

# JHFニュートリノ実験 と超低エネルギービーム

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# JHF Neutrino Working Group

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Dec.99: Working group formed.

Mar.00: Letter of Intent prepared (<http://neutrino.kek.jp/jhfnu>)

Now : Working to prepare a proposal

# Neutrino Oscillation

Neutrino Mixing  $|\nu_l\rangle = \sum U_{li} |\nu_i\rangle$

Weak eigenstates      Mass eigenstates

## Maki-Nakagawa-Sakata Matrix

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} = \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \cdot \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \cdot \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & e^{-i\delta} \end{pmatrix} \cdot \begin{pmatrix} c_{13} & 0 & s_{13} \\ 0 & 1 & 0 \\ -s_{13} & 0 & c_{13} \end{pmatrix}$$

## Oscillation Probability

$$s_{ij} = \sin\theta_{ij}, \quad c_{ij} = \cos\theta_{ij}$$

$$P_{l \rightarrow m} = |\langle \nu_m(t) | \nu_l(0) \rangle|^2 = \delta_{ml} - 2 \sum_{i < j} \text{Re} \left[ (U_{mi}^* U_{li}) \cdot (U_{mj} U_{lj}^*) \cdot \left\{ 1 - \exp\left( -i \frac{\Delta m_{ij}^2 L}{2E} \right) \right\} \right]$$

$L$ : flight length,  $E$ : neutrino energy,  $\Delta m_{ij}^2 \equiv m_i^2 - m_j^2$ ,  $m_i$ : mass eigenvalues

$$P_{l \rightarrow m} \neq \delta_{ml} \Leftrightarrow \Delta m_{ij} \neq 0$$

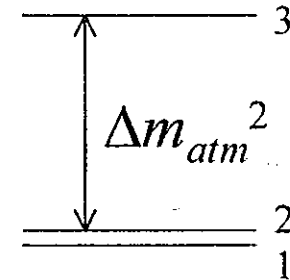
LFV

# Oscillation probabilities

when  $\begin{cases} \Delta m_{12}^2 \ll \Delta m_{23}^2 \approx \Delta m_{13}^2 \equiv \Delta m_{atm}^2 & \text{contribution from } \Delta m_{12} \text{ is small} \\ E_\nu \approx \Delta m_{atm}^2 \cdot L \end{cases}$

$\nu_e$  appearance

$$P_{\mu \rightarrow e} \approx \underbrace{\sin^2 \theta_{13}}_{\sim 0.5} \cdot \underbrace{\sin^2 2\theta_{13}}_{\sin^2 2\theta_{\mu e}} \cdot \sin^2 \left( 1.27 \Delta m_{atm}^2 / E_\nu \right)$$



$\nu_\mu$  disappearance

$$P_{\mu \rightarrow x} = 1 - (P_{\mu \rightarrow e} + P_{\mu \rightarrow \tau} + P_{\mu \rightarrow \text{sterile}})$$

$$P_{\mu \rightarrow \tau} \approx \underbrace{\cos^4 \theta_{13}}_{\sim 1} \cdot \sin^2 2\theta_{23} \cdot \sin^2 \left( 1.27 \Delta m_{atm}^2 / E_\nu \right)$$

Neutral Current (NC) measurement

$$N_{NC} \propto P_{\mu \rightarrow \text{active}} = 1 - P_{\mu \rightarrow \text{sterile}}$$

Cf. Chooz ( $\nu_e$  disappearance)

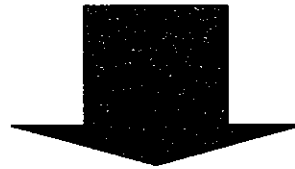
$$P_{e \rightarrow x} \approx 1 - \sin^2 2\theta_{13} \cdot \sin^2 \left( 1.27 \Delta m_{atm}^2 / E_\nu \right)$$

# CP violation

No CPV in disappearance from unitarity

$$P_{\mu \rightarrow e}(CPV) = \sin \delta \cdot (\sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13}) \cdot \cos \theta_{12} \\ \times \left( \sin \frac{\Delta m_{12}^2 L}{2E} + \sin \frac{\Delta m_{23}^2 L}{2E} - \sin \frac{\Delta m_{31}^2 L}{2E} \right)$$

- LMA solution for solar neutrino  
 $\Delta m_{12}^2 \sim 10^{-4} \text{eV}^2$ ,  $\sin^2 \theta_{12} \sim 0.8$
- $\theta_{13}$  is non zero



CPV effect in lepton sector could be detectable

# Physics motivation

## 1. Test our current picture of 3 flavor neutrino oscillation

- Spectrum shape of  $\nu_\mu$  disappearance
  - Test exotic models (decay, extra dimensions,....)
- Appearance of  $\nu_e$  at the same  $\Delta m^2$  as  $\nu_\mu$  disappearance
- NC measurements
  - No additional “neutrino”?

## 2. Precise measurement of $\Delta m^2$ and mixing angles ( $\theta_{23}$ , $\theta_{13}$ )

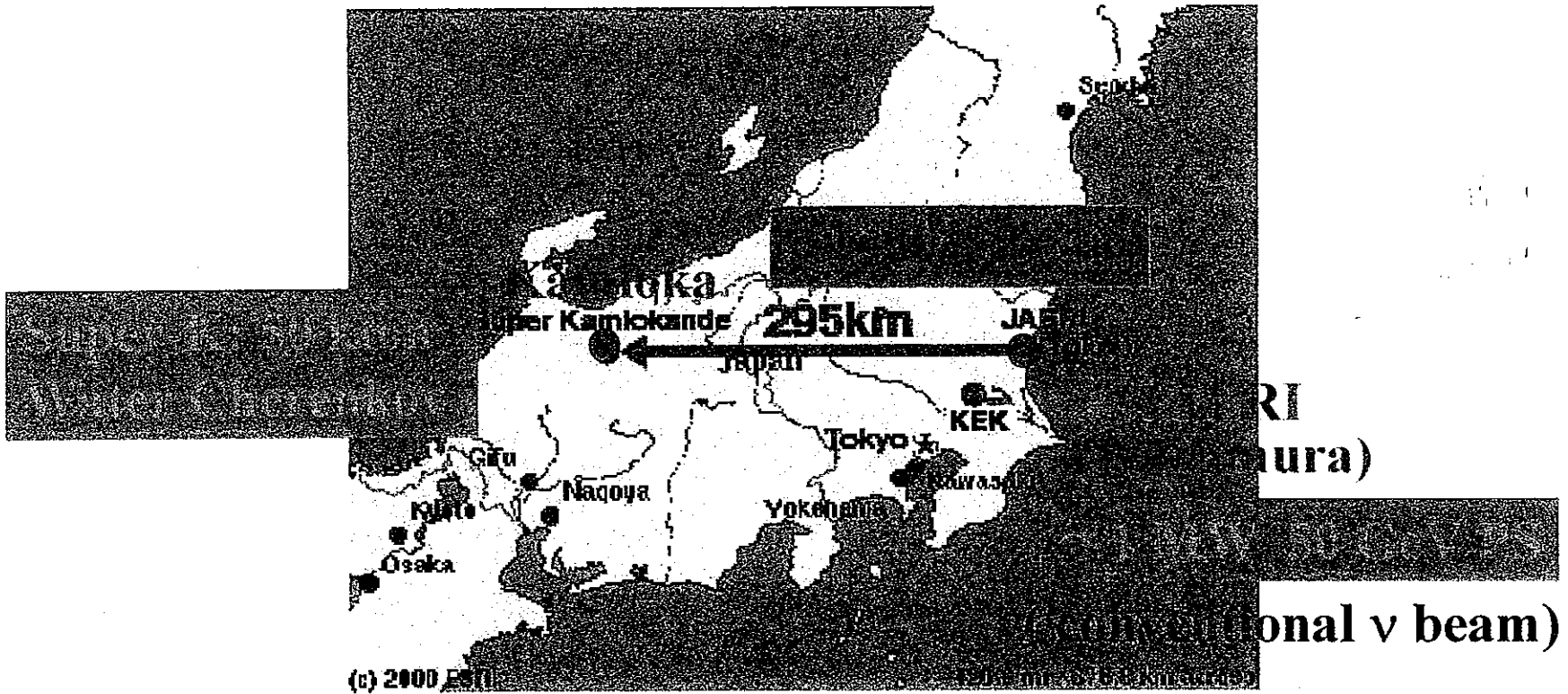
- mixing matrix in quark sector: well known
- understanding of mixing in lepton sector
- understanding of mass structure
  - hints on physics beyond the SM (GUTs,...)

## 3. Discovery of $\nu_e$ appearance

- Open possibility to detect CPV effect in lepton sector

# JHF v experiment

# Overview



- $\nu_{\mu} \rightarrow \nu_{\tau}$  disappearance
- $\nu_{\mu} \rightarrow \nu_{e}$  appearance
- NC measurement

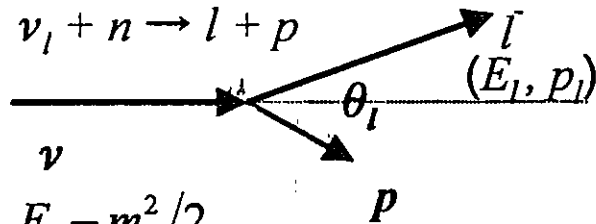


# Principle

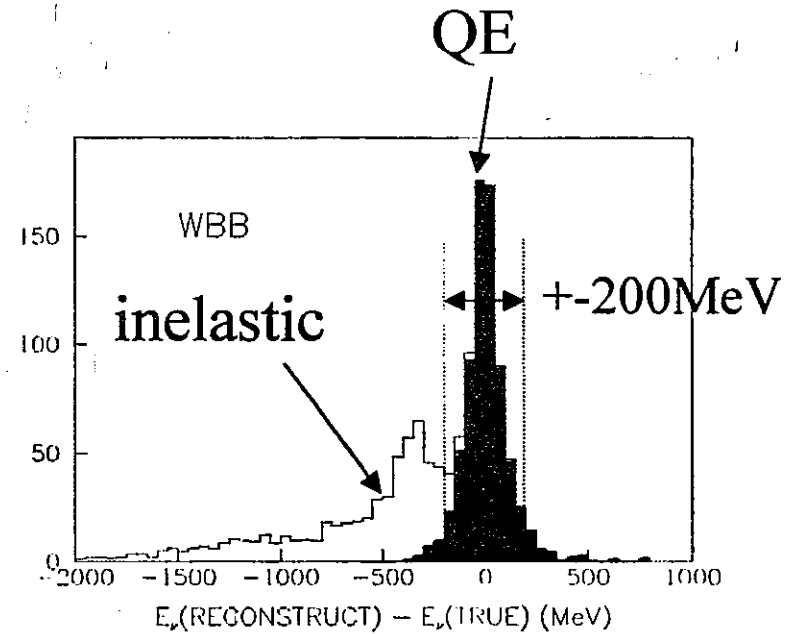
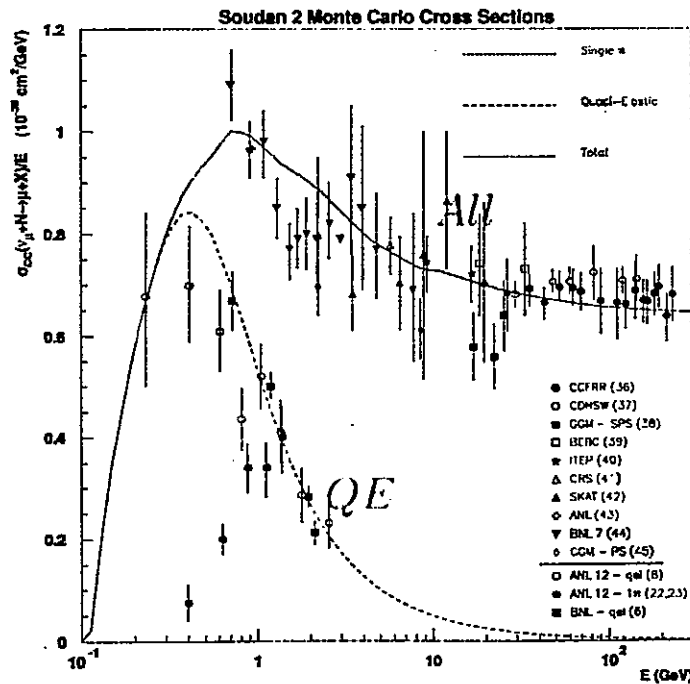
- Super-Kamiokande at 295km as far detector
- Beam energy is tuned to be at the oscillation maximum.
  - High sensitivity  $\Delta m^2 = 2 \sim 5 \times 10^{-3} \text{eV}^2$
  - Less background  $E_\nu = 0.5 \sim 1.2 \text{GeV}$
- Neutrino energy reconstruction by using **Quasi-elastic (QE)** interaction.
  - Oscillation pattern measurement
  - BG due to miss-reconstruction of inelastic interaction
    - Greatly improved by using narrow band beam (NBB)

