

# Leptogenesis

**Takuya Morozumi**

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*Collaboration with*

*T. Endoh and S. K. Kang.*

*(Hiroshima Univ.)*

*S. Kaneko and M. Tanimoto*

*(Niigata)*

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## Where did anti-matter disappear ?

$$n_B/n_\gamma = 10^{-9\sim-10}$$

- Baryon density of the universe

$$\rho_B = 5 \times 10^{-30\sim-31} (g/cm^3)$$

$$\rightarrow n_B = N_A \times \rho_B = 3 \times 10^{-6\sim-7} /cm^3$$

- Photon number density of the universe

$$n_\gamma = 2.404 \frac{T^3}{\pi^2} = O(750/cm^3)$$

$T = 3$  K background black body radiation.

# **Particle Physics Model for generating Baryon number asymmetry**

**→ Three Ingredients:** Sakharov

**1. B-L violation at high energy**

**2. CP violation**

**3. Thermal non-equilibrium:**

## Why Baryogenesis through Leptogenesis ?

(Fukugita and Yanagida)

A simple baryogenesis scenario based on GUT does not work under the presence of anomaly.  $B + L$  is not conserved. (The effect is significant for  $T > 1(\text{TeV})$ .)

$$\partial_\mu(j_B^\mu + j_L^\mu) \neq 0 \quad (1)$$

Primordial  $B + L$  will be washed out as  $(B + L)_{\text{primordial}} \exp[-t/\tau]$  and the present baryon number is proportional to  $B - L$ .

$$\begin{aligned} B_{\text{now}} &= \frac{8N_g + 4N_H}{22N_g + 13N_H} (B - L)_{\text{primordial}} \\ &\sim \frac{1}{3} (B - L)_{\text{primordial}} \end{aligned}$$

Particle Physics model must break  $B - L$  at high energy.  $B - L$  is gauge symmetry spontaneously broken or it must be explicitly broken.

## **CP violation**

Standard model can not account for baryon number though it is good at explaining the present measured CP violation phenomena observed in Kaon and B meson system.

→ some new source of CP violation other than Kobayashi Maskawa phase is required.

## **The Seesaw Models (=the standard model + heavy Majorana neutrinos),**

$B-L$  is broken and the other two conditions (CP violation and thermal non-equilibrium) can be satisfied.

This talk is about:

- 1. How well can we predict the baryon number asymmetry based on a specific model ?**
- 2. How can we test the model in the laboratory experiments ?**

## Three conditions in the seesaw baryogenesis

Heavy Majorana neutrinos  $N_R$  with  $M_R = (M_1, M_2, \dots, M_N)$ .

$\nu_{Li}$  ( $i = 1 \sim 3$ ): left-handed neutrinos.

$l = e, \mu, \tau$ : charged leptons with  $m_l = (m_e, m_\mu, m_\tau)$ .

$$\mathcal{L} = -[\overline{\nu}_L m_D N_R + \frac{1}{2} \overline{(N_R)^c} M_R N_R + \overline{l}_L m_l l_R] - h.c. [\overline{N}_R m_D^\dagger \nu_L \dots] \quad (2)$$

### 1. Lepton number violating mass term.

$$N_R \rightarrow \exp(i\theta) N_R :$$

$$\overline{(N_R)^c} M_R N_R \rightarrow \exp(i2\theta) \overline{(N_R)^c} M_R N_R.$$

### 2. CP violating Dirac mass terms

$$m_D$$

$\text{Im}(m_D) \neq 0 \rightarrow$  CP violation:

$$CP[\overline{\nu}_L m_D N_R] CP^{-1} = \overline{N}_R m_D^T \nu_L.$$

$$L_{\text{primordial}} \sim \Gamma[N \rightarrow l^- \phi^+] - \Gamma[N \rightarrow l^+ \phi^-]$$

$L_{\text{primordial}}$  is generated with the mechanism known as "Direct CP violation" in K and B physics.

