



# Updates of sensitivity study

10/6/2005

2km meeting

1. Current status
2. Sensitive curve
3. Conclusions and Future Plans

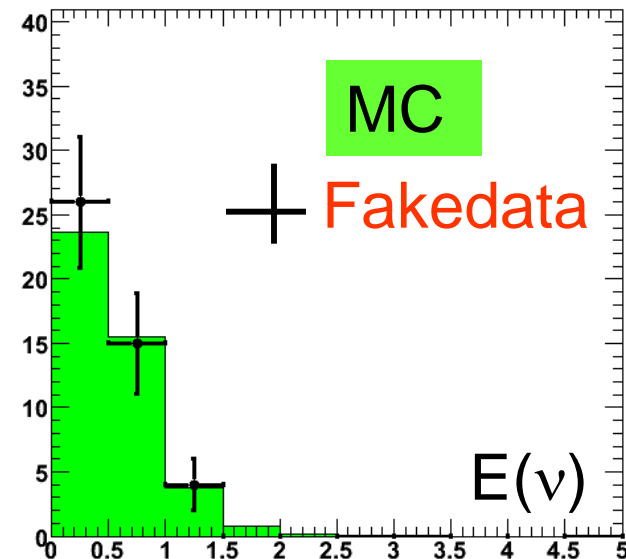
Naho Tanimoto

Maximilien Fechner

Chris Walter

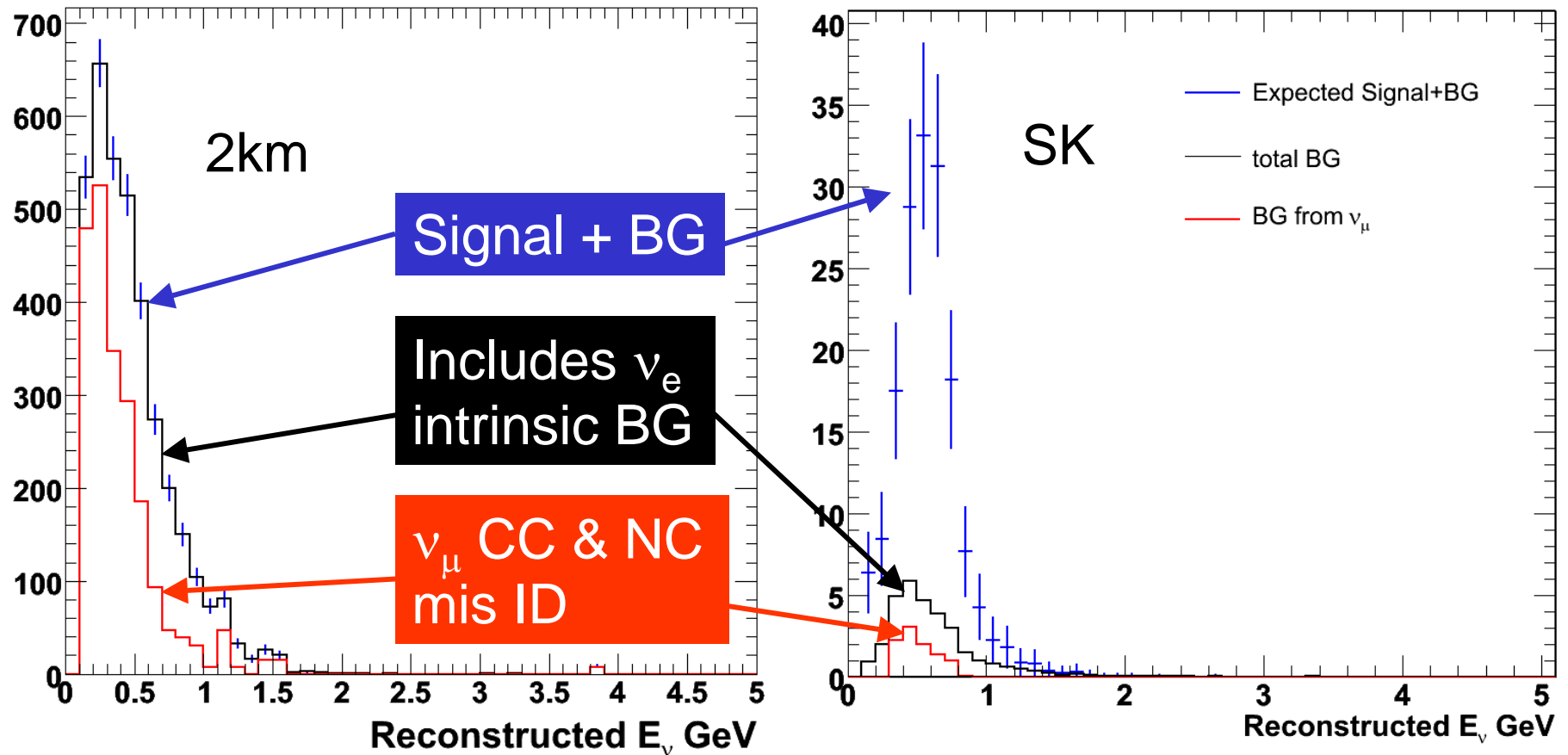
**Goal** : Do full  $\nu_e$  sensitivity analysis using both the 2km and SK MC with realistic systematic errors

