

# **MOON Detector**

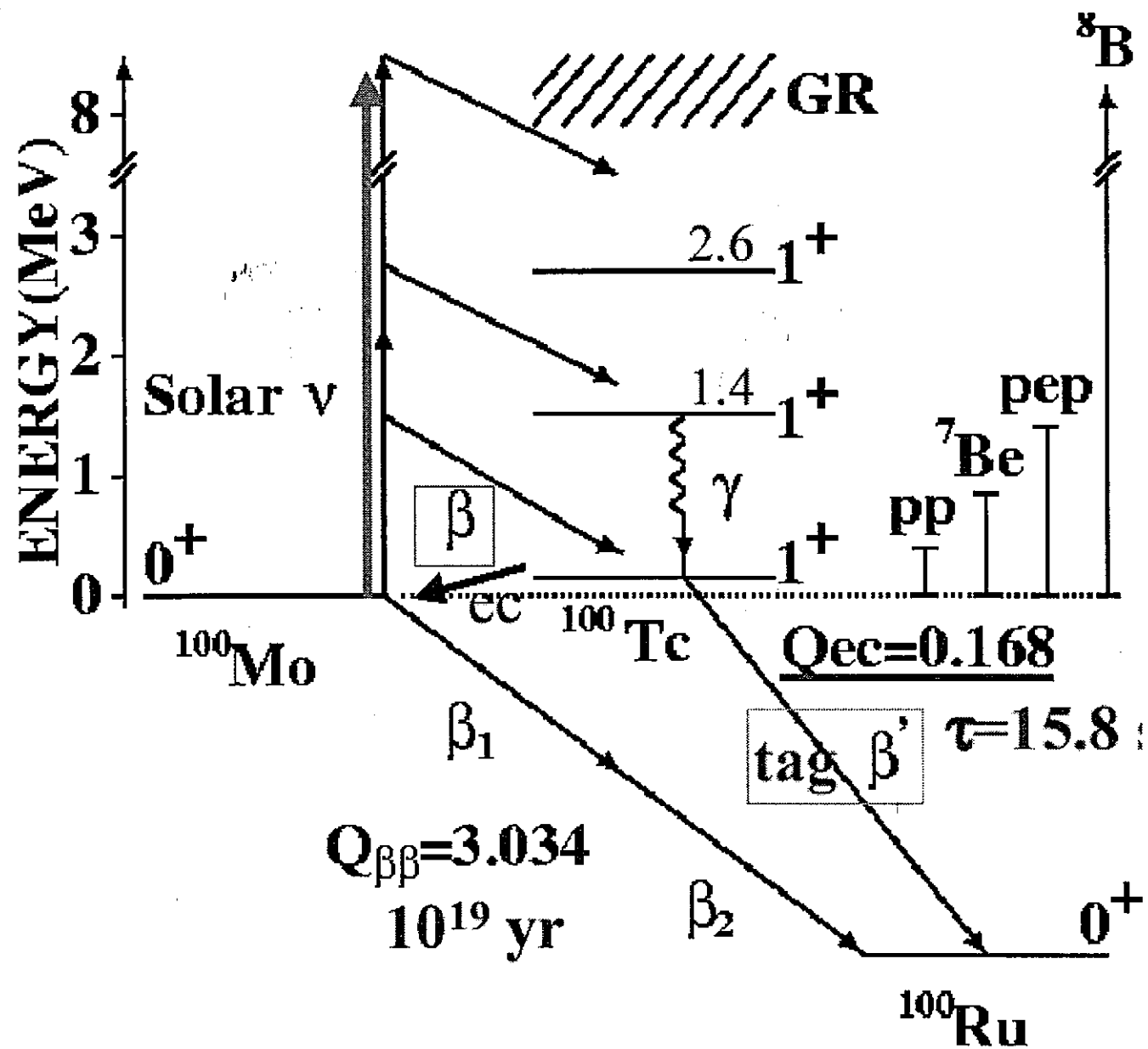
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# MOON Objectives

- Double beta ( $\beta\beta$ ) decays with  $m_\nu \sim 0.03$  eV.
- Low energy pp& $^7\text{Be}$  solar  $\nu_e$  and supernova  $\nu_e$  by inverse  $\beta$  followed by successive  $\beta$

## Spectroscopy of two $\beta$ rays from 1ton of $^{100}\text{Mo}$

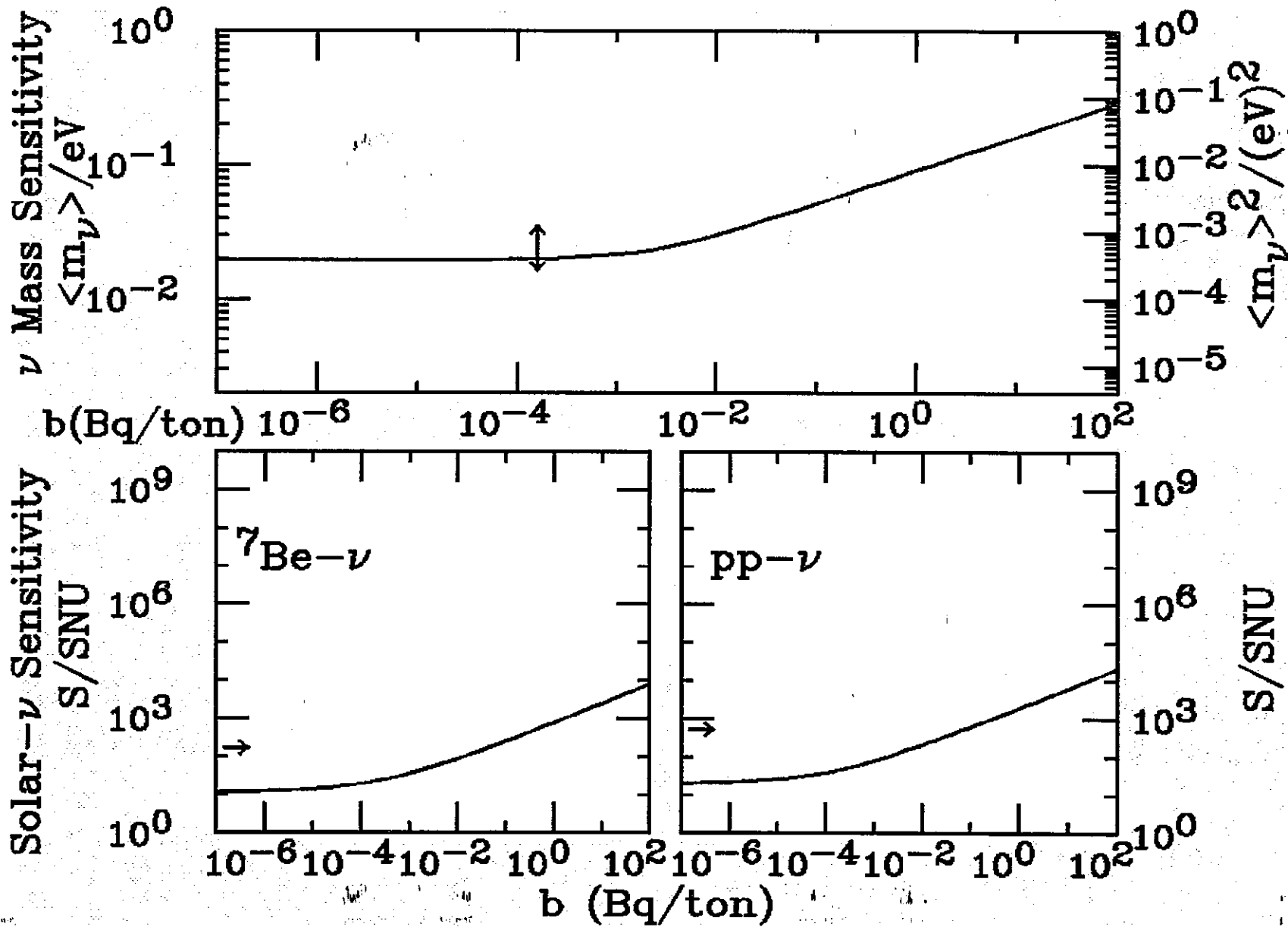
- large responses for  $\beta\beta$ - $\nu$  and low energy solar/supernova  $\nu_e$
- low threshold ( $Q_\beta$ )



