

# The Quest for the Origin of Cosmic Rays

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Heidelberg







Sagittarius  
Region

Hubble  
Space  
Telescope





Galaxy M 51  
T. & M. Hallas



Cone Nebula  
ACS NASA

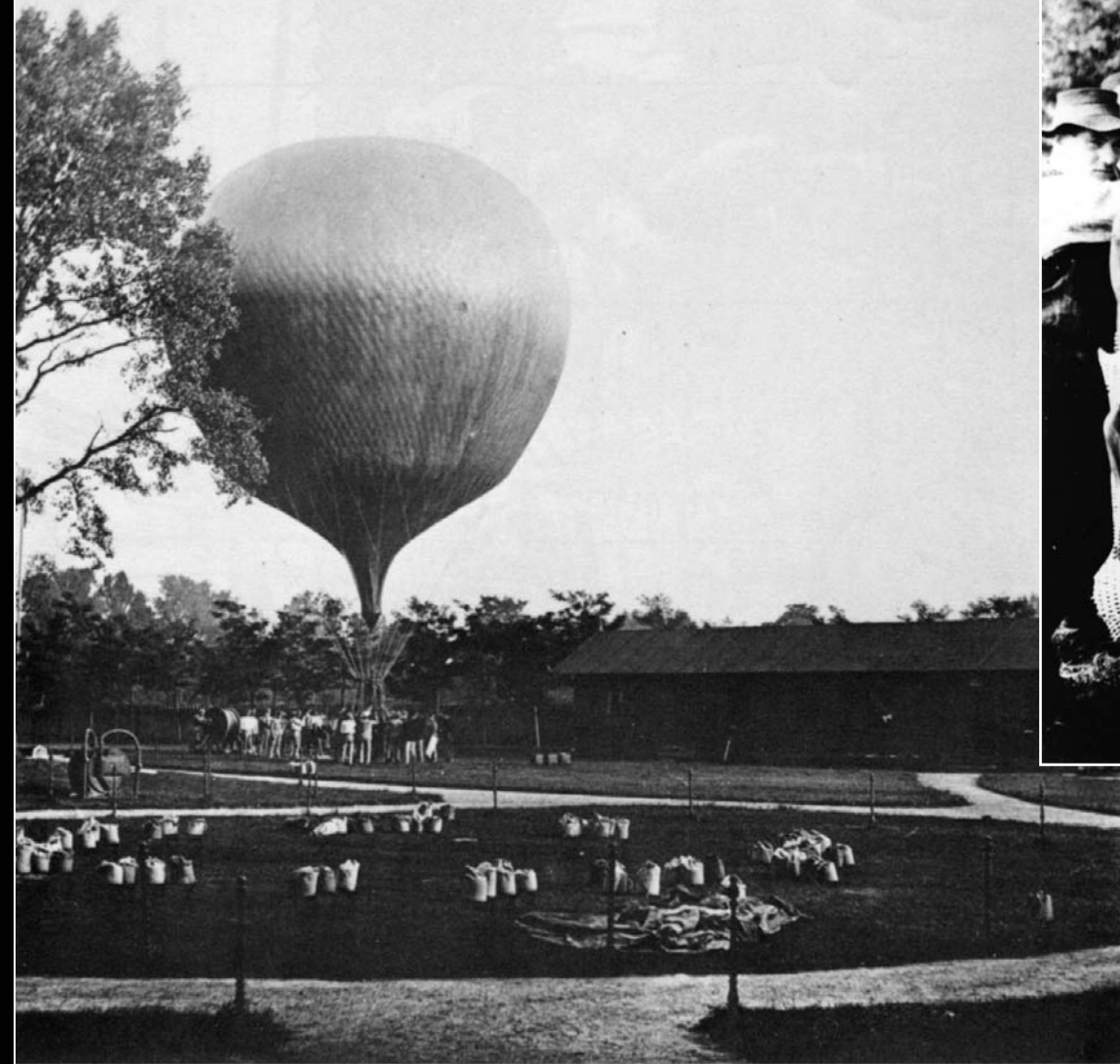


Eagle Nebula  
T.A. Rector &  
B.A. Wolpa  
NOAO





# Cosmic rays



Discovery:  
Victor Hess 1912  
(Nobel Prize 1932)



# A detective story in science

## The case:

A 90-year mystery –  
the origin of Cosmic Rays

## The suspects:

Black holes & exploding stars

## The evidence:

Blue flashes from space -  
Astrophysics with Cherenkov  
telescopes

## The judgement





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A 90-year mystery –  
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The suspects:  
Black holes & dying stars



