Since 2002

JAPAN: Hiroshima University / Osaka University

CANADA: TRIUMF / University of British Columbia

Italy: Rome **France:** Saclay **Spain:** Barcelona / Valencia **Switzerland:** Geneva

RUSSIA: INR-Moscow

Overview



2. K2K experiment overview

~1 event/2days

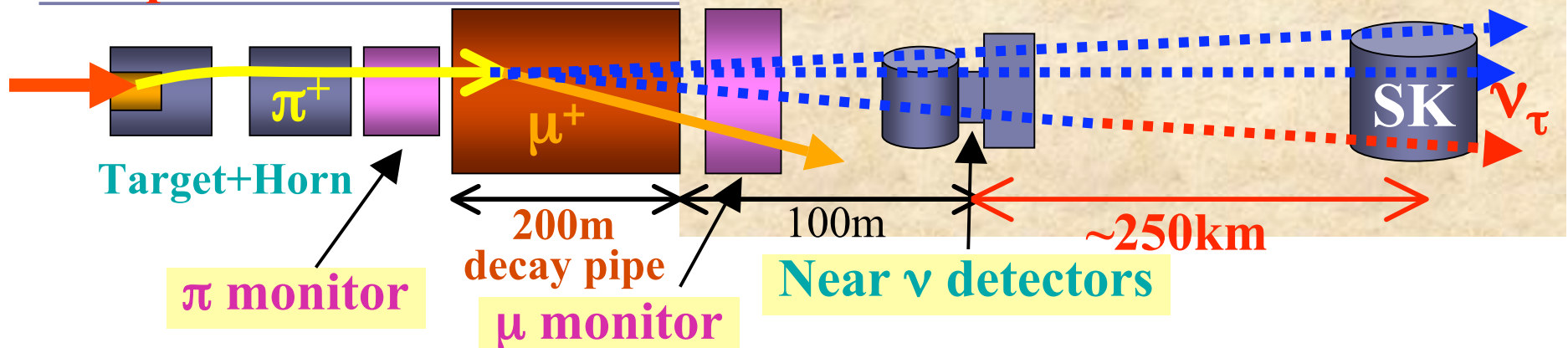
$\sim 10^{11} \nu_\mu / 2.2\text{sec}$

$\sim 10^6 \nu_\mu / 2.2\text{sec}$

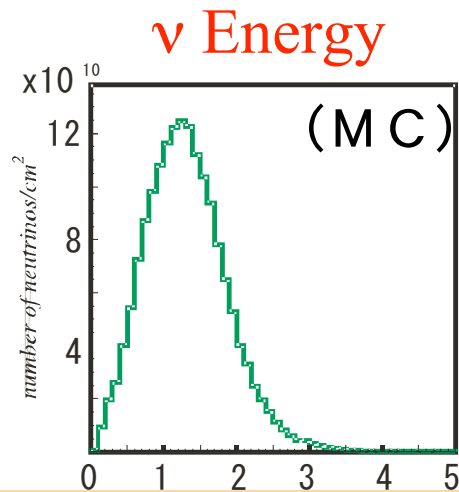
(/10m×10m) ν_μ

(/40m×40m)

12GeV protons



$$\text{prob.} = \sin^2 2\theta \cdot \sin^2 \left(\frac{1.27 \Delta m^2 L}{E_\nu} \right)$$

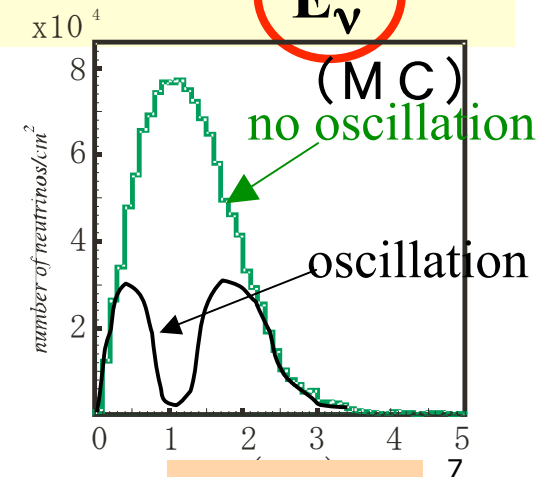


Near detectors at KEK

250km

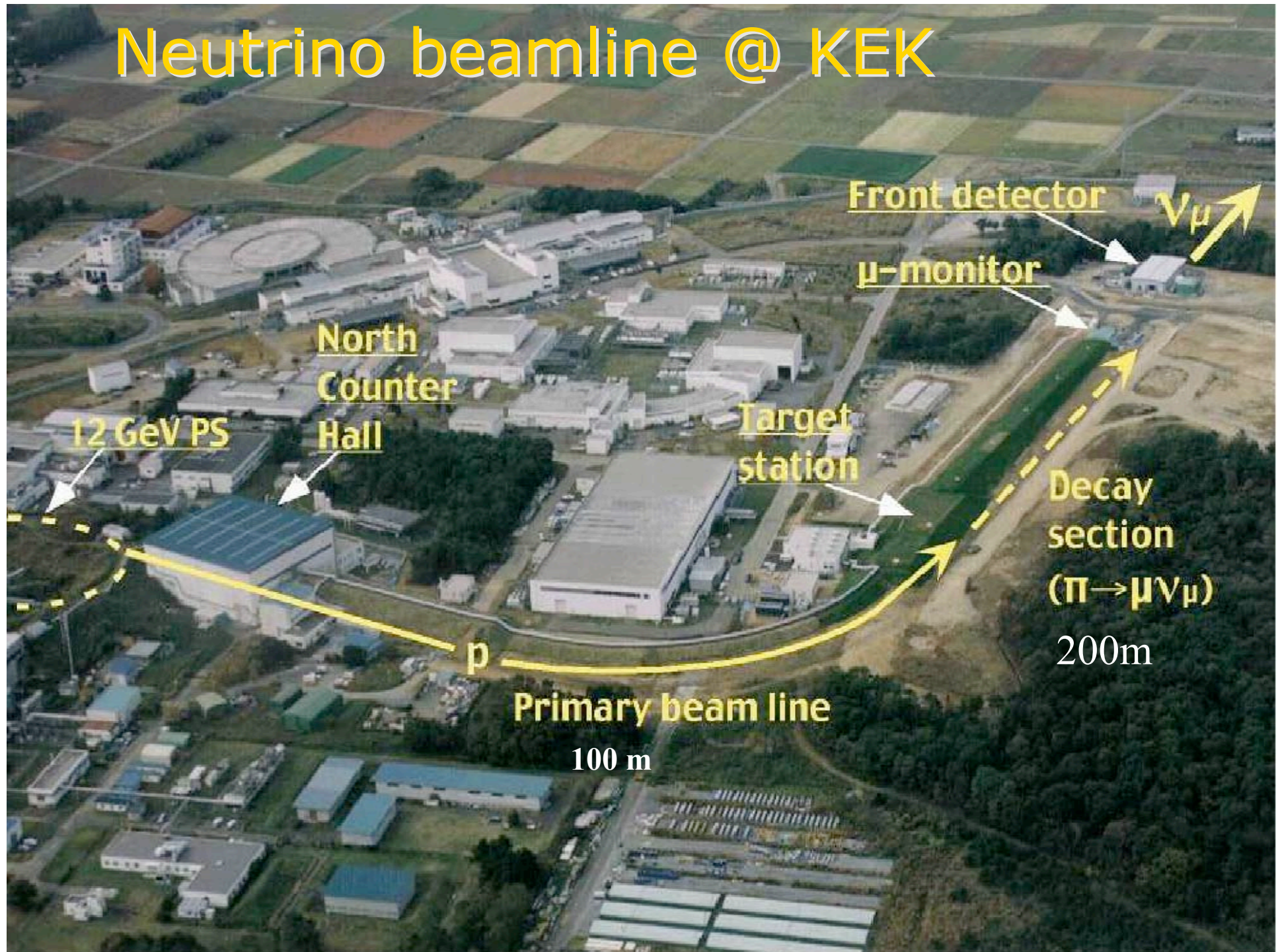
Extrapolation

π monitor + simulation



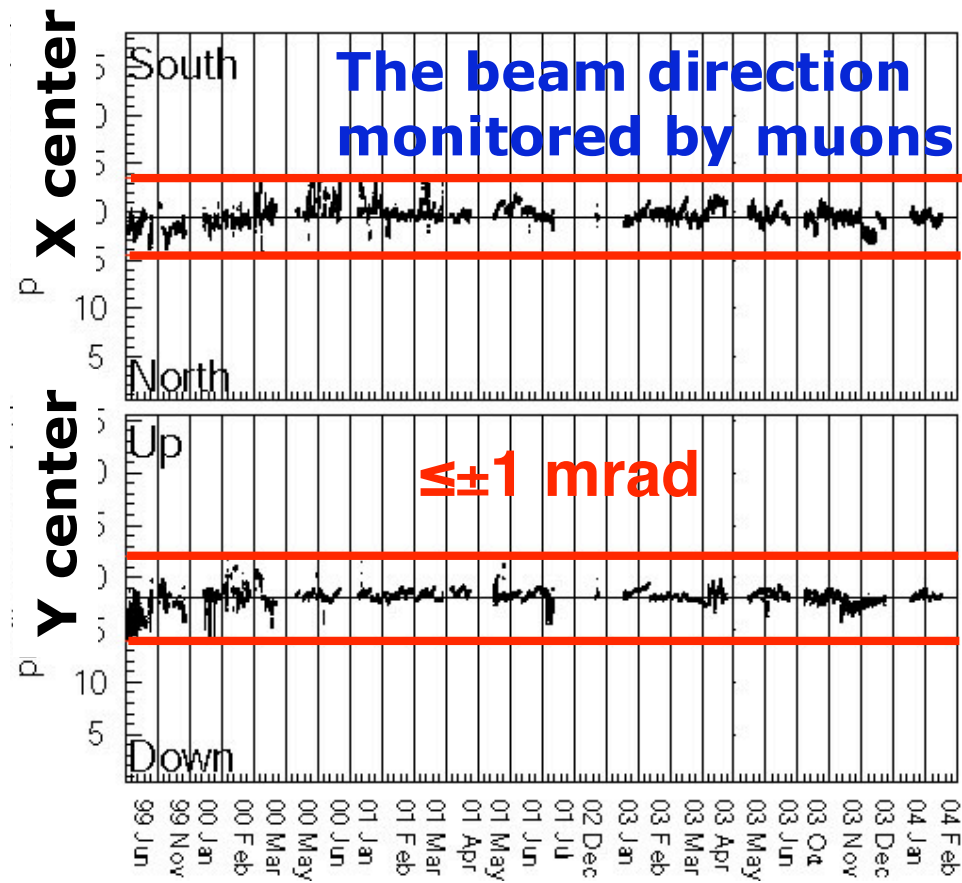
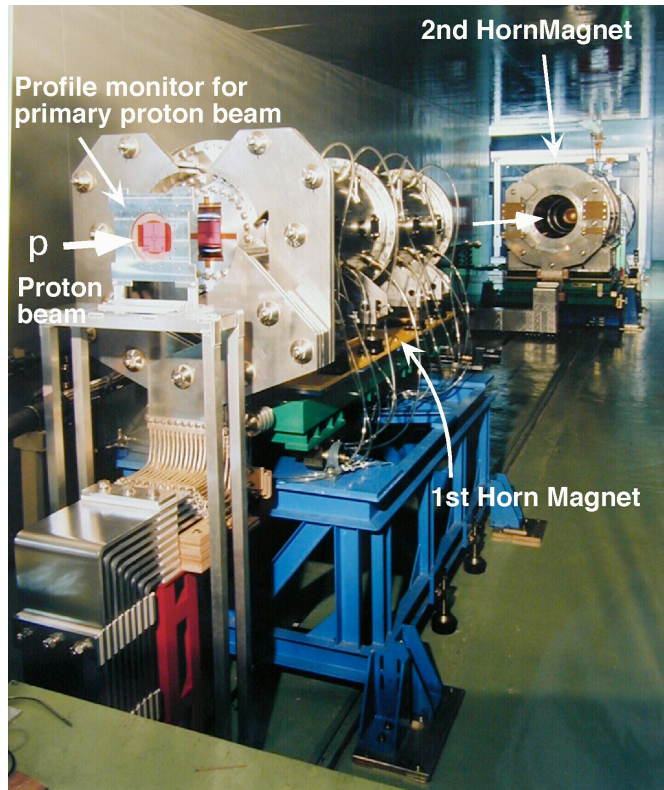
Super-K

Neutrino beamline @ KEK



Neutrino beam and the directional control

- $\sim 1\text{GeV}$ neutrino beam by a dual horn system with 250kA.

