

Atmospheric neutrino observations in Super-Kamiokande I

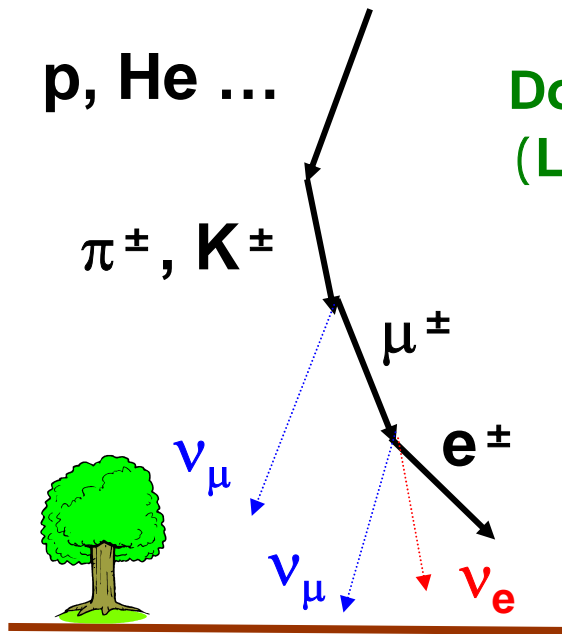
M.Ishitsuka (ICRR)

SK Iにおける大気ニュートリノの観測
(最終結果に近い段階の解析の現状)

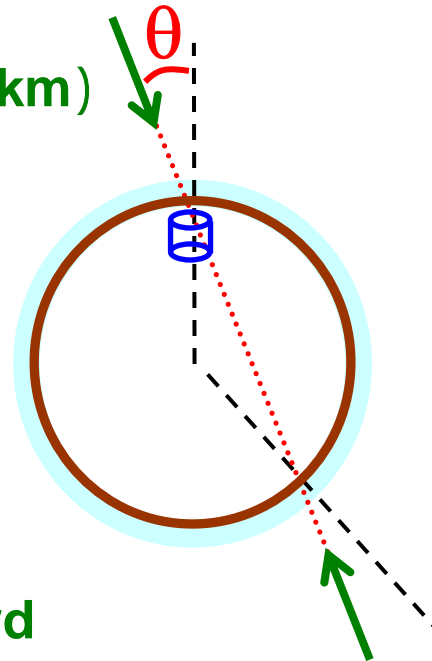
石塚正基(東大宇宙線研)

Atmospheric neutrino

Zenith angle

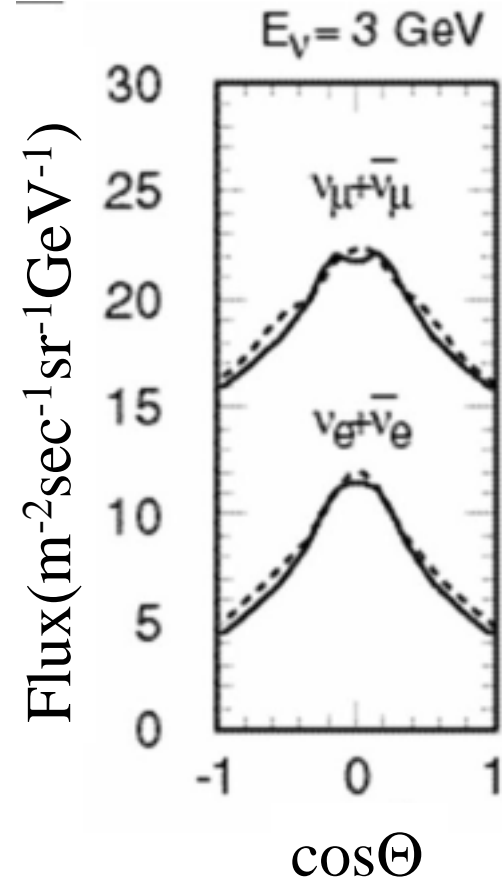


Downward
($L=10\sim 100$ km)



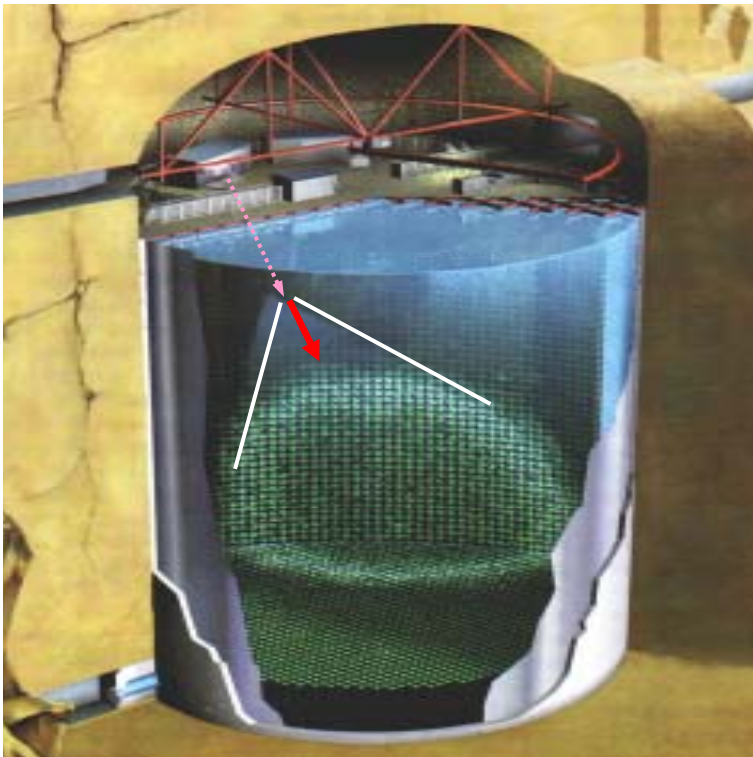
Upward
($L=\text{up to } 13000$ km)