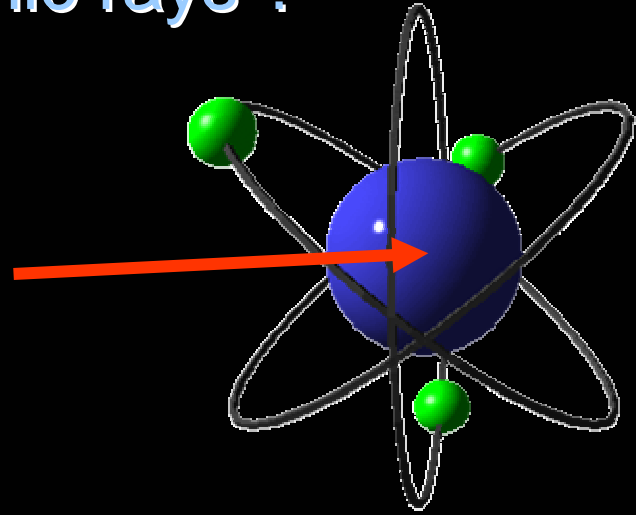
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# What do we know about cosmic rays ?

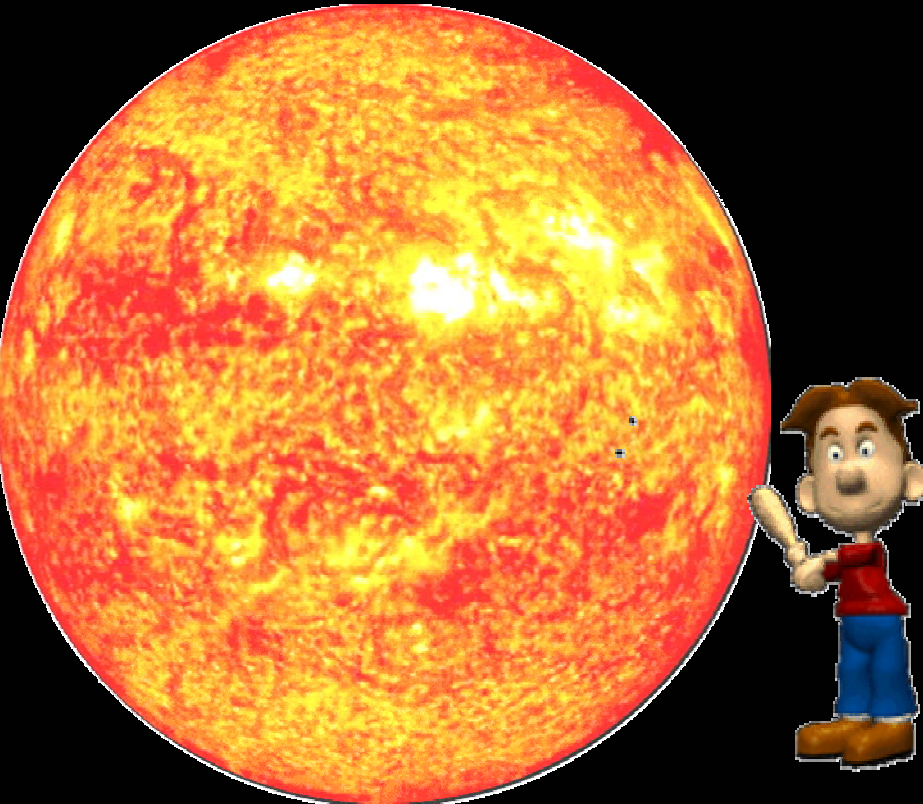
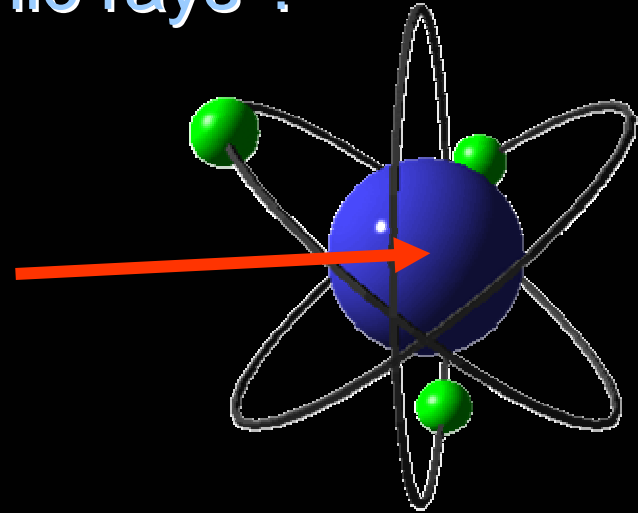
- They consist mostly of atomic nuclei, like those found in the sun