# Neutrino Energy Reconstruction

Assume CC quasi elastic (CCQE) reaction

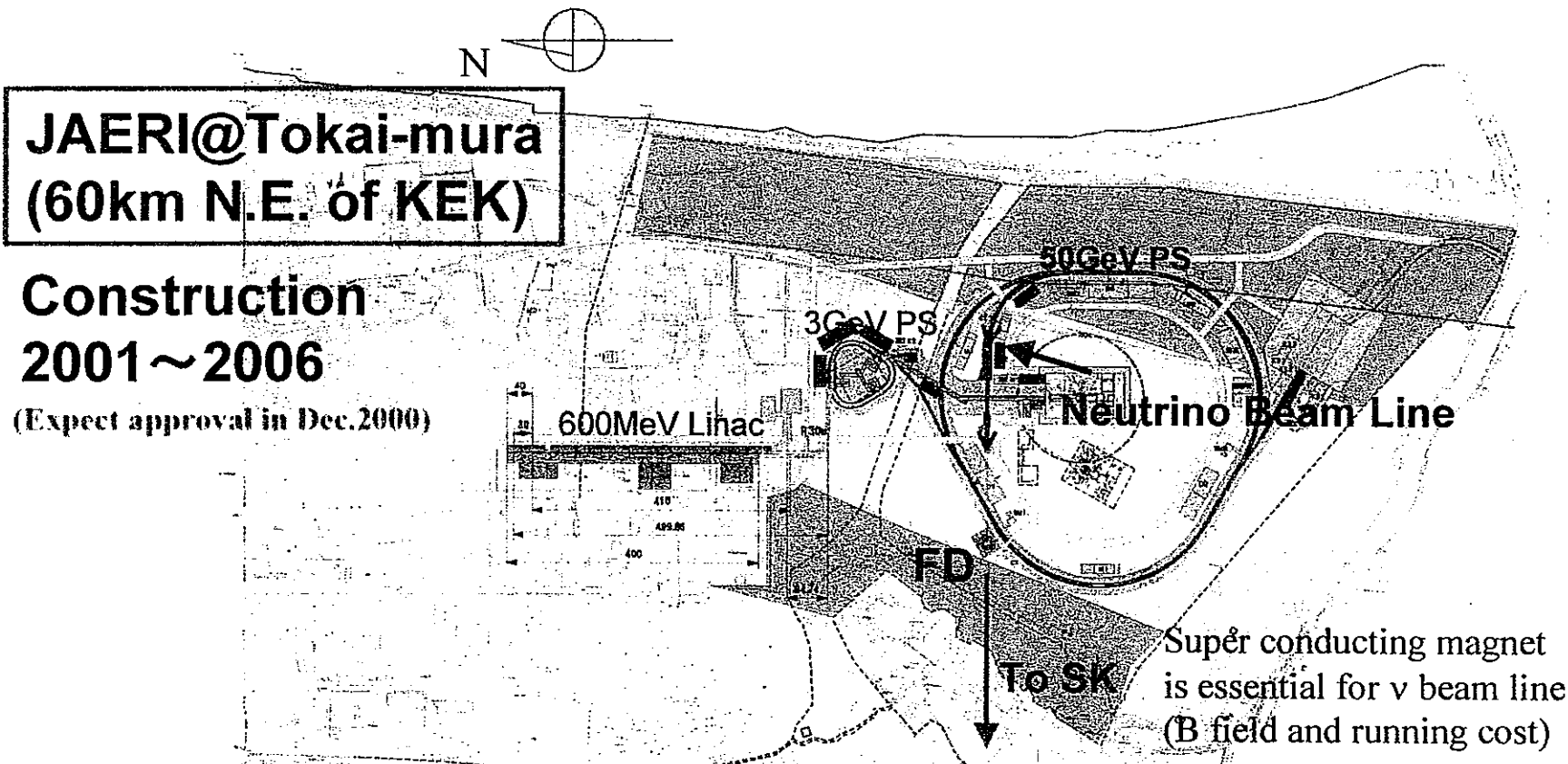


$$E_\nu = \frac{m_N E_l - m_l^2 / 2}{m_N - E_l + p_l \cos \theta_l}$$



QE dominate at ~1 GeV

# JHF project and neutrino beam line



	JHF	MINOS	K2K
E(GeV)	50	120	12
Int.( $10^{12}$ ppp)	330	40	6
Rate(Hz)	0.292	0.53	0.45
Power(MW)	0.77	0.41	0.0052

$10^{21}$  POT (1300day)  $\equiv$  '1 year'

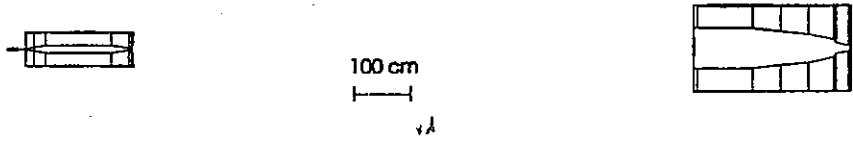
# Neutrino Beam @ JHF

## Three beam configurations

- **Wide Band Beam (WBB)**
  - 2 Horns almost the same as K2K
  
- **Narrow Band Beam (NBB)**
  - Horn(s) + Bending
  
- **Off Axis Beam (OAB)**
  - Another option of NBB

# Wide Band Beam

Target : Cu 1cm $\phi$  x 30cm  
 Horn : 250kA  
 Decay Pipe : 50m x 1.5m $\phi$   
 Gcalor



2 horns (almost same design as K2K)

$\sim 4200 \nu_{\mu}$  int./22.5kt/yr

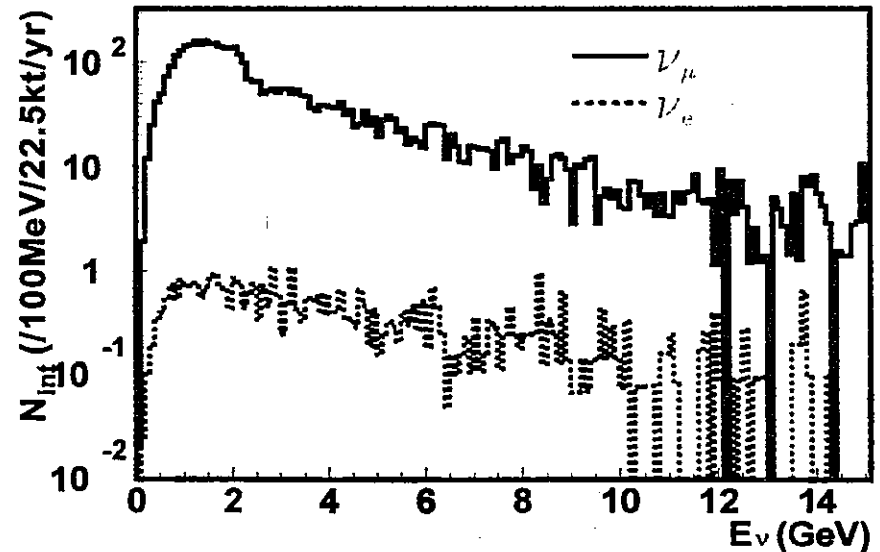
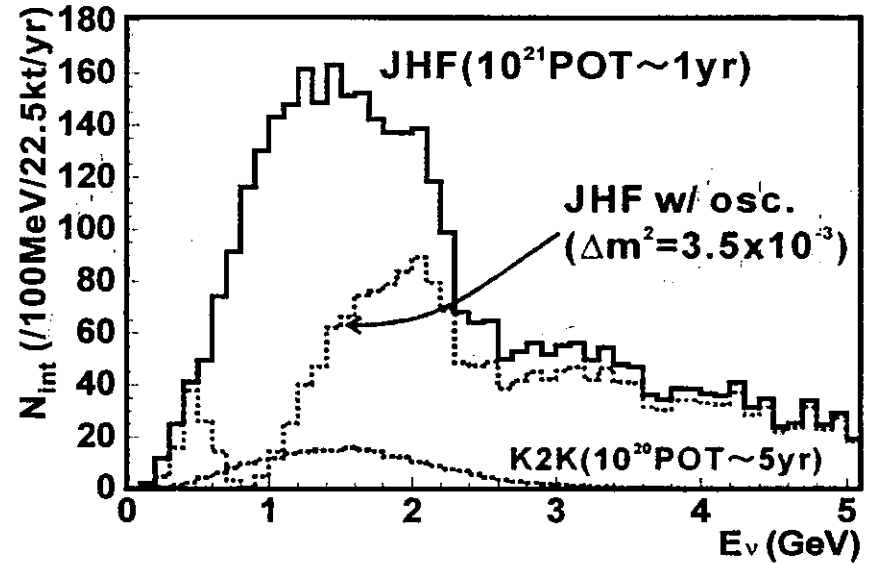
$\nu_e$ :0.8%

Intense

Wide sensitivity in  $\Delta m^2$

BG from HE tail

Syst. err from spectrum extrapolation

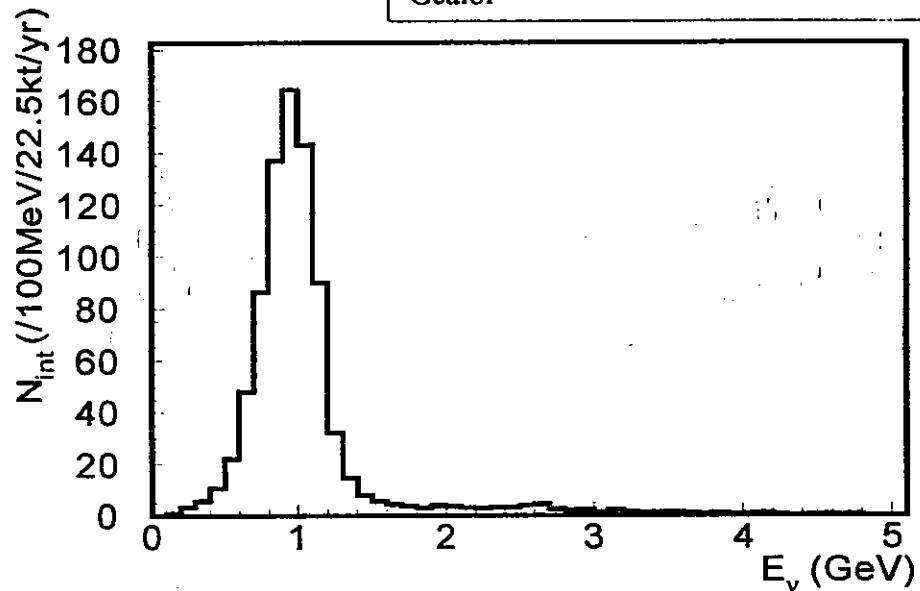
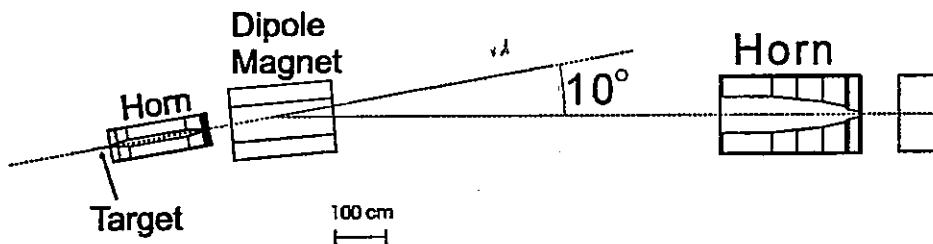


# Narrow Band Beam

Updated from LOI

(factor  $\sim 2$  increased by adding 2nd horn)

Target	: Cu $1\text{cm}^\phi \times 30\text{cm}$
Horn	: 250kA
Decay Pipe	: $155\text{m} \times 1.5\text{m}^\phi$
Dipole	: $50\text{cm}(V) \times 70\text{cm}(H) \times 2\text{m}(L)$
Gcalor	: 0.58T (10deg@2GeV/c)



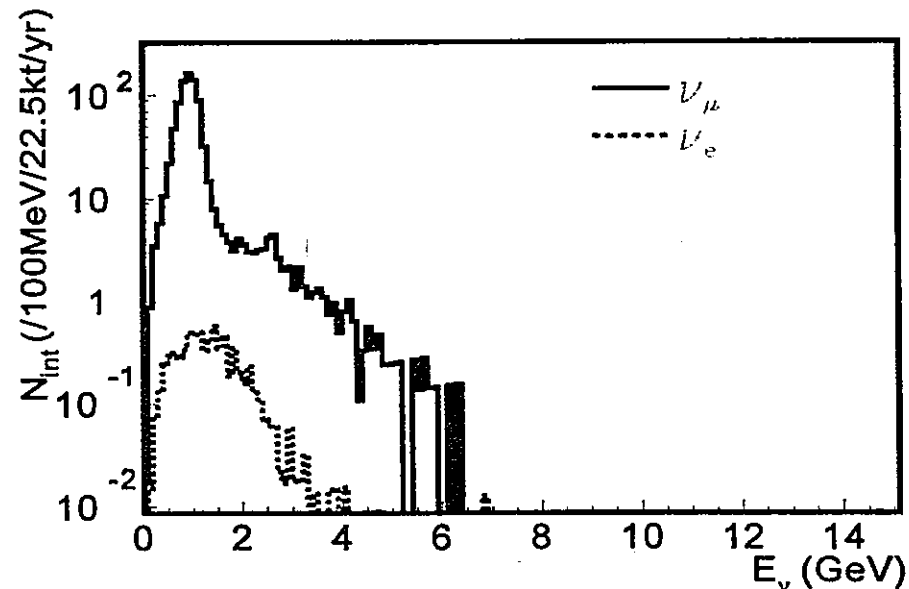
**$\sim 830$  int./22.5kt/yr**  
 $\nu_e$ : 0.8% (0.3% @ peak)