# Major BG at the $0\nu\beta\beta$ window

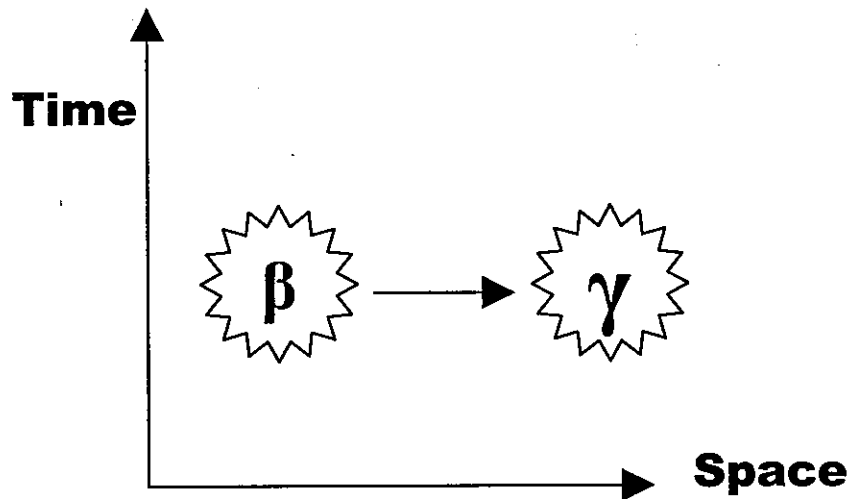
- **Radio Activities in Detector/Source**
  - 0.1ppt of U-Th: 1.25-0.45 mBq/t .
- **Cosmogenic Radio Activities**
  - Negligible at underground lab.
- **$2\nu\beta\beta$** 
  - $T_{1/2} \sim 10^{19}$  y

# $\nu$ -mass and solar $\nu$ sensitivities and U-Th impurities,



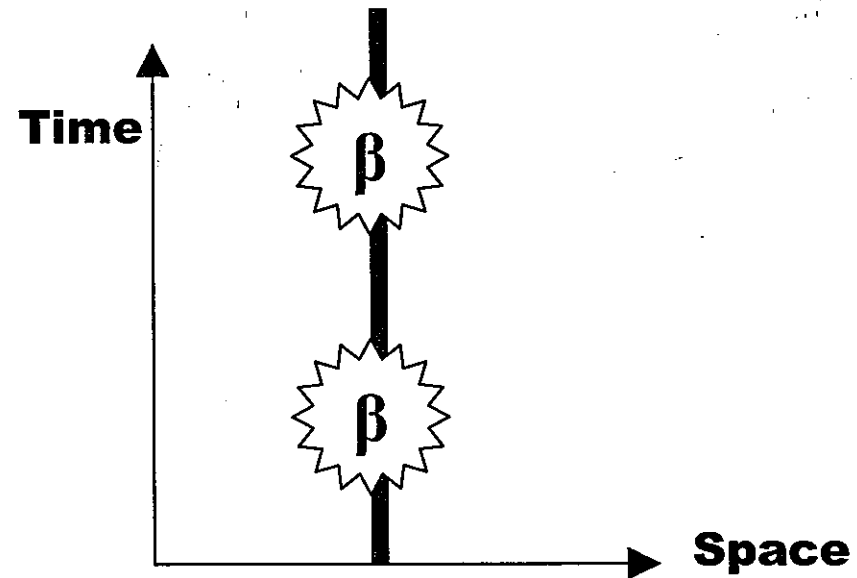
# Background Identification

- Most of alpha/beta decays are followed by gamma decay, which may cause second energy deposit in the scintillator.



$\beta$   $\gamma$  Successive decay  
Two energy deposit  
Distance  
Two energy  
Coincidence

- Many of alpha activities causes another signal at the same position by a following decay.



$\beta$   $\beta$  Successive decay  
Same position  
Two energy  
Interval

**Space-Time topology**

# Major BG at the $0\nu\beta\beta$ window

- **Radio Activities in Detector/Source**

- 0.1ppt of U-Th: 1.25-0.45 mBq/t .

- **Cosmogenic Radio Activities**

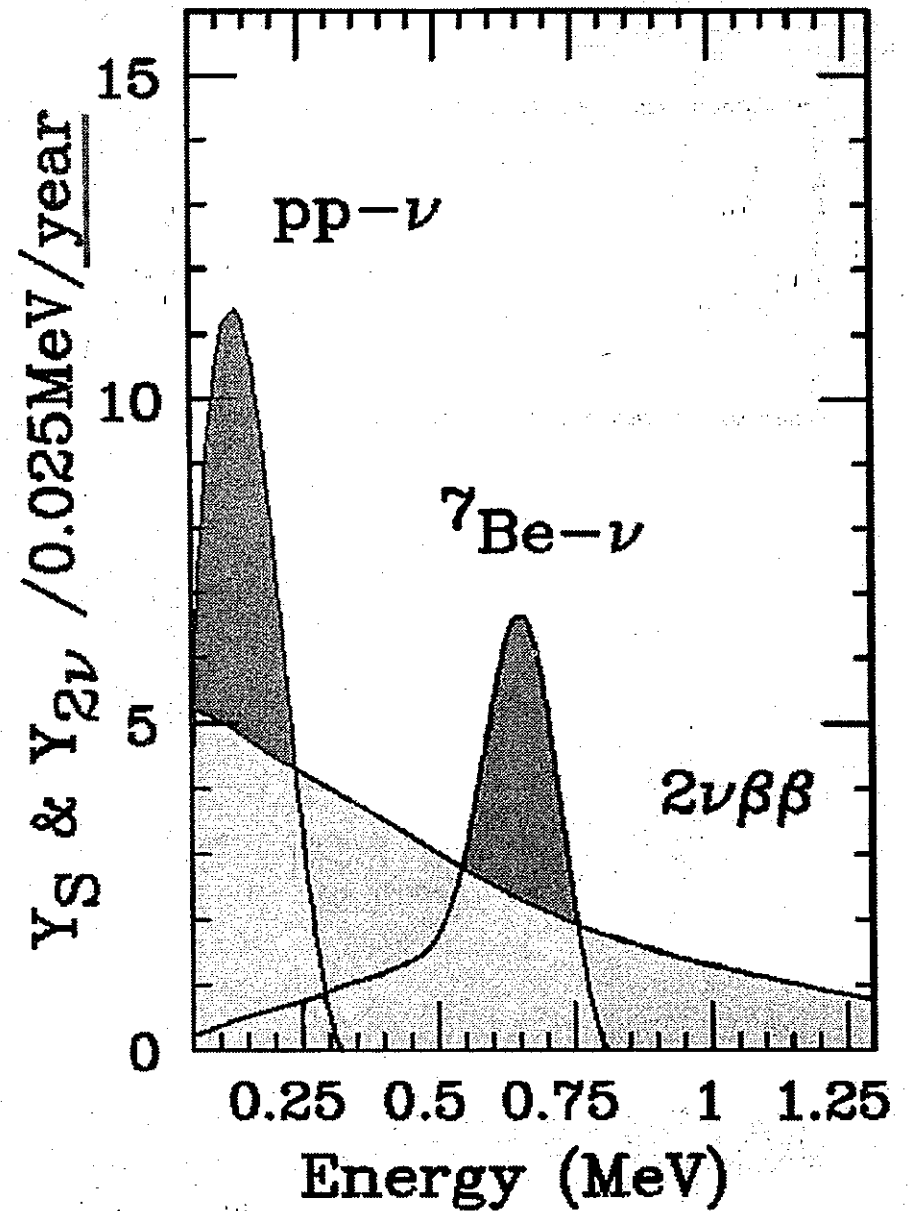
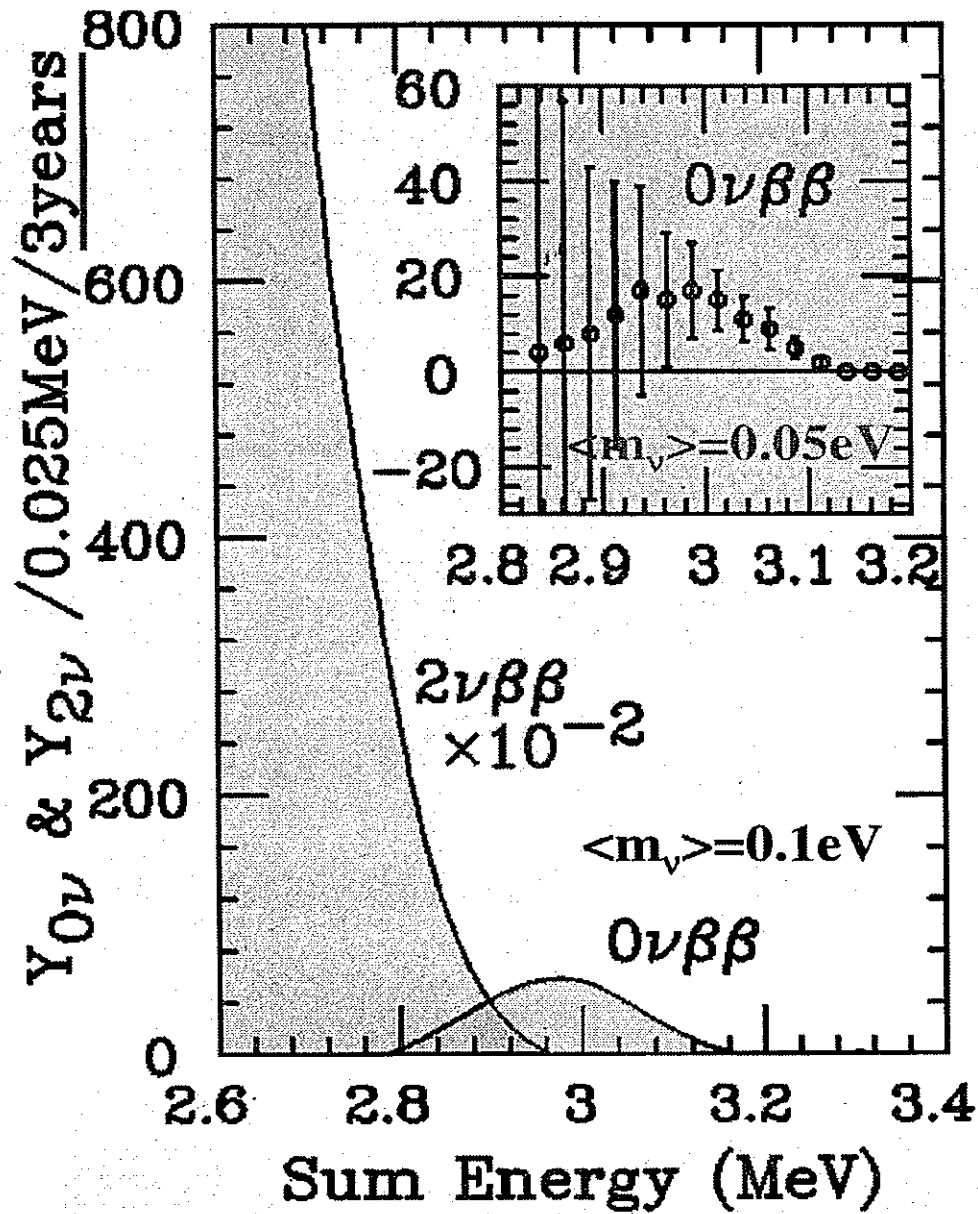
- Negligible at underground lab.