### Direct CP violation

$$\begin{aligned} & \Gamma[B^+ \rightarrow K^+ \pi^0] - \Gamma[B^- \rightarrow K^- \pi^0] \\ & \sim \text{Im}(a_0 a_1^*) \sin(\delta_0 - \delta_1) \end{aligned}$$

$$\text{Amp.}(B^+ \rightarrow K^+ \pi^0) = a_0 \exp(i\delta_0) + a_1 \exp(i\delta_1)$$

$$\text{Amp.}(B^- \rightarrow K^- \pi^0) = a_0^* \exp(i\delta_0) + a_1^* \exp(i\delta_1)$$

$\text{Im}(a_0 a_1^*)$  : weak interaction CP violating phase (In SM, KM phase).

$\delta$ : strong phase.(phase due to threshold effect of intermediate or final states interactions).



The formulae for  
 "Direct CP violation" (asymmetry).  
 $N = 2$  case.

$$\epsilon_1 = \frac{\Gamma_1 - \overline{\Gamma}_1}{\Gamma_1 + \overline{\Gamma}_1} = -\frac{3M_1}{2M_2V^2} \frac{\text{Im}[\{(m_D^\dagger m_D)_{12}\}^2]}{(m_D^\dagger m_D)_{11}},$$

where  $V = \sqrt{4\pi}v$  with  $v = 246$  GeV.

$$-\mathcal{L} = y_D \bar{\psi}_L \tilde{\phi} N_R$$

$$y_D \sim \frac{m_D}{v}$$

$L_{\text{primordial}}$  may be only generated in the early hot universe.  $T > M_R$ .

**3.  $L_{\text{primordial}}$  is produced under the non-equilibrium condition.** In the thermal equilibrium, the decay process and production process are balanced and the net production of lepton number is zero.  $N \leftrightarrow l^- \phi^+$   
 In the expanding universe, e.g., in the Friedmann Universe:

$$ds^2 = dt^2 - a(t)^2(dx^2 + dy^2 + dz^2).$$

Time evolution of the number density of heavy Majorana neutrino  $n(t)$  and lepton number density  $L(t)$ .

$$\begin{aligned} \frac{dn(t)}{dt} + 3Hn(t) &= -\Gamma[N](n - n_{eq}). \\ \frac{dL(t)}{dt} + 3HL(t) &= (\Gamma[N \rightarrow l^- \phi^+] - \Gamma[N \rightarrow l^+ \phi^-])(n - n_{eq}) \quad (4) \end{aligned}$$

$$t = \sqrt{\frac{45}{16\pi^3 g^*}} \frac{M_{pl}}{T^2}, \quad \text{Initial condition: } \odot \quad T = 10^{16} \text{ (GeV)}$$

$$n = n_{eq}, L = 0$$

The result depends on  $\Gamma[N]$ ,

" Direct CP violation"  $\epsilon = \frac{\Gamma[N \rightarrow l^- \phi^+] - \Gamma[N \rightarrow l^+ \phi^-]}{\Gamma[N \rightarrow l^- \phi^+] + \Gamma[N \rightarrow l^+ \phi^-]}$