Up/Down Symmetry



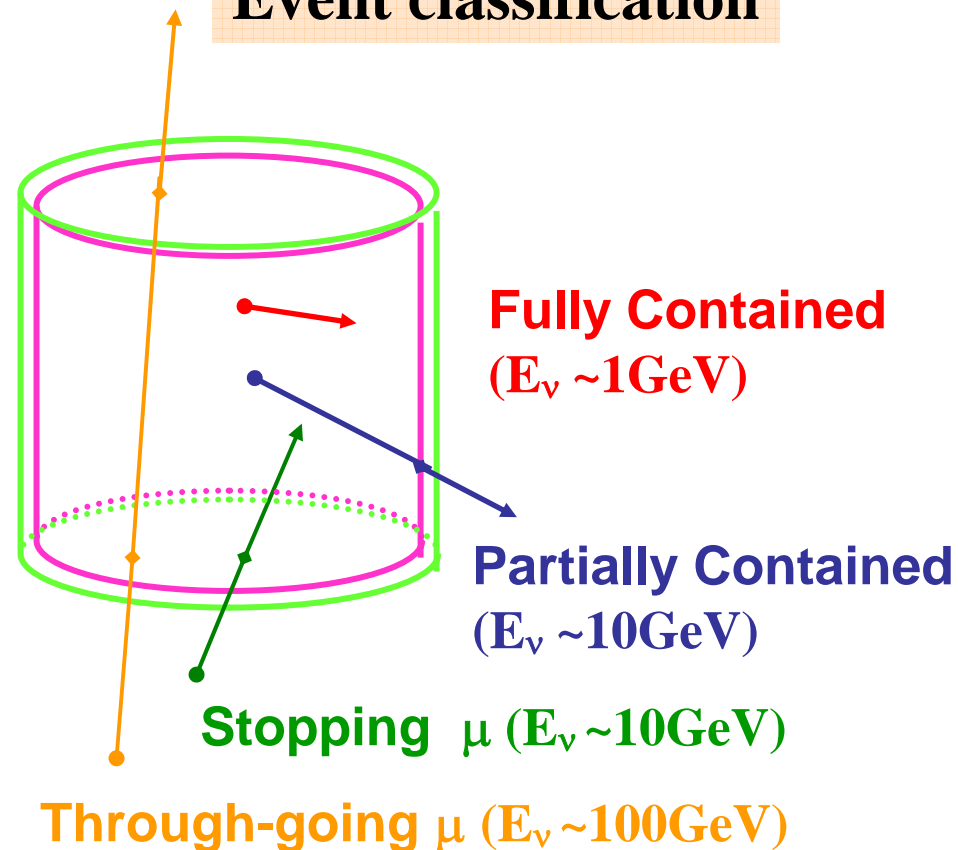
Super-Kamiokande detector & atmospheric neutrino data

Water Cherenkov detector



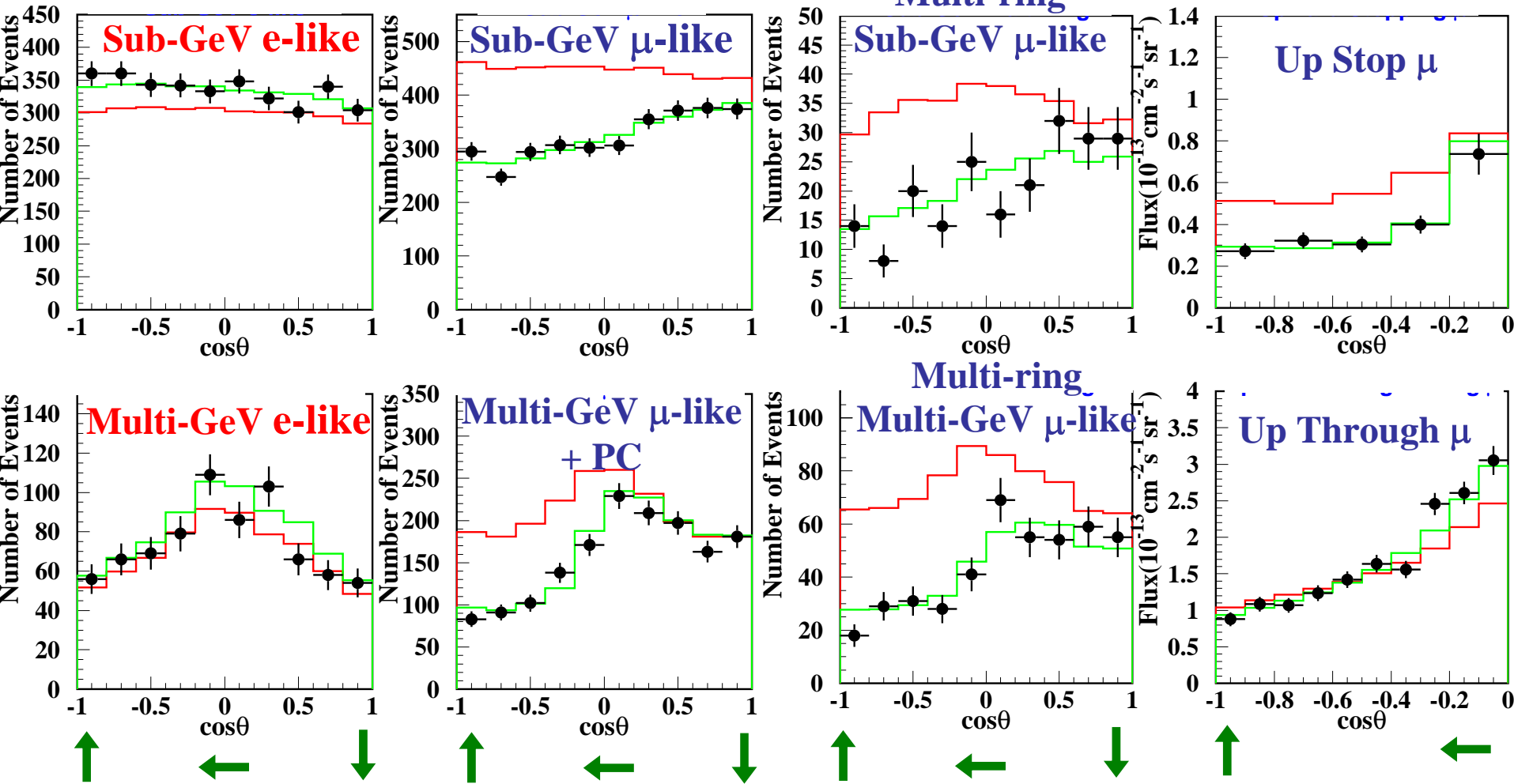
- 1000 m underground
- 50,000 ton (22,500 ton fid.)
- 11,146 20 inch PMTs
- 1,885 anti-counter PMTs