- MC samples : SK [  $\nu_e$ ,  $\nu_\mu$  ] and 2km [  $\nu_e$ ,  $\nu_\mu$  ]
- Applied  $\nu_e$  appearance cuts +  $0.35 < E(\nu) < 0.85$  GeV  
(See next slide)
- For the sensitivity study, generated 'fake data' at both of SK and 2km based on non-oscillated  $E(\nu)$ 
  - Generate Poisson random number in each energy bin  
→ Fakedata
  - Centered value of random number is based on rescaled MC with fluctuation given by proper statistics



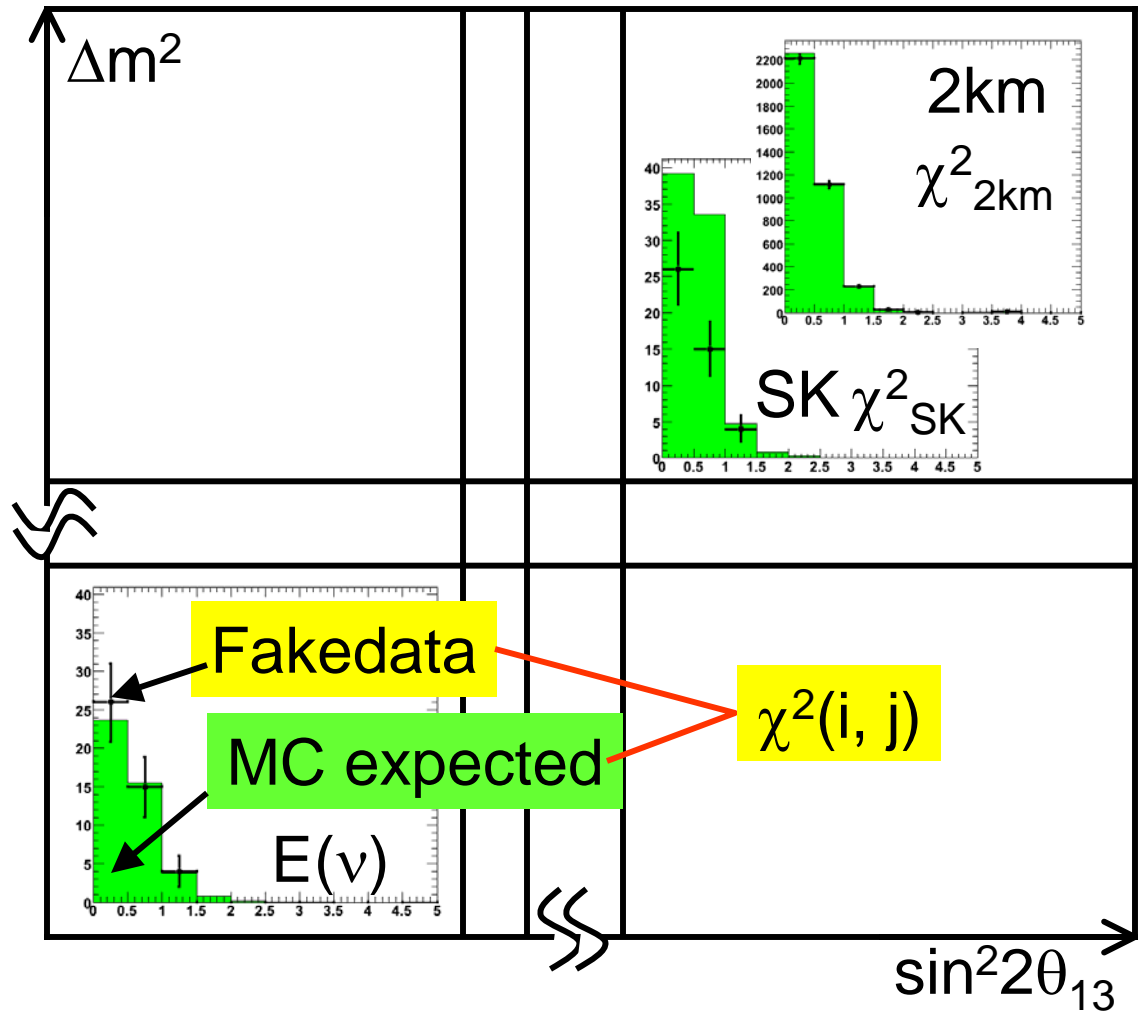
# $E(\nu)$ with all $\nu_e$ appearance cuts