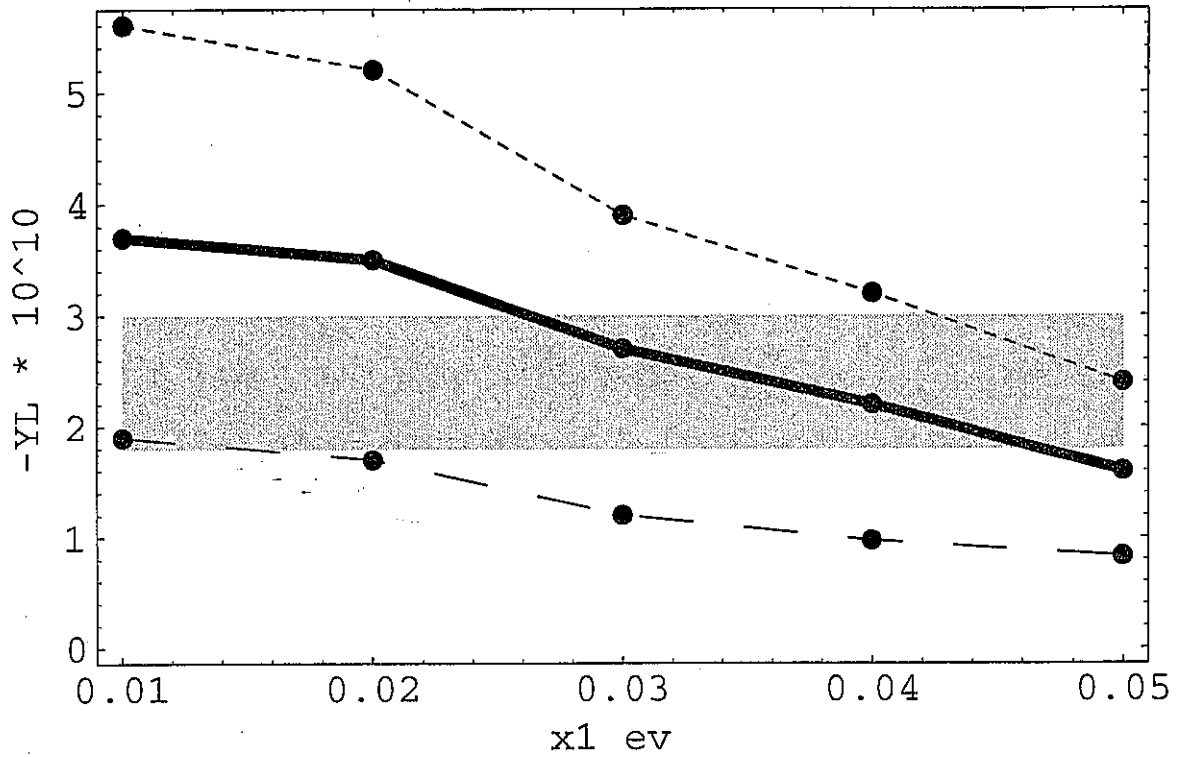
and expansion rate of the universe (Hubble)

$H = \frac{\dot{a}}{a}$ .  $L_{primordial}$  may be easily produced if the non-equilibrium condition is satisfied.

$$H > \Gamma[N] \quad (5)$$

Fig.1  $\rightarrow -Y_L = \frac{L}{s}$  ( $s$  entropy density).

$$x_1 = \frac{(m_D^\dagger m_D)_{11}}{M_1} = \frac{\Gamma[N] V^2}{M_1^2}. \quad V = \sqrt{4\pi v}, \quad v = 256 \text{ GeV}.$$



The maximum possible lepton number density ( $-Y_L$ ) as a function of  $x_1$  for three different  $M_1$ . From up to bottom,  $M_1 = 3, 2, 1 \times 10^{11}$  GeV.

Is the CP violation of baryogenesis related to CP violation in low energy ?

$$P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) = 4J \left[ \sin\left(\frac{\Delta m_{12}^2 L}{2E}\right) + \sin\left(\frac{\Delta m_{23}^2 L}{2E}\right) + \sin\left(\frac{\Delta m_{31}^2 L}{2E}\right) \right].$$

$$J = \text{Im} \left( U_{e1} U_{\mu 1}^* U_{e2}^* U_{\mu 2} \right).$$

Branco Morozumi Nobre and Rebelo (Nucl.Phys.B617) show, in general, there may be some relation.

$$\mathcal{L} = \bar{l}^i m_{l_i} l^i + \bar{\nu}^i m_{Dij} N_{Rj} + \frac{1}{2} (N_{Rj}^-)^c M_j N_{Rj}, \quad (6)$$

(3, N)model.