Event classification

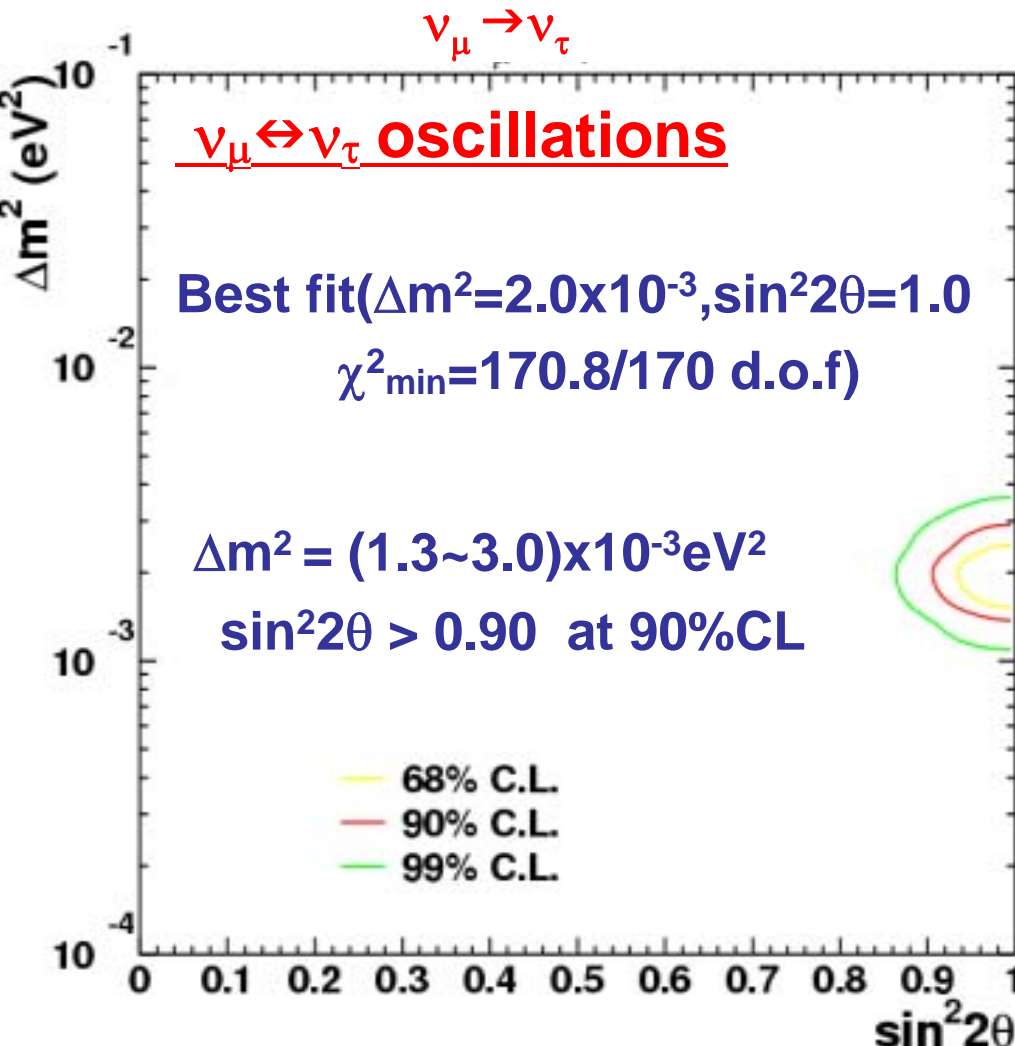


Zenith angle distributions (FC+PC+up- μ)

- No oscillation
- Oscillation best fit ($\Delta m^2=2.0 \times 10^{-3} \text{eV}^2$, $\sin^2 2\theta=1.00$)



Allowed region (FC+PC+up- μ)



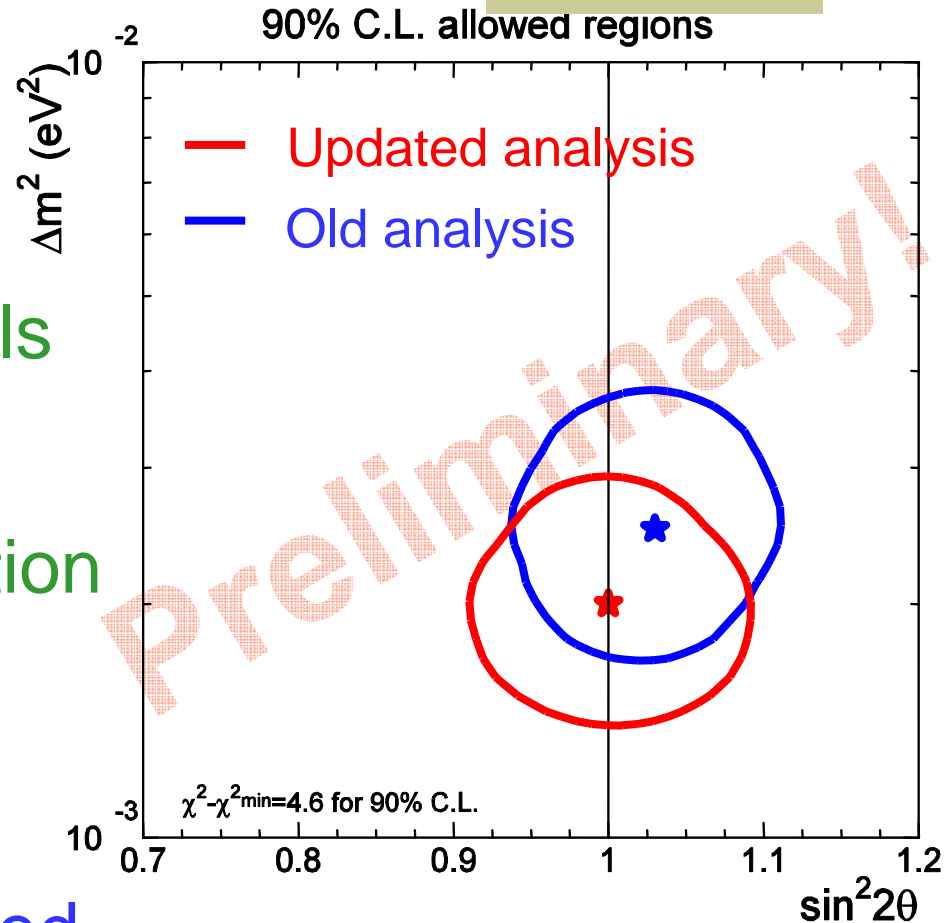
Combined Systematic Errors

		α_i	best fit
α	Absolute Normalization Uncertainty	Free	8.70 %
α_L	SubGeV Multi-ring Absolute Normalization Uncertainty	Free	-23.8 %
α_H	MultiGeV Multi-ring Absolute Normalization Uncertainty	Free	-26.1 %
δ	Ev Spectrum Index	0.05	0.005
β_L	SubGeV μ/e Ratio	8 %	-6.1 %
β_H	MultiGeV μ/e Ratio	12 %	-12.5 %
ρ	FC/PC Relative Normalization	8 %	4.3 %
η_L	SubGeV Up/Down Asymmetry	2.4 %	-2.2 %
η_H	MultiGeV Up/Down Asymmetry	2.7 %	-1.1 %
β_1	FC+PC/Stop \uparrow_{μ} Relative Normalization	7 %	-2.8 %
β_2	Through \uparrow_{μ} /Stop \uparrow_{μ} Relative Normalization	7 %	8.2 %
	FC+PC Horizontal/Vertical Uncertainty	4 %	0.2 %
	\uparrow_{μ} Horizontal/Vertical Uncertainty	3 %	1.5 %
	L/E Uncertainty	15 %	-5.8 %