- $\nu_e$  analysis cuts applied
- Fully reconstructed  $\nu_e$  energy at SK and 2km for 5 yr with  $\sin^2 2\theta_{13}=0.1$  and  $\Delta m^2=0.0025$



Now we apply cuts correctly

- Made map of  $E(\nu)$  histograms at both of SK and 2km in  $\sin^2 2\theta_{13}$  and  $\Delta m^2$  space
- Calculate  $\chi^2$  between Fakedata and MC,  $\chi^2 = \chi^2_{SK} + \chi^2_{2km}$
- Calculate  $\Delta\chi^2(i,j) = \chi^2(i,j) - \chi^2(\min)$



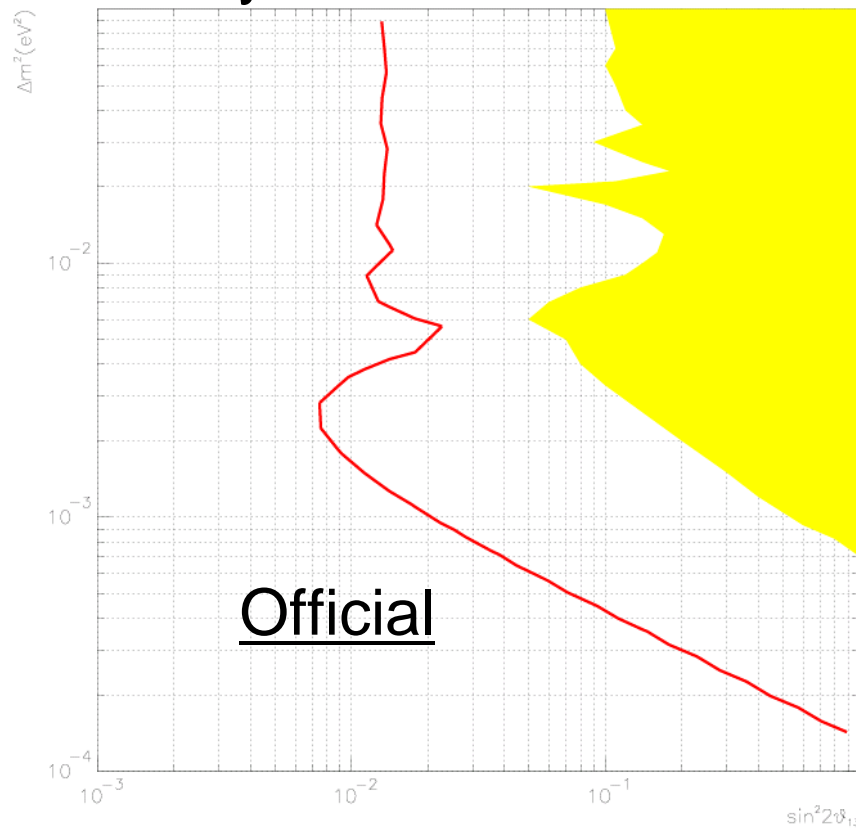


# Sensitivity curve at 90% C.L.

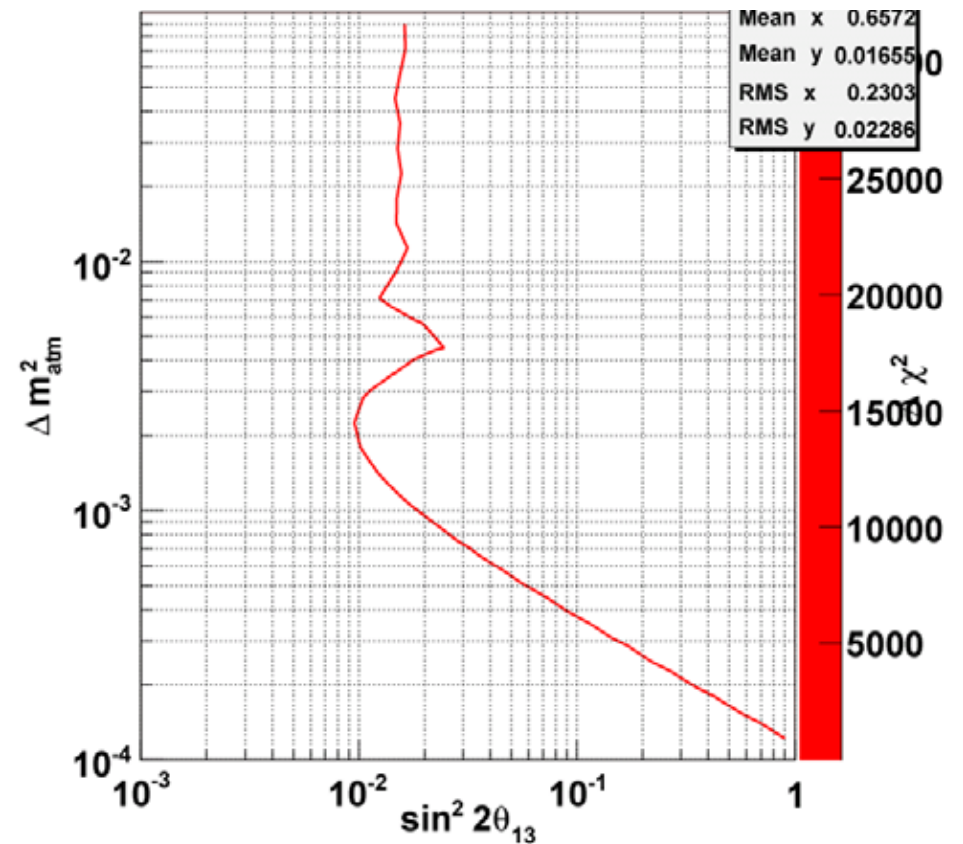


With 10% systematic uncertainty  
SK only

No systematic uncertainty  
SK + 2km



Based only on # of events



Based on of  $\chi^2$  energy distribution at SK and 2km

# List of Systematic errors

SK atmospheric  
 $\nu$  group's pull fit  
 method

$$\chi^2 = \sum_{n=1} \frac{\left( N_{obs}^n - N_{exp}^n \left( 1 + \sum_{k=1}^K f_k^n \cdot \epsilon^k \right) \right)^2}{\sigma_n^2} + \sum_{k=1}^K \left( \frac{\epsilon^k}{\sigma_k} \right)^2$$

$N_{obs}^n$  : observed,  $N_{exp}^n$  : expected,  $\epsilon^k$  : systematics

SK:

- ~~Neutrino flux~~
- ~~flux absolute normalization~~
- ~~flavor ratios ( $E_\nu < 1.33 \text{ GeV}, E_\nu > 1.33 \text{ GeV}$ )~~
- ~~anti-neutrino/neutrino ratio ( $\nu_e, \nu_\mu$ )~~
- ~~Up/down ratio~~
- ~~Horizontal/vertical ratio (3D calc.,  $K/\pi$ )~~
- ~~Neutrino flight length~~
- ~~Energy spectrum~~
- ~~Sample-by-sample normalization (FC multi-GeV, PC+up stop  $\mu$ )~~

## Neutrino interaction

- $M_A$  in quasi-elastic and single-pi
- Quasi elastic scattering (model dependence)