- **$2\nu\beta\beta$**

- $T_{1/2} \sim 10^{19} \text{y}$

# $^{100}\text{Mo}$ $\beta\beta$ and solar $\nu$

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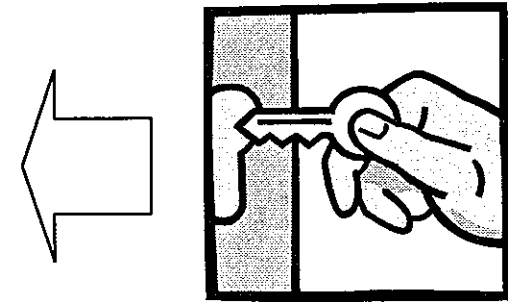




# Ultra-low background

- Extremely good event selection is necessary
  - Energy deposit
  - Event topology (Space-time)

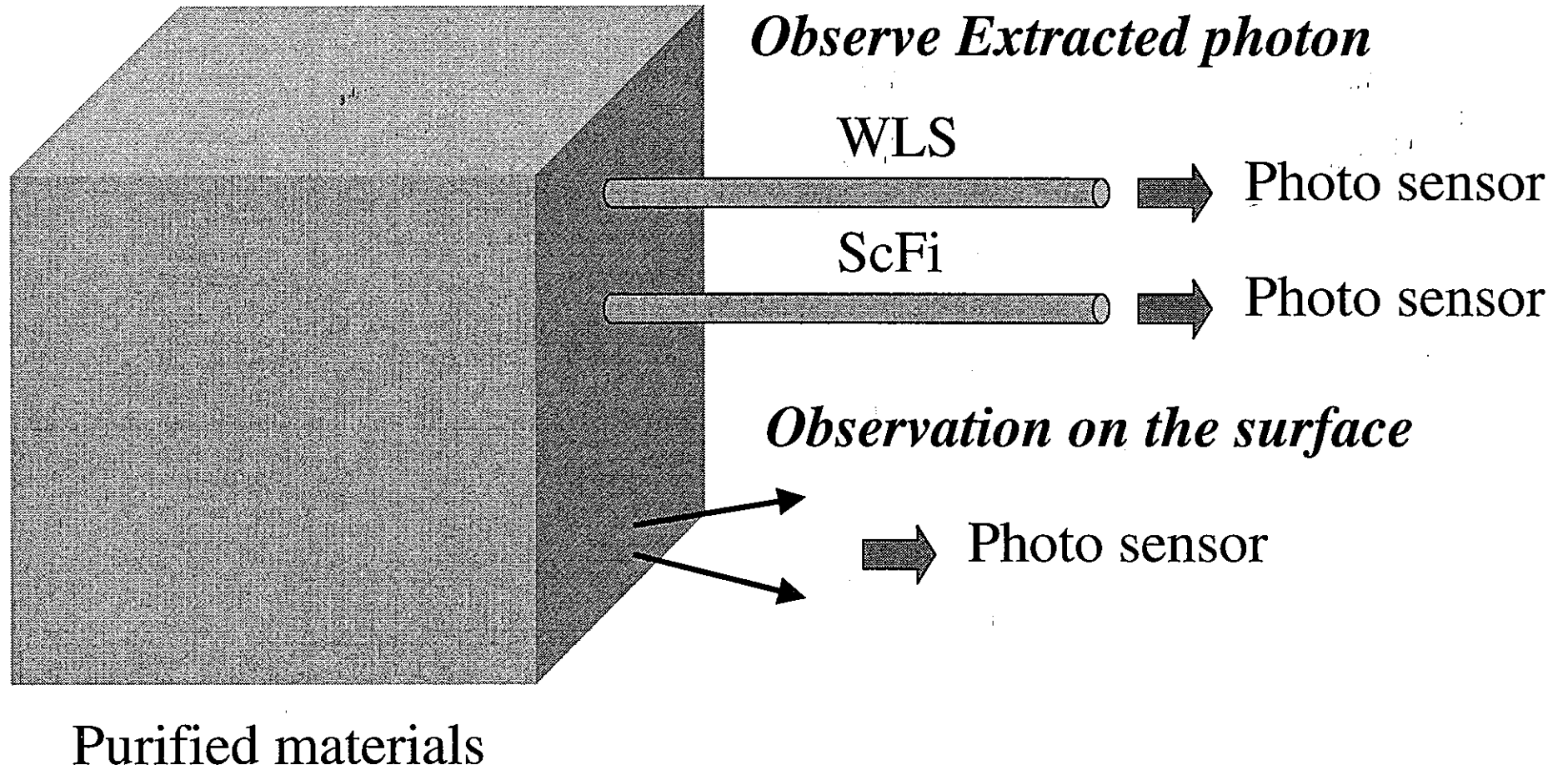
- Efficient photon collection
- High granularity



- Extremely good purification is necessary
  - Purified material (less kind of material)