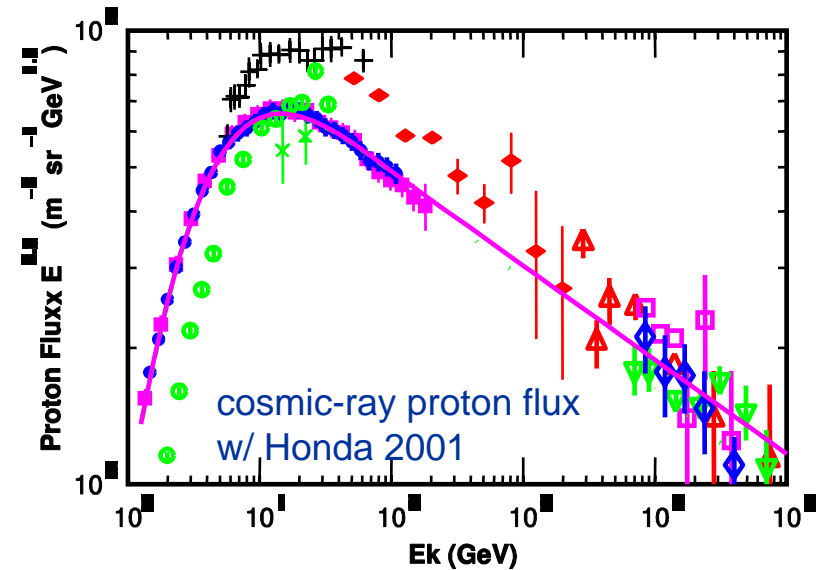
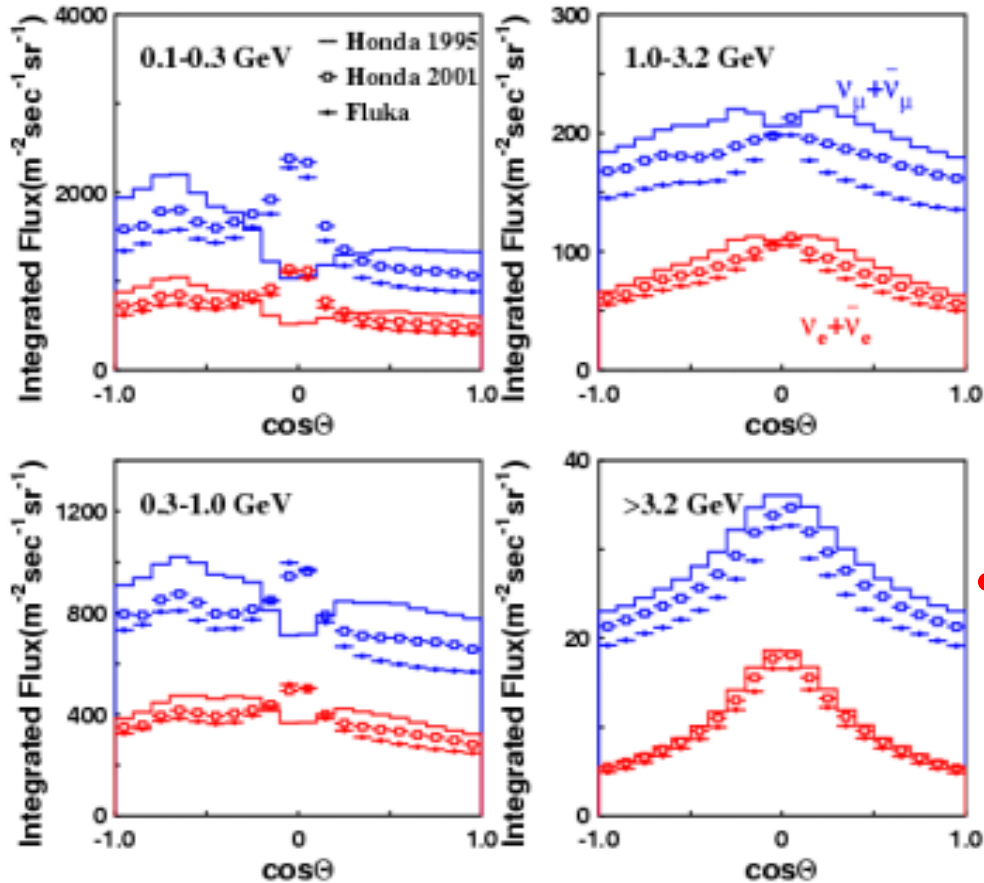
Comparison with old result

- Neutrino flux
(Honda 1995 Honda 2001)
- Neutrino interaction models
(several improvements,
agree with K2K near data)
- Improved detector simulation
- Improved event
reconstruction tools