Less HE tail

Less sys err from spectrum

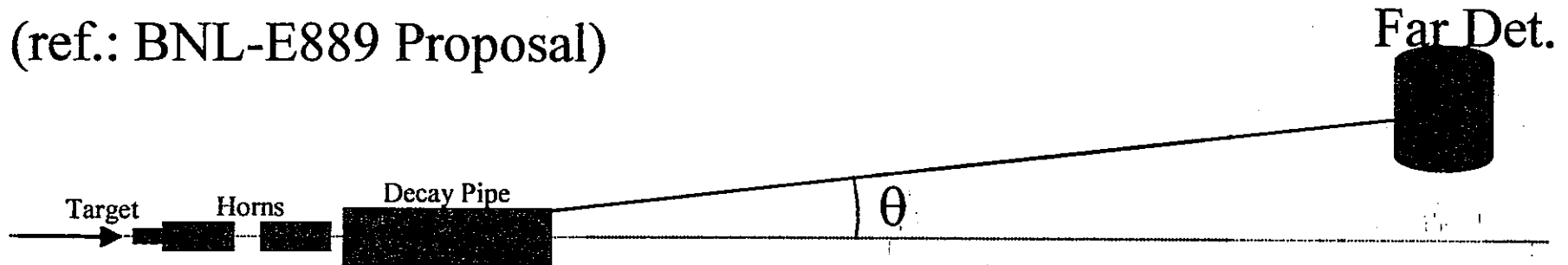
“counting experiment”

Easy to tune  $E_\nu$



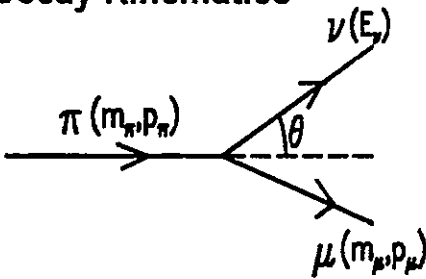
# Off Axis Beam (another NBB option)

(ref.: BNL-E889 Proposal)

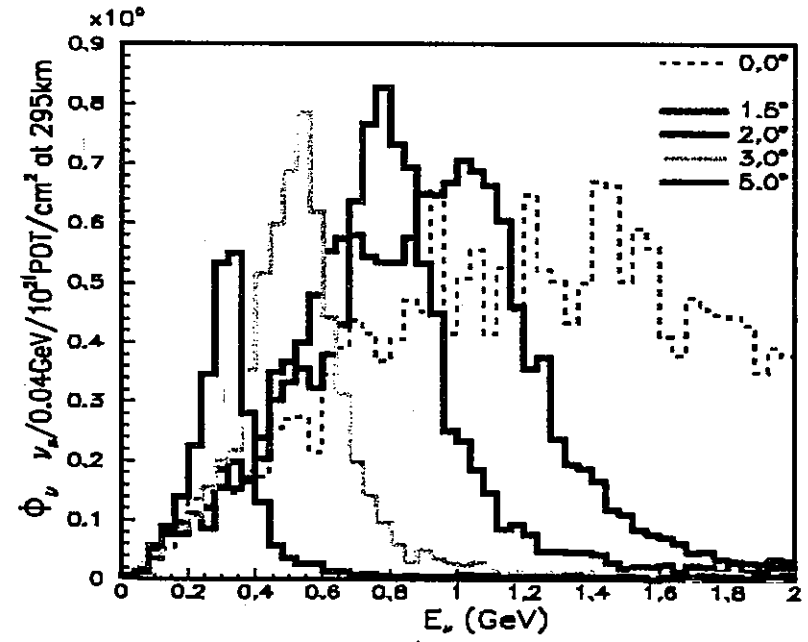
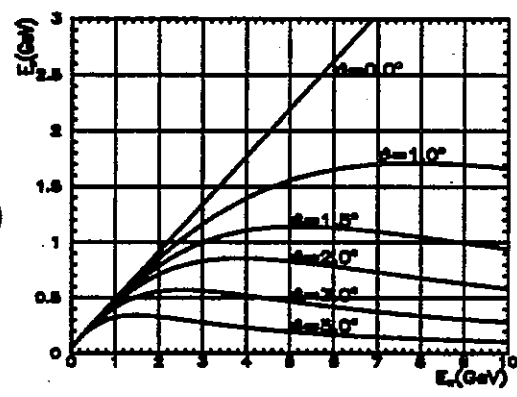


WBB w/ intentionally misaligned beam line from det. axis

Decay Kinematics



$$E_\nu = \frac{m_\pi^2 - m_\mu^2}{2(E_\pi - p_\pi \cos\theta)}$$



## Quasi Monochromatic Beam

263

# Off axis beam

~2200 int./22.5kt/yr

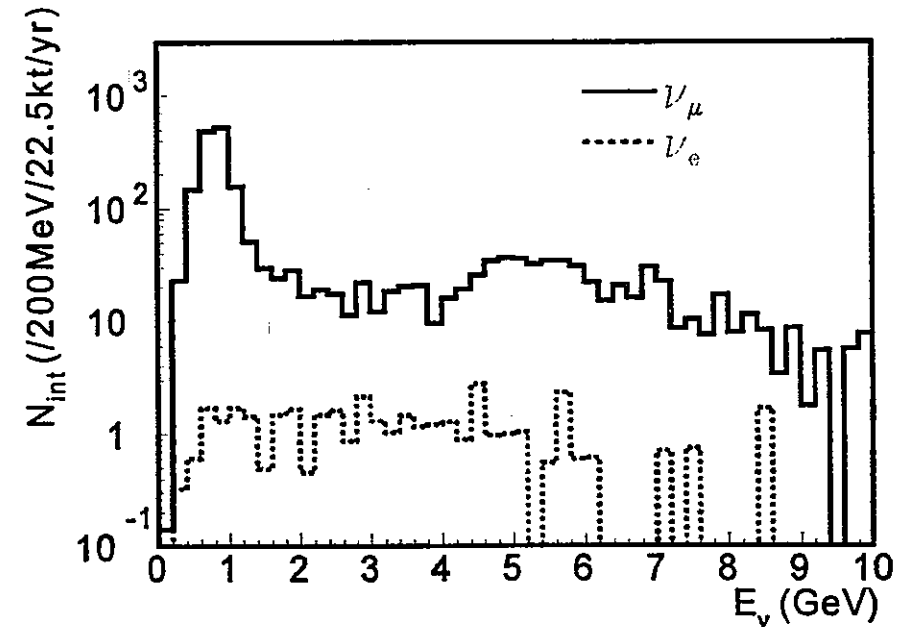
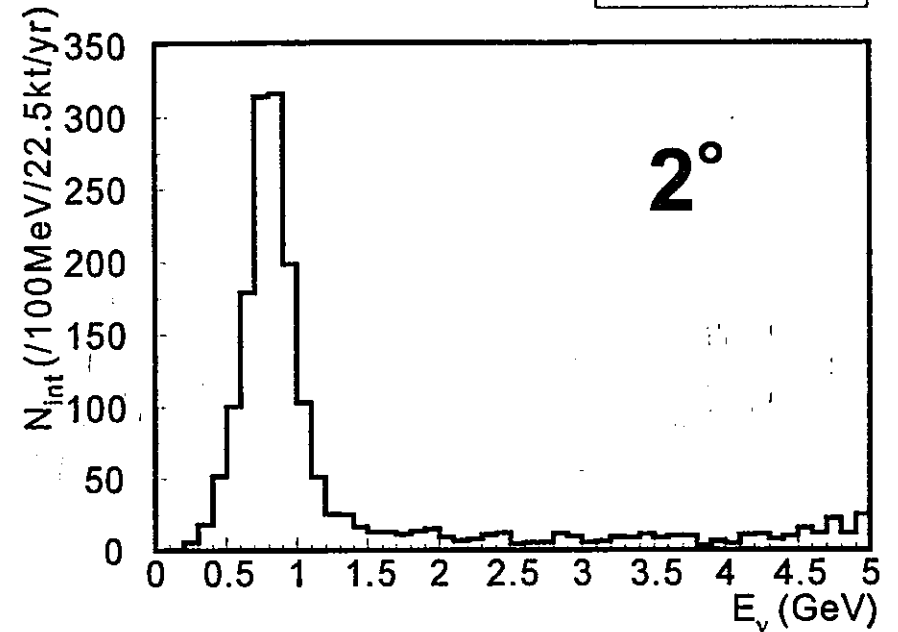
$\nu_e$ : 0.8% (0.2% @ peak)

High int. narrow band beam

More HE tail than NBB

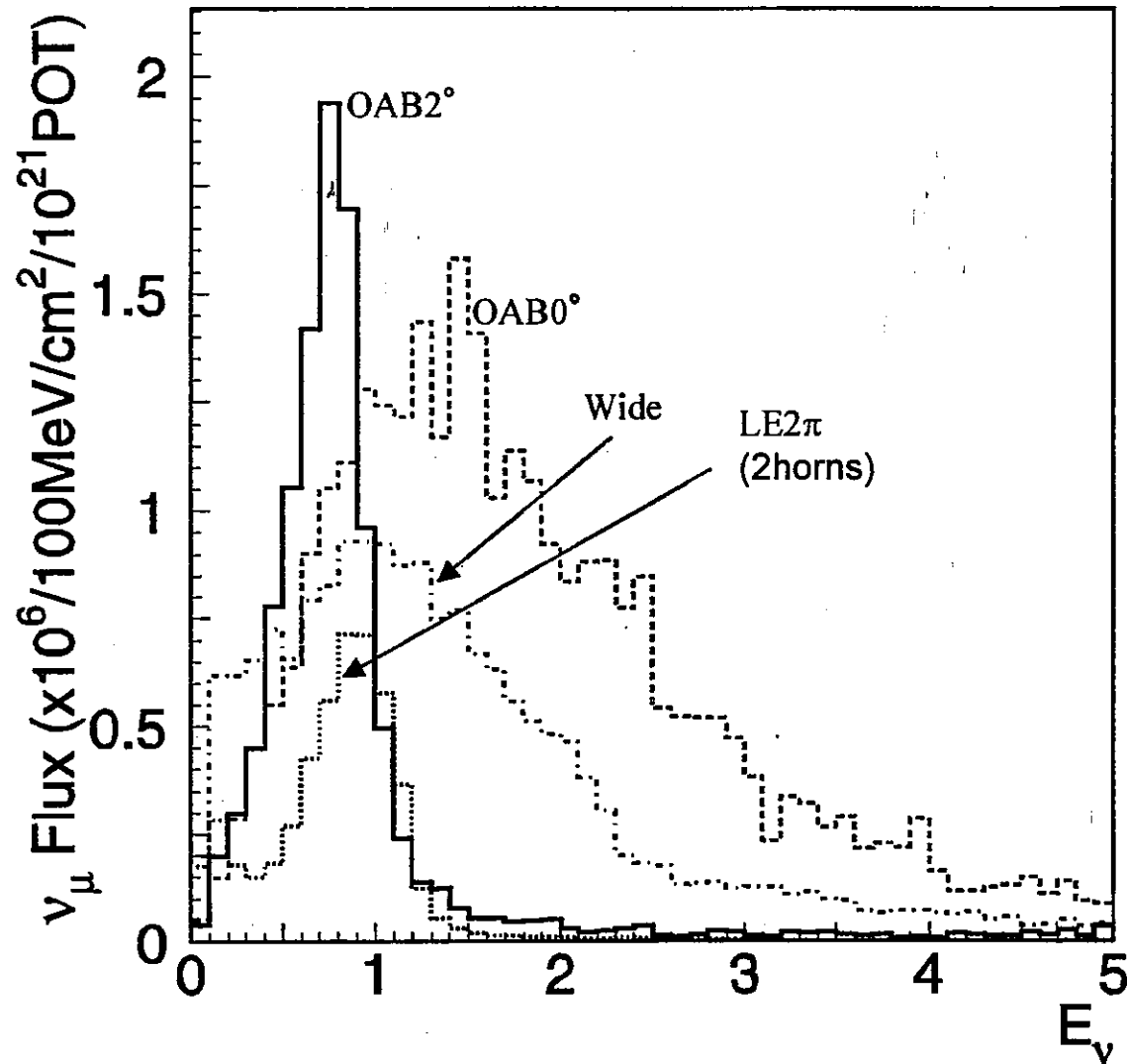
Hard to tune  $E_\nu$

BNL-E889 Horns  
90m decay pipe



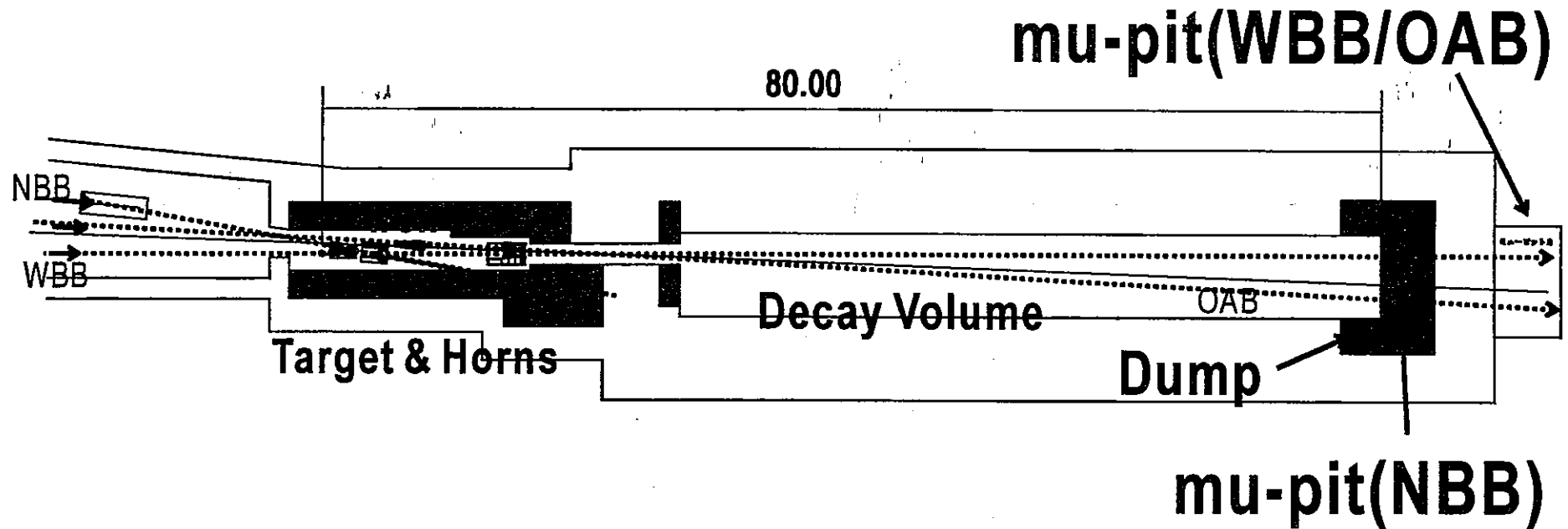


# Comparison of Beams



(same decay pipe length=50m)

# Current design of target station and decay volume



266

WBB/NBB/OAB can be switched by replacing optics  
Decay volume is shared (flat pipe)

Design being optimized: flux, radiation shielding, cost

# Physics Sensitivity

# Strategy and Goal

- First 1 year WBB
  - pin down  $\Delta m_{23}^2$  to  $\pm 10\%$  level
  - NC measurement
- 5year NBB or OAB
  - precise measurement of  $\theta_{23}$  and  $\theta_{13}$ .