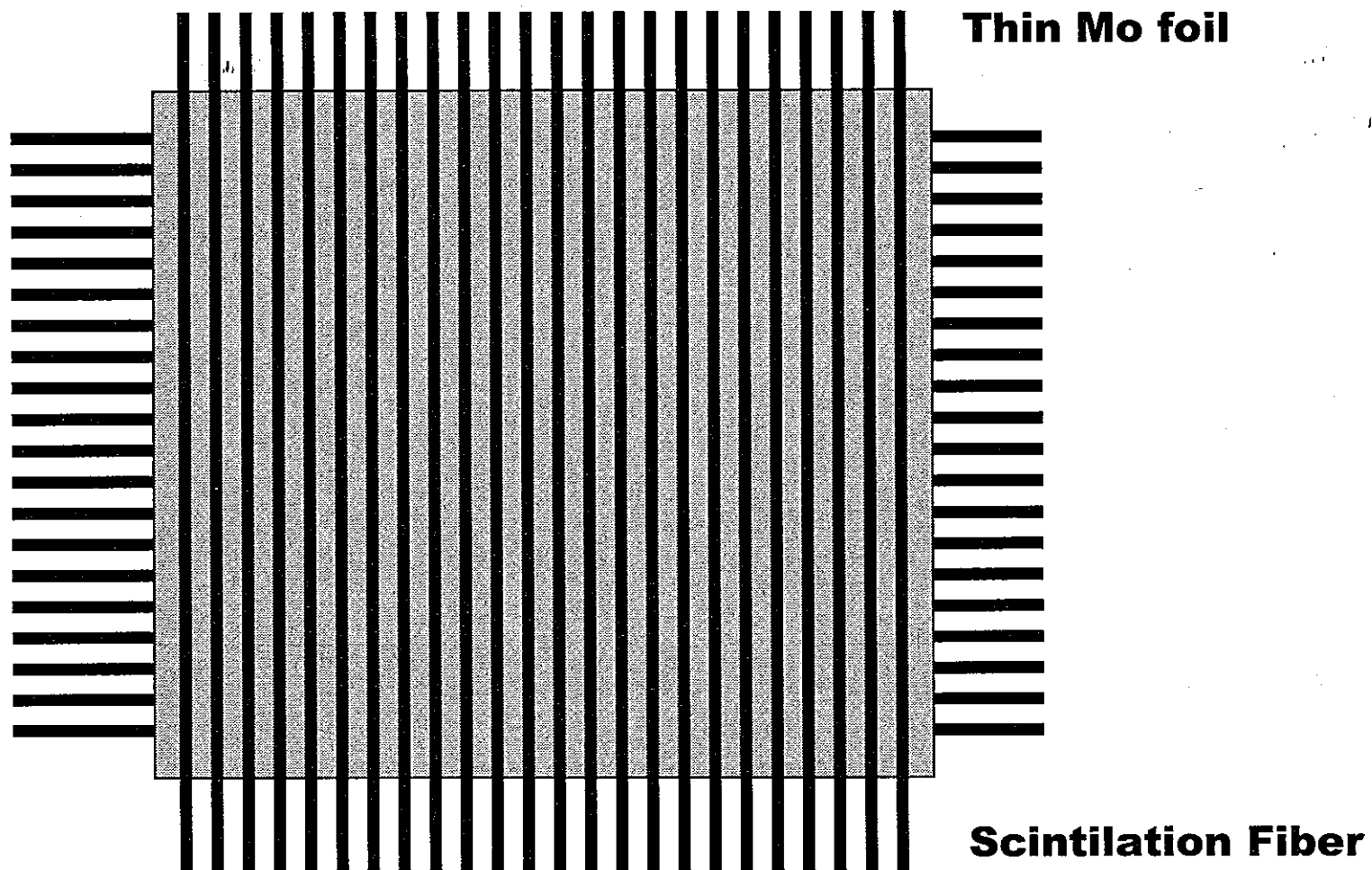
# Requirements for MOON

- Large mass of  $^{100}\text{Mo}$  ~1 ton
- Two  $\beta$  coin.  $\Delta t \sim \text{ns}$  for  $\beta\beta$ ,  
 $\Delta t \sim 1\text{-}30\text{s}$  solar- $\nu$ .
- Dynamic range  $E_\beta \sim 0.1\text{-}40 \text{ MeV}$
- Energy resolution 7% for 3MeV  $0\nu\beta\beta$
- Granularity  $\sim 10^9$
- Purity 0.1 ppt  $10^{-3} \text{ Bq/ton}$   
for U, Th isotopes.



# MOON

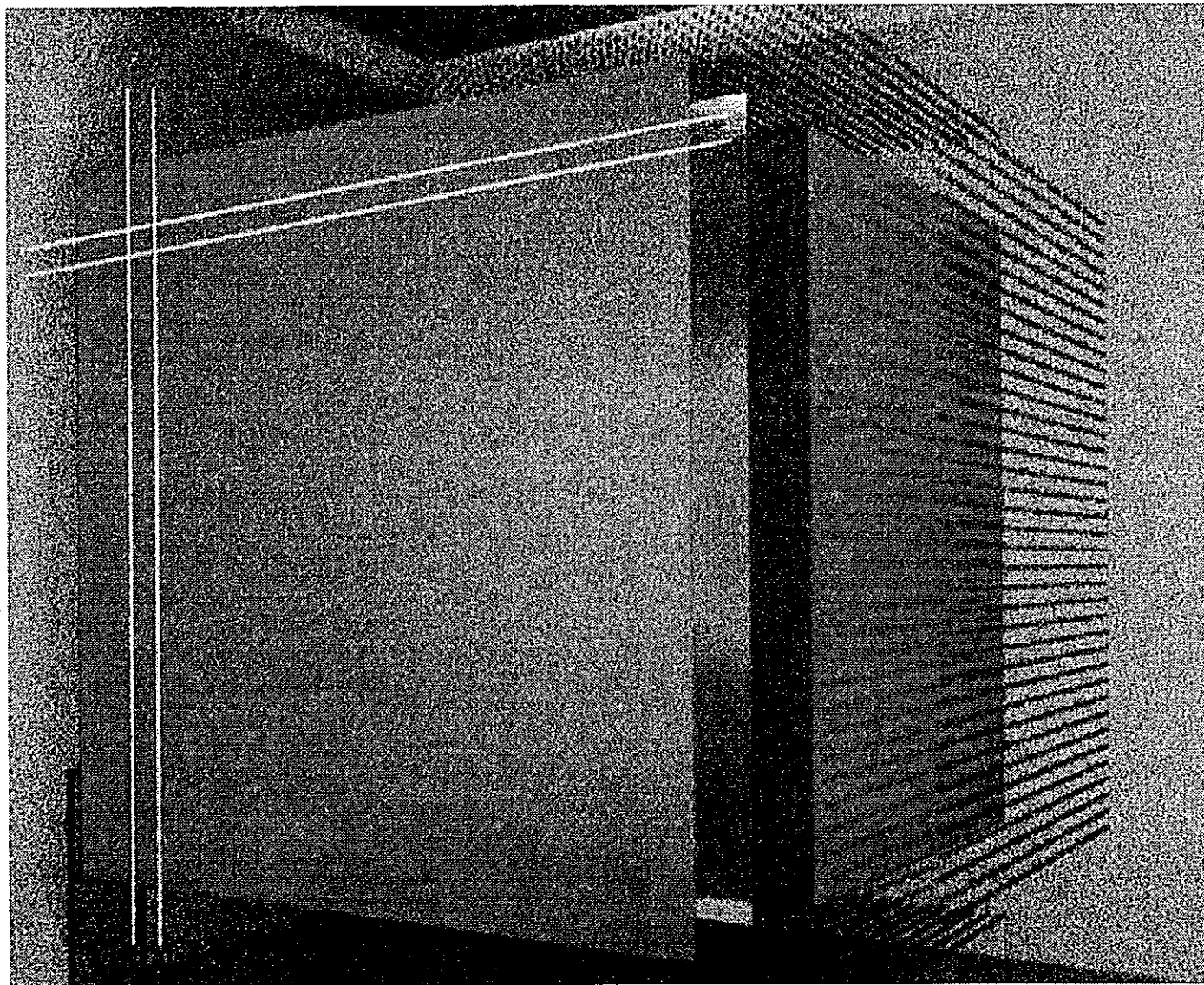
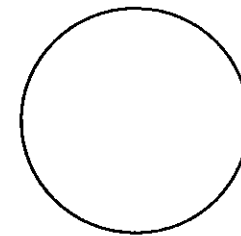
## Plastic fiber-Mo Ensemble



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# MOON

## Plastic fiber-Mo Ensemble



# A. Plastic Scintillation Fiber Detector

- Scintillator  $\lambda \sim 430$  nm

Transmission (both end)	$t \sim 0.144$
One unit = 250 modules,	2.5m-2.5m-0.5m
One module	2.5m-2.5m-2mm
Number of fibers /unit with	$n_f \sim 250K$

- PMT Hamamatsu M6A 16 anodes

With one anode: 4 fibers.  $N_{pm} \sim 10$  K

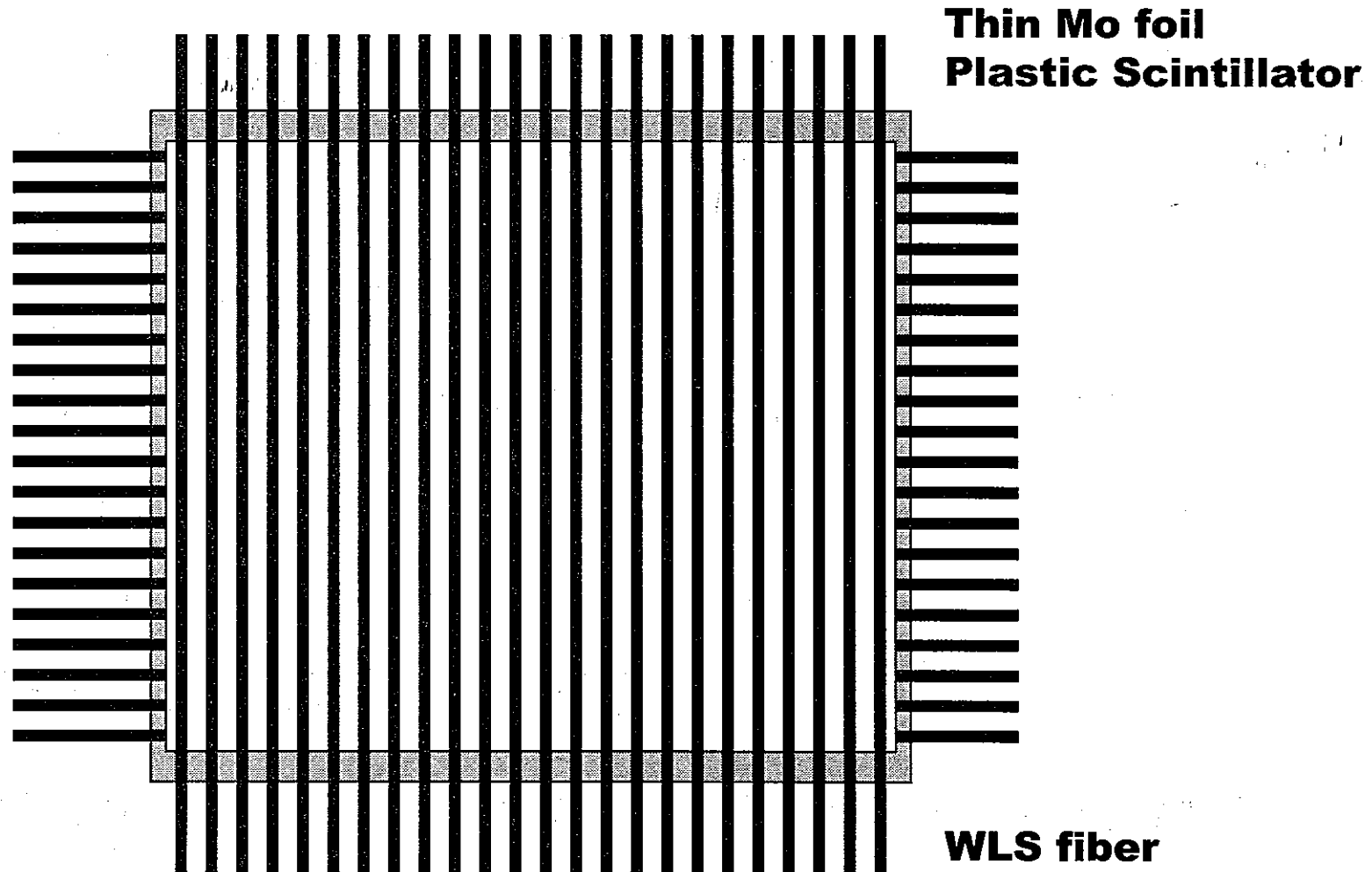
- Mo 30 mg/cm<sup>2</sup> <sup>100</sup>Mo (9.6%) <sup>100</sup>Mo 0.045 ton