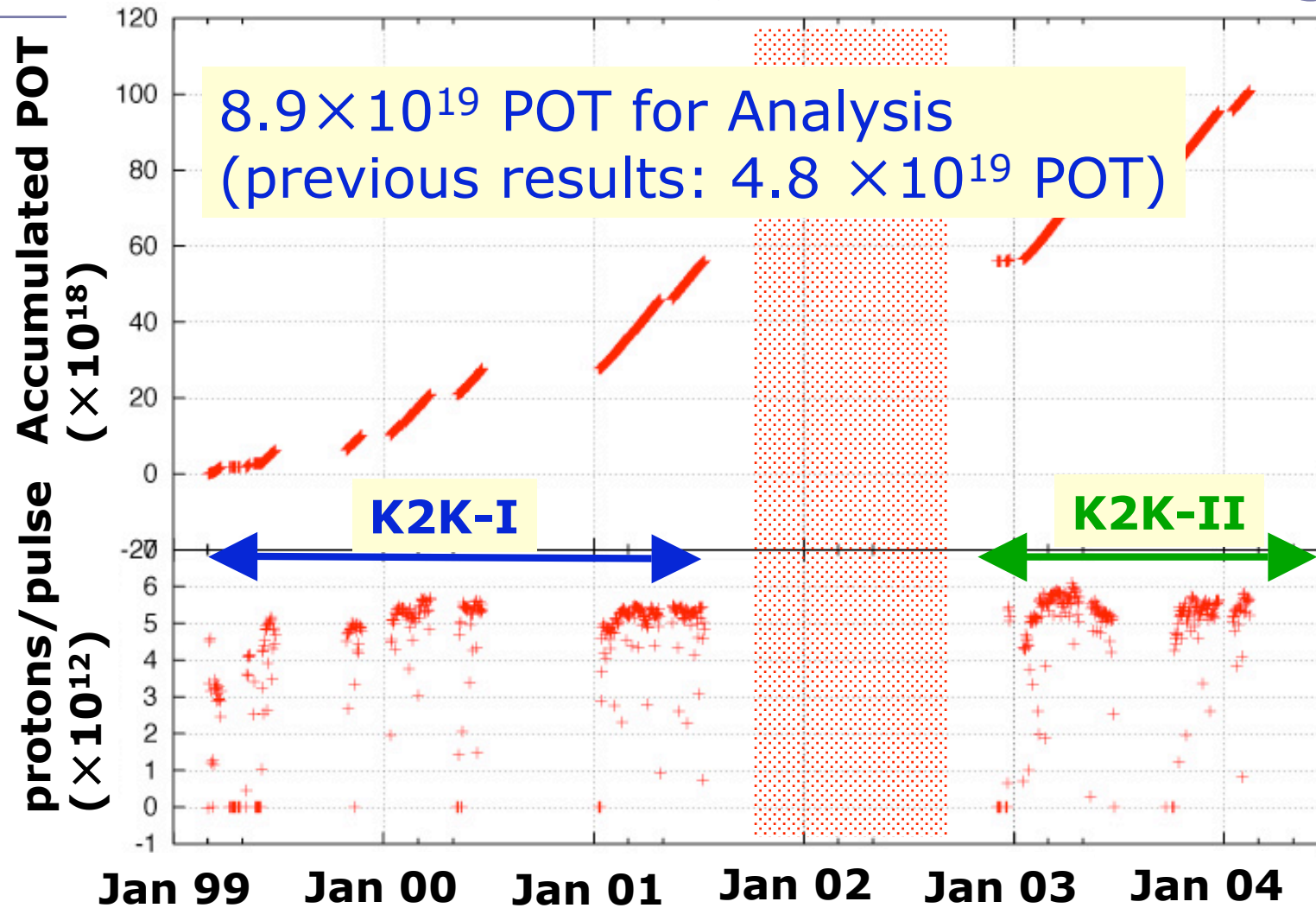


99 Jun

~ 5 years

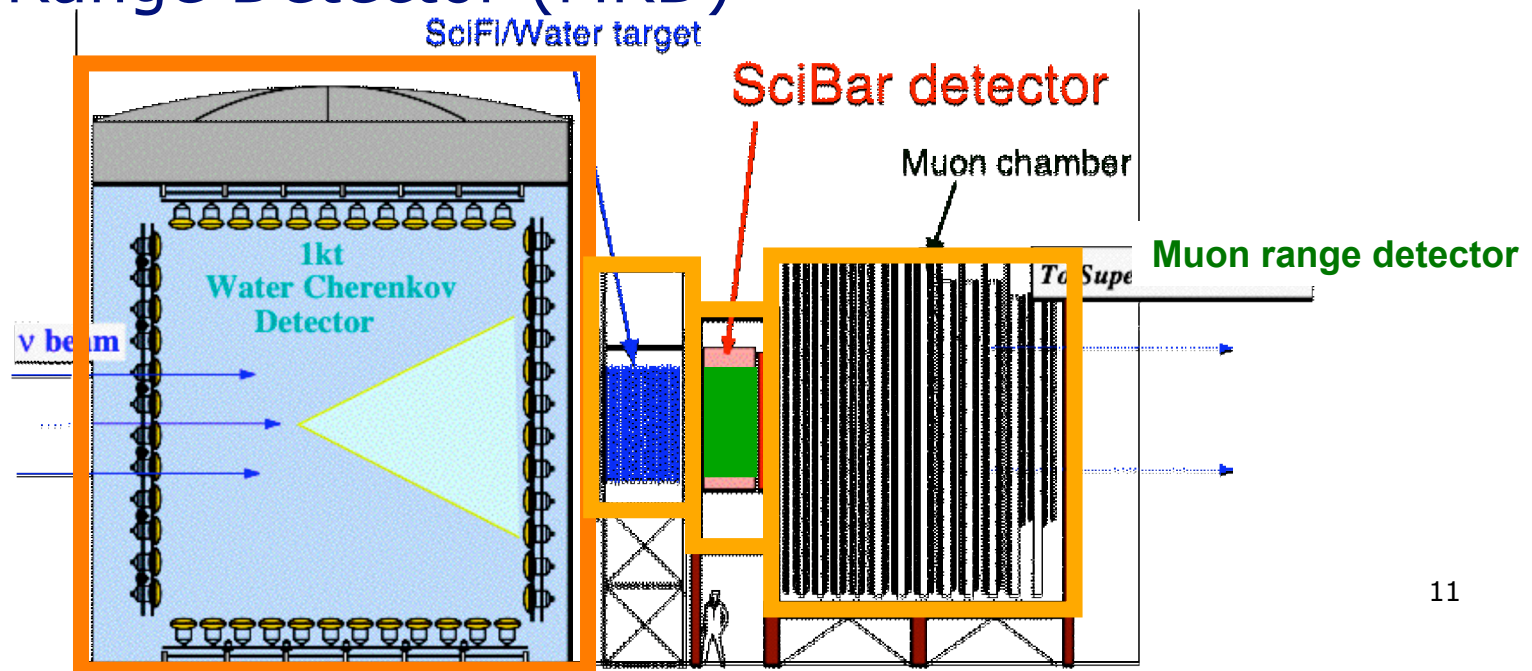
04 Feb

Accumulated POT (Protons On Target)



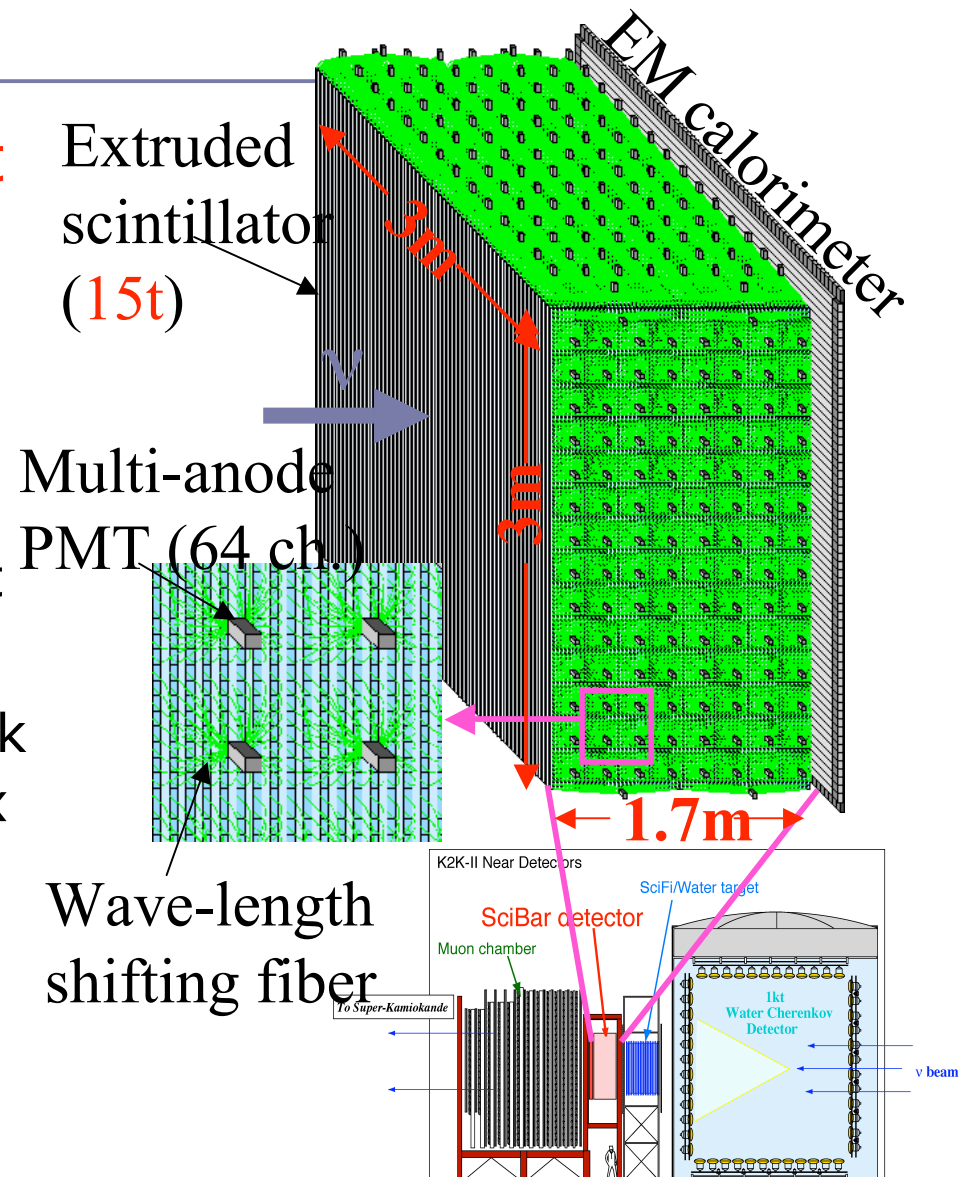
Near detector system at KEK

- ❑ 1KT Water Cherenkov Detector (1KT)
- ❑ Scintillating-fiber/Water sandwich Detector (SciFi)
- ❑ Lead Glass calorimeter (LG) before 2002 (K2K-I)
- ❑ **Scintillator Bar Detector (SciBar)** from 2003 (K2K-II)
- ❑ Muon Range Detector (MRD)



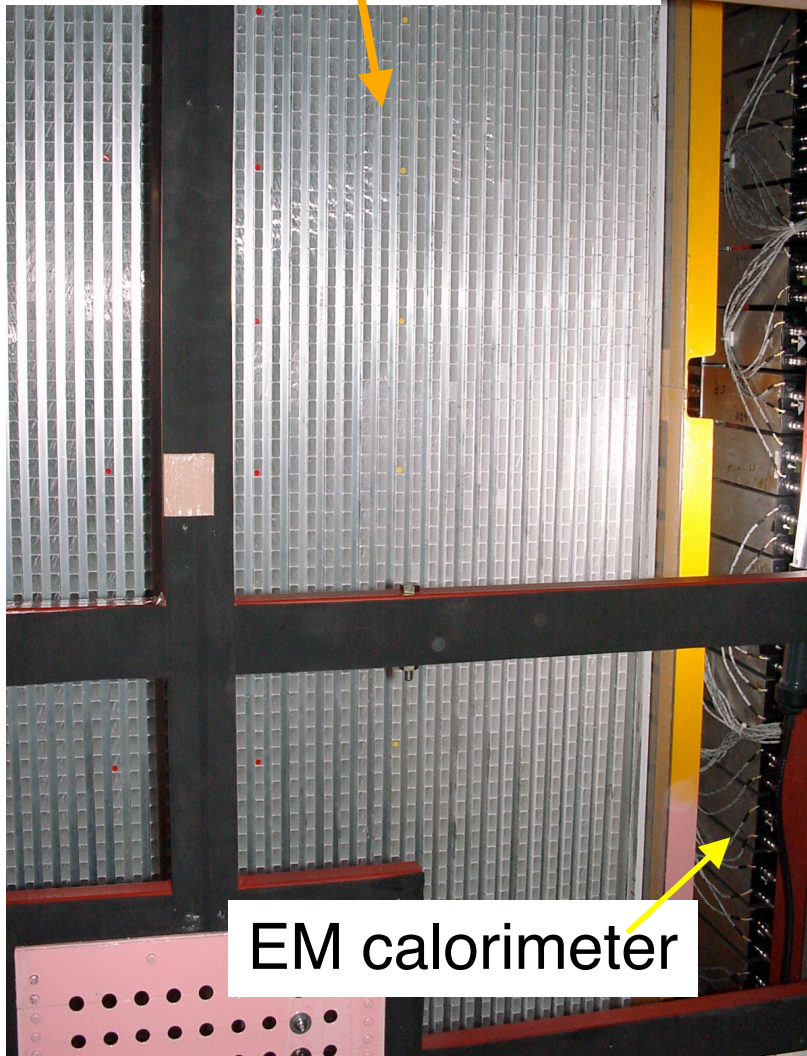
SciBar Detector

- Full-active, fine-segment detector made of Scintillator Bars
 - $2.5 \times 1.3 \times 300 \text{ cm}^3$ cell
 - ~ 15000 channels (dead-ch: < 10)
 - WLS fiber+MAPMT readout
 - Detect short ($\sim 10 \text{ cm}$) track
 - p/π separation using dE/dx
- Precise ν spectrum measurement
- ν interaction study

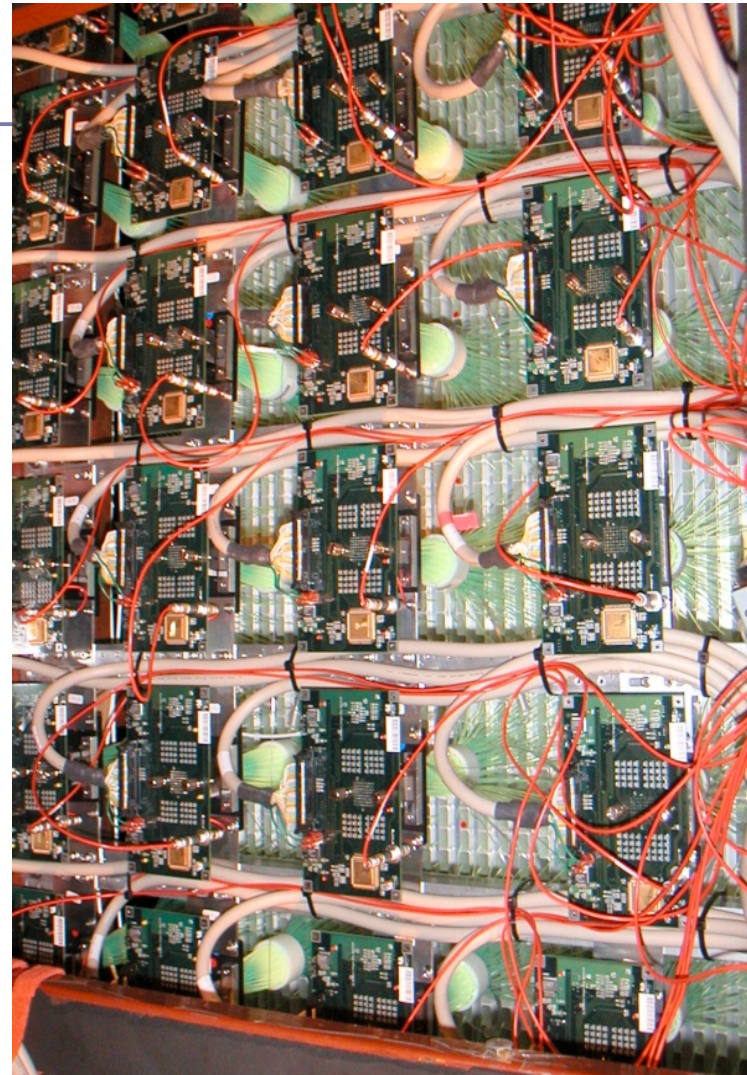


Detector Photos

Scintillators (64layers)