- Quasi elastic scattering (cross section)
- single-pion production (cross section)
- multi-pion production (model dependence)
- multi-pion production (cross section)
- coherent pion production (cross section)
- NC/CC ratio
- Nuclear effect in  $^{16}\text{O}$
- Charged current  $\nu_\tau$  interaction

## (C) Event selection

- a. FC reduction
- b. PC reduction
- c. Up- $\mu$  detection efficiency
- d. FC/PC relative normalization
- e. Hadron simulation
- f. Non- $\nu$  BG (e-like,  $\mu$ -like)
- g. Through-going/stopping  $\mu$  separation

+ T2K related :

- 2km detector
- ....

## (D) Event reconstruction(6)

- a. 1-ring/multi-ring separation
- b. Particle ID (single-ring, multi-ring)
- c. Energy calibration for FC
- d. Energy cut for upward stopping  $\mu$
- e. Up-down asymmetry of energy calibration

Total number of errors: 36

**Treated as independent error term in  $\chi^2$  calculation**

Saji's talk at Noon2004

## Conclusion

- Made ROOT based analysis framework
- Fake data generator is implemented
- Made sensitivity curve without systematic uncertainty

## Plan

- Started to incorporate systematic uncertainties using SK Atmospheric  $\nu$  group's pull fit method
- Systematic uncertainties (two ways)
  - Calculate correlations between systematic errors and solve equations analytically
  - Use MINUIT to minimize  $\chi^2$  for systematic errors
- Generate a lot of fakedata histograms  $\rightarrow$  Average of  $\chi^2$
- Flat beam intensity profile in time  $\rightarrow$  Use time dependent intensity
- Add more histograms to constrain systematic errors (i.e.  $\pi^0$  BG, intrinsic  $\nu_e$  BG)