Each changes slightly shifted
the allowed region to lower Δm^2

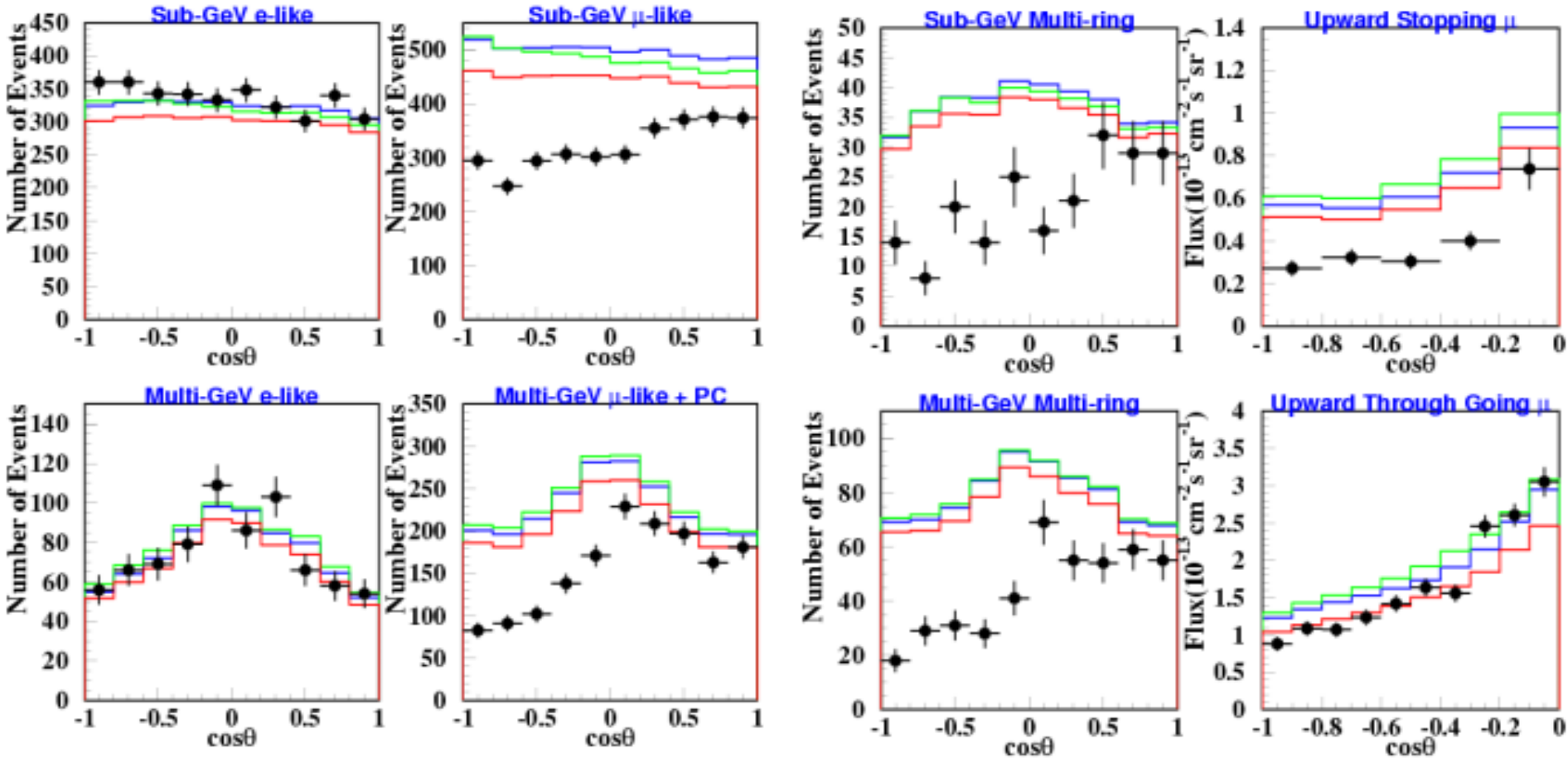


Flux changes



- Honda1D \rightarrow Honda 3D
 - Absolute normalization lower
 - “3D” enhancement
 - At low energies
 - Near the horizon

Different flux in Super-Kamiokande



— Honda 2001 (3D)
— Honda 1995 (1D)

— Bartol 1996
+ Data

Other improvements

Neutrino interactions

Agree with K2K near data

Q.E. $M_A=1.0$ \longrightarrow 1.1

Single π $M_A=1.0$ \longrightarrow 1.1

Analysis tools

SK detector simulator

Light scattering parameters in water (laser measurement)

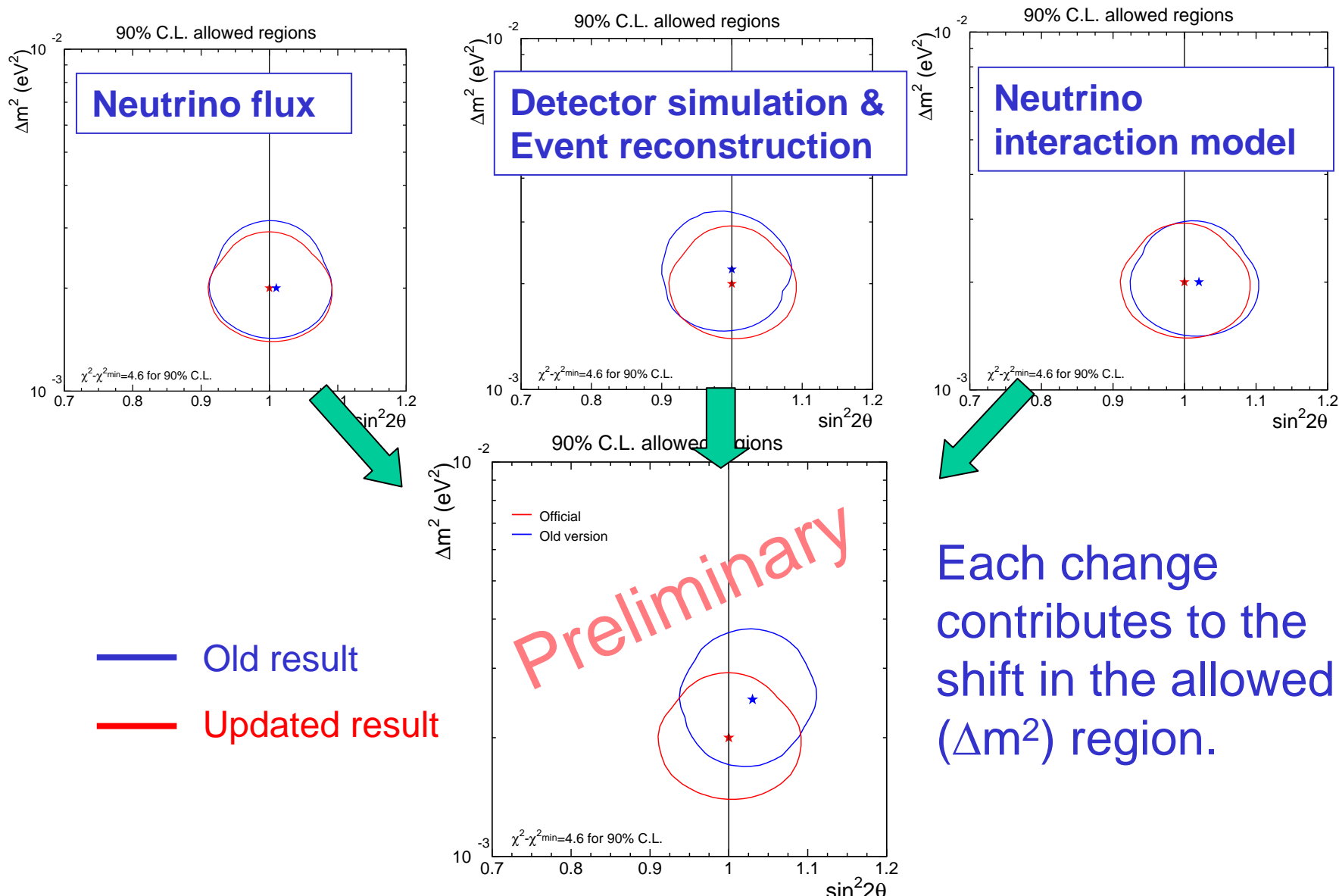
Outer detector PMT gain parameter (single p.e. distribution)

dE/dx in water \rightarrow decrease 3% for μ ($>300\text{MeV}/c$)