The basis all flavor mixing and CP origin are in the Dirac mass term of neutrino  $m_D$ .

$$M_j = \text{Diag.}(M_1, M_2, \dots, M_N), m_l = \text{Diag.}(m_e, m_\mu, m_\tau).$$

$$L_{ini} \sim \text{Im}(m_D^\dagger m_D)_{ij}^2, \quad (i \neq j) :$$

$$-U^\dagger m_D \frac{1}{M} m_D^T U^* = \begin{pmatrix} n_1 & 0 & 0 \\ 0 & n_2 & 0 \\ 0 & 0 & n_3 \end{pmatrix}$$

$$L \leftarrow \text{Im}(m_D) \rightarrow \text{Im}(U) \rightarrow J.$$

Both  $J$  and  $L$  are related to CP phases in  $m_D$ .

- How are  $J$  and  $B(L)$  related ? Identifying the low energy CP phase which is related to leptogenesis.
  - For quantitative analysis on the correlation, we must fix the parameters of the model. What kind of physical observables are needed to fix all parameters ? Write  $J$  in the parametrization which apply for the most general case.
  
- Give a specific example which shows a direct link " correlation" between  $J$  and  $B$ .
  - The general case and numerical analysis:

**Independent number of parameters in neutrino sector of seesaw model.**

<i>model</i>	(3,N)	(3,3)	(3,2)
$M$	N	3	2
$Re(m_D)$	3 N	9	6
$Im(m_D)$	3N-3	6(3)	3 (1)
<i>total</i>	7N-3	18	11

(blue) Independent number of CP phases for leptogenesis.  $(m_D^\dagger m_D)_{ij} (i \neq j)$  Low energy observables:  $= 7 \ll 11$  (Minimal model:  $N = 2$ ).

<i>mixing angles</i>	$2 +  U_{e3} $
$\Delta m^2$	2 (solar, atm.)
<i>neutrinoless double <math>\beta</math></i>	$ (m_{eff})_{ee} $
<i>CP violation in oscillation</i>	1 ( $\Delta P$ )
<i>total</i>	7

We adopt four high energy physical quantities as input. Heavy Majorana masses ( $M_1, M_2$ ) and their decay widths ( $\Gamma_1, \Gamma_2$ .)



We can fix 11 parameters of the minimal seesaw (3,2) model.

7 (low energy observables) + 4 (high energy observables) = 11. Besides  $M_1, M_2$ , 9 parameters in  $m_D$  is parametrized as:

$$m_D = \begin{pmatrix} m_{11} & m_{12} \\ m_{21} & m_{22} \\ m_{31} & m_{32} \end{pmatrix} \\ = U_L \begin{pmatrix} 0 & 0 \\ m_2 & 0 \\ 0 & m_3 \end{pmatrix} V_R (= U_L m V_R).$$

$U_L$ :  $3 \times 3$ ,  $V_R$ :  $2 \times 2$  unitary matrix.

$$U_L = O(\theta_{L23}) U(\theta_{L13}, \delta_L) O(\theta_{L12}) \\ \times \text{diag.}(1, \exp[-i\frac{\gamma_L}{2}], \exp[i\frac{\gamma_L}{2}])$$

$$V_R = \begin{pmatrix} c_R & s_R \\ -s_R & c_R \end{pmatrix} \begin{pmatrix} \exp(-i\frac{\gamma_R}{2}) & 0 \\ 0 & \exp(i\frac{\gamma_R}{2}) \end{pmatrix}.$$

Among 3 CP phases( $\delta_L, \gamma_L, \gamma_R$ )  $\gamma_R \rightarrow$  leptogenesis.

$$\begin{aligned}\epsilon_1 &\sim \Gamma[N_1 \rightarrow l^- \phi^+] - \Gamma[N_1 \rightarrow l^+ \phi^-] \\ &\sim -\text{Im}[(m_D^\dagger m_D)_{12}^2] \\ &\sim -(m_3^2 - m_2^2)^2 s_R^2 c_R^2 \sin 2\gamma_R.\end{aligned}$$

$\gamma_R$  has impact on low energy CP  $J$  through MNS,  $U$ .  $U$  is determined as a unitary matrix diagonalizing  $m_{eff} = -m_D \frac{1}{M} m_D^T$ .