## Sensitivity (goal):

$$\delta \sin^2 2\theta_{23} \sim 0.01$$

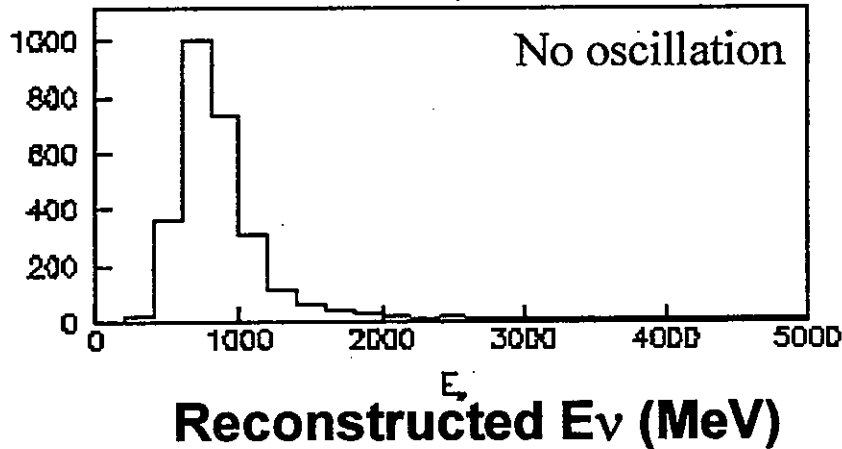
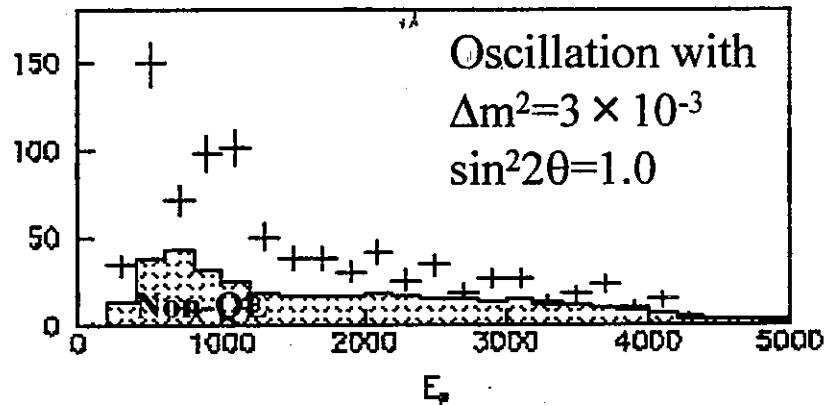
$$\sin^2 2\theta_{13} \sim 5 \times 10^{-3} \text{ (90\% CL)}$$

$$\delta \Delta m_{23}^2 \sim 1.5 \times 10^{-4} \text{eV}^2$$

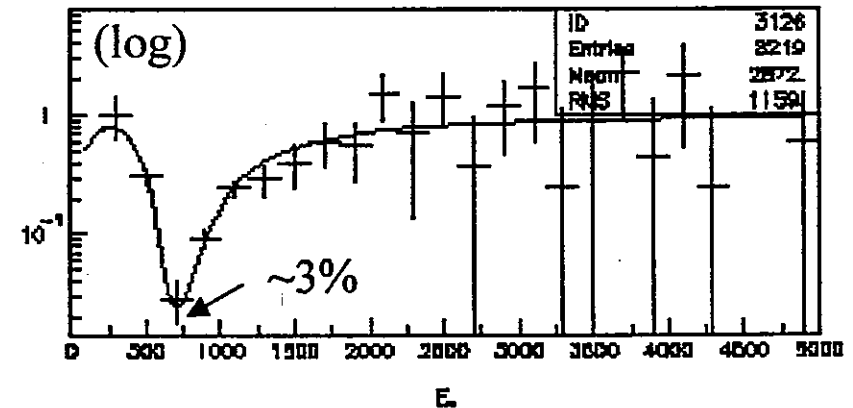
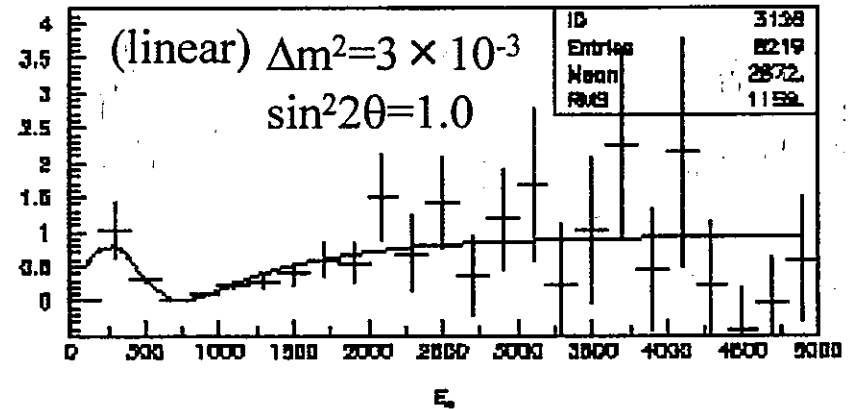
$$\text{at } (\sin^2 2\theta = 1.0, \Delta m^2 = 3.2 \times 10^{-3} \text{eV}^2)$$

# $\nu_\mu$ disappearance

1ring FC  $\mu$ -like



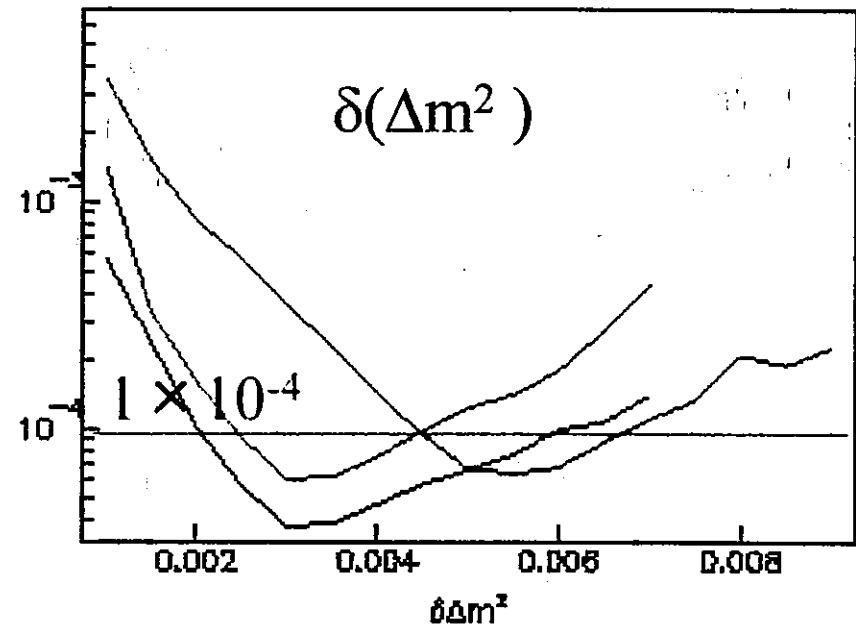
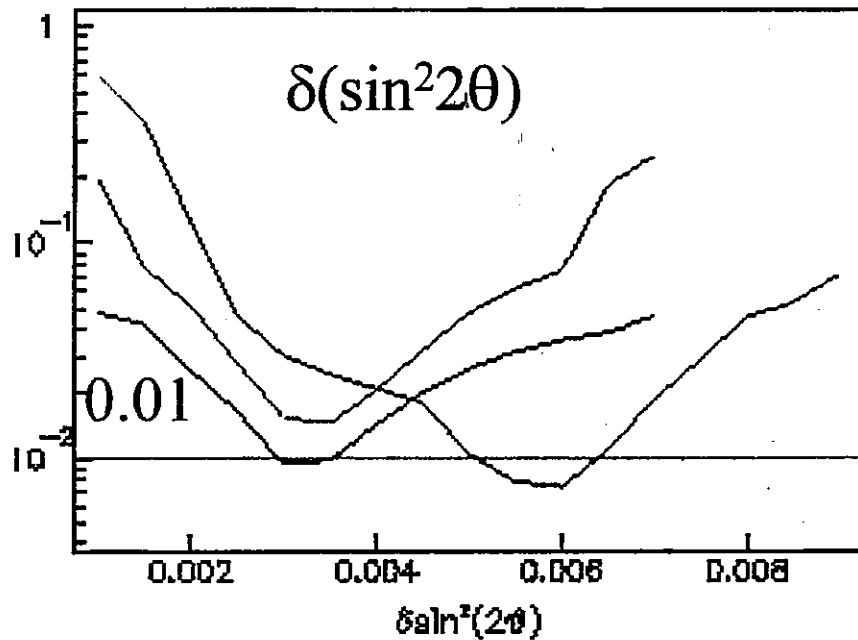
Ratio after BG subtraction



Fit with  $1 - \sin^2 2\theta \cdot \sin^2(1.27 \Delta m^2 L/E)$

# 5 years precision

NBB-3GeV $\pi$ , OAB-2degree, NBB-1.5GeV $\pi$



$\delta(\sin^2 2\theta) \sim 0.01$  in 5 years

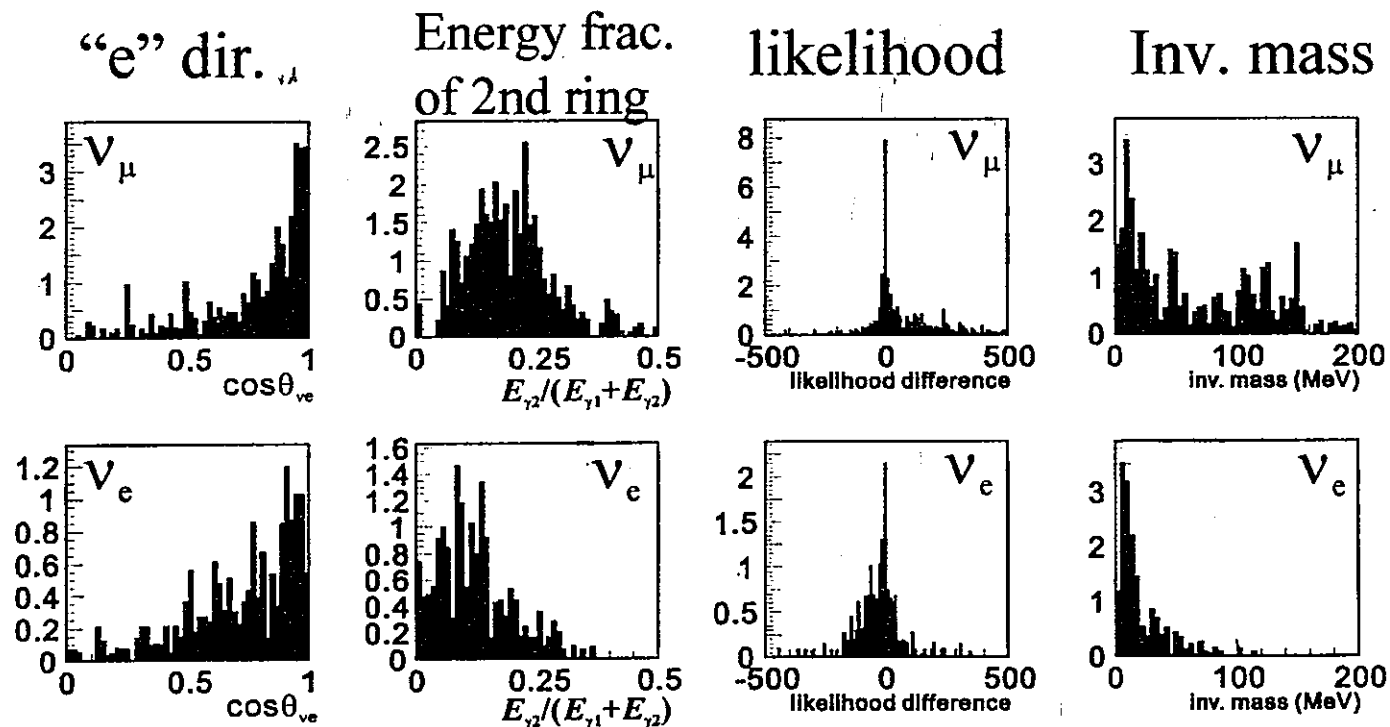
$\delta(\Delta m^2) \sim < 1 \times 10^{-4}$  in 5 years