85% enriched <sup>100</sup>Mo 0.4 ton

- MOON with 1 ton <sup>100</sup>Mo uses 2 units, 20K PMT

# MOON

## WLS fiber - Mo Ensemble



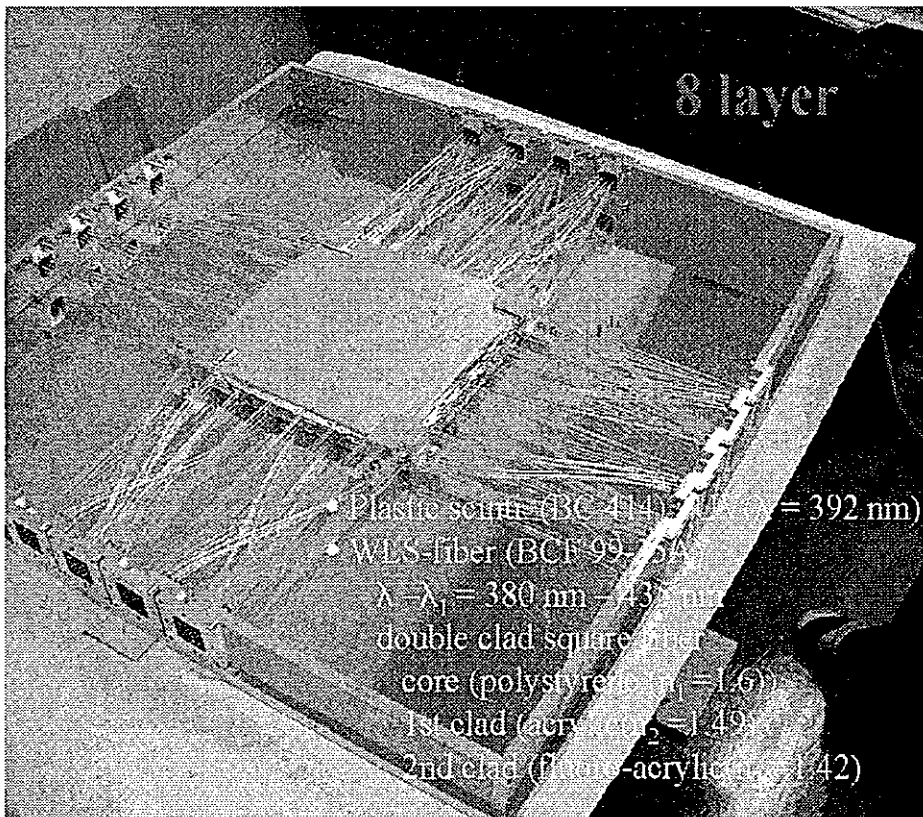
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# B. WLS (WaveLengthShifter) Read Out for Energy and Position

- Scintillator BS 414  $\lambda = 392 \text{ nm}$ 
  - One unit  $\approx 200$  modules, 2m-2m-0.5m.
  - One module 2m-2m-2.5mm.
- WLS BCS 99-35A  $\lambda - \lambda_1 = 380\text{nm}-435\text{nm}$ 
  - 2.7 cm interval x - y directions on PL
  - Double clad with polystyrene( $n_1=1.6$ ), acrylic( $n_2=1.49$ ), fluoro-acrylic( $n_3=1.42$ )  $t = 14.6 \%$
  - Number of fibers /unit with  $\sim 30\text{K}$ .
- PMT Hamamatsu M6A 16 anodes
  - Number of PMT  $\sim 620$ .

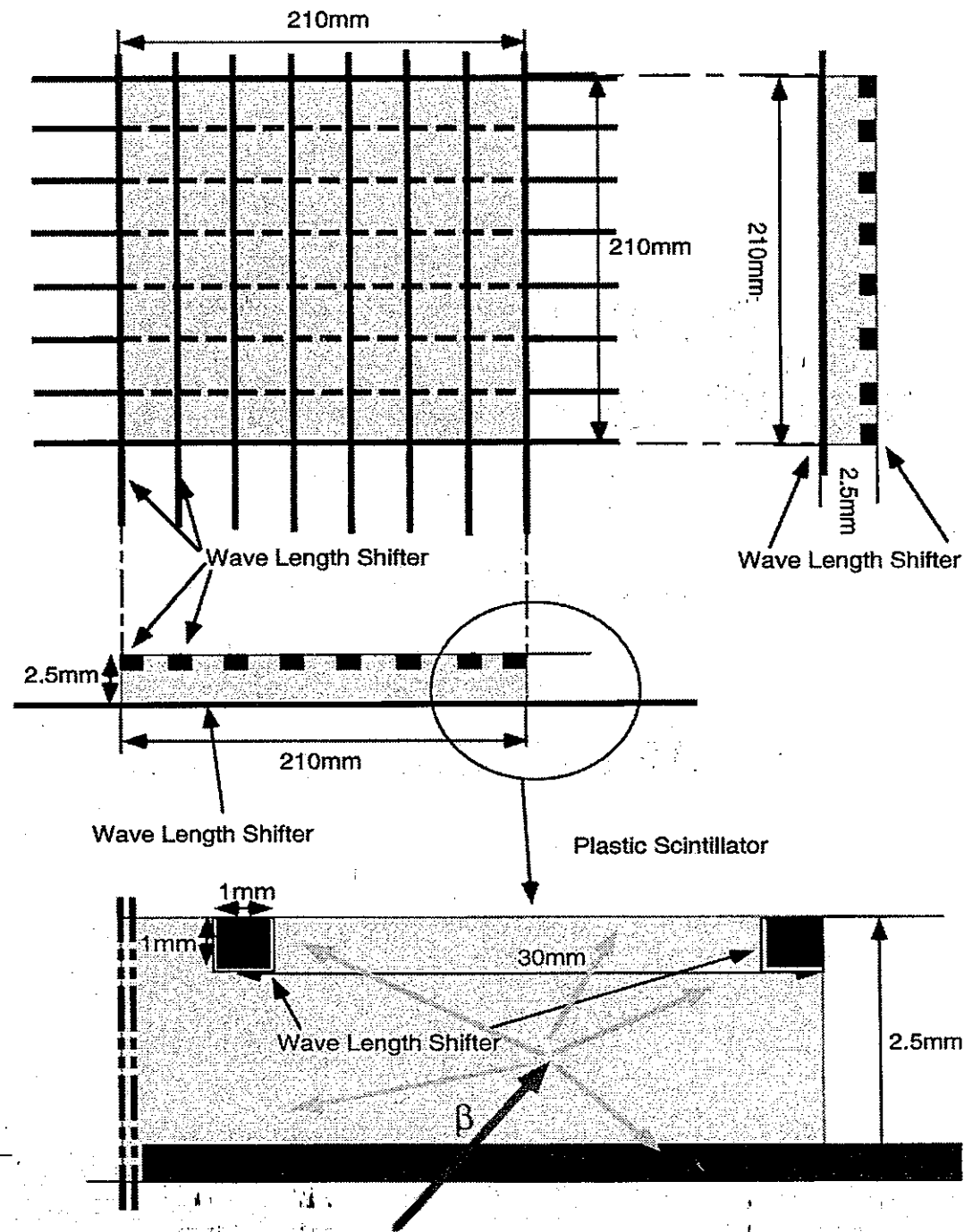


# WLS Test at OSAKA



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Good Energy and Position Resolution