Data reduction

Fully automated FC and PC data reduction

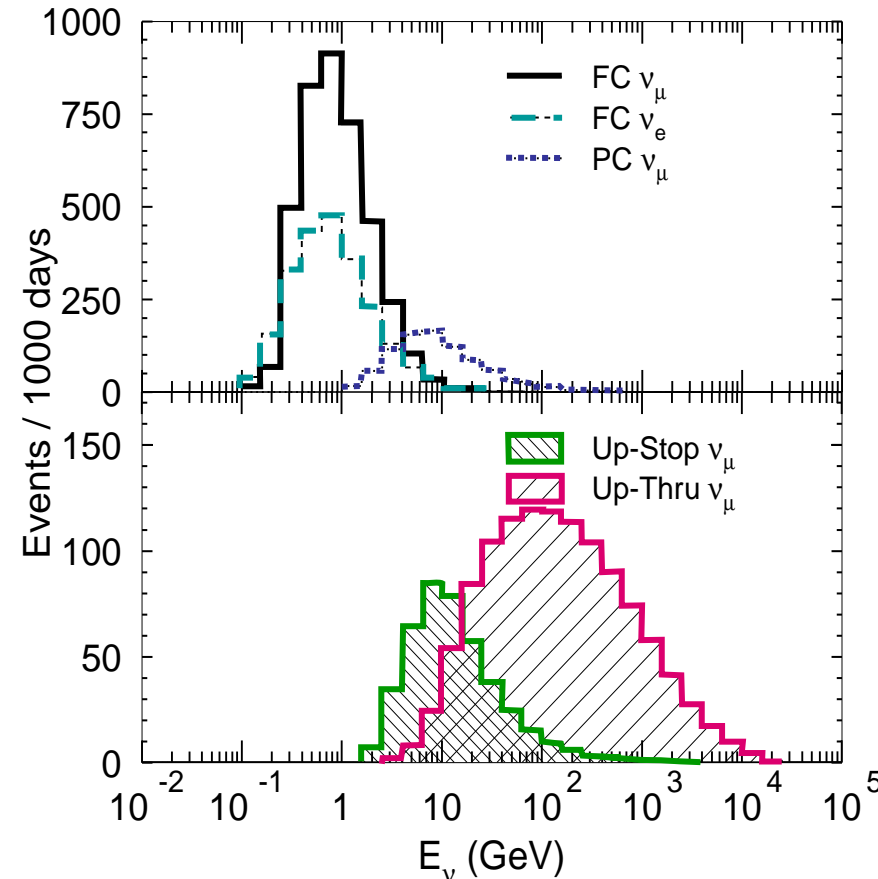
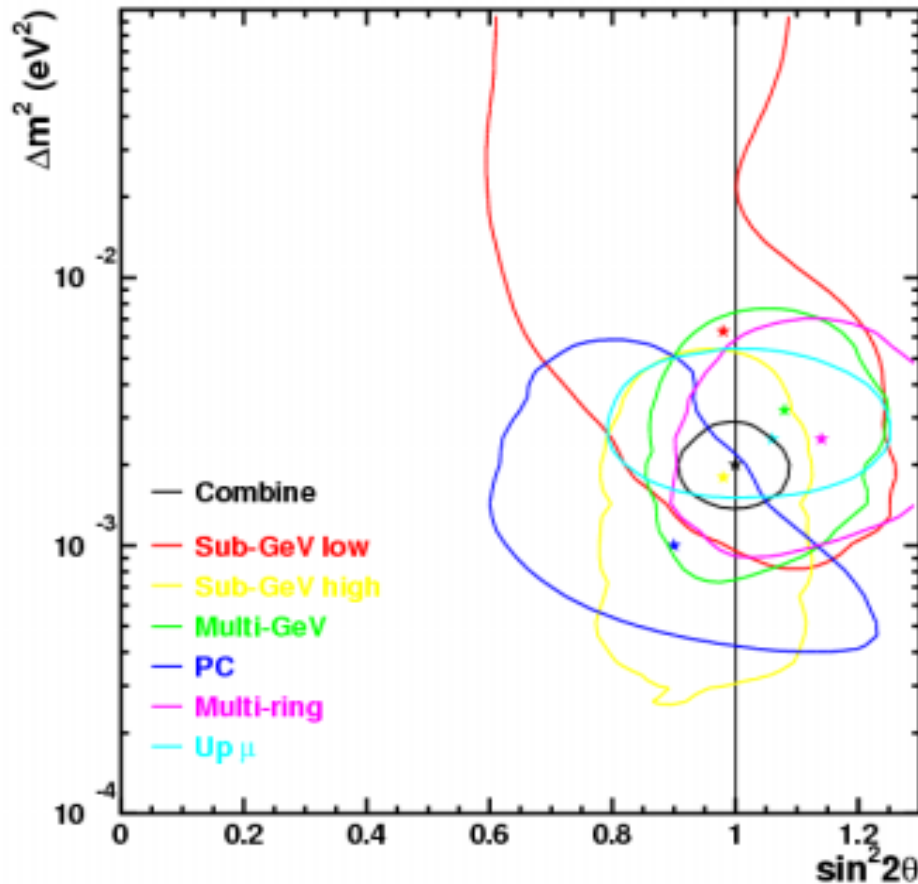
Effect from each change



Sub-sample consistency

Check the result using different types of data

→ all allowed regions overlap best fit



Further improvement for SK I final result

✓ Honda flux updated

Dipole approximation → Multi pole for geomagnetic field

✓ Upward going muon prediction

Analytical calculation → Monte Carlo method

- Consider angular correlation $\nu \rightarrow \mu$

✓ re-estimation of systematic errors

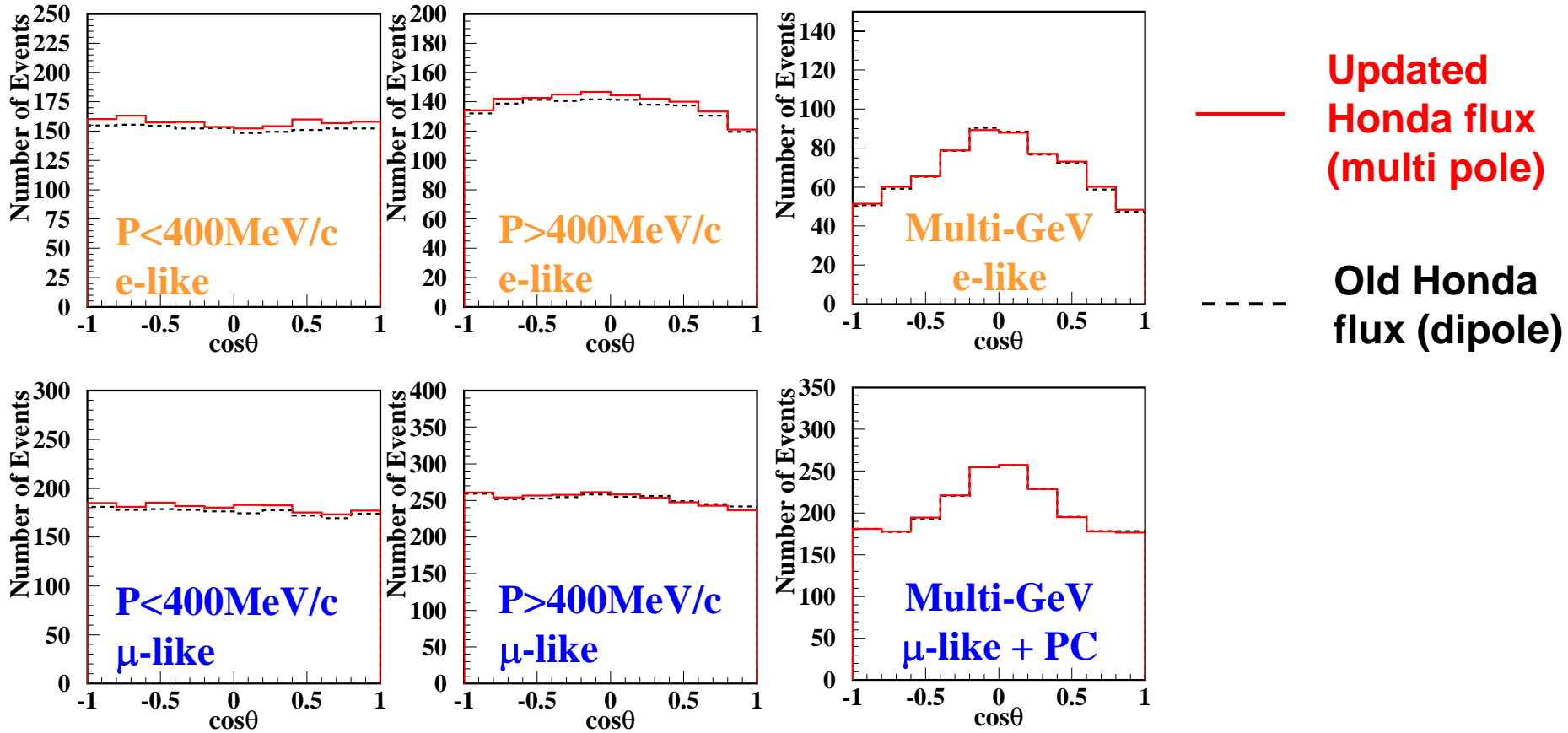
→ more fundamental components with mathematical method