# What do we know about cosmic rays ?

- They consist mostly of atomic nuclei, like those found in the sun

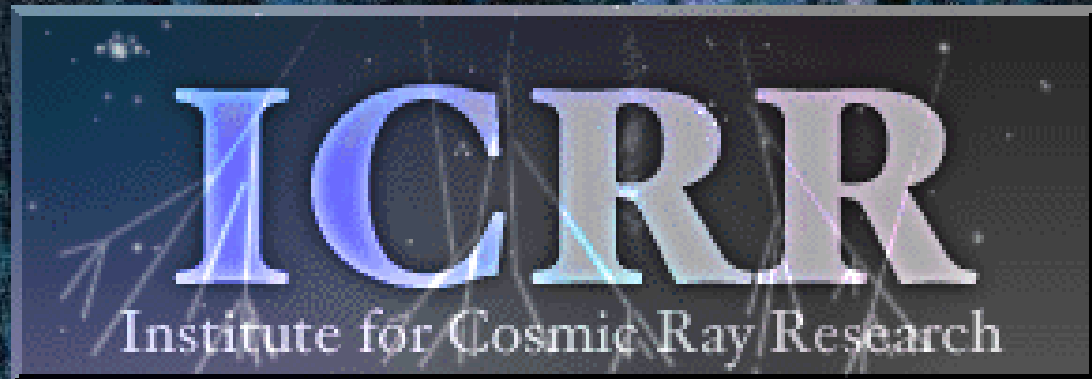


- They cover a wide spectrum of energies; the highest energies are 100000000 times higher than those of the largest man-made accelerator
- They cover the whole sky more or less uniformly ... they come from everywhere



# Why do we care ?

The sources of cosmic rays must be some of the most interesting and most violent objects in the Universe



Exploding stars: Supernovae

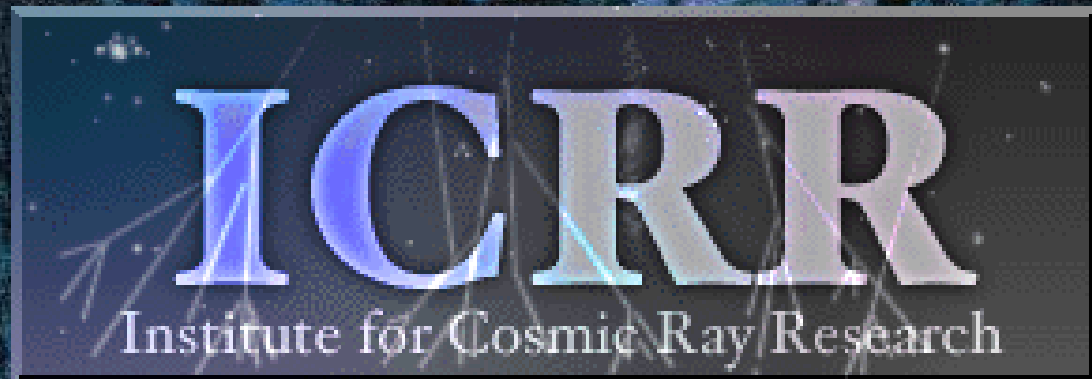


Giant Black  
Holes at the  
Cores of Galaxies



# Why do we care ?

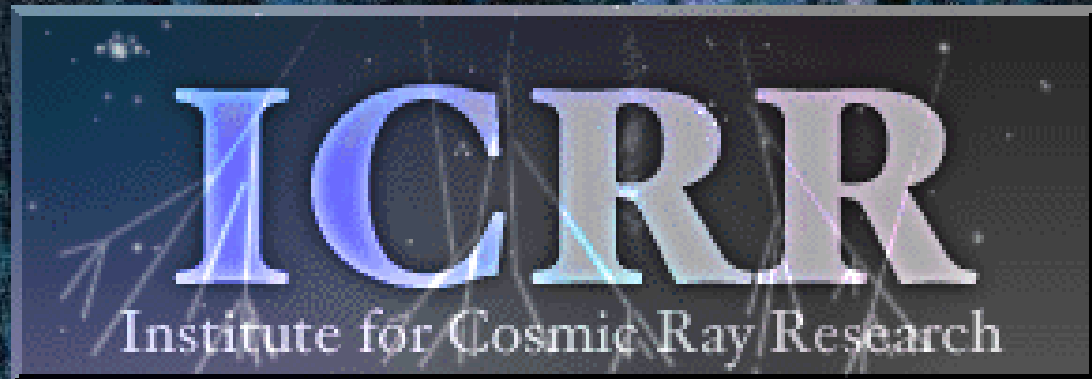
Cosmic rays are  
“responsible” for the  
evolution of species