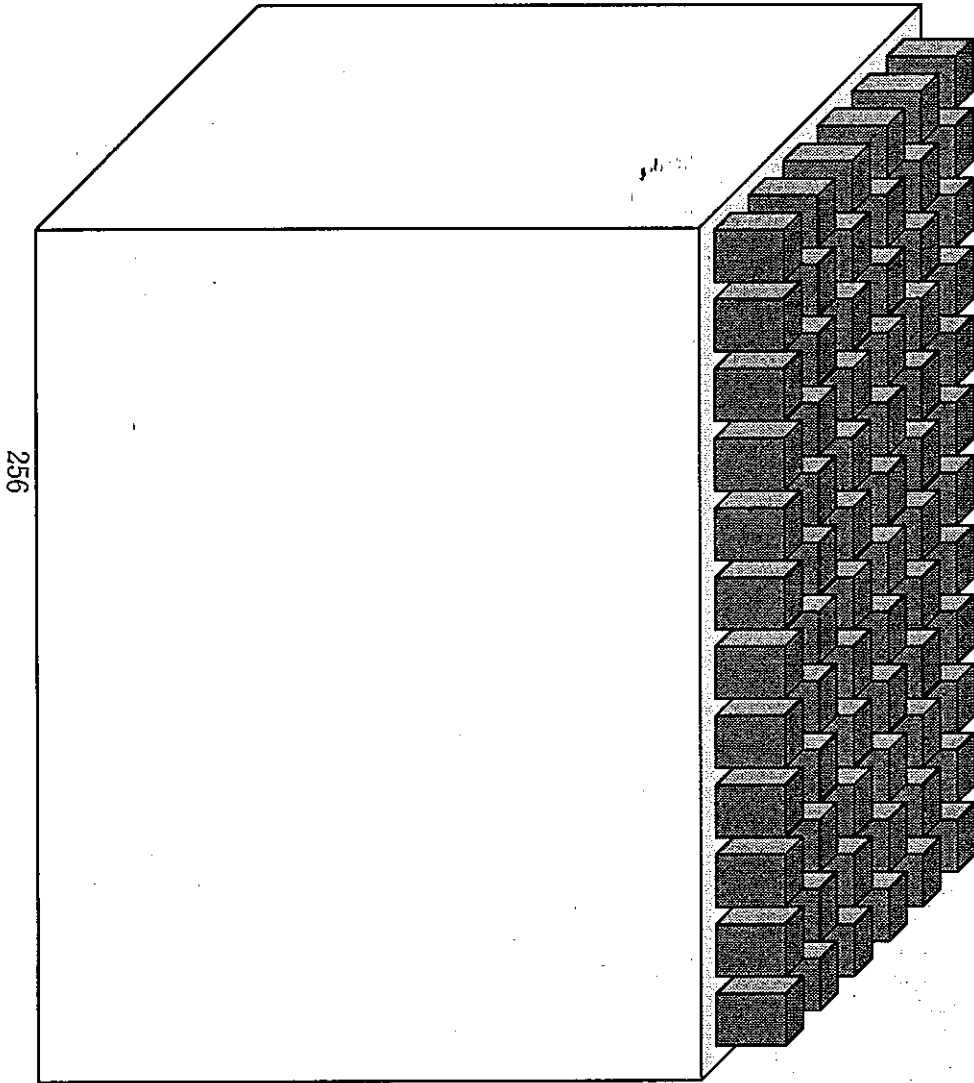
Efficiency for light collection of WLS

$\epsilon_w \sim 0.14$  (Expected to be 1.0)