- Check the sources clearly
- Treat the correlations of errors correctly

→ re-estimate systematic uncertainties
estimate additional errors

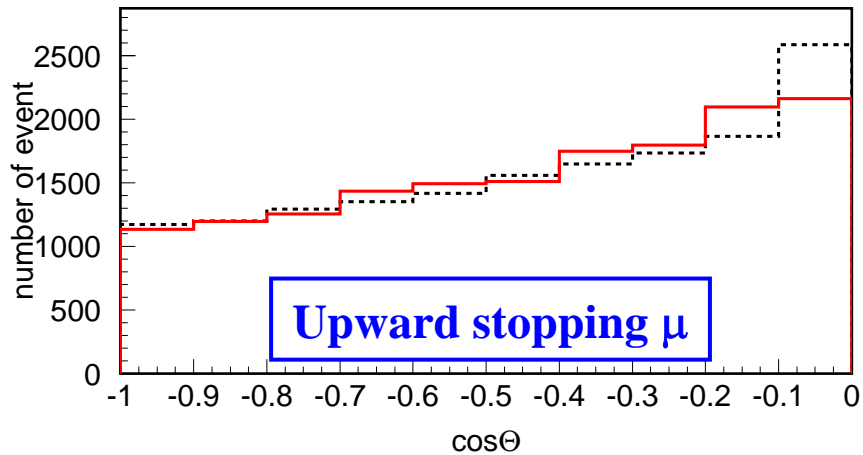
SK I oscillation analysis will be finalized soon (with full paper)

Updated Honda flux



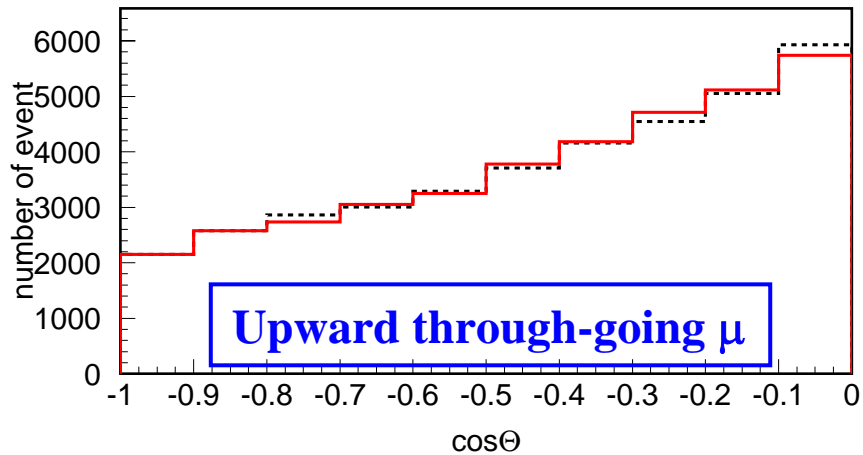
The difference is small

Upward-going μ Monte Carlo



--- Neutrino direction

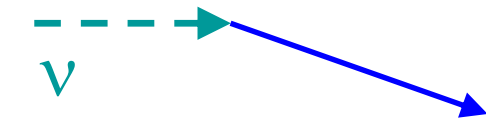
— Muon direction



Analytical calculation



MC



$\mu \rightarrow$ fitter

The difference is small in upward stopping μ 5 bin analysis

\rightarrow becomes large in 10 bin analysis

Calculation methods, basic idea

(based on Fogli et al., hep-ph/0206162)

Definition of χ^2

$$\chi^2 = \sum_{n=1} \frac{\left(N_{\text{obs}}^n - N_{\text{exp}}^n \left(1 + \sum_{k=1}^K f_k^n \cdot \varepsilon^k \right) \right)^2}{\sigma_n^2} + \sum_{k=1}^K \left(\frac{\varepsilon^k}{\sigma_k} \right)^2$$

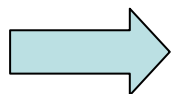
N_{obs}^n : observed, N_{exp}^n : expected, ε^k : systematics

In our current analysis, minimum χ^2 is obtained by optimizing ε_j ($j = 1, 2, \dots$) for each $(\sin^2 2\theta, m^2)$ point.

But, mathematically,

χ^2 is minimum

$$\frac{\partial \chi^2}{\partial \varepsilon_j} = 0 \quad \text{for any } j.$$



Minimizing χ^2

Solving n_{sys} linear equation.

Re-estimation of systematic errors

	Sub-GeV	Multi-GeV
Reduction	$\ll 1$	3
Neutrino flux	< 5	< 5
$E^{-\gamma}$	0.6	1.6
PID	2	3
Ring counting	3	6
Vertex position	0.6	2.4
Energy calibration	1.0	4.1
Non- ν BG		
flasher	< 0.5	< 0.1
cosmic m	< 0.1	< 0.2
neutron BG	< 0.1	< 0.2
Neutrino interaction		
CC	3.5	4.3
NC	3	4
Hadron simulator	0.5	1.0
FC/PC separation	< 0.1	0.5



Treat as independent error term and re-estimate the values

Error in μ/e ratio 7.8% 11.8%

List of updated systematic errors

(A) Neutrino flux

- flux absolute normalization
- ν_e/ν_μ ratio
- anti-neutrino/neutrino ratio
- geomagnetic effect(up-down ratio)
- horizontal-vertical ratio (3D effect)
- horizontal-vertical ratio (K/pi)
- Neutrino flight length (production height)
- Energy spectrum
- Sample-by-sample normalization

(B) Neutrino cross section & kinematics

- Quasi elastic scattering
- low energy neutrino scattering in QE
- single-pion production
- multi-pion production
- Bodek correction in DIS
- coherent pion
- NC/CC ratio
- neutrino-lepton angular correlation