# $\nu_e$ appearance ( $\theta_{13}$ )

- Signal
  - 1 ring e-like ring
  - At energy of  $\nu_\mu$  disappearance dip
- Backgrounds
  - $\nu_\mu$  NC  $\pi^0$  production
    - Lower  $E$  photon is missed
  - Beam  $\nu_e$  contamination
    - Broad  $E$  dist. Can be reduced w/ energy window.
    - 0.2-0.3% of  $\nu_\mu$  at peak of NBB/OAB

# $\pi^0$ BG rejection (updated from LOI)

Force to find 2nd ring in 1-ring e-like sample



**Factor  $\sim 10$  improvement in BG rejection  
while  $\nu_e$  eff. decrease is only 30%**

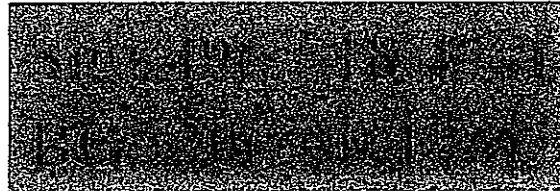
**Preliminary**



# Expected signal

$$\sin^2 2\theta_{\mu e} = 0.05 \text{ (Chooz limit)}$$

WBB



$e/\pi^0$  cut tightened to reduce BG

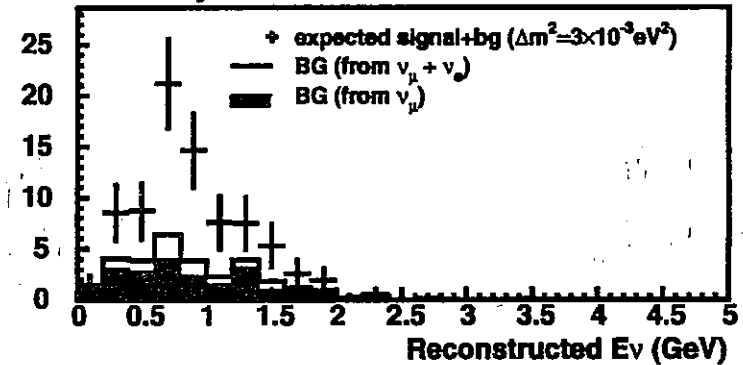
NBB



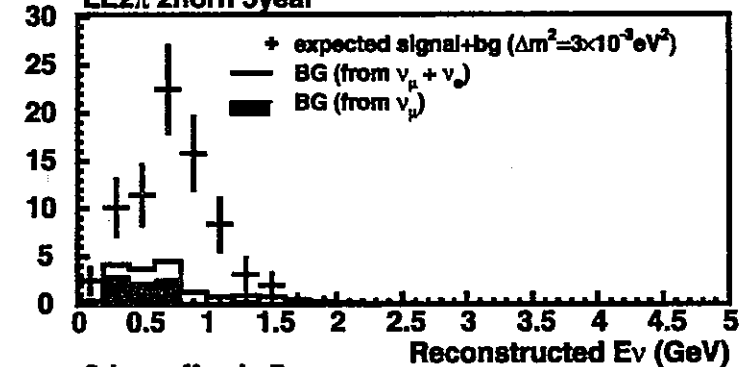
OAB(2deg)



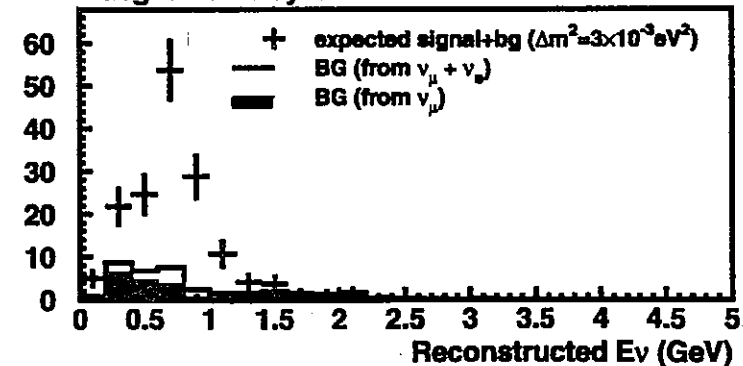
WIDE 5year



LE2π 2horn 5year

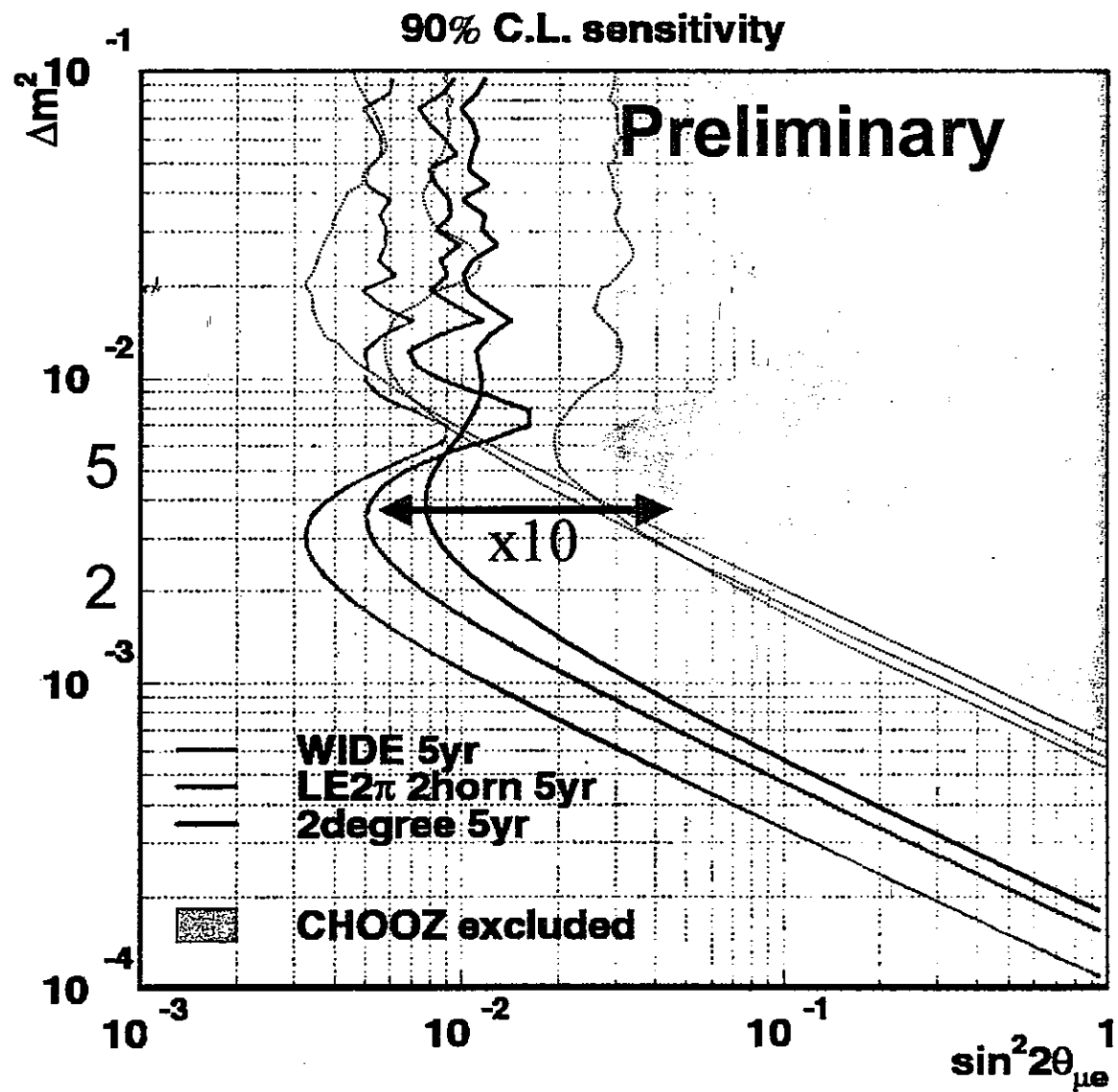


2deg. off axis 5year



Preliminary

# Sensitivity on $\nu_\mu \rightarrow \nu_e$ appearance



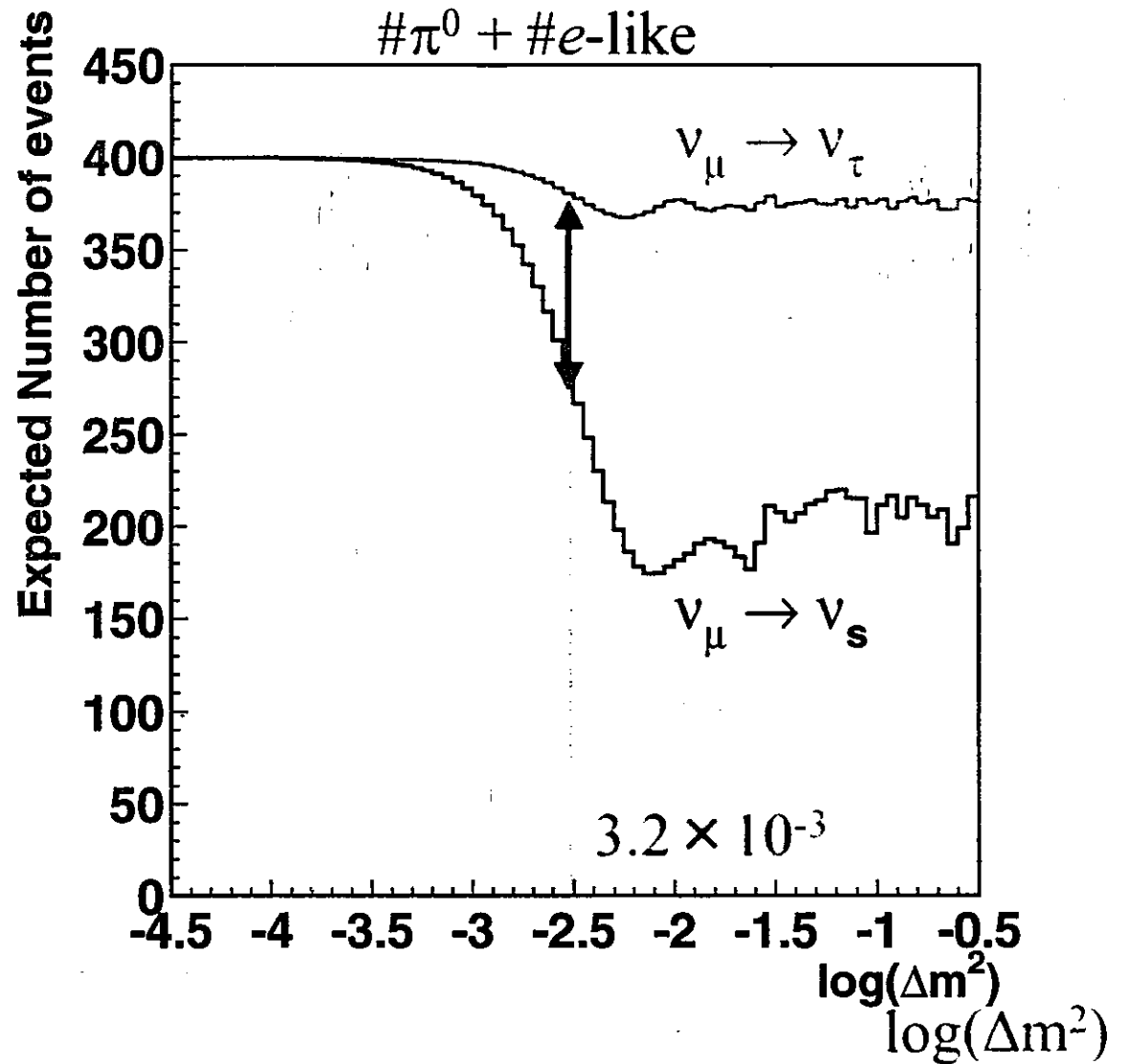
Dashed lines: MINOS Ph2le, Ph2me, Ph2he from right  
(A.Para, hep-ph/0005012)

# NC measurement

# of NC events

$$N_{NC} \propto P_{\mu \rightarrow \text{active}} = 1 - P_{\mu \rightarrow \text{sterile}}$$

- NC/CC sensitive to  $\nu_s$
- NC Enriched Sample



# Comparison with other LBL projects

- ICANOE (2005~)

- CERN SPS(400GeV) → Gran Sasso LBL (732km)

- $E_\nu \sim 20\text{GeV}$

- Optimized for  $\nu_\tau$  search

**Complementary to JHF $\nu$**

- MINOS (2003~)

- Fermilab Main Injector (120GeV) → Soudan mine (730km)

- $E_\nu > 3\text{GeV}$

- $\nu_\mu$  disappearance:  $\delta(\Delta m^2) \sim 2.4 \times 10^{-4} \text{eV}^2$ ,  $\delta(\sin^2 2\theta_{23}) \sim 0.06$

- $\nu_e$  appearance :  $\delta(\sin^2 2\theta_{\mu e}) > 0.04$  @  $\Delta m^2 = 3 \times 10^{-3} \text{eV}^2$

(read from A.Para, ICHEP2000 by T.K.)