Need increase to get a required energy resolution

# MOON

## $^{100}\text{Mo}$ loaded Liquid Scintillator

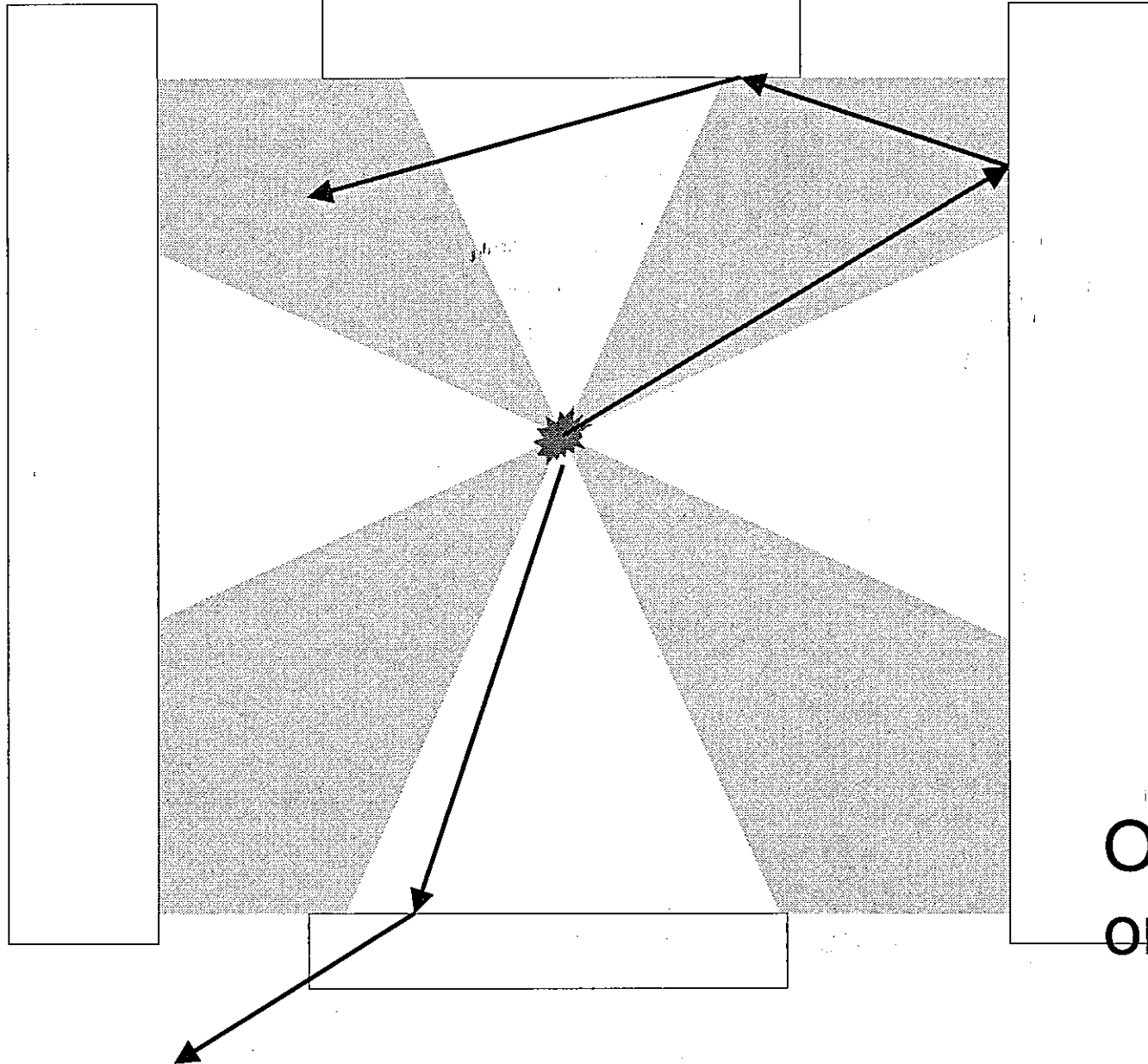


Photon detector array  
on all surfaces

1. Photon sensor is far from source.
2. No structure near source.
3. Less kind of material

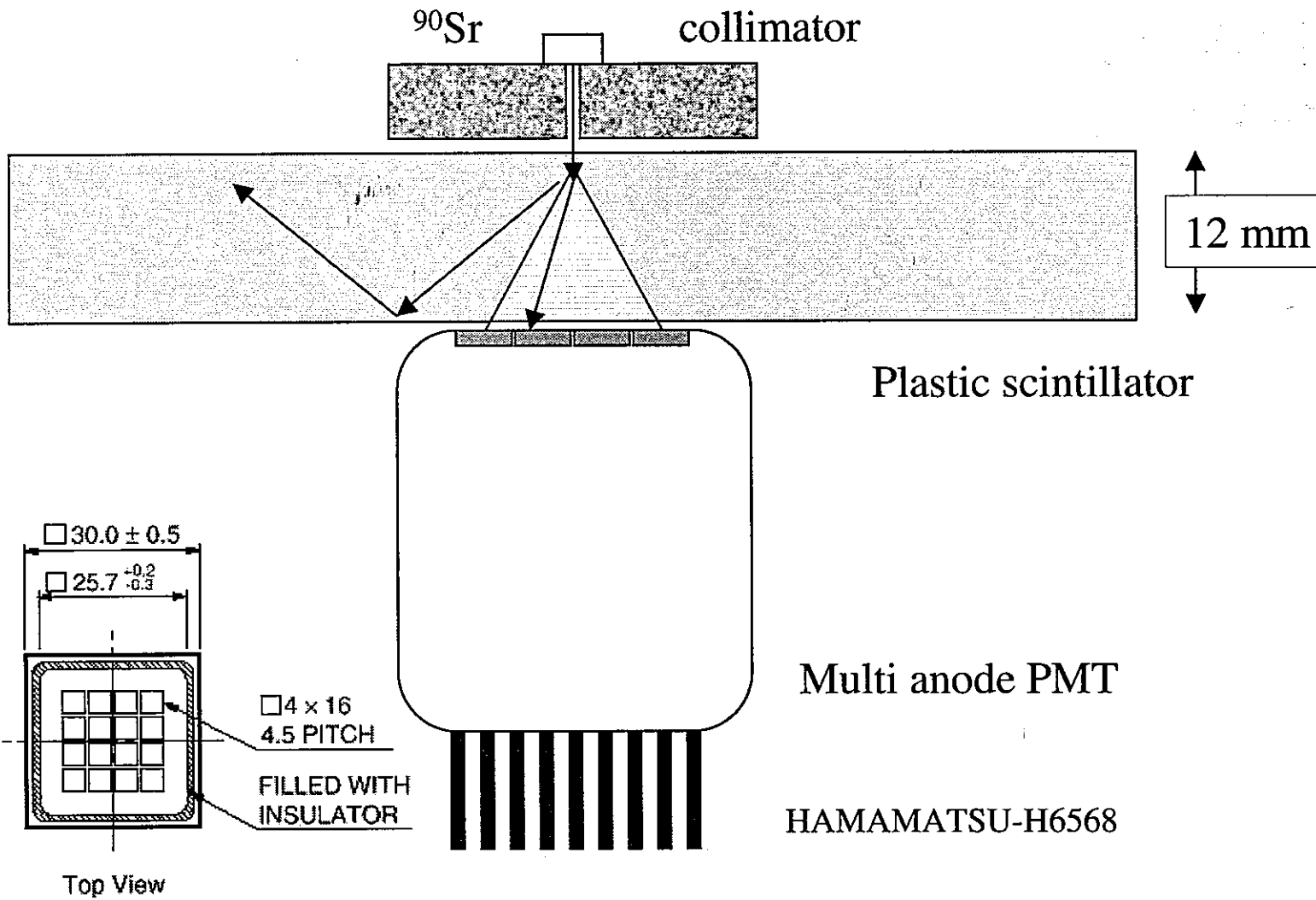
**9m<sup>3</sup> Liquid Scintillator**

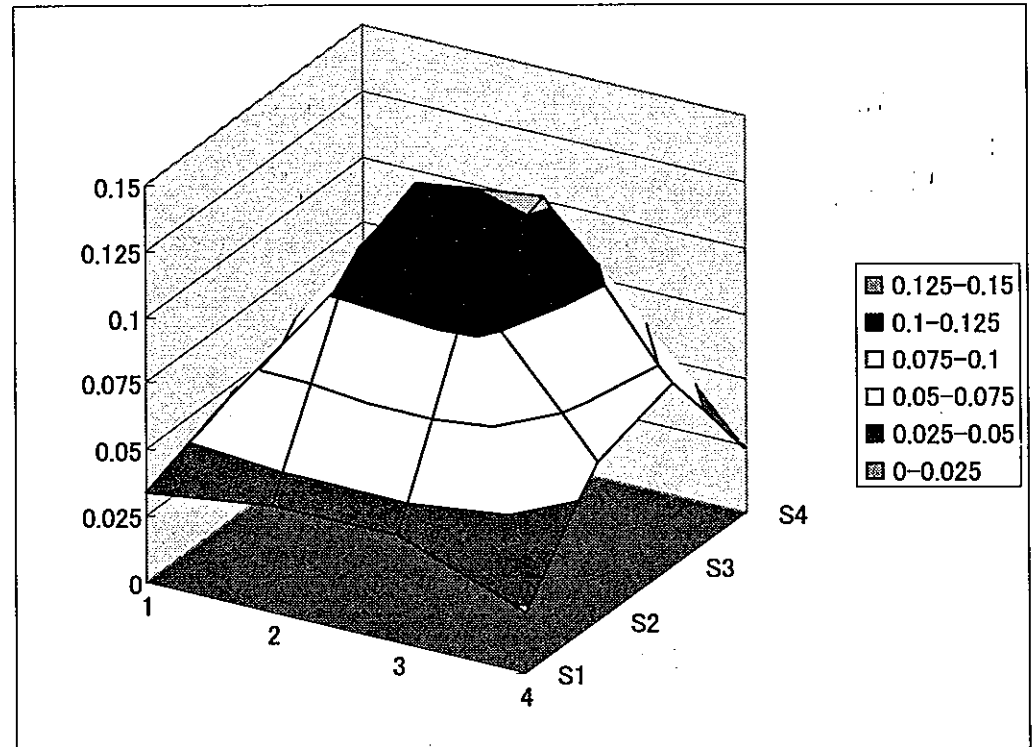
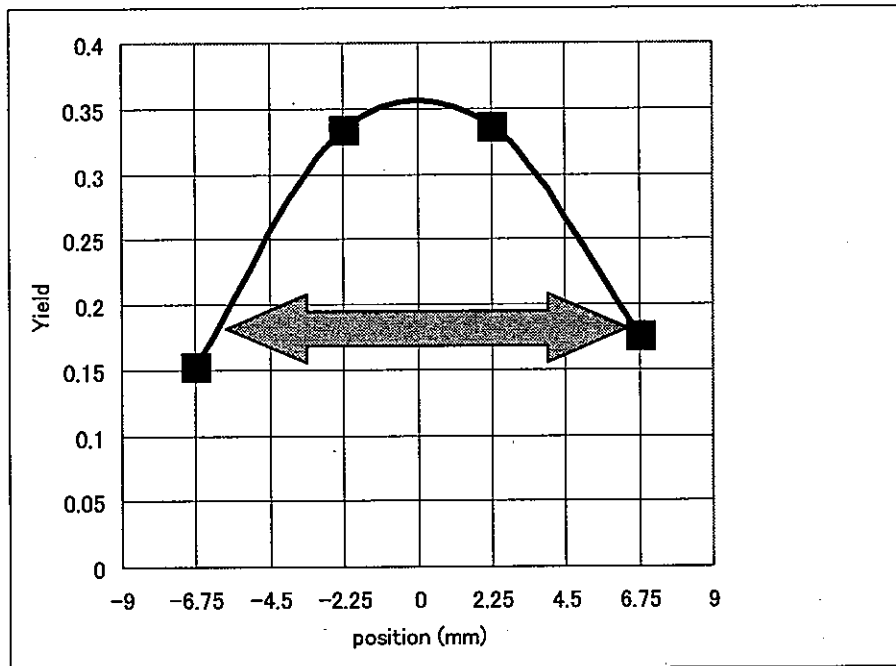
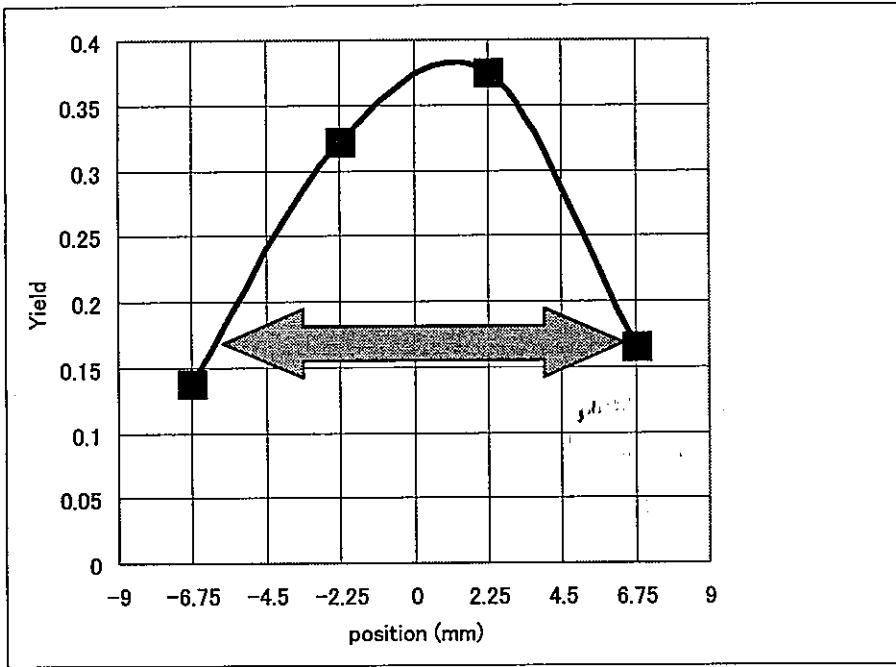
# Total reflection