$$U^\dagger m_{eff} U^* = \text{diag.}(n_1, n_2, n_3)$$

$$m_{eff} = -U_L (m V_R \frac{1}{M} V_R^T m^T) U_L^T$$

$$U = U_L K_R,$$

$$K_R^\dagger \begin{pmatrix} 0 & 0 & 0 \\ 0 & Z_{22} & Z_{23} \\ 0 & Z_{32} & Z_{33} \end{pmatrix} K_R^* = \begin{pmatrix} 0 & 0 & 0 \\ 0 & n_2 & 0 \\ 0 & 0 & n_3 \end{pmatrix}.$$

$Z_{ij} = Z_{ij}(M_i, m_i, \gamma_R, \theta_R)$   $n_1 = 0$  one neutrino is massless.

$K_R$  low energy manifestation of high energy physics.  $(\gamma_R, \theta_R, M_i, m_i)$ .

$$K_R = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta & \sin \theta \exp[-i\phi] \\ 0 & -\sin \theta \exp[i\phi] & \cos \theta \end{pmatrix} \\ \times \begin{pmatrix} 1 & 0 & 0 \\ 0 & \exp[i\alpha] & 0 \\ 0 & 0 & \exp[-i\alpha] \end{pmatrix}.$$

$$2\theta = \text{Arc.tan} \left( \frac{2|Z_{22}^* Z_{23} + Z_{23}^* Z_{33}|}{|Z_{33}|^2 - |Z_{22}|^2} \right) \Leftrightarrow \theta_R$$

$$\phi = \text{Arg.}(Z_{22}^* Z_{23} + Z_{23}^* Z_{33}) \Leftrightarrow \gamma_R$$

$$2\alpha = \text{Arg.}(\cos^2 \theta Z_{22} + \sin^2 \theta Z_{33} \exp(-2i\phi) \\ - \sin 2\theta Z_{23} \exp(-i\phi)) \Leftrightarrow \gamma_R$$

All three CP phases contribute to  $J = \text{Im}[U_{e1} U_{\mu 1}^* U_{e2}^* U_{\mu 2}]$ .

$$J = \frac{1}{8} \sin 2\theta_{L12} \sin 2\theta_{L13} [c_{L13} \cos 2\theta \sin \delta_L \sin 2\theta_{L23} \\ + c_{L12} \sin 2\theta \sin(\delta_L - \gamma_L - \phi) \cos 2\theta_{L23} \\ - \frac{1}{2} s_{L12} s_{L13} \sin 2\theta \sin 2\theta_{L23} \sin(2\delta_L - \gamma_L - \phi)] \\ + \frac{1}{8} \sin 2\theta \sin 2\theta_{L23} \sin(\gamma_L + \phi) \\ \times (\sin 2\theta_{L12} c_{L13} s_{L12} - \sin 2\theta_{L13} s_{L13} c_{L12}).$$

We can get the values of  $\theta$  and  $\phi$  from  
 1) high energy physics input  $(M_1, M_2, \Gamma_1, \Gamma_2)$  +  
 2) neutrino mass eigenvalue eq.

$$\det.(m_{eff} m_{eff}^\dagger - n^2) = 0, \quad (n_2 = \sqrt{\Delta m_{sol.}^2} = 7 \times 10^{-3} (eV), n_3 = \sqrt{\Delta m_{atm.}^2} = 5 \times 10^{-2}).$$

With the input we can determine  $(m_2, m_3, \gamma_R, \theta_R)$  which leads to the prediction of  $\theta, \phi$  and also lepton number asymmetry  $\epsilon_1$ .

$$\cos 2\gamma_R = \frac{n_2^2 + n_3^2 - x_1^2 - x_2^2}{2(x_1 x_2 - n_2 n_3)}, \quad x_i = \frac{(m_D^\dagger m_D)_{ii}}{M_i} = \Gamma_i \left( \frac{V}{M_i} \right)^2$$