EM calorimeter



Fibers and front-end elec₁₃

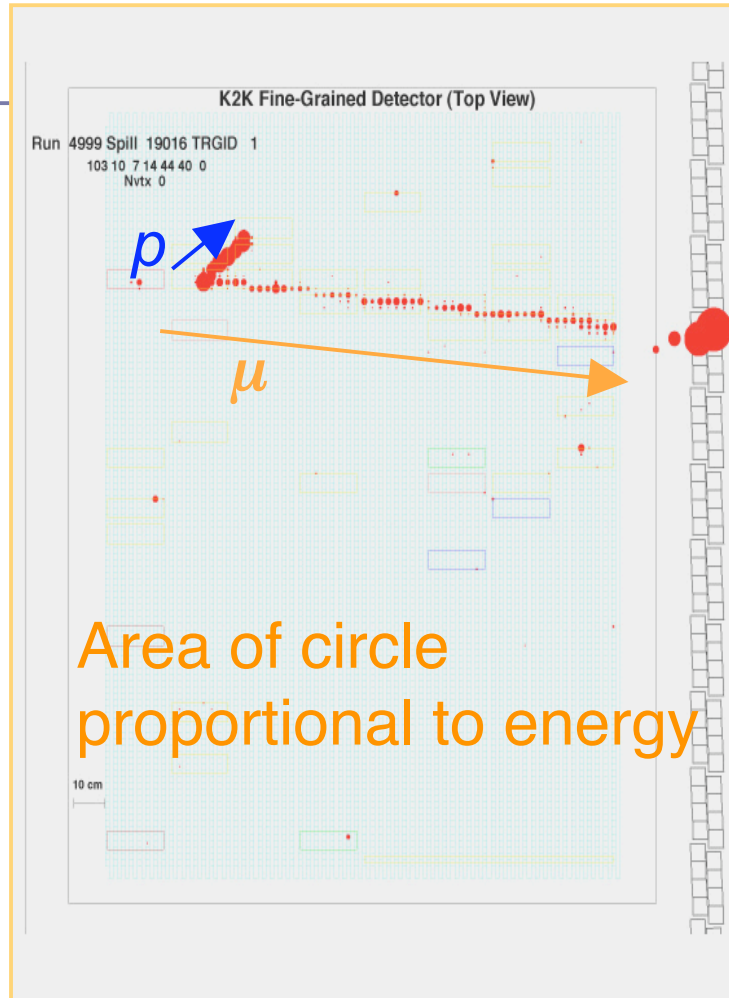
Just Completed!



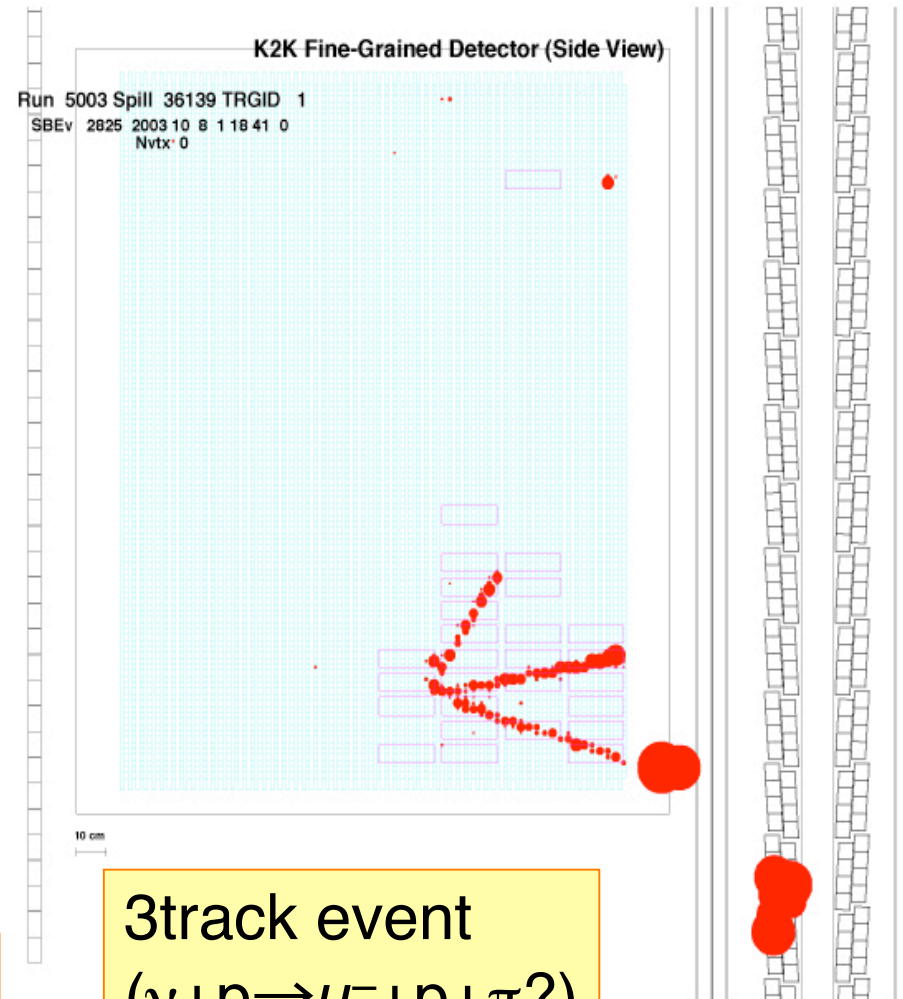
Aug. 22, 2003

Event Display

Area of red circles
Proportional to energy



CCQE ($\nu+n \rightarrow \mu^-+p$) candidate



3track event
($\nu+p \rightarrow \mu^-+p+\pi?$)

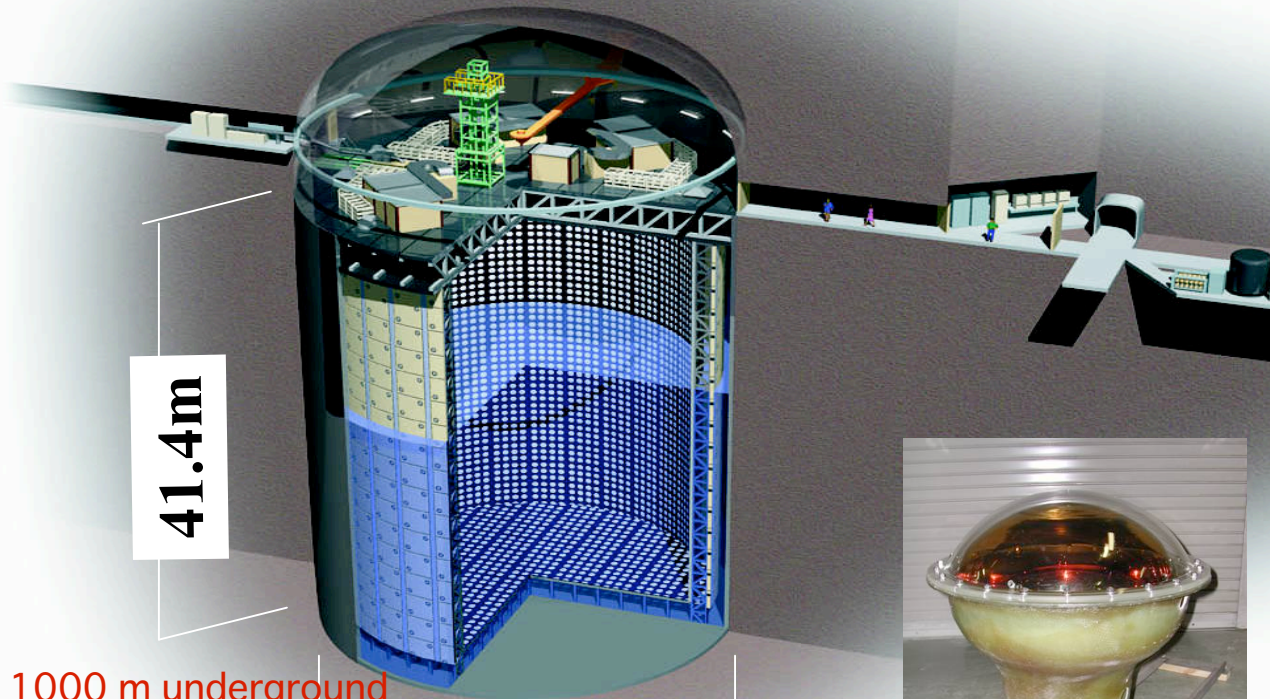
Far Detector -Super-Kamiokande-

50,000 ton water Cherenkov detector (22.5 kton fiducial volume)

Optically separated **INNER** and **OUTER** detector

SK is back!

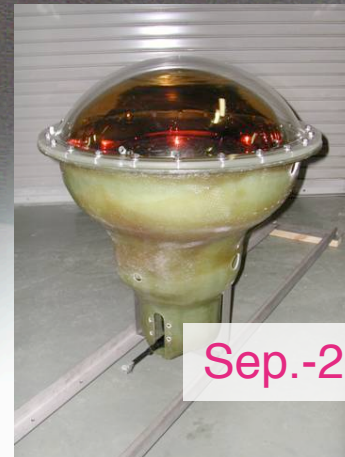
w/ half PMT density



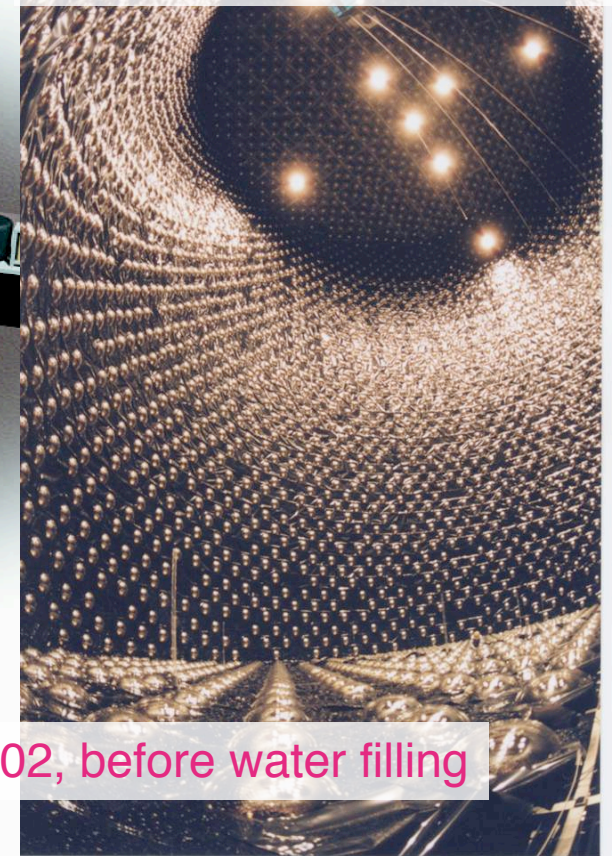
41.4m

40m

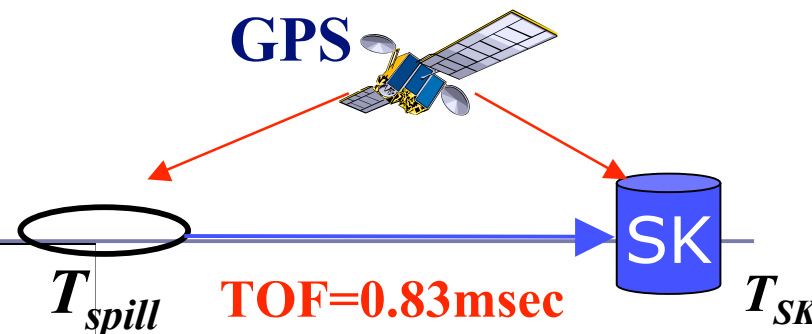
- 1000 m underground
- 50,000 ton (22,500 ton fid.)
- 11,146 20 inch PMTs
- 1,885 anti-counter PMTs