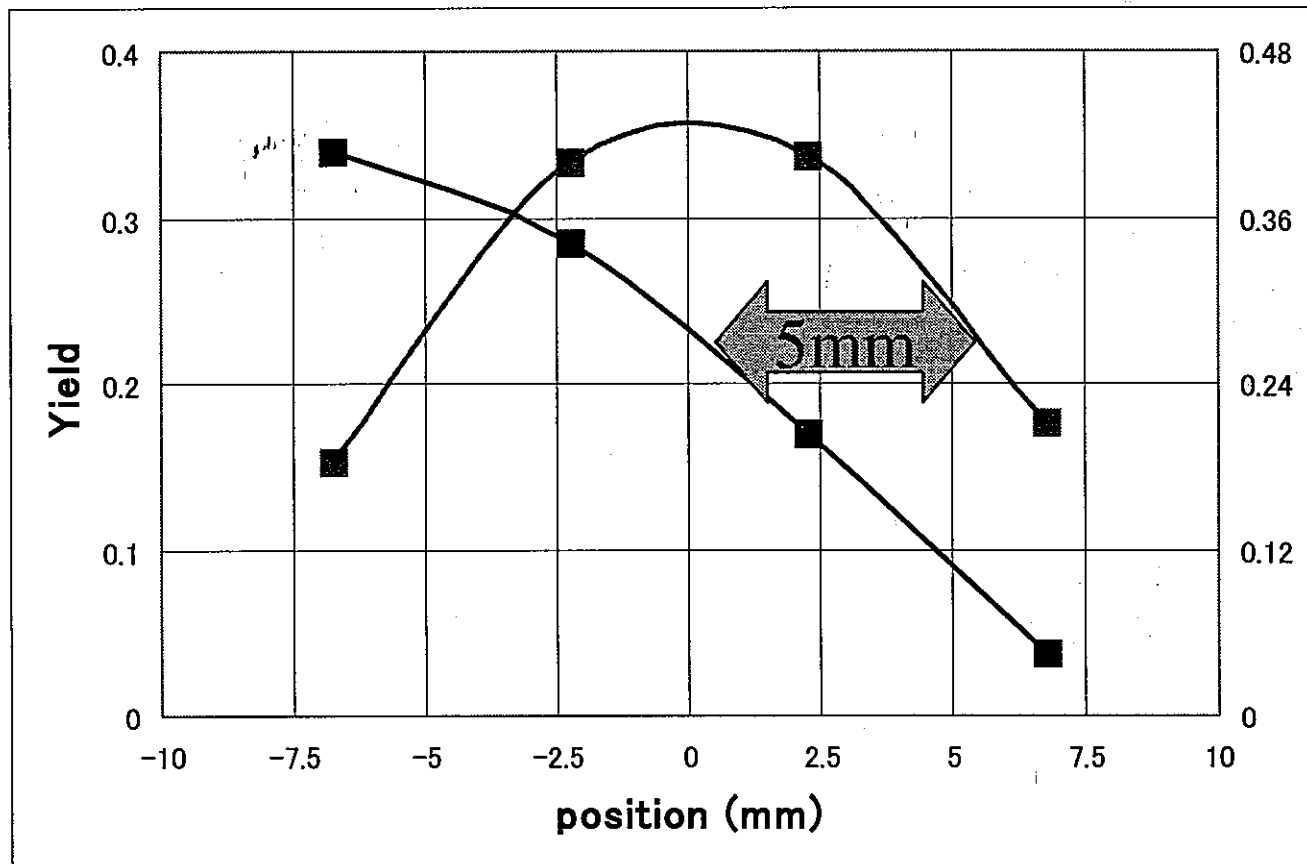
Observation  
on the surface

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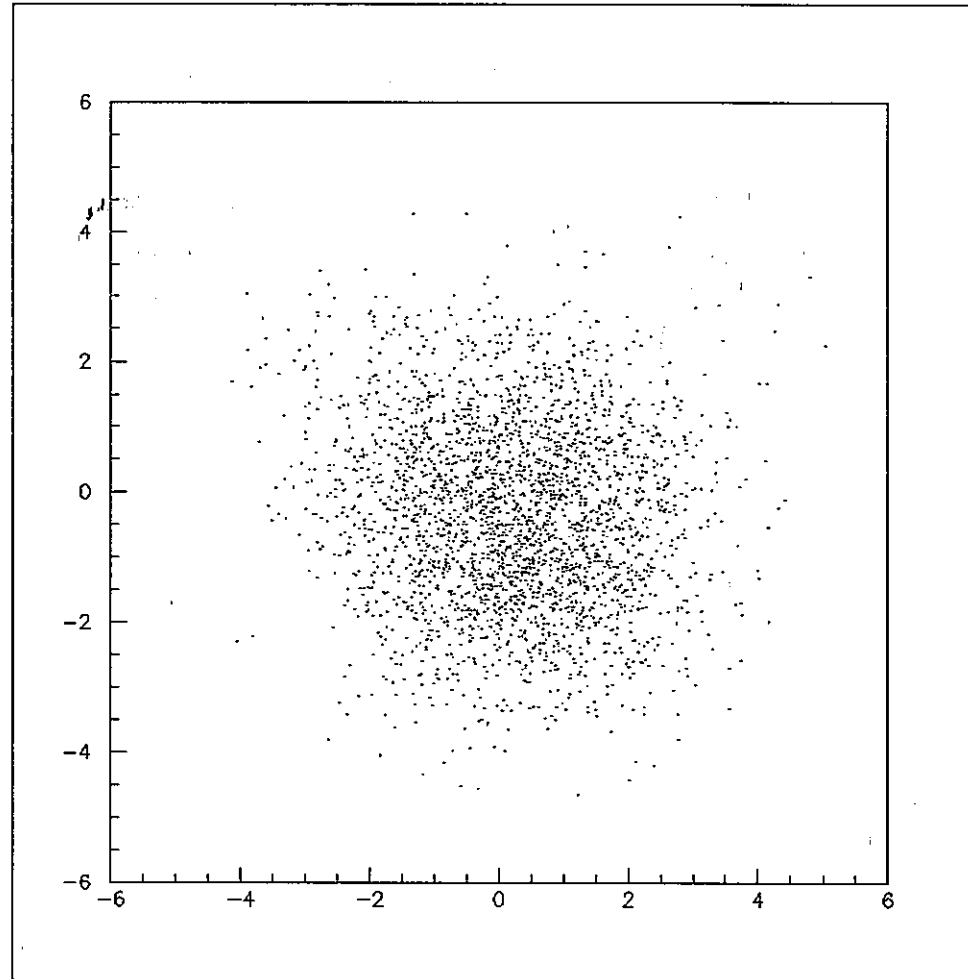




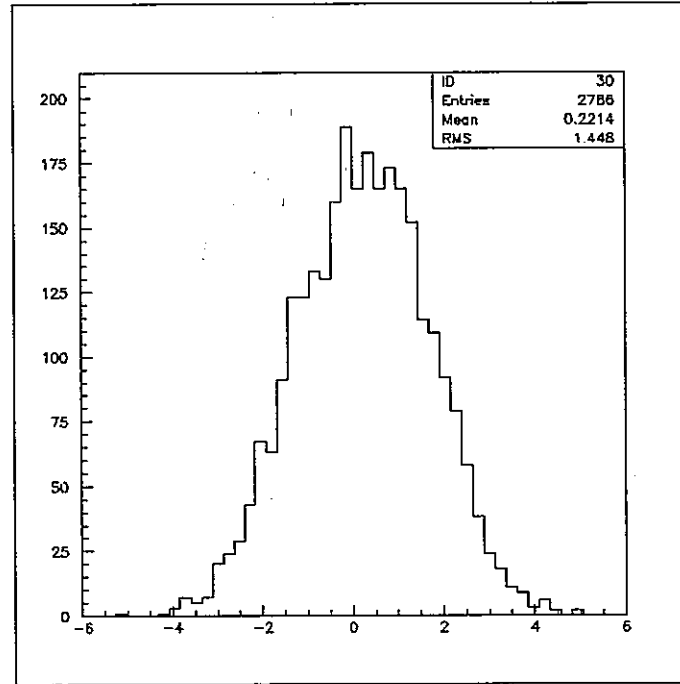
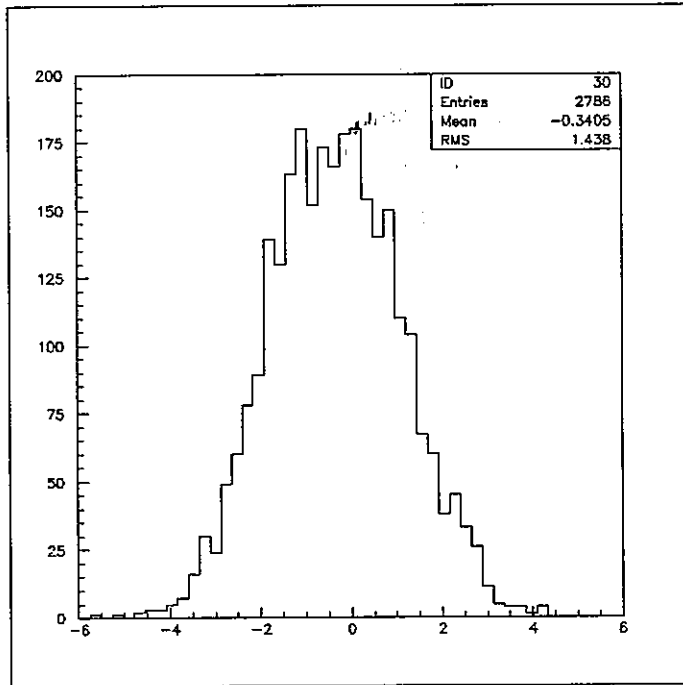
Photon distribution  
~ 12mm (diameter)



# Distribution of calculated X-Y position



X-Y position is calculated by weighted mean.



Position resolution (RMS) : 1.4 mm



# Successive $\alpha$ -decay

- Two successive alpha decay in GSO crystal



- Time Interval                      Half life 145 ms
- Position                              At the same position

96.0 $\pm$ 8.6 mBq/kg is clearly identified

# Energy spectrum

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GSO  
3.8cmx3.8cmx18cm  
1.744kg

# Time Interval

265