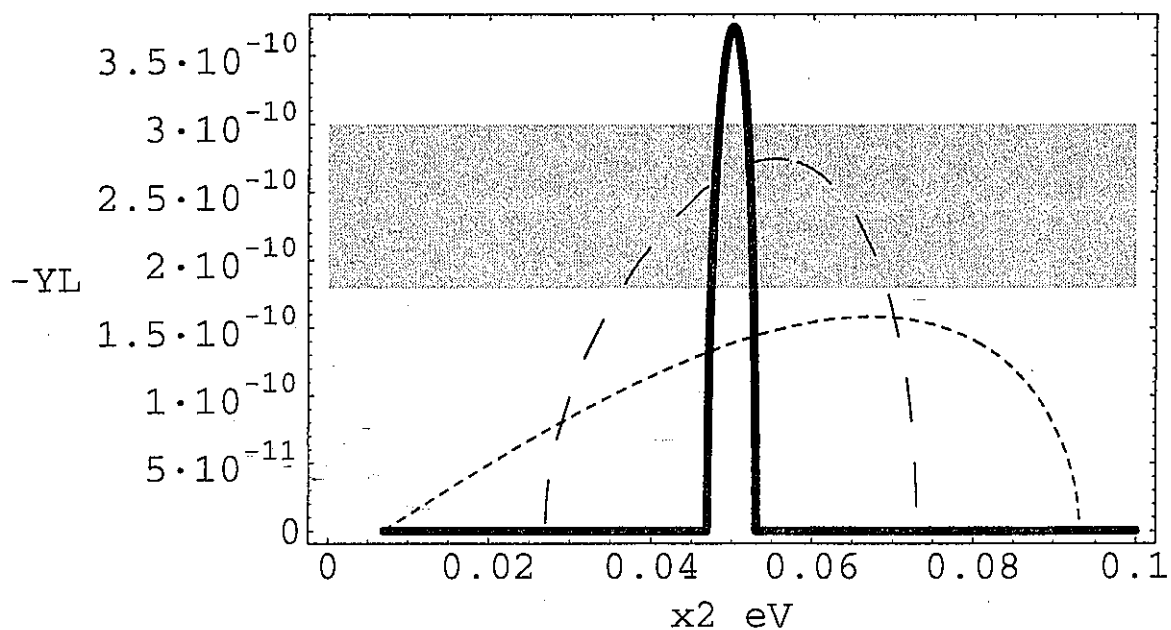
$$\epsilon_1 = -\frac{3M_1}{4x_1 V^2} \sqrt{((n_-)^2 - (x_-)^2) ((x_+)^2 - (n_+)^2)},$$

where  $n_\pm = n_3 \pm n_2$  and  $x_\pm = x_1 \pm x_2$ .

The range of  $(\theta, \phi)$  which is consistent with baryogenesis is limited.

The lower bound of  $M_1$  is obtained from lepton number asymmetry.

The decay width  $\Gamma_1$  (or  $x_1$ ) has upper bound from non-equilibrium condition (obtained from solving Boltzman equation.)



Lepton number density ( $-Y_L$ ) as a function of  $x_2$  for  $M_1 = 2 \times 10^{11}$  GeV. Solid line corresponds to  $x_1 = 0.01$  eV. The long dashed line corresponds to  $x_1 = 0.03$  eV and the dotted line corresponds to  $x_1 = 0.05$  eV.

$$U_{MNS} \simeq \begin{pmatrix} c_{L12} & s_{L12} & s_{L13}e^{-i\delta_L} + s_{L12}s_{\theta}e^{-i\phi'} \\ -s_{L12}c_{L23} & c_{L12}c_{L23} & s_{L23} \\ s_{L12}s_{L23} & -c_{L12}s_{L23} & c_{L23} \end{pmatrix} \times P(\alpha', -\alpha')$$

$$|(U_{MNS})_{e3}| \simeq |s_{L13}e^{-i\delta} + \frac{s_{\theta}}{\sqrt{2}}e^{-i\phi'}|,$$

$$J \simeq \frac{1}{4} \left( s_{L13} \sin \delta_L + \frac{s_{\theta}}{\sqrt{2}} \sin \phi' \right),$$

$$|(m_{eff})_{ee}| \simeq \left| \frac{n_2}{2} e^{4i\alpha'} + n_3 (s_{L13}e^{-i\delta_L} + \frac{s_{\theta}}{\sqrt{2}}e^{-i\phi'})^2 \right|.$$