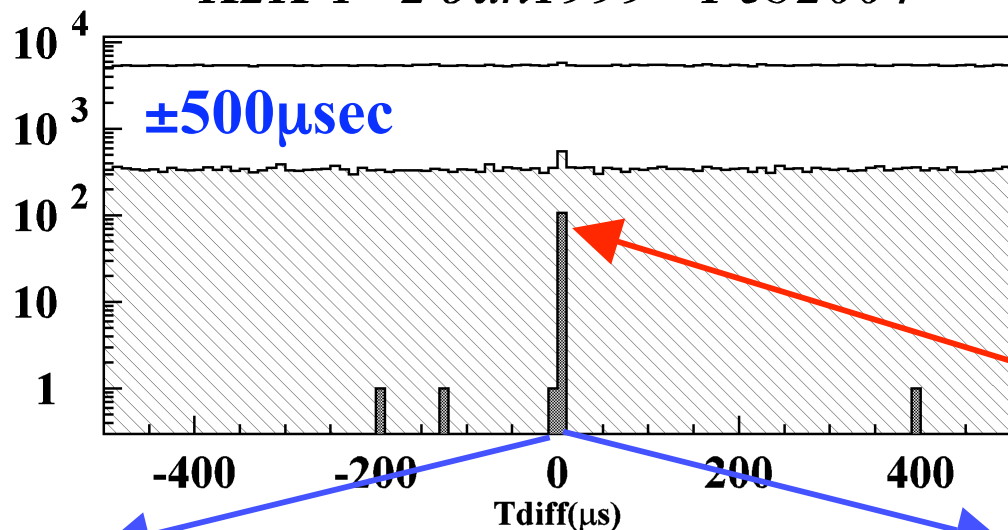
Sep.-2002, before water filling



SK Events



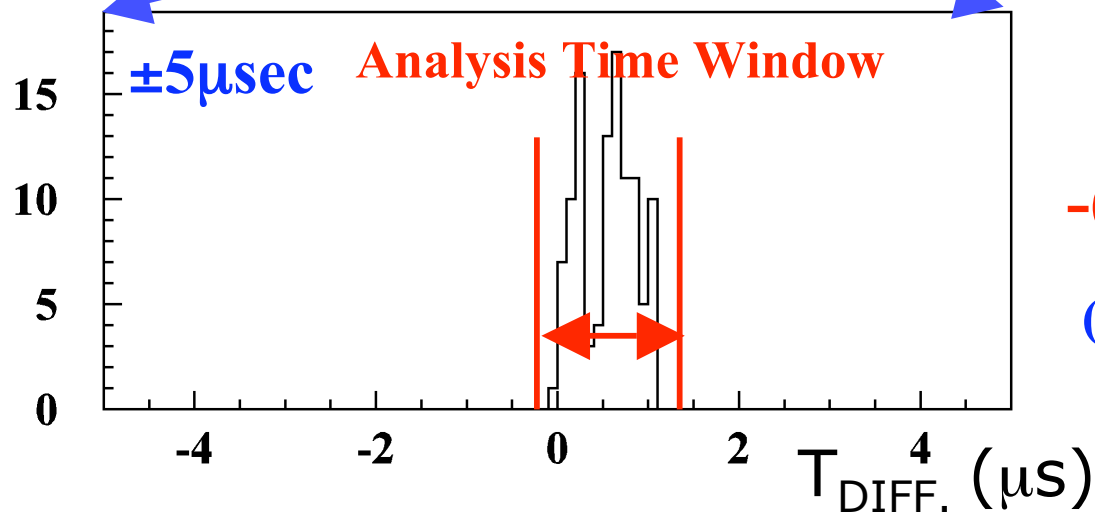
K2K-1+2 Jun1999 - Feb2004



← Decay electron cut.

← $\geq 20\text{MeV}$ Deposited Energy

**No Activity in Outer Detector
Event Vertex in Fiducial Volume
More than 30MeV Deposited Energy**



108 events (K2K-1:56)

$-0.2 < T_{SK} - T_{spill} - \text{TOF} < 1.3\mu\text{sec}$

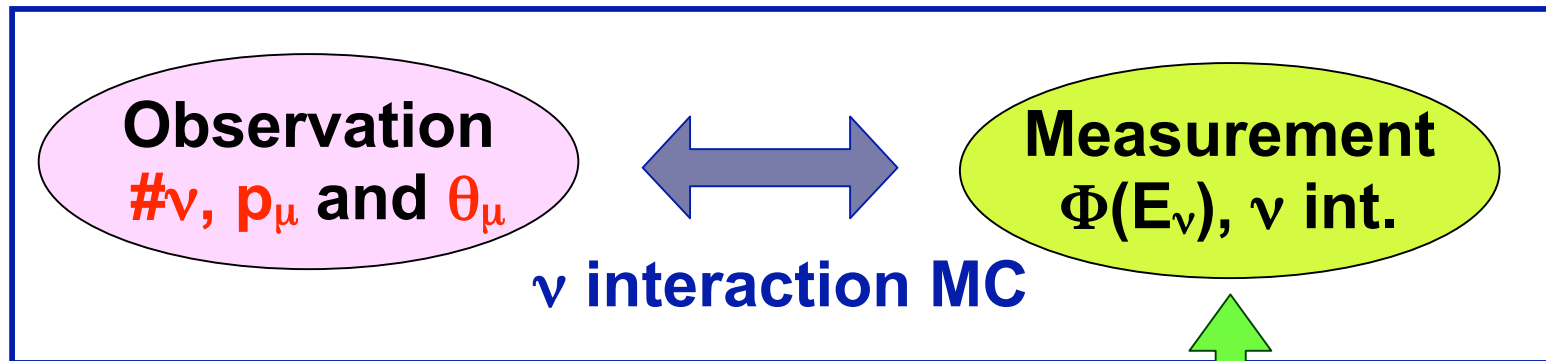
**(BG: 1.6 events within $\pm 500\mu\text{s}$
 2.4×10^{-3} events in $1.5\mu\text{s}$)**

Near Detector Measurements



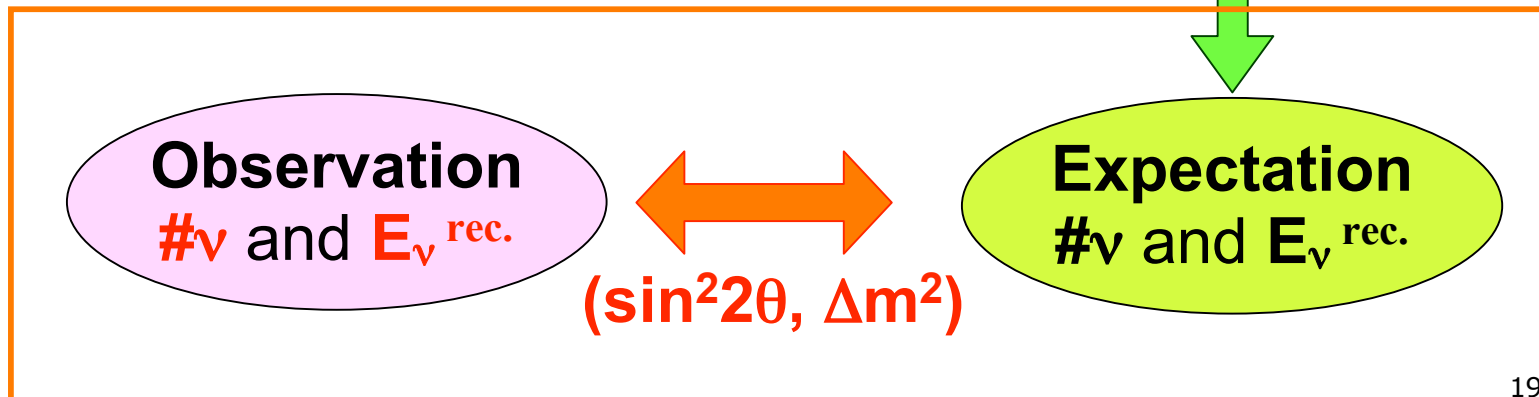
3. Analysis Overview

KEK



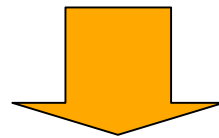
Far/Near Ratio
(beam MC with π mon.)

SK



4. Near Detector measurements

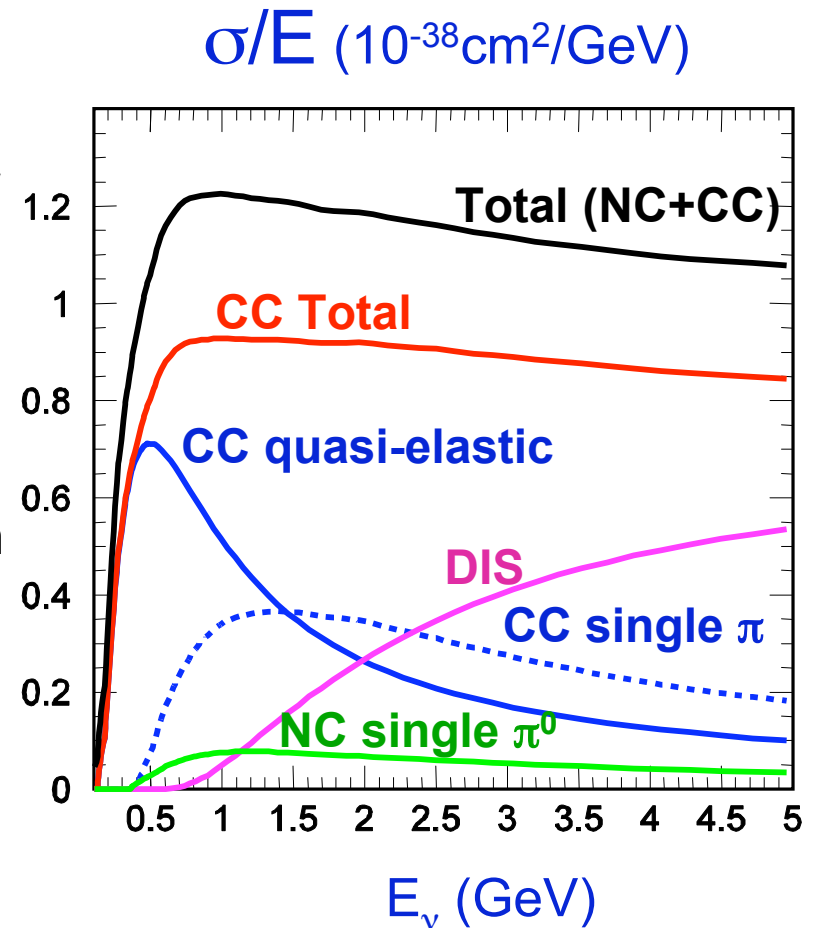
- Event rate measurement (# of ν int.)
 - Measurement w/ 1KT
 - Cross-checked by other detectors
- Spectrum shape measurement
 - 1KT, SciFi, SciBar (p_μ, θ_μ)
 - Measure spectrum and nQE/QE (ν interaction model)



Predict **number of event** and **spectrum shape** at SK

NEUT: K2K Neutrino interaction MC

- **CC quasi elastic (CCQE)**
 - Smith and Moniz with $M_A=1.1\text{GeV}$
 - **CC (resonance) single π (CC-1 π)**
 - Rein and Sehgal's with $M_A=1.1\text{GeV}$
 - **DIS**
 - GRV94 + JETSET with Bodek and Yang correction.
 - **CC coherent π**
 - Rein&Sehgal with the cross section rescale by J. Marteau
 - **NC**
- + Nuclear Effects**



4.1 Event rate measurement @1KT

- The same detector technology as Super-K.
- Sensitive to low energy neutrinos.

$$N_{SK}^{exp} = N_{KT}^{obs} \cdot \frac{\int \Phi_{SK}(E_\nu) \sigma(E_\nu) dE_\nu}{\int \Phi_{KT}(E_\nu) \sigma(E_\nu) dE_\nu} \cdot \frac{M_{SK}}{M_{KT}} \cdot \frac{\epsilon_{SK}}{\epsilon_{KT}}$$

≡ Far/Near Ratio (by MC) $\sim 1 \times 10^{-6}$

M: Fiducial mass $M_{SK}=22,500\text{ton}$, $M_{KT}=25\text{ton}$

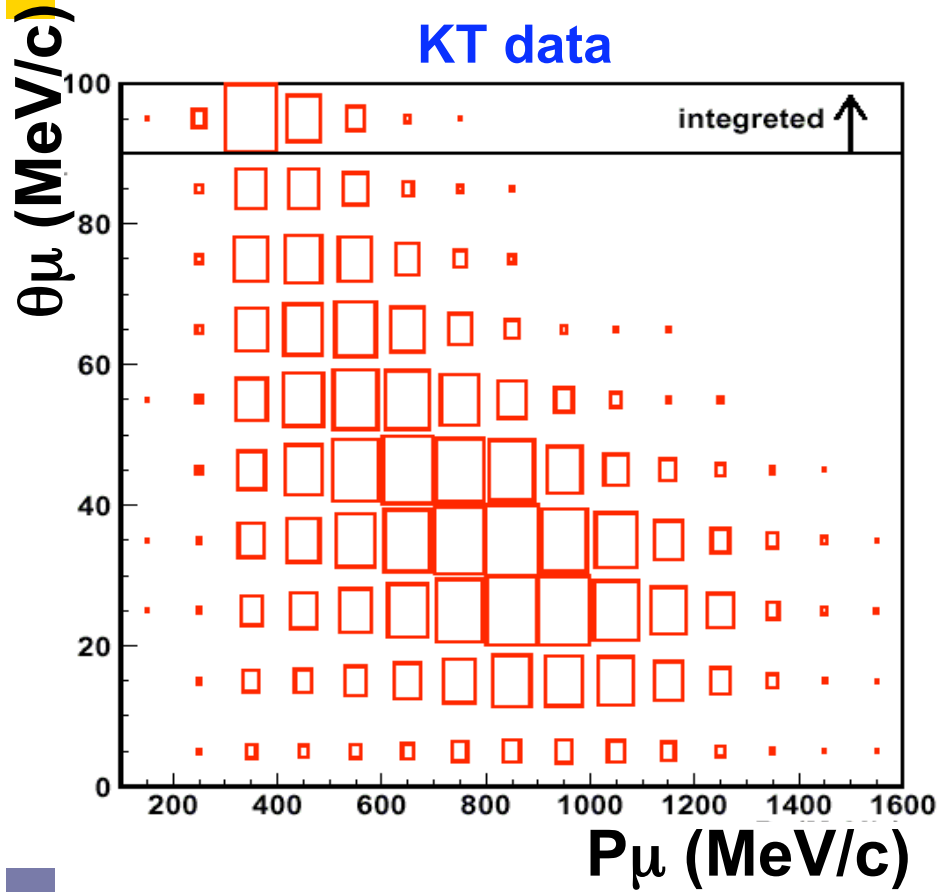
ε: efficiency $\epsilon_{SK-I(II)}=77.0(78.2)\%$, $\epsilon_{KT}=74.5\%$

$$N_{SK}^{exp} = 150.9^{+11.6}_{-10.0}$$



$$N_{SK}^{obs} = 108$$

4.2 Near Detector Spectrum Measurements

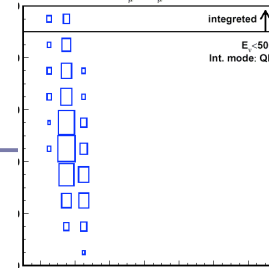


E_ν

QE (MC)