	Beam	$E_\nu$	$(E/L)(\pi/2)/1.27$	Det.	$E_\nu \text{ rec}^{\text{nst}}$	CC event
JHF $\nu$	NBB/WBB/OAB	$\sim 1\text{GeV}$	$3.8 \times 10^{-3} \text{eV}^2$	Water Cherenkov	QE	3200/yr(WBB)
MINOS	WBB	$> 3\text{GeV}^*$	$5 \times 10^{-3} \text{eV}^2$	Iron cal.	Hadr. Cal.	2500/yr*

**JHF project has much higher potential.**

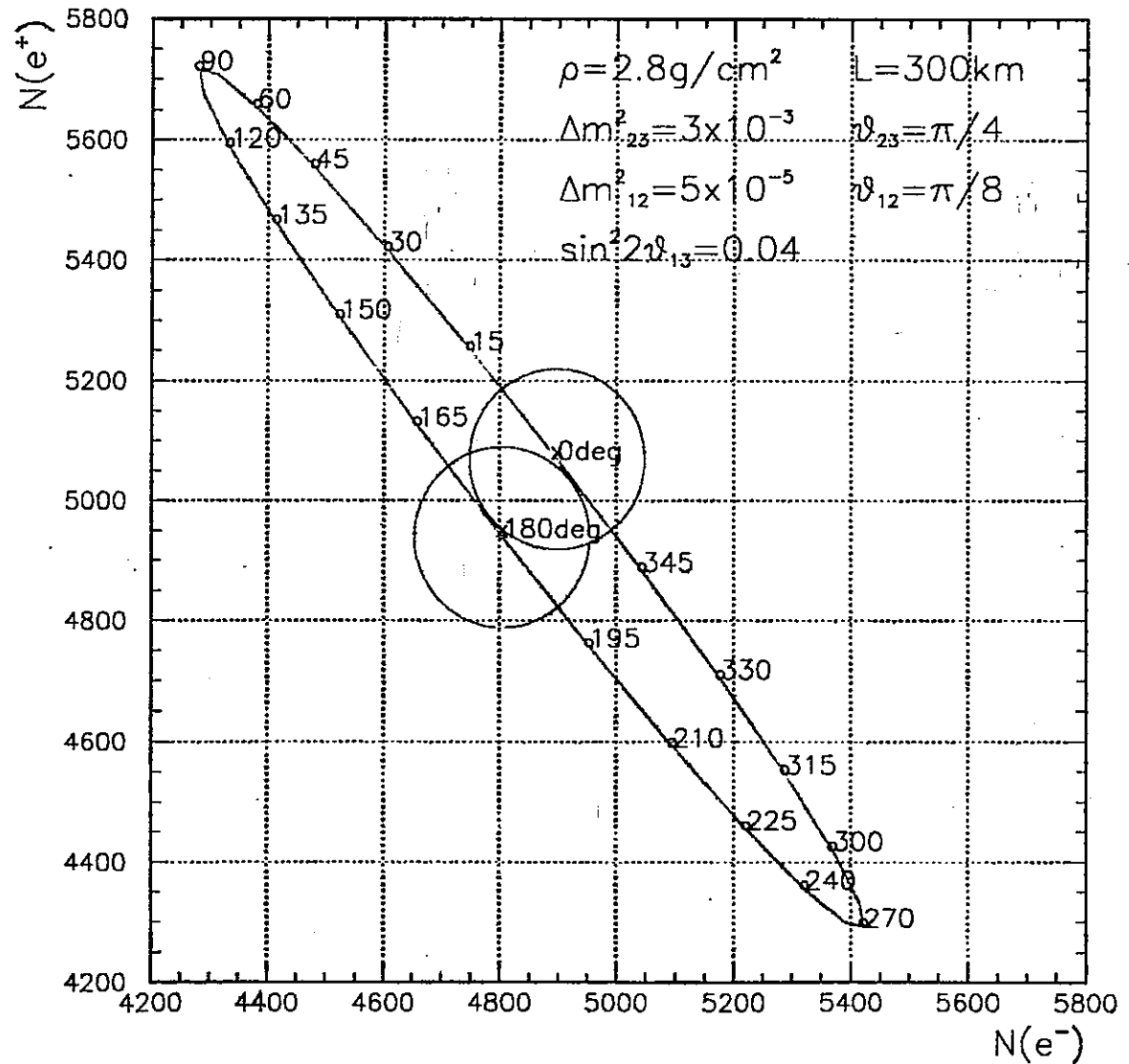
\*:PH2(Low) option

# Future Extensions

- **PS upgrade to 4MW and 1Mton water Cherenkov detector**
  - 2 order increase in statistics
  - CPV if  $\nu_e$  appearance discovered in the 1st phase
    - $O(100)$   $\nu_e$  events/year if  $\theta_{13}=0.1x(\text{Chooz limit})$
  - (Proton decay)
- **Very LBL experiment (1000-2000km)**
  - $\sim 300(1200)$ CC events/100kt/yr @ 2000(1000)km w/ 6GeV NBB
  - Sign of  $\Delta m^2 s$
  - Matter effect
  - CPV

# Sensitivity to CPV

4MW PS  
 1Mton Hyper Kam.  
 2years for  $\nu_\mu$   
 6years for  $\nu_\mu$  bar



# Summary

## ● JHF $\nu$ project

- ✓  $\sim$ MW 50GeV PS @ JHF
- ✓ Super-Kamiokande@ 295km as far detector
- ✓ Low energy( $\sim$ 1GeV) conventional  $\nu_\mu$  beam tuned at osc. max.
- ✓ Energy reconstruction by using QE
- ✓ NBB/OAB to reduce background and syst. err.

## ● Physics sensitivity

- ✓  $\delta\sin^2 2\theta_{23} \sim 0.01$
- ✓  $\sin^2 2\theta_{13} \sim 5 \times 10^{-3}$  (90% CL)
- ✓  $\delta\Delta m_{23}^2 \sim 1 \times 10^{-4} \text{eV}^2$
- ✓  $\nu_s$  existence can be tested.

- Design and R&D work have just been started.
- JHF $\nu$  is not included in current budget request.
- Full approval within a few years

Data taking in 2006-7

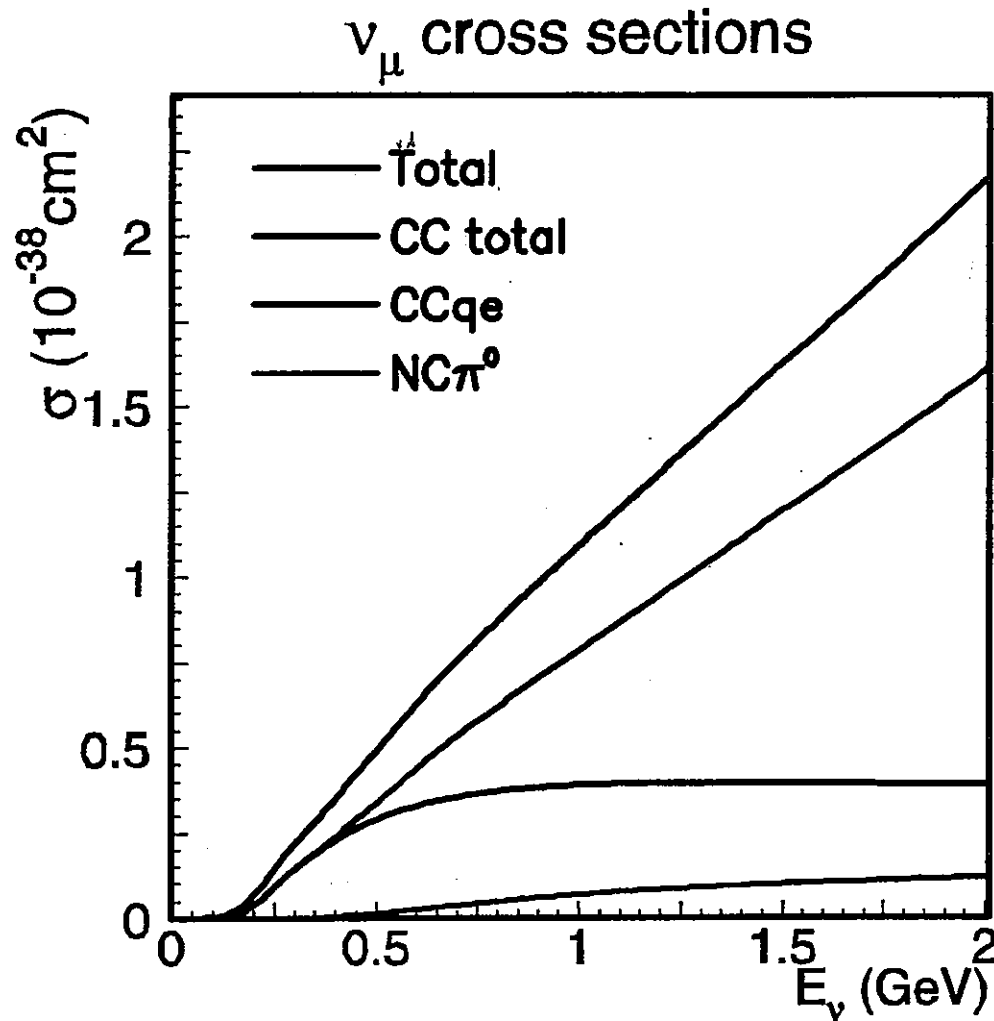
# A study on possibility of Super LowE Super Conventional Beam

- Possible advantages
  - Below  $\pi^0$  threshold  $\rightarrow$  expect low BG
  - 2<sup>nd</sup> osc. max.  $\rightarrow$  expect large CP asym.
  - Small contribution from matter effect
- Questions
  - Statistics
    - neutrino flux enough? Spectrum?
  - Background
    - Really no  $\pi^0$  BG?
    - $e/\mu$  separation at low energy?
    - $\nu_e$  BG?
  - Systematics
    - cross section, .....

24/Feb/2001  
T.Kobayasi



# $\nu_\mu$ cross sections



## SLE Region (<500MeV)

- Small cross section
- Rapidly changing
- CCqe dominates CCint.
- Small  $\pi^0$  prod. cross sect.

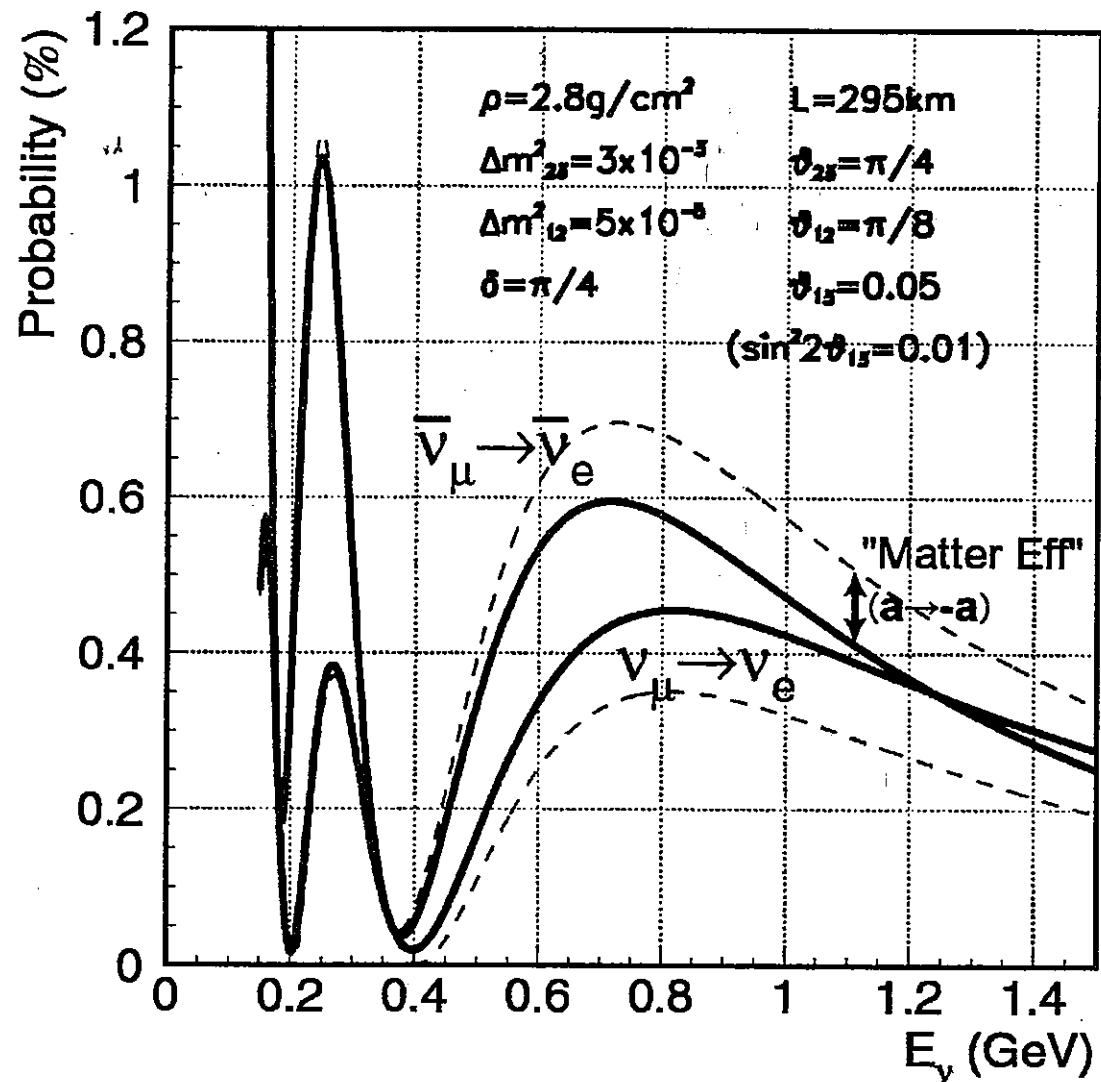
# $\nu_\mu \rightarrow \nu_e$ oscillation probability