(D) Event selection

- FC reduction
- PC reduction
- Up- μ detection efficiency
- 1.6 GeV cut (\rightarrow 2% energy calibration)
- Through-going/stopping μ separation

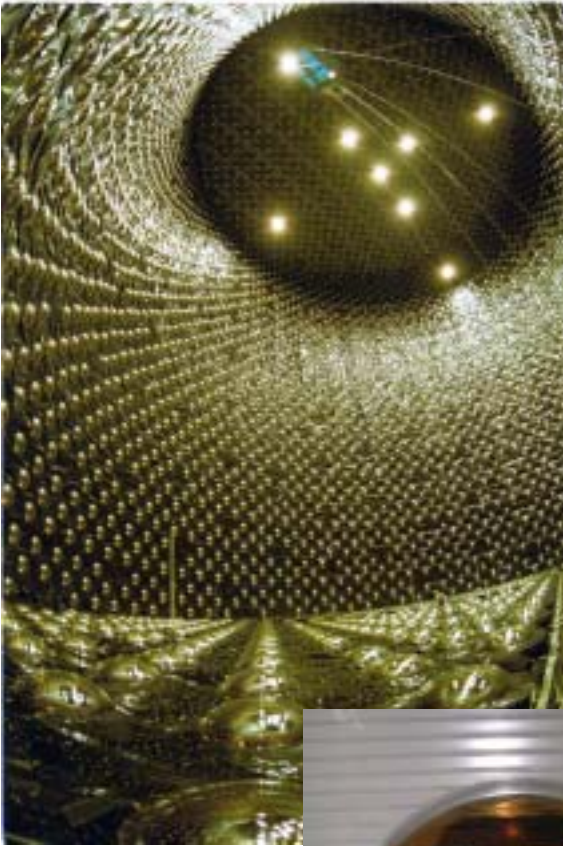
(E) Event reconstruction

- 1-ring/multi-ring separation
- Particle ID
- Energy calibration
- Up-down asymmetry of energy calibration

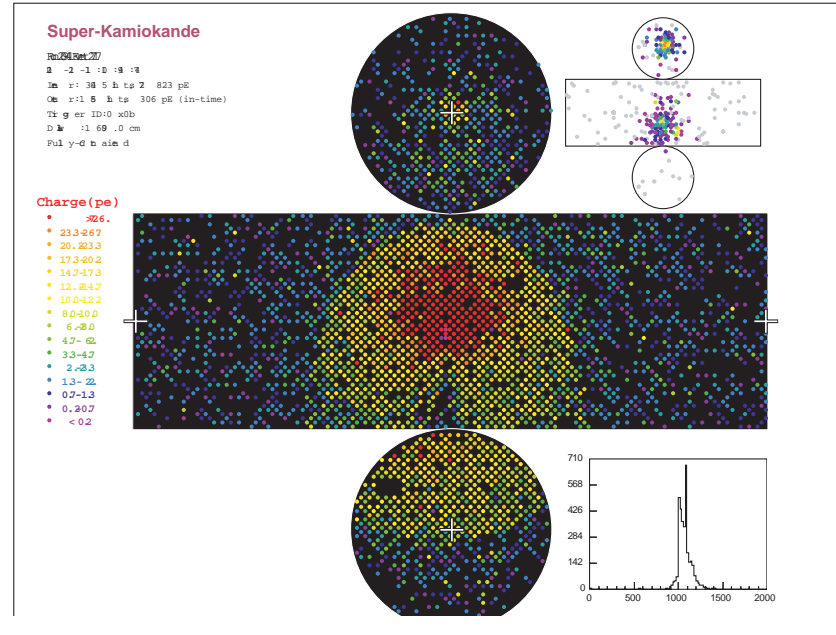
(F) Other sources

- m/e ratio
- up/down ratio
- FC/PC relative normalization
- τ from oscillation

SK-II is working



20inch PMT with
Acrylic + FRP vessel



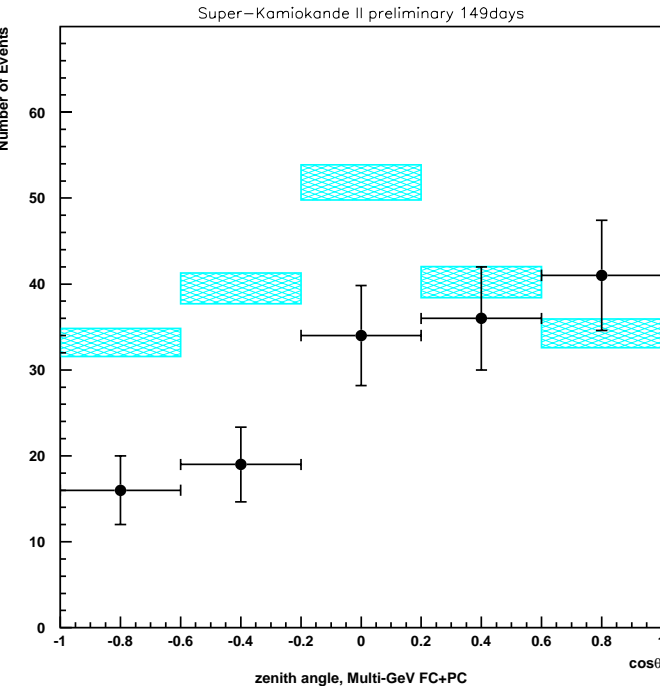
SK-II Cosmic ray muon sample

With about half of original ID PMTs

- 20-inch PMTs in acrylic shells to prevent future chain implosions
- OD at full complement (1885) of 8-inch PMTs

First report from SK-II

Number of Events / Event rate in SK-II



	SK-II 149 days	SK-I 1489 days
Fully-contained events	1245 (8.33 ± 0.24)	12180 (8.18 ± 0.07)
Partially-contained events	80 (0.54 ± 0.06)	911 (0.62 ± 0.02)

The event rate is consistent with SK-I

SK-II zenith angle for
Multi-GeV μ -like + PC

Clear deficit in upward-going events

Summary

➤ $\nu_\mu \leftrightarrow \nu_\tau$ 2 flavor oscillations

➤ Updated result

- Best fit value is ($\Delta m^2 = 2.0 \times 10^{-3} \text{ eV}^2$, $\sin^2(2\theta) = 1.0$)
- $-1.3 < \Delta m^2 < 3.0 \times 10^{-3} \text{ eV}^2$, $\sin^2(2\theta) > 0.90$ @ 90% C.L.
- Each change contributes to the shift of allowed region

➤ final SK-I result

- To be published soon with SK-I paper after modifications
 - updated Honda flux calculation
 - prediction of upward-going μ by Monte Carlo method
 - re-estimated systematic errors

➤ Status of SK-II

➤ SK-II analysis has started

- Event rate is consistent with SK-I
- deficit in upward-going event in Multi-GeV μ -like + PC sample