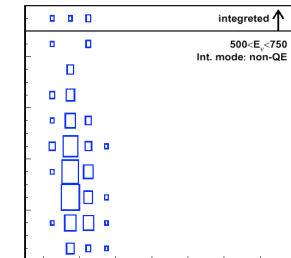
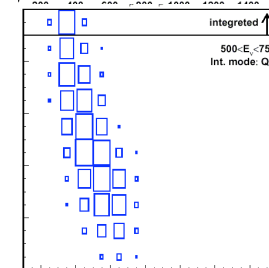
nQE(MC)

0-0.5 GeV

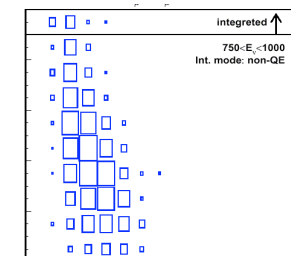
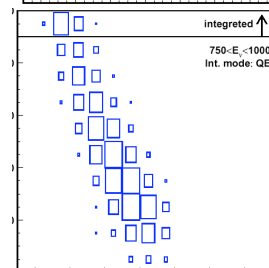


MC templates

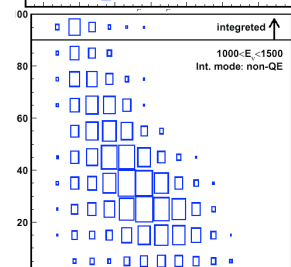
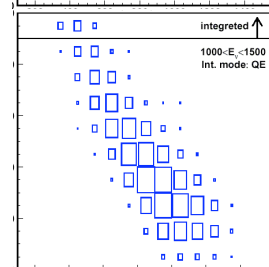
0.5-0.75 GeV



0.75-1.0 GeV



1.0-1.5 GeV



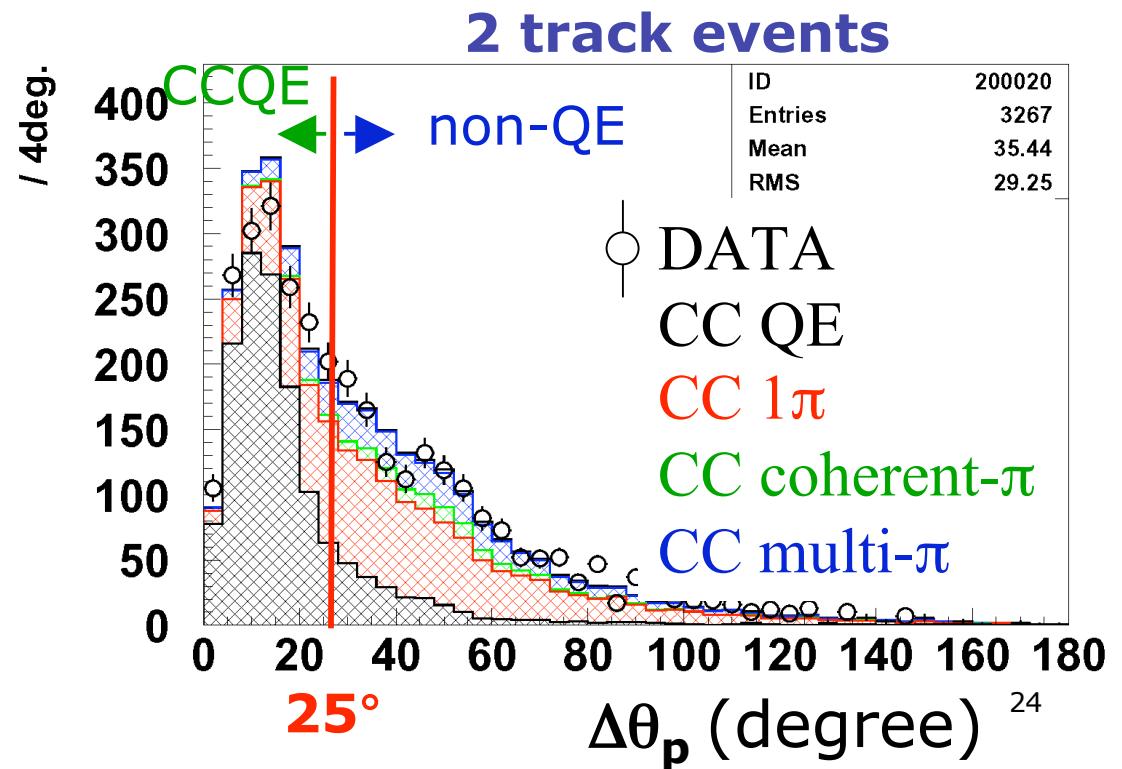
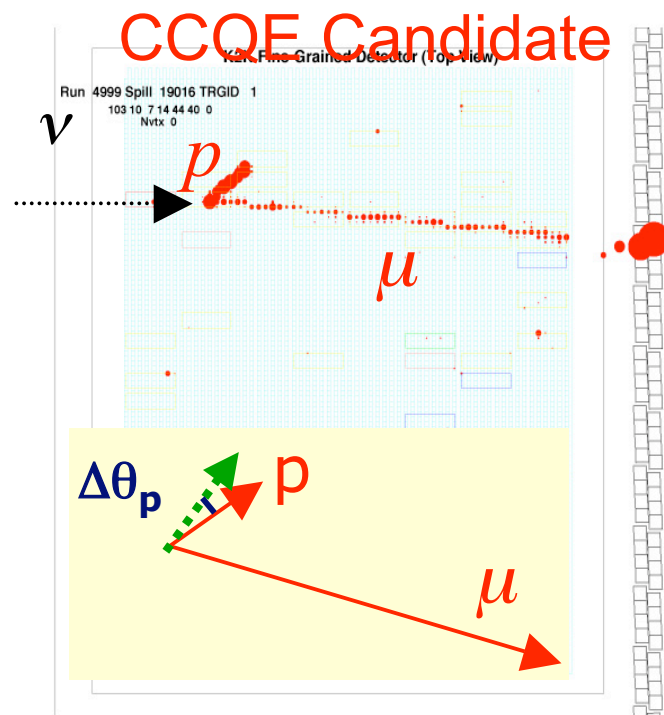
- ν flux $\Phi_{KEK}(E_\nu)$ (8 bins)
- ν interaction (nQE/QE)

•
•

•
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Measurement with SciBar

- Full Active Fine-Grained detector (target: CH).
 - Sensitive to a low momentum track.
 - Identify CCQE events and other interactions (non-QE) separately.



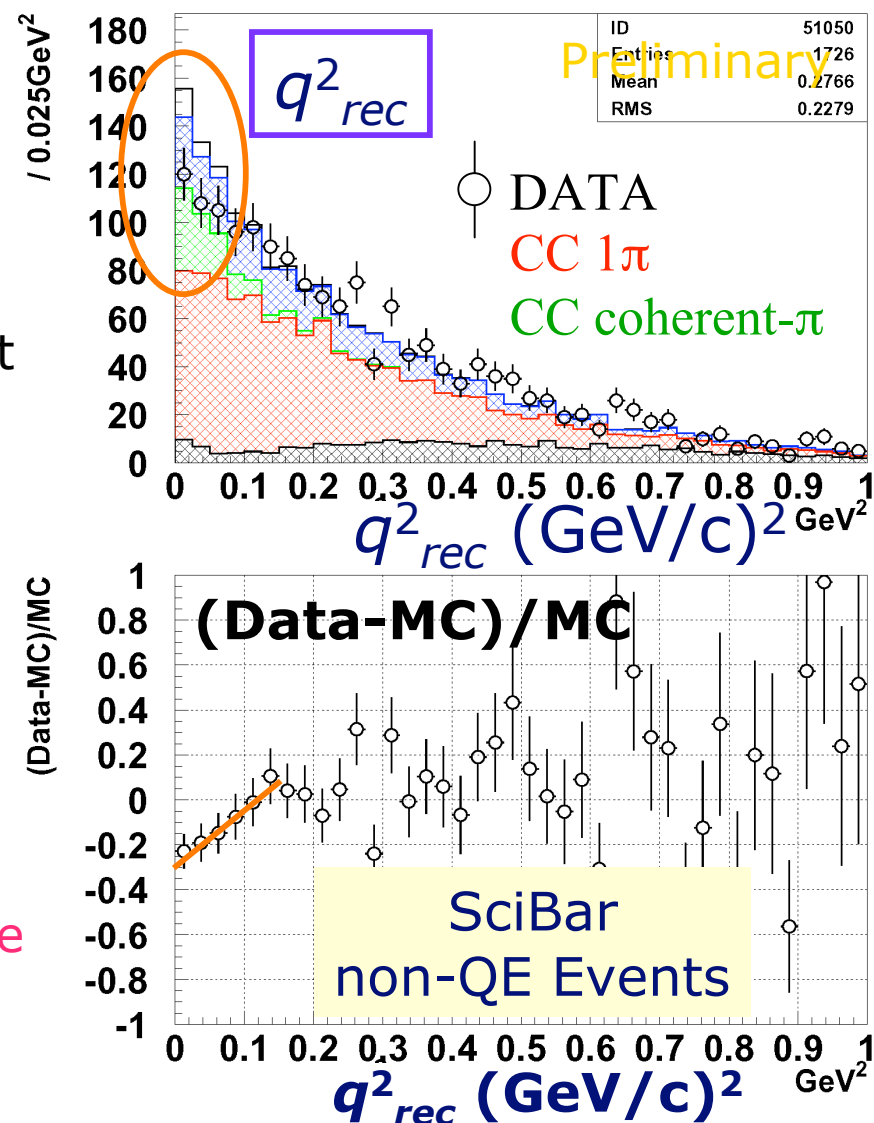
A hint of K2K forward μ deficit.

K2K observed forward μ deficit in **all ND** (KT, SciFi and SciBar).

- A source is non-QE events.
- For CC- 1π ,
 - Suppression of $\sim q^2/0.1[\text{GeV}^2]$ at $q^2 < 0.1[\text{GeV}^2]$ may exist.
- For CC-coherent π ,
 - The coherent π may not exist.

We do not identify which process causes the effect. The MC CC- 1π (coherent π) model is corrected to be consistent with data.

Oscillation analysis is insensitive to the choice.



Near Detector Data

□ 1KT

- Fully Contained 1 ring μ (FC1R μ) sample.

□ SciBar

- 1 track, 2 track QE ($\Delta\theta_p \leq 25^\circ$), 2 track nQE ($\Delta\theta_p > 25^\circ$) where one track is μ .

□ SciFi

- 1 track, 2 track QE ($\Delta\theta_p \leq 25^\circ$), 2 track nQE ($\Delta\theta_p > 30^\circ$) where one track is μ .

With the low q^2 suppression of nQE in SciBar, angular distributions of all other samples are reasonably reproducible with the correction.

Near Detectors combined measurements

(p_μ, θ_μ) for 1track ($1R_\mu$), 2trackQE and 2track nQE samples
→ $\Phi(E_\nu)$, nQE/QE

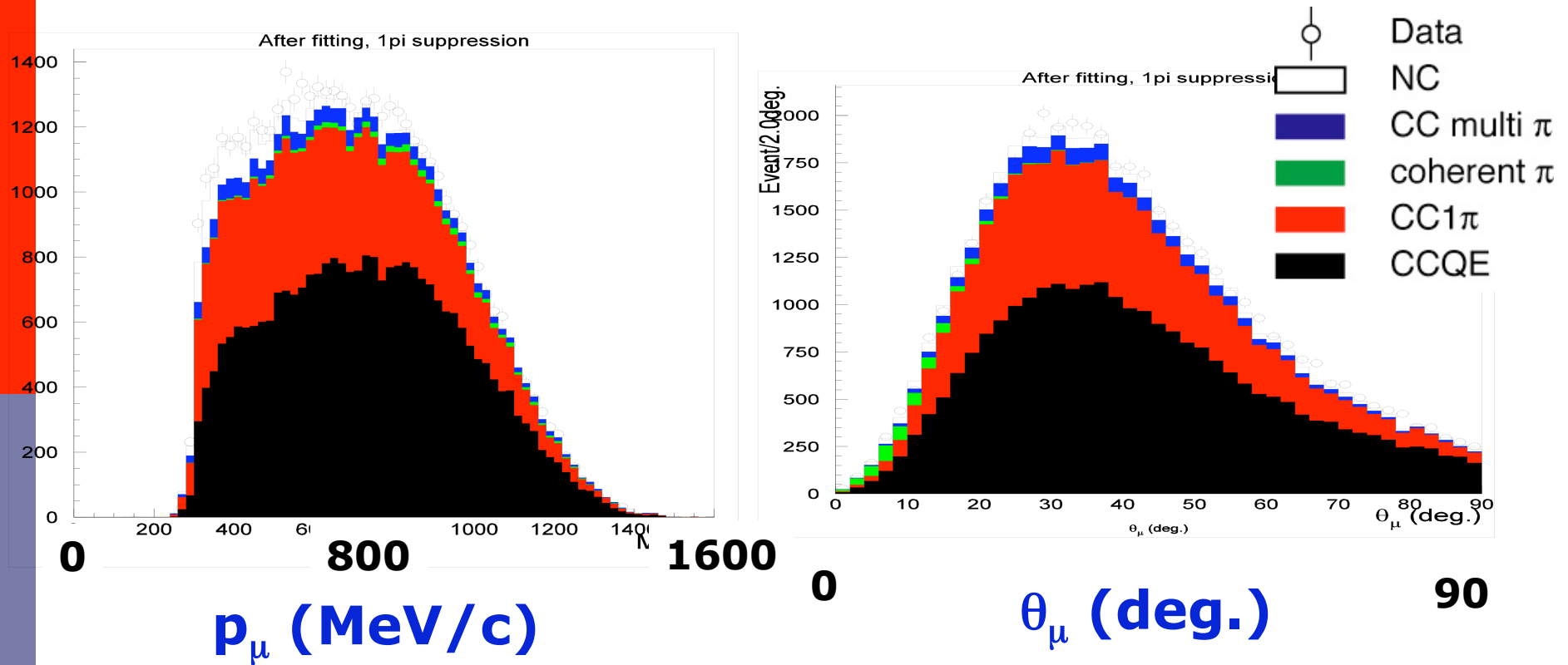
□ Fitting parameters

- $\Phi(E_\nu)$, nQE/QE ratio
- Detector uncertainties on the energy scale and the track counting efficiency.
- The change of track counting efficiency by nuclear effect uncertainties; proton re-scattering and π interactions in a nucleus ...