$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) = & 4C_{13}^2 S_{13}^2 S_{23}^2 \sin^2 \frac{\Delta m_{31}^2 L}{4E} \times \left( 1 + \frac{2a}{\Delta m_{31}^2} (1 - 2S_{13}^2) \right) \\
 & + 8C_{13}^2 S_{12} S_{13} S_{23} (C_{12} C_{23} \cos \delta - S_{12} S_{13} S_{23}) \cos \frac{\Delta m_{32}^2 L}{4E} \sin \frac{\Delta m_{31}^2 L}{4E} \sin \frac{\Delta m_{21}^2 L}{4E} \\
 & - 8C_{13}^2 C_{12} C_{23} S_{12} S_{13} S_{23} \sin \delta \sin \frac{\Delta m_{32}^2 L}{4E} \sin \frac{\Delta m_{31}^2 L}{4E} \sin \frac{\Delta m_{21}^2 L}{4E} \\
 & + 4S_{12}^2 C_{13}^2 \{ C_{12}^2 C_{23}^2 + S_{12}^2 S_{23}^2 S_{13}^2 - 2C_{12} C_{23} S_{12} S_{23} S_{13} \cos \delta \} \sin^2 \frac{\Delta m_{21}^2 L}{4E} \\
 & - 8C_{13}^2 S_{13}^2 S_{23}^2 \cos \frac{\Delta m_{32}^2 L}{4E} \sin \frac{\Delta m_{31}^2 L}{4E} \frac{aL}{4E} (1 - 2S_{13}^2)
 \end{aligned}$$

$$a = 7.56 \times 10^{-5} [\text{eV}^2] \cdot \left( \frac{\rho}{[\text{g/cm}^3]} \right) \cdot \left( \frac{E}{[\text{GeV}]} \right)$$

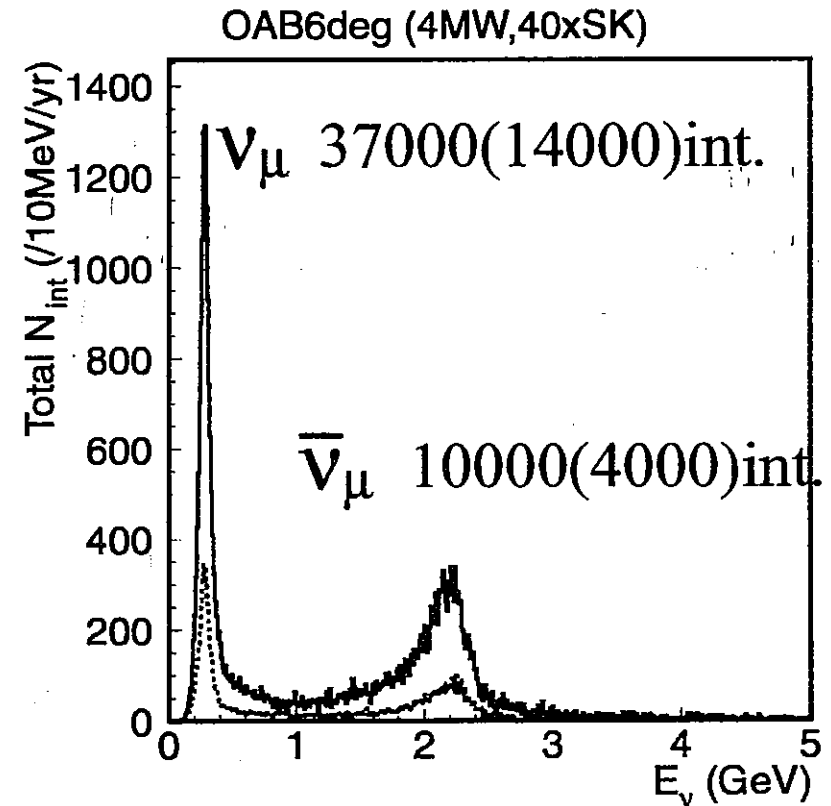
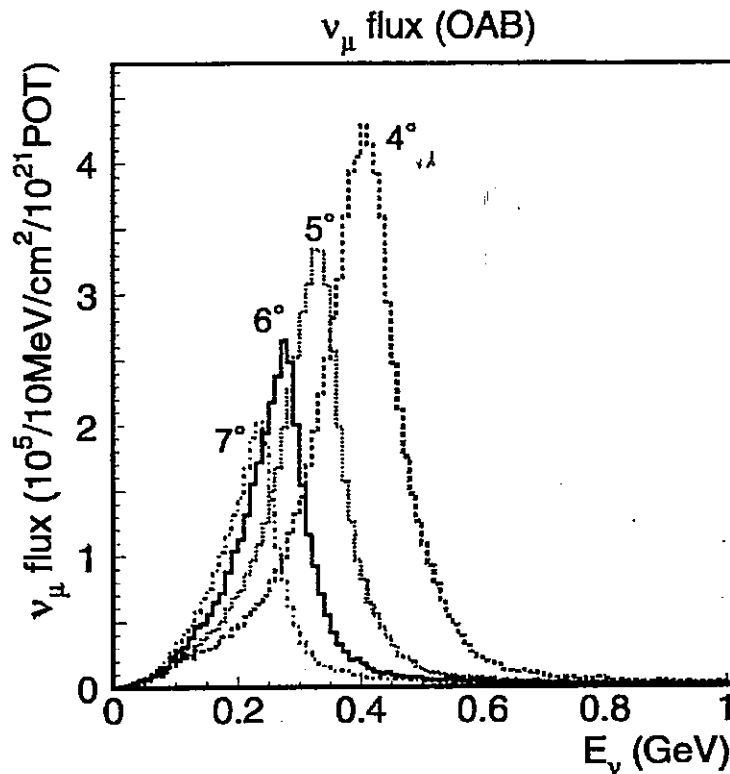
$$\delta \rightarrow -\delta, a \rightarrow -a \text{ for } \bar{\nu}_\mu \rightarrow \bar{\nu}_e$$

# Oscillation Probabilities



# Simulation of Hypothetical SLE Beam

Large angle OAB(150m decay pipe), 4MWx40SK → 200xJHF1



Very small statistics!(c.f.~GeV beam)

( ):  $N_{\text{int}} < 500\text{MeV}$

Need ~3.5times running for  $\bar{\nu}_\mu$

2nd peak from Kaon → serious BG?

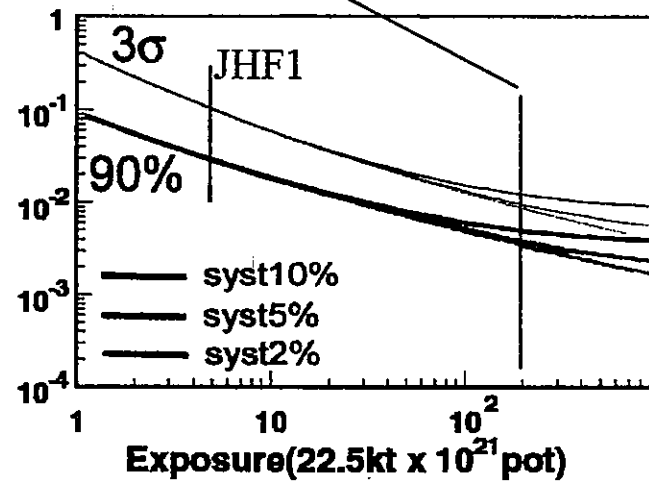
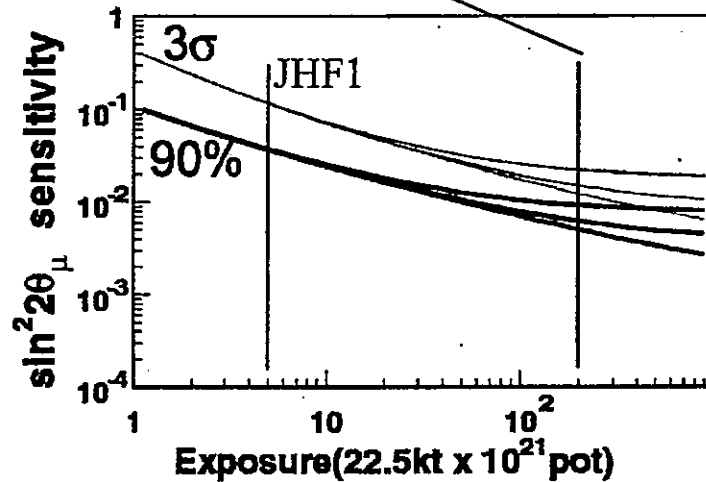
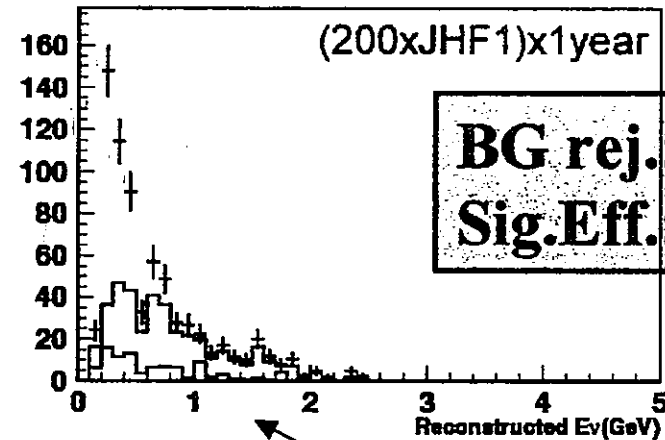
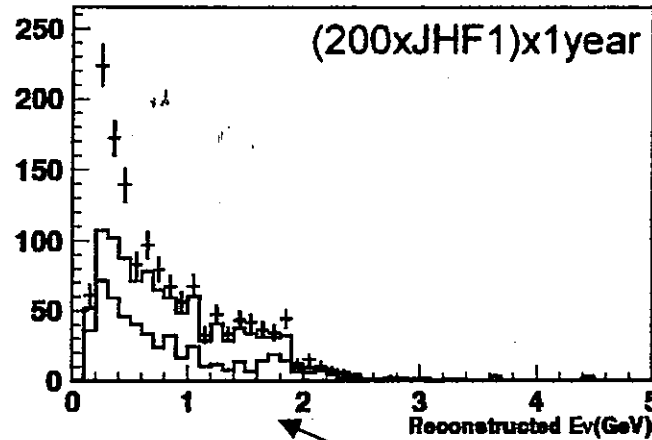
# Sensitivity on $\nu_e$ appearance

## OAB6deg Beam

1ring e-like

+e/ $\pi^0$  separation

$\sin^2 2\theta_{13} = 0.1$   
 $(\sin^2 2\theta_{\mu e} = 0.05)$



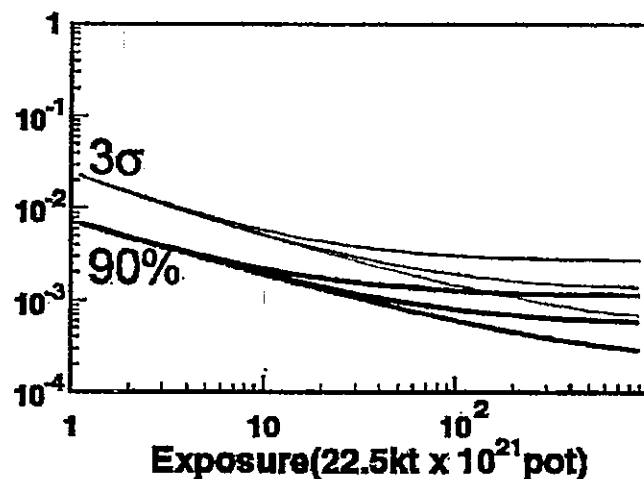
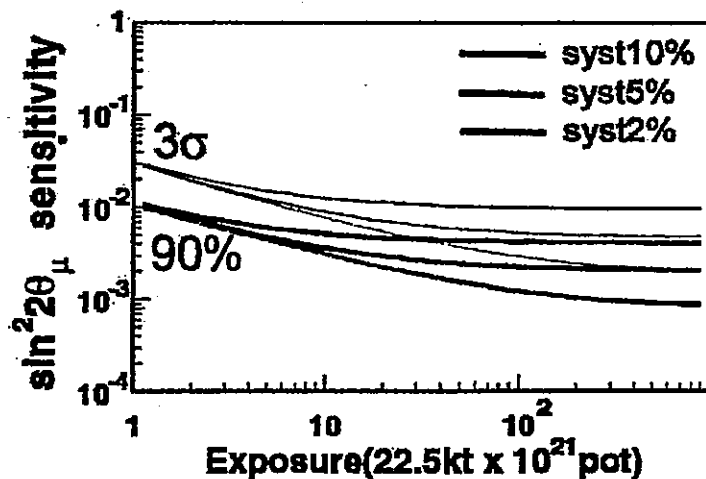
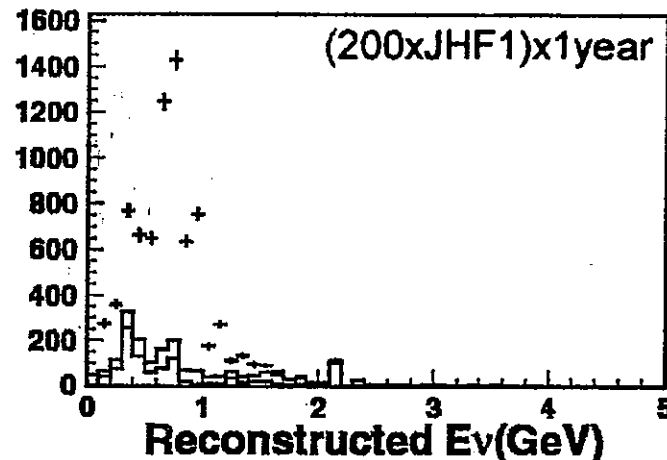
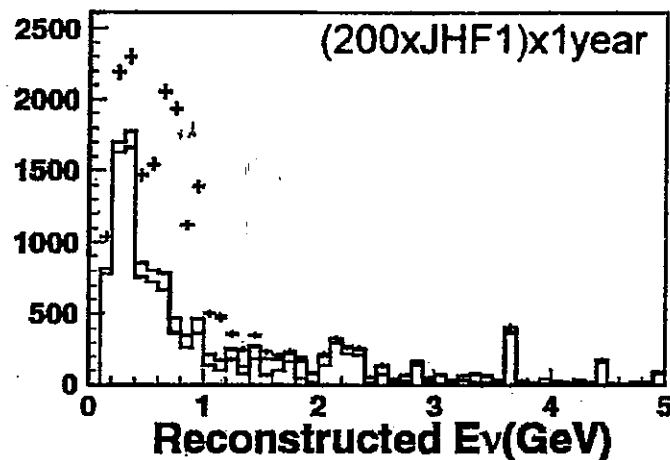
(Beam  $\nu_e$  contamination is not negligible.)