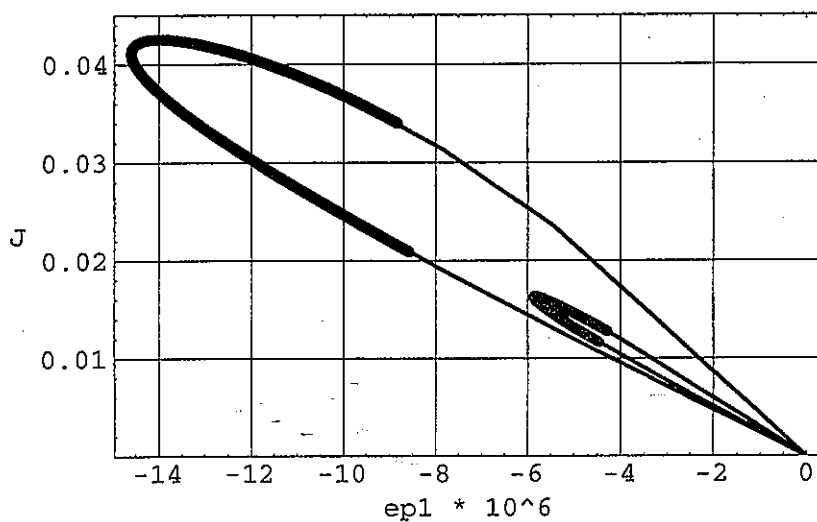
Here we give an example for the model in which  $J$  is determined by leptogenesis phase  $\gamma_R$ . Suppose  $U_L$  is a real orthogonal matrix as  $U_L = O_{23}O_{12}$ . MNS matrix  $U$  has the following form; Taking  $\theta_{L23} = \pi/4$ ,  $\theta_{L12} = \pi/4$ , we have;

$$U = \begin{pmatrix} \frac{1}{\sqrt{2}} & \frac{\cos \theta}{\sqrt{2}} & \frac{\sin \theta \exp[-i\phi]}{\sqrt{2}} \\ -\frac{1}{2} & \frac{\cos \theta - \sqrt{2} \sin \theta \exp[i\phi]}{2} & \frac{\sin \theta \exp[-i\phi] + \sqrt{2} \cos \theta}{2} \\ \frac{1}{2} & \frac{-\cos \theta + \sqrt{2} \sin \theta \exp[i\phi]}{2} & \frac{-\sin \theta \exp[-i\phi] + \sqrt{2} \cos \theta}{2} \end{pmatrix} \times P.(1, \alpha, -\alpha),$$

$$J = \sin 2\theta \sin \phi \frac{1}{8\sqrt{2}}$$

$$\gamma_R \neq 0 \Leftrightarrow J \neq 0.$$

The sign (excess of matter (antimatter)) has also correlation in the case.



Correlation be-

tween lepton number asymmetry  $\epsilon_1$  and neutrino CP violation  $J$ , for  $M_1 = 2 \times 10^{11}$  GeV. The large (small) contour corresponds to  $x_1 = 0.03(0.01)$  eV. The thick solid lines show the allowed region from baryogenesis.



## Summary

- leptogenesis phase ( $\gamma_R$ ) certainly affects the neutrino oscillation CP violation through ( $\phi$ )
- However, in general ground, we can not distinguish the phase coming from leptogenesis ( $\phi$ ) and another phase  $\gamma_L$  in  $U_L$  because only a combination appear as  $\phi + \gamma_L$ . The isolation must be done using some other quantities, double  $\beta$  decay etc.
- Leptogenesis formule is written in terms of physical quantities: Heavy neutrino mass and decay width and light neutrino mass.