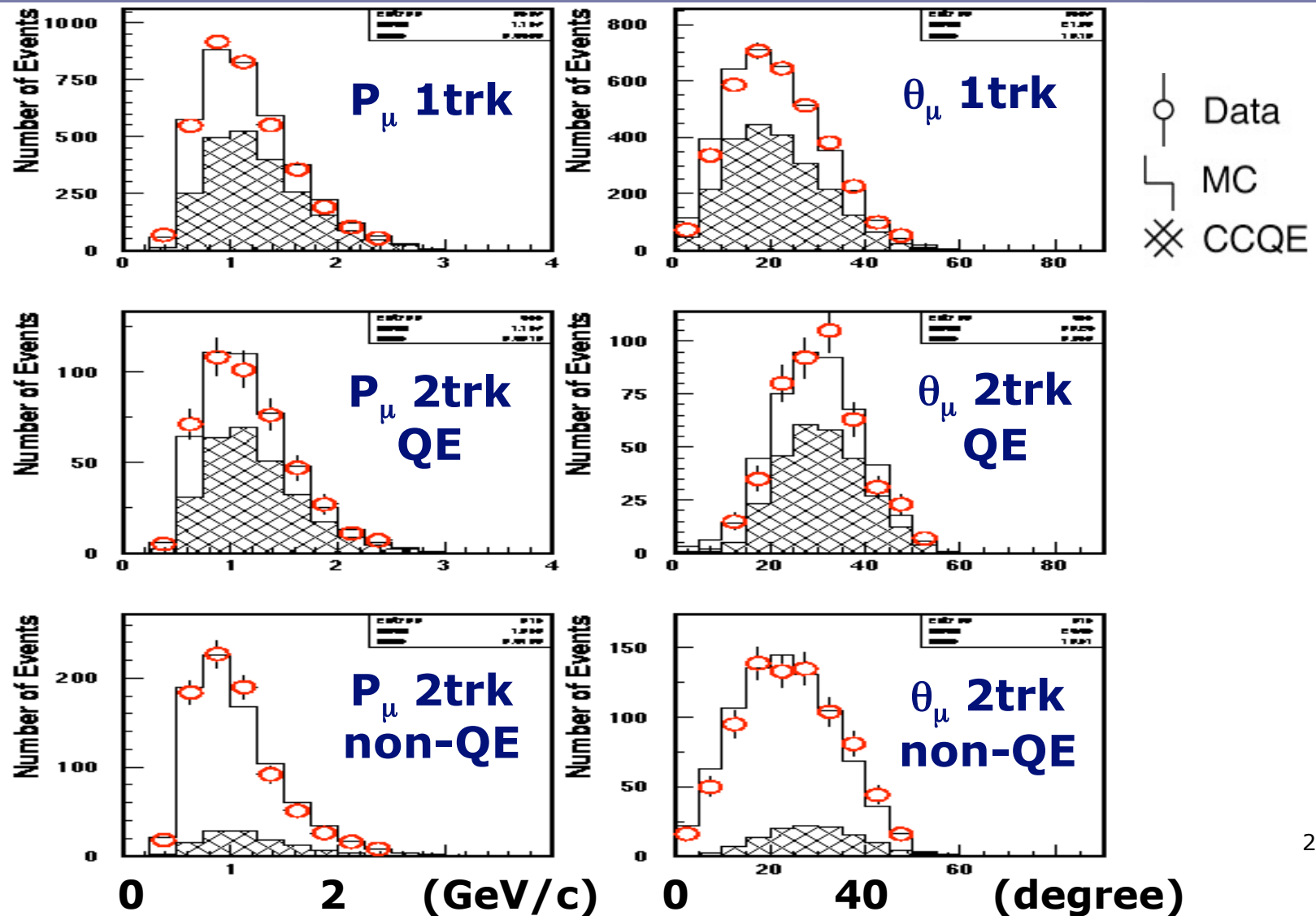
□ Strategy

- ① Measure $\Phi(E_\nu)$ in the more relevant region of $\theta_\mu \geq 20^\circ$ for 1KT and $\theta_\mu \geq 10^\circ$ for SciFi and SciBar.
- ② Apply a low q^2 correction factor to the CC- 1π model (or coherent π) in MC.
- ③ Measure nQE/QE ratio for the entire θ_μ range.

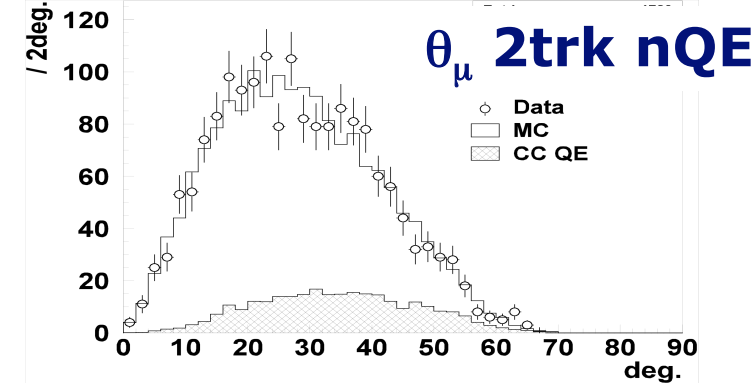
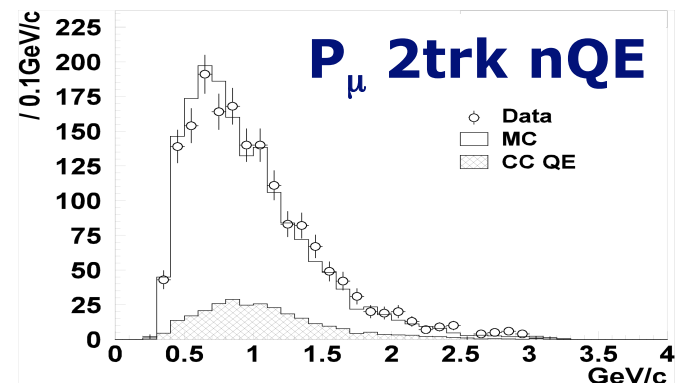
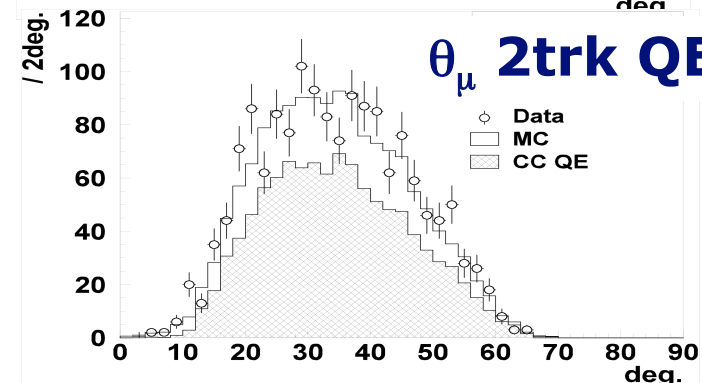
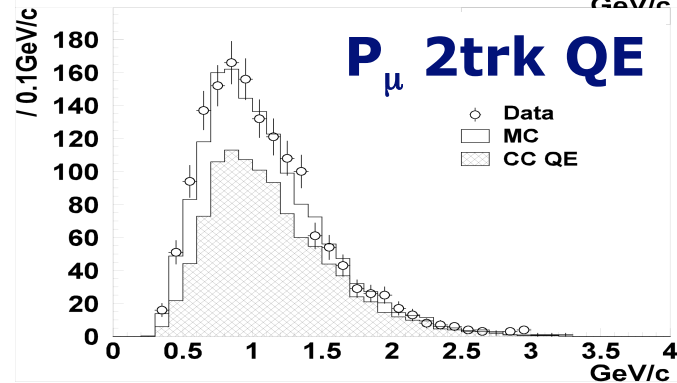
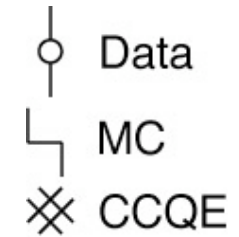
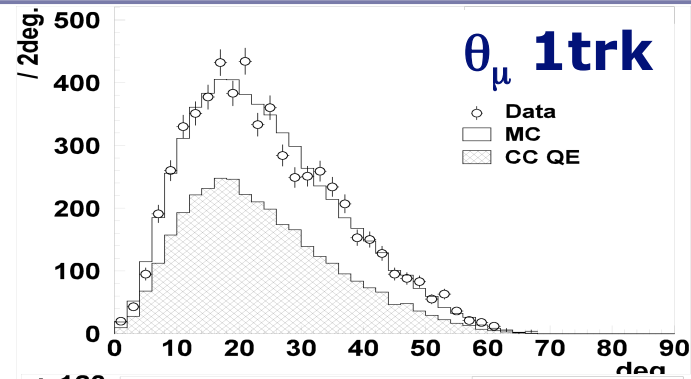
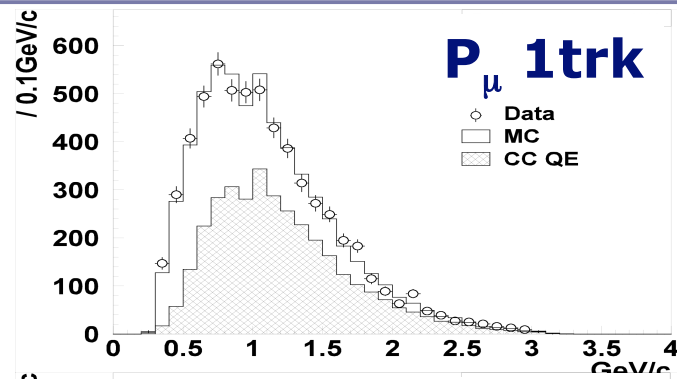
1KT: μ momentum and angular distributions with measured spectrum



SciFi (K2K-IIa with measured spectrum)



SciBar (with measured spectrum)



ND measurement results

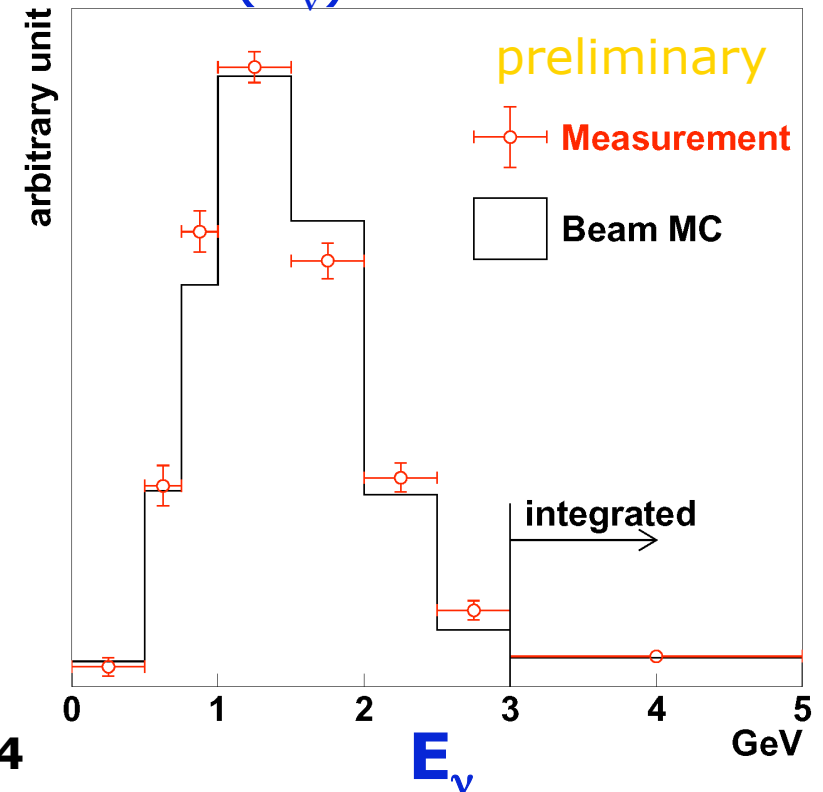
$\chi^2=638.1$ for 609 *d.o.f*

- $\Phi 1$ ($E_\nu < 500$) = 0.78 ± 0.36
- $\Phi 2$ ($500 \leq E_\nu < 750$) = 1.01 ± 0.09
- $\Phi 3$ ($750 \leq E_\nu < 1000$) = 1.12 ± 0.07
- $\Phi 4$ ($1500 \leq E_\nu < 2000$) = 0.90 ± 0.04
- $\Phi 5$ ($2000 \leq E_\nu < 2500$) = 1.07 ± 0.06
- $\Phi 5$ ($2500 \leq E_\nu < 3000$) = 1.33 ± 0.17
- $\Phi 6$ ($3000 \leq E_\nu$) = 1.04 ± 0.18
- nQE/QE = 1.02 ± 0.10

The nQE/QE error of 10% is assigned based on the variation by the fit condition.