# Sensitivity of $\sim$ GeV beam for comparison

## OAB2deg Beam

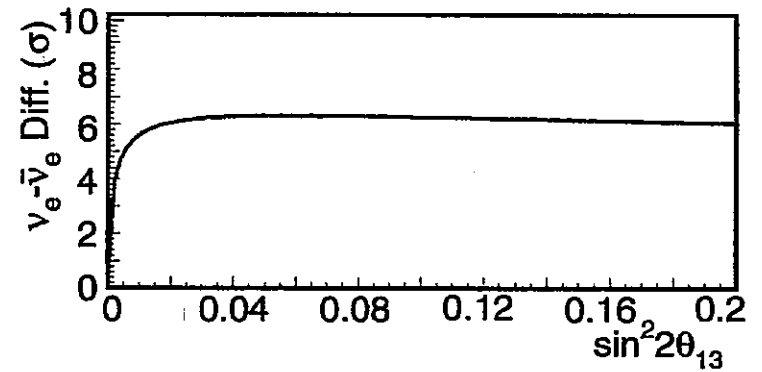
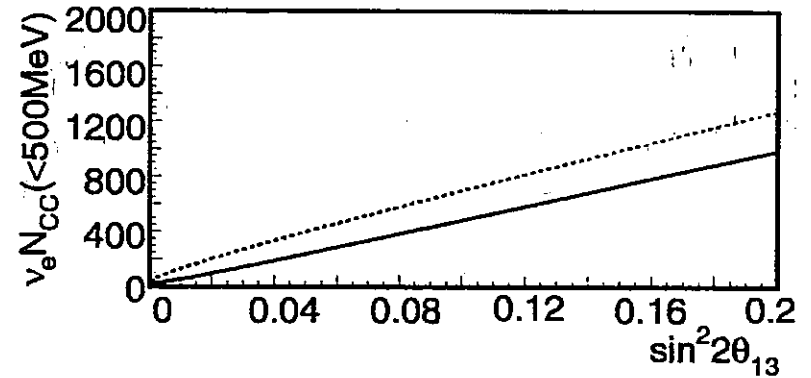
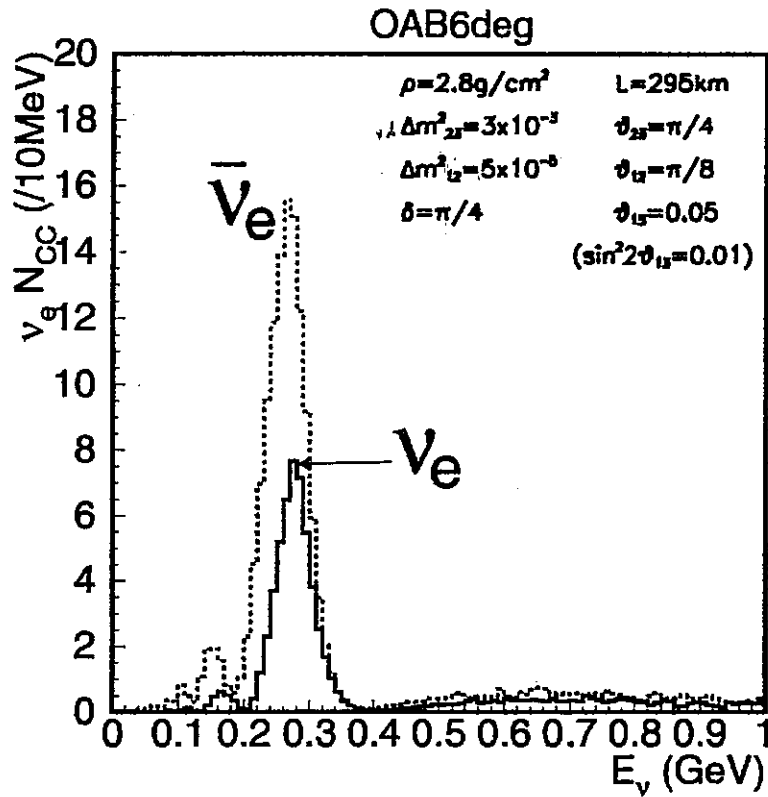
1ring e-like

+e/ $\pi^0$  separation



**OAB2deg(w/ e/ $\pi^0$ ) much higher sens. than SLE beam**

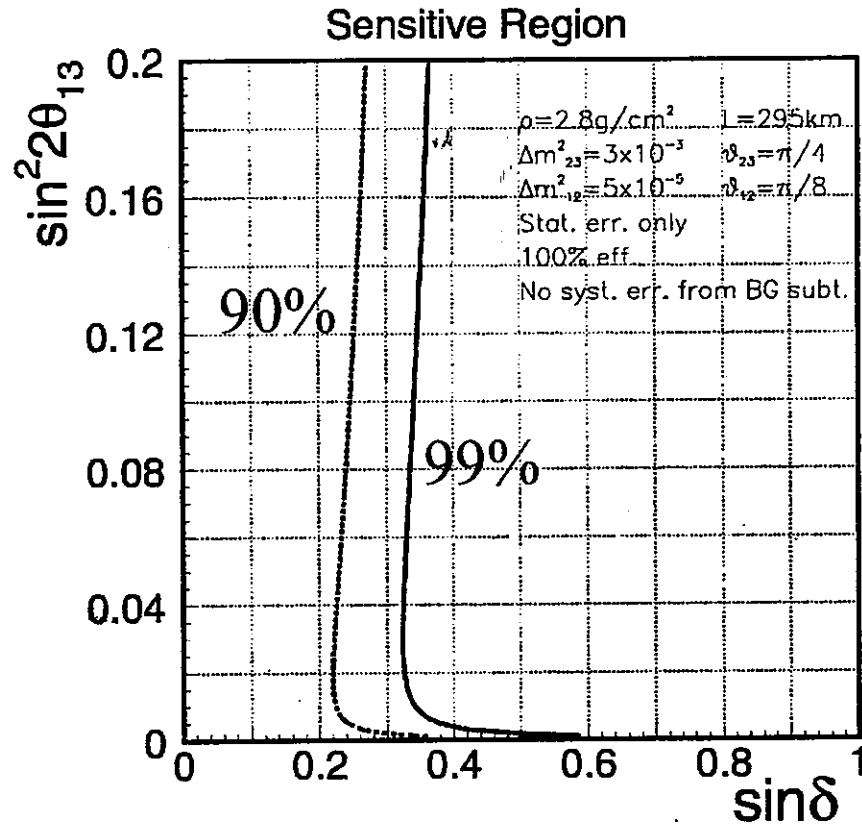
# CPV Measurement



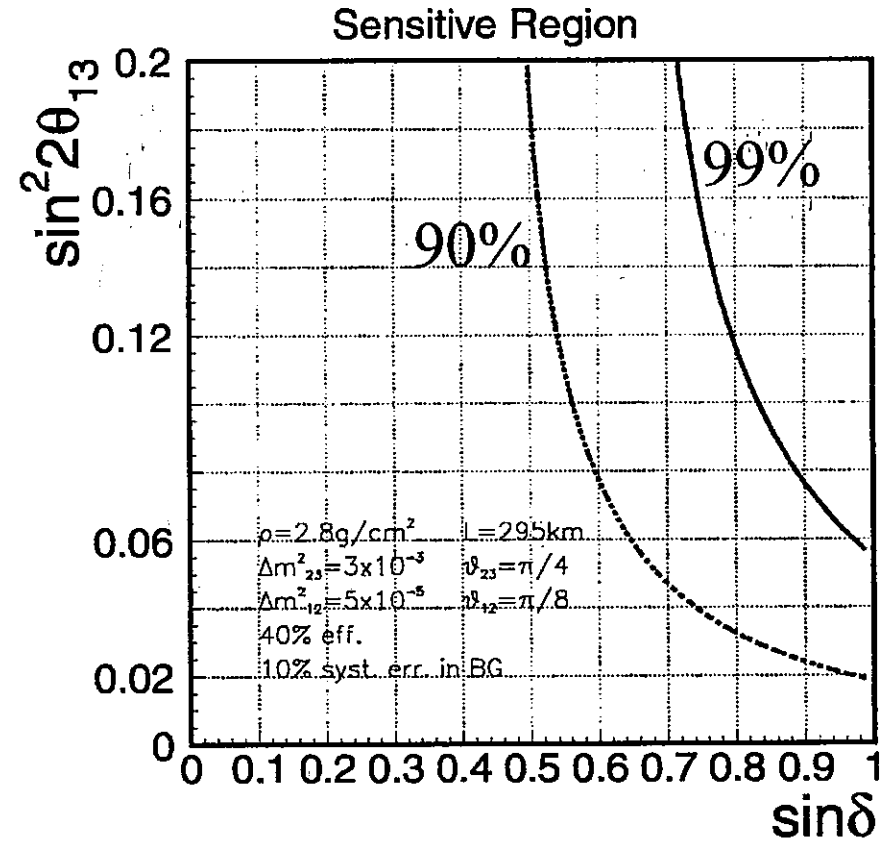
$\nu_\mu$  beam: 2years  
 $\nu_\mu$  beam: 7.5years

# Sensitivity on CPV

OAB6deg Ideal Case



OAB6deg Realistic



200xJHF1

2years for  $\nu_\mu$

7.5years for  $\bar{\nu}_\mu$

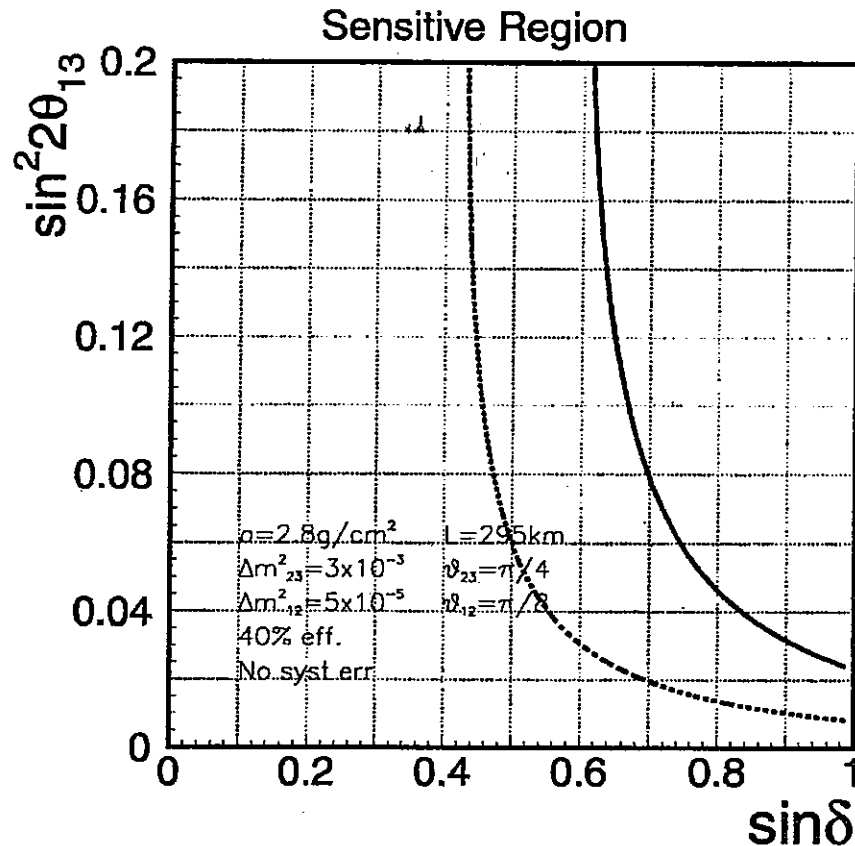
$\epsilon = 40\%$ , BG rej: 0.3%

BG syst.: 10%

$\nu_\mu \text{ BG} = \bar{\nu}_\mu \text{ BG}$



# Sensitivity on CPV



Slightly better than  
w/ BG syst error.

But, essentially,  
sensitivity is limited by  
stat. error due to  
bad S/N ratio.

$e = 40\%$ ,  $BG_{rej} = 0.3\%$   
No BG syst. err

# Summary of SLE beam study

- Studied potential of SLE Super Conventional beam
- Worse sensitivity in  $\nu_e$  appearance due to small statistics than  $\sim$ GeV beam
- Sensitivity in CPV is limited by BAD S/N ratio.
- Possible sources of systematic error
  - rapid rise of cross section
  - $\nu_{\mu}$   $\bar{\nu}_{\mu}$  BG difference
  - $\pi/K$  production ratio at target(affect both  $\nu_{\mu}$   $\nu_e$  flux)
  - .....
- Possible improvements....
  - go further
  - compare 1st and 2nd osc. max (by Konaka)
  - .....