- $\theta > 10^\circ$ (20°) cut: $nQE/QE = 0.95 \pm 0.04$
- standard (CC- 1π low q^2 corr.): $nQE/QE = 1.02 \pm 0.03$
- No coherent: $\pi = nQE/QE = 1.06 \pm 0.03$

$\Phi(E_\nu)$ at KEK



Oscillation Analysis



K2K-SK events

preliminary

K2K-all (K2K-I, K2K-II)	DATA (K2K-I, K2K-II)	MC (K2K-I, K2K-II)
FC 22.5kt	108 (56, 52)	150.9 (79.1, 71.8)
1ring	66 (32, 34)	93.7 (48.6, 45.1)
μ-like for E_{ν}^{rec}	57 (56) (30, 27)	84.8 (44.3, 40.5)
e-like	9 (2, 7)	8.8 (4.3, 4.5)
Multi Ring	42 (24, 18)	57.2 (30.5, 26.7)

Ref; K2K-I(47.9×10^{18} POT), K2K-II(41.2×10^{18} POT) 33

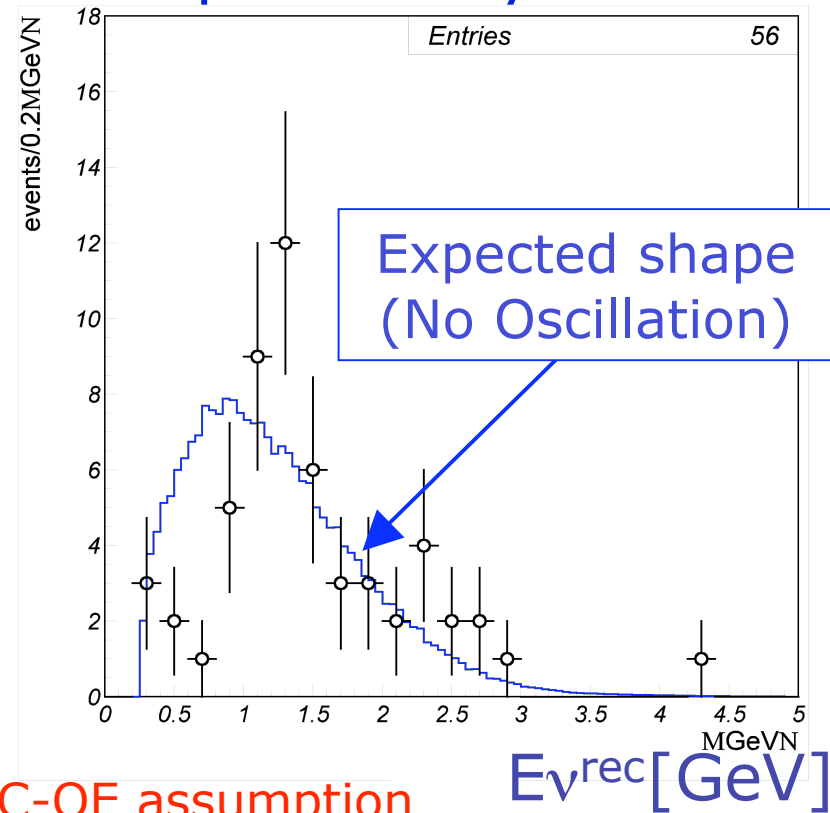
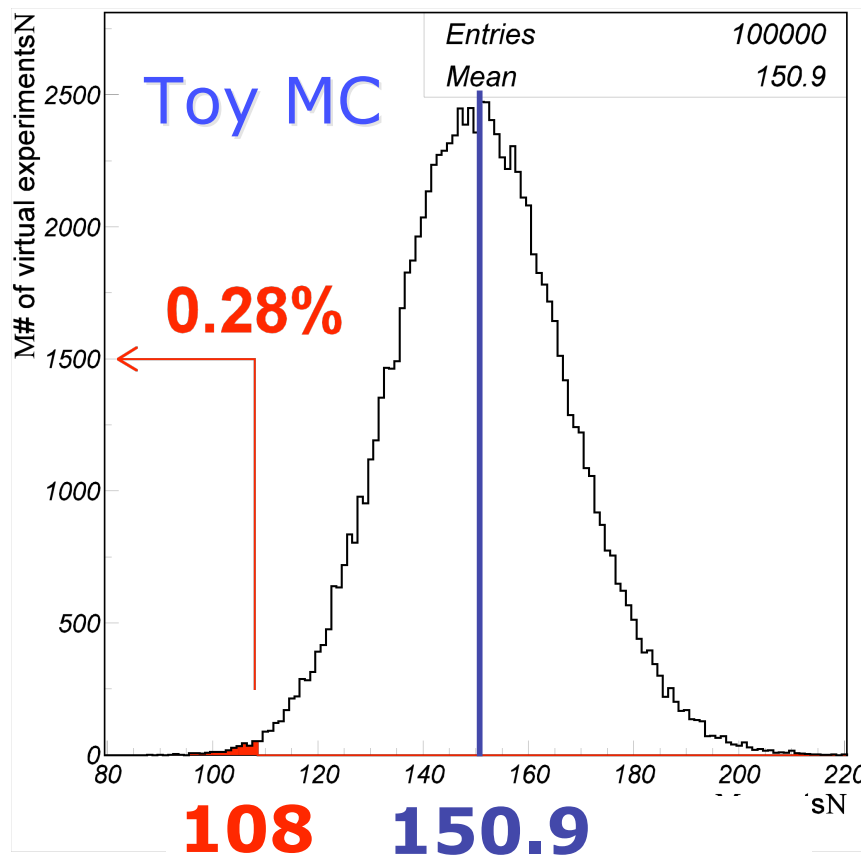
Compared with null-oscillation...

$$L_{norm}(\Delta m^2, \sin 2\theta, f^x)$$

$$L_{shape}(\Delta m^2, \sin 2\theta, f^x)$$

KS probability=0.11%

#SK Events



CC-QE assumption

$$E_{\nu}^{rec} = \frac{(m_N - V)E_{\mu} - m_{\mu}^2/2 + m_N V - V^2/2}{(m_N - V) - E_{\mu} + p_{\mu} \cos \theta_{\mu}}$$

V: Nuclear potential

5. Super-K oscillation analysis

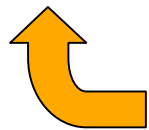
- Total Number of events
- E_ν^{rec} spectrum shape of FC-1ring- μ events
- Systematic error term

$$L(\Delta m^2, \sin 2\theta, f^x)$$

$$= \underline{L_{\text{norm}}(\Delta m^2, \sin 2\theta, f^x)} \cdot \underline{L_{\text{shape}}(\Delta m^2, \sin 2\theta, f^x)} \cdot \underline{L_{\text{syst}}(f^x)}$$

f^x : Systematic error parameters

Normalization, Flux, and nQE/QE ratio are in f^x



Near Detector measurements, Pion Monitor constraint, beam MC estimation, and Super-K systematic uncertainties.

6. Likelihood fit results

preliminary

□ Best fit values.

- $\sin^2 2\theta = 1.53$
- $\Delta m^2 [\text{eV}^2] = 2.12 \times 10^{-3}$

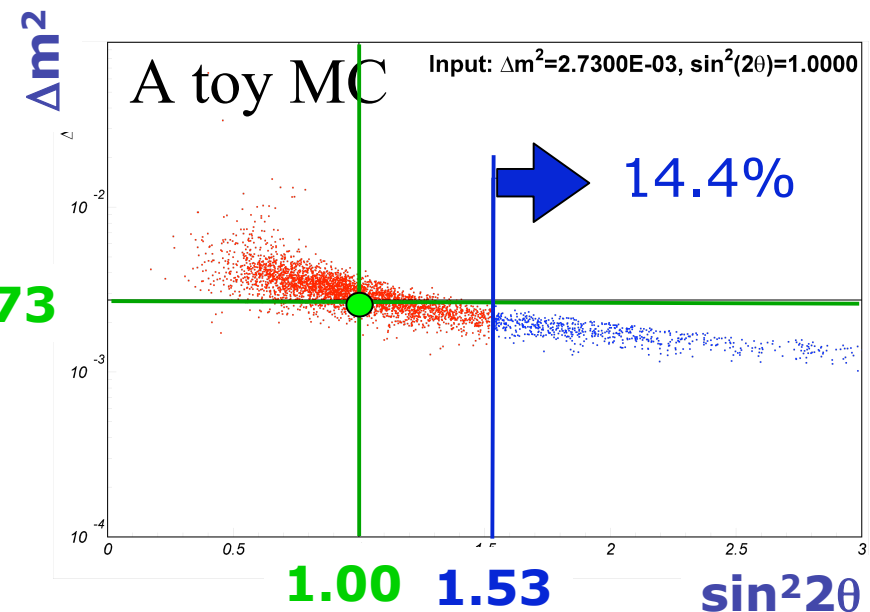
□ Best fit values in the physical region.

- $\sin^2 2\theta = 1.00$
- $\Delta m^2 [\text{eV}^2] = 2.73 \times 10^{-3}$

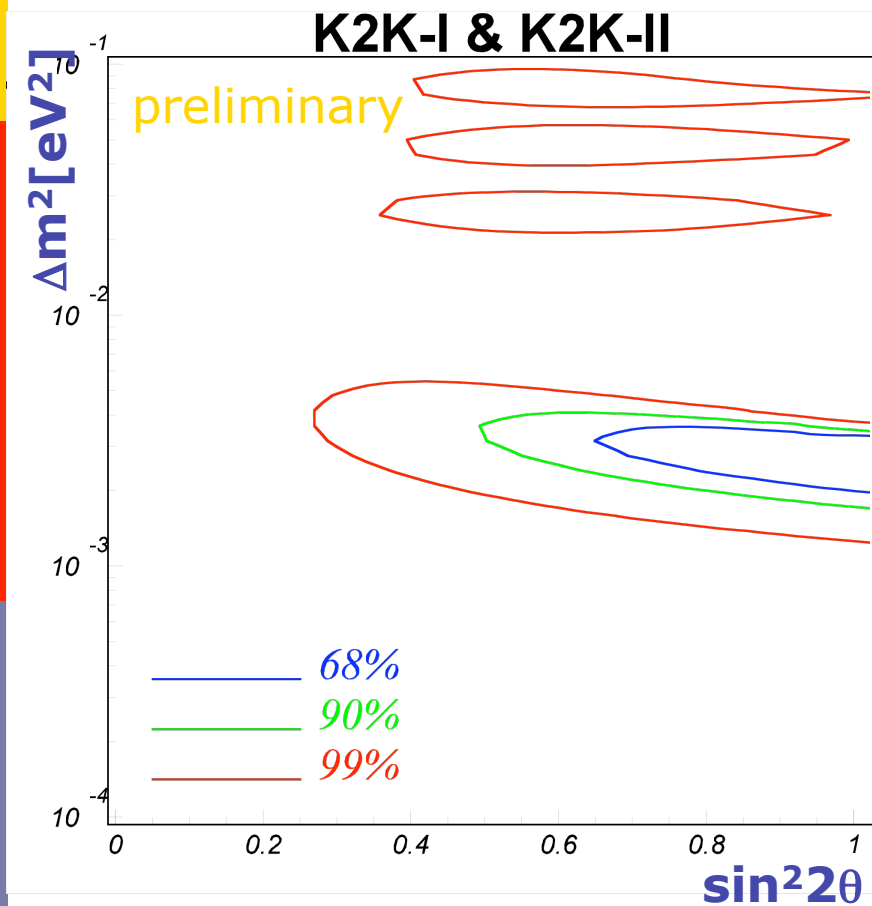
$\Delta \log L = 0.64$

2.73

$\sin^2 2\theta = 1.53$ can occur by statistical fluctuation with 14.4% probability.

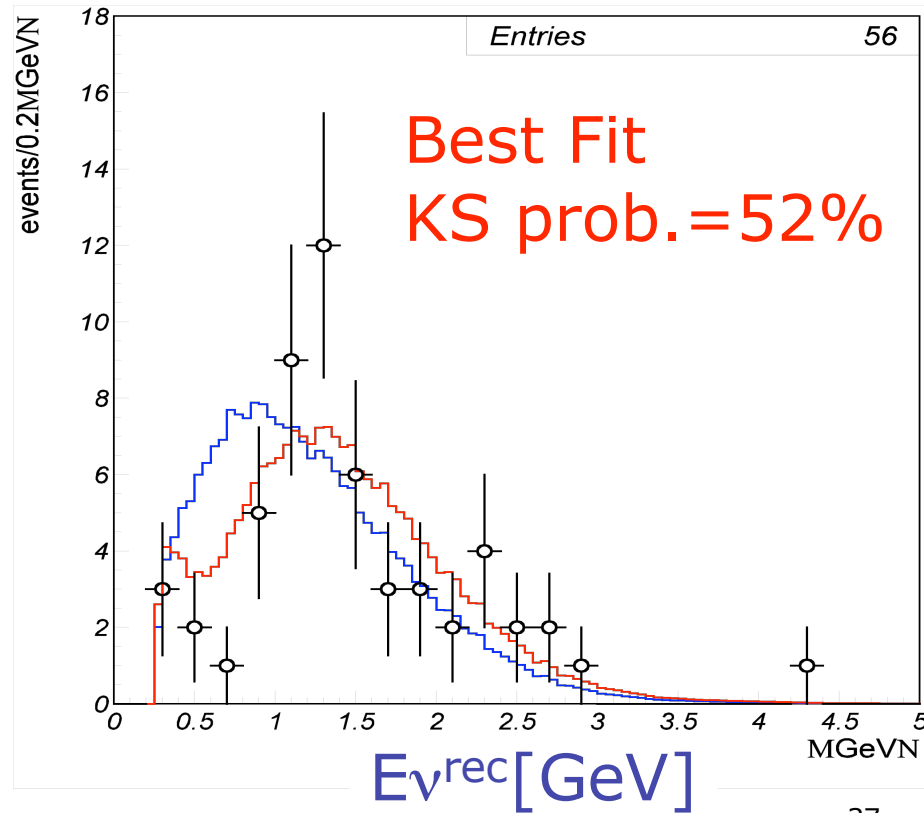


Data are consistent with the oscillation.

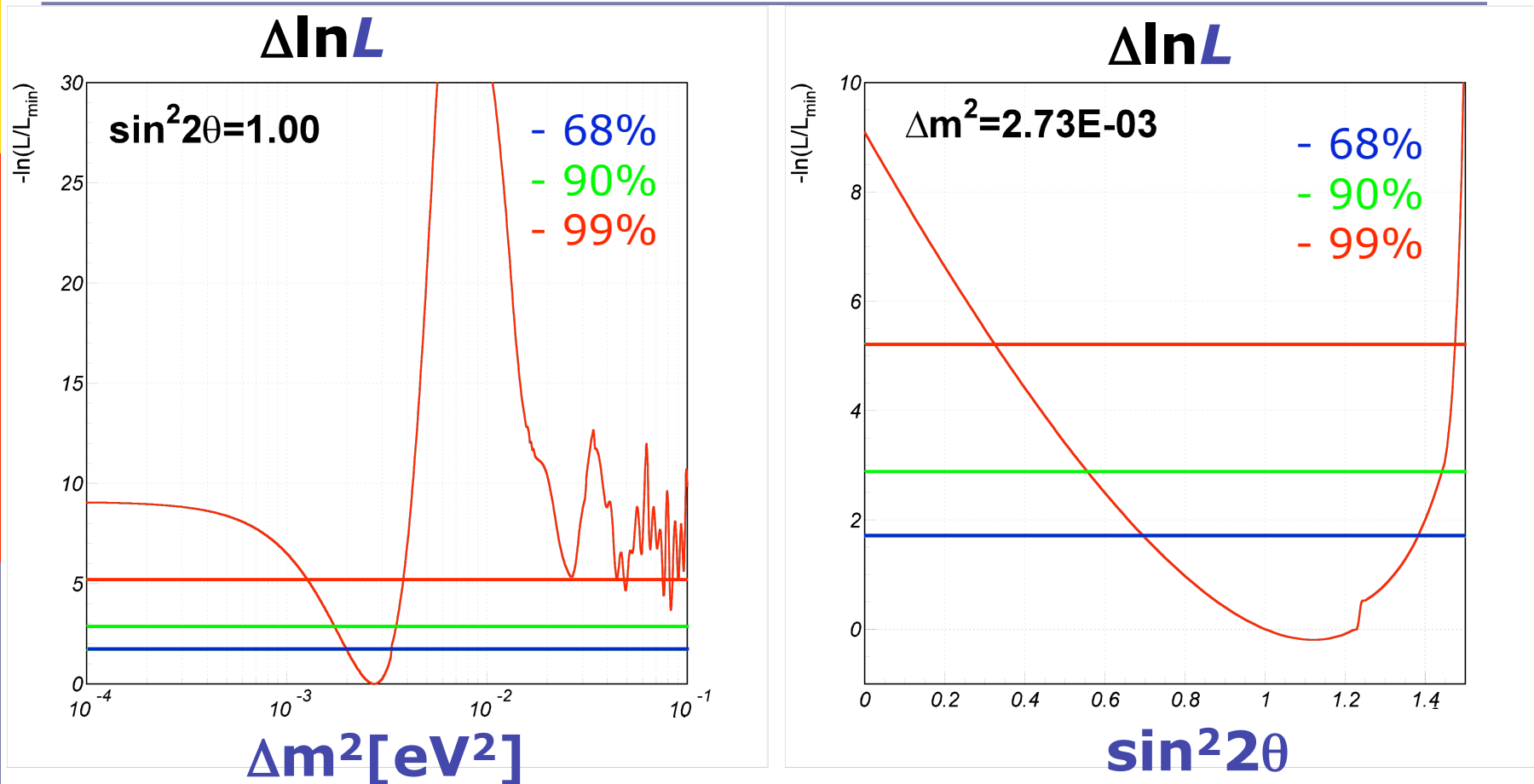


Based on $\Delta \ln L$

- $N_{\text{SK}}^{\text{obs}} = 108$
- $N_{\text{SK}}^{\text{exp}} (\text{best fit}) = 104.8$

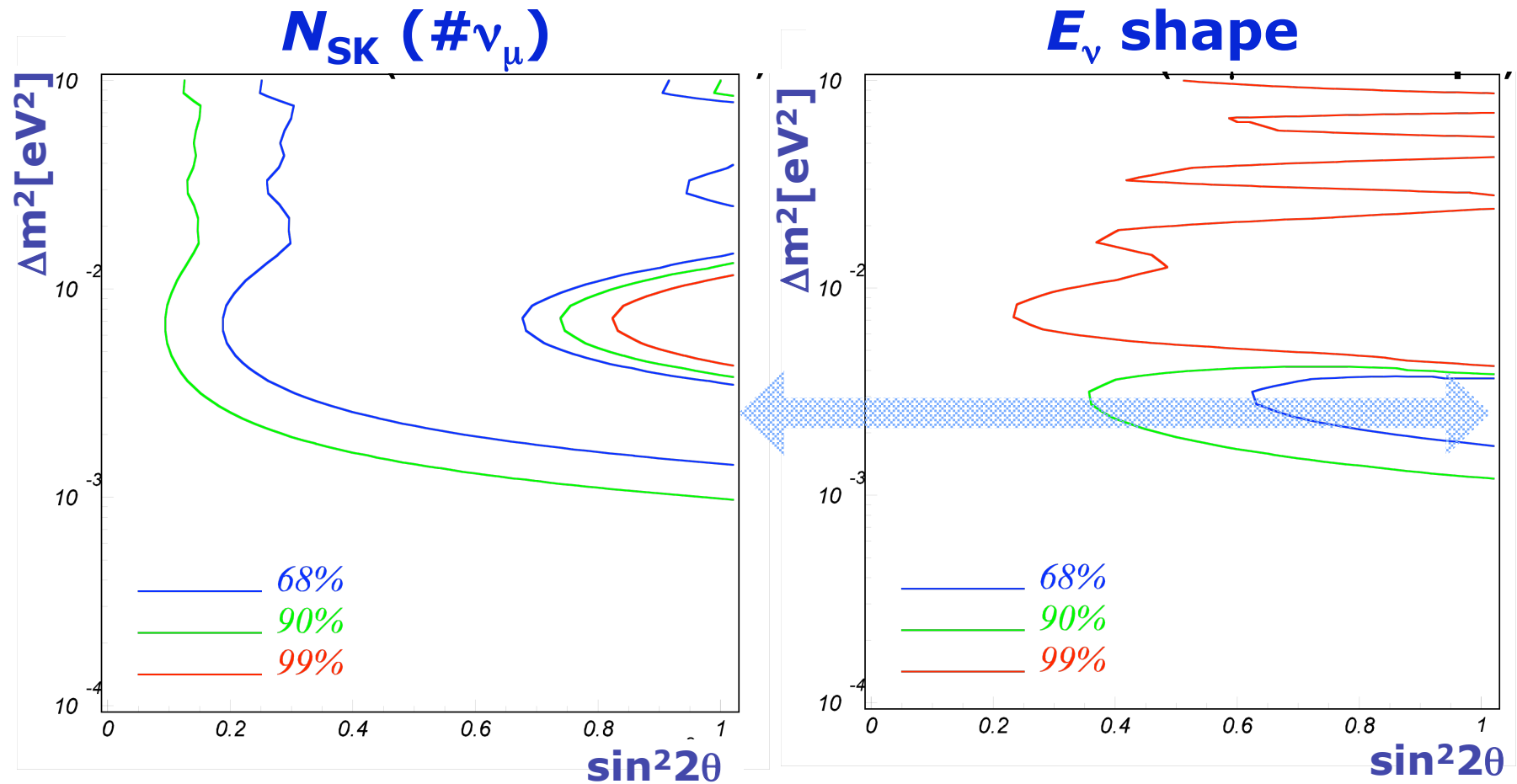


Log Likelihood difference from the minimum.



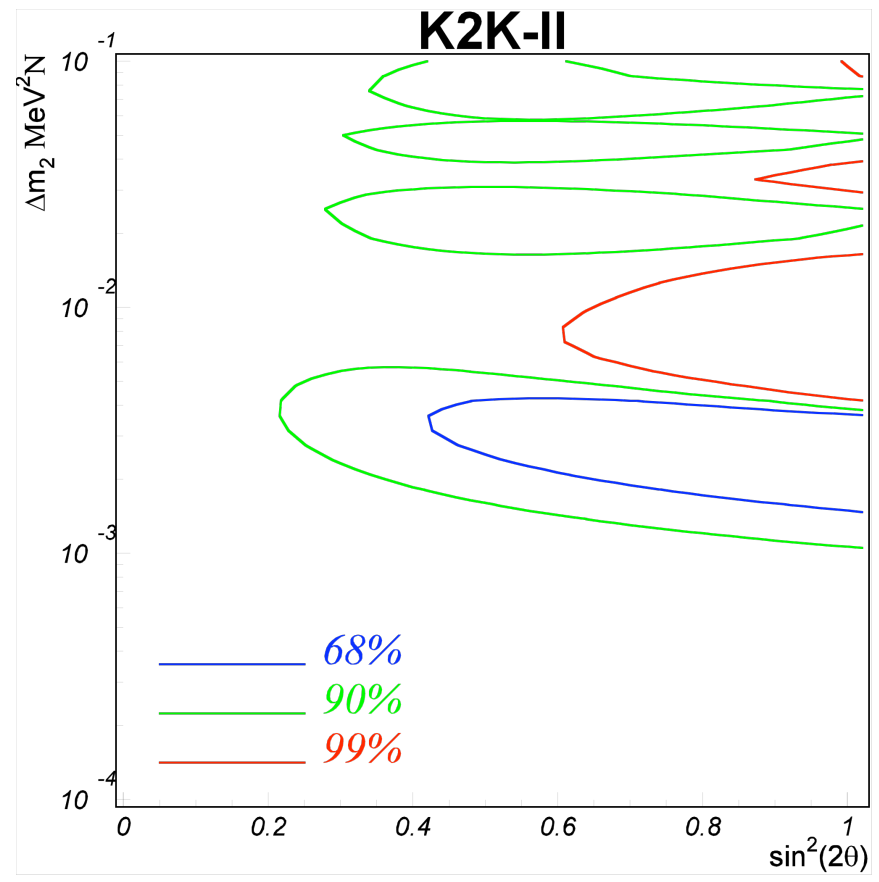
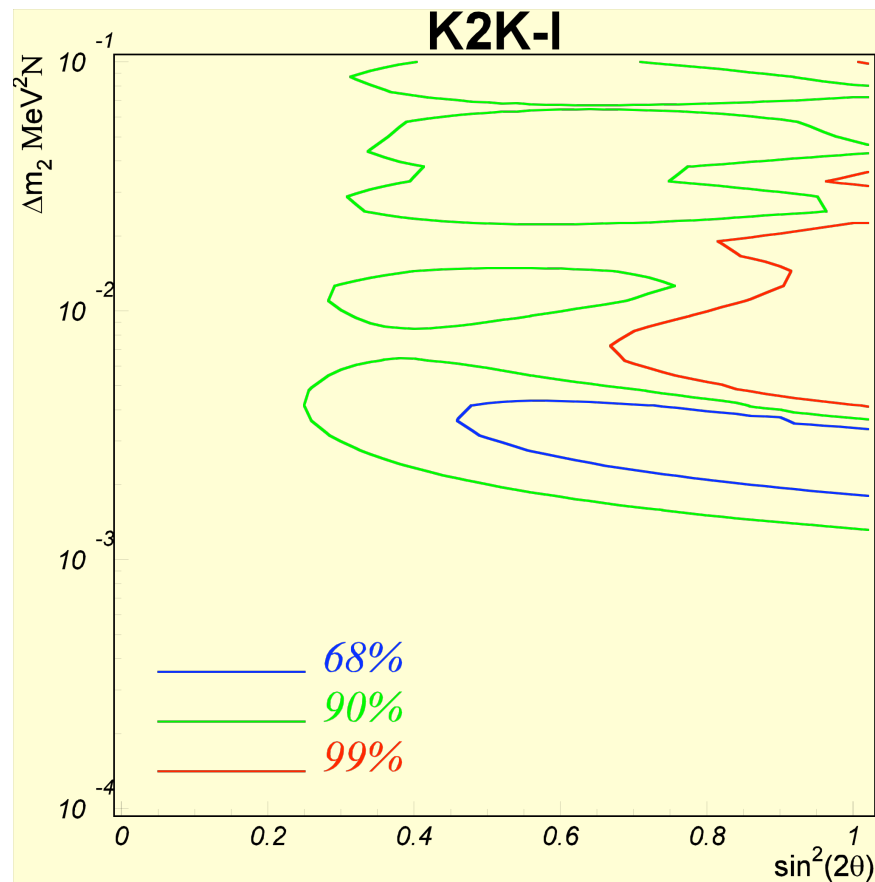
□ $\Delta m^2 < (1.7 \sim 3.5) \times 10^{-3} \text{ eV}^2$ at $\sin^2 2\theta = 1.0$ (90% C.L.)

ν_μ disappearance versus E_ν shape distortion



Both disappearance of ν_μ and the distortion of E_ν spectrum give consistent results.

K2K-I vs K2K-II



Consistent results from both sub-samples.

Null oscillation probability

preliminary

The null oscillation probabilities are calculated based on $\Delta \ln L$.

	K2K-I	K2K-II	K2K-all
ν_μ disappearance	2.0%	3.7%	0.33%(2.9 σ)
E_ν spectrum distortion	19.5%	5.4%	1.1% (2.5 σ)
Combined	1.3% (2.5 σ)	0.56% (2.8 σ)	0.011% (3.9 σ)

Disappearance of ν_μ and distortion of the energy spectrum as expected in neutrino oscillation.

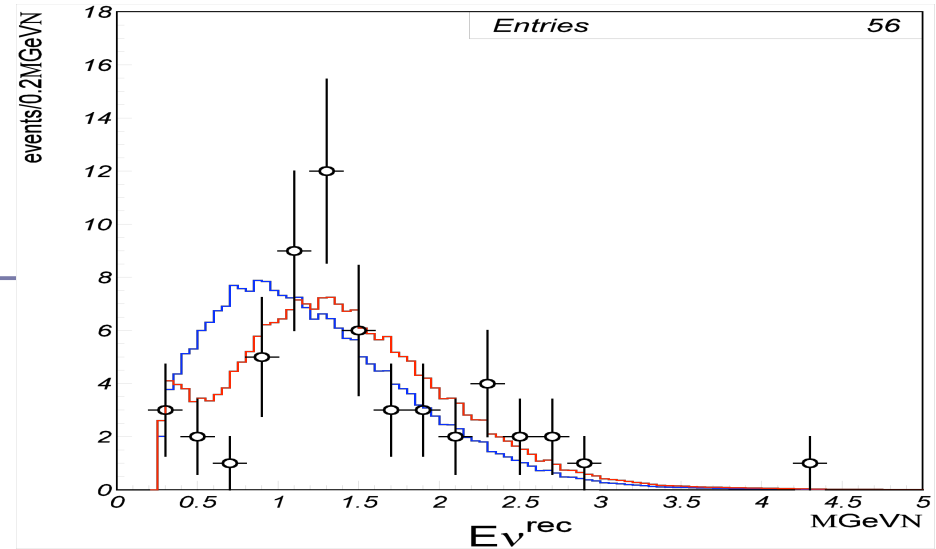
K2K confirmed neutrino oscillation discovered in Super-K atmospheric neutrinos.

8. Summary

With 8.9×10^{19} POT,

K2K has confirmed neutrino oscillation with 3.9σ .

- Disappearance of ν_μ **2.9σ**
- Distortion of E_ν spectrum **2.5σ**



K2K-I & K2K-II

Δm^2 [eV²]

preliminary

0.006

0.004

0.002

0.0

0.2

0.4

0.6

0.8

1.0

$\sin^2 2\theta$

99%
90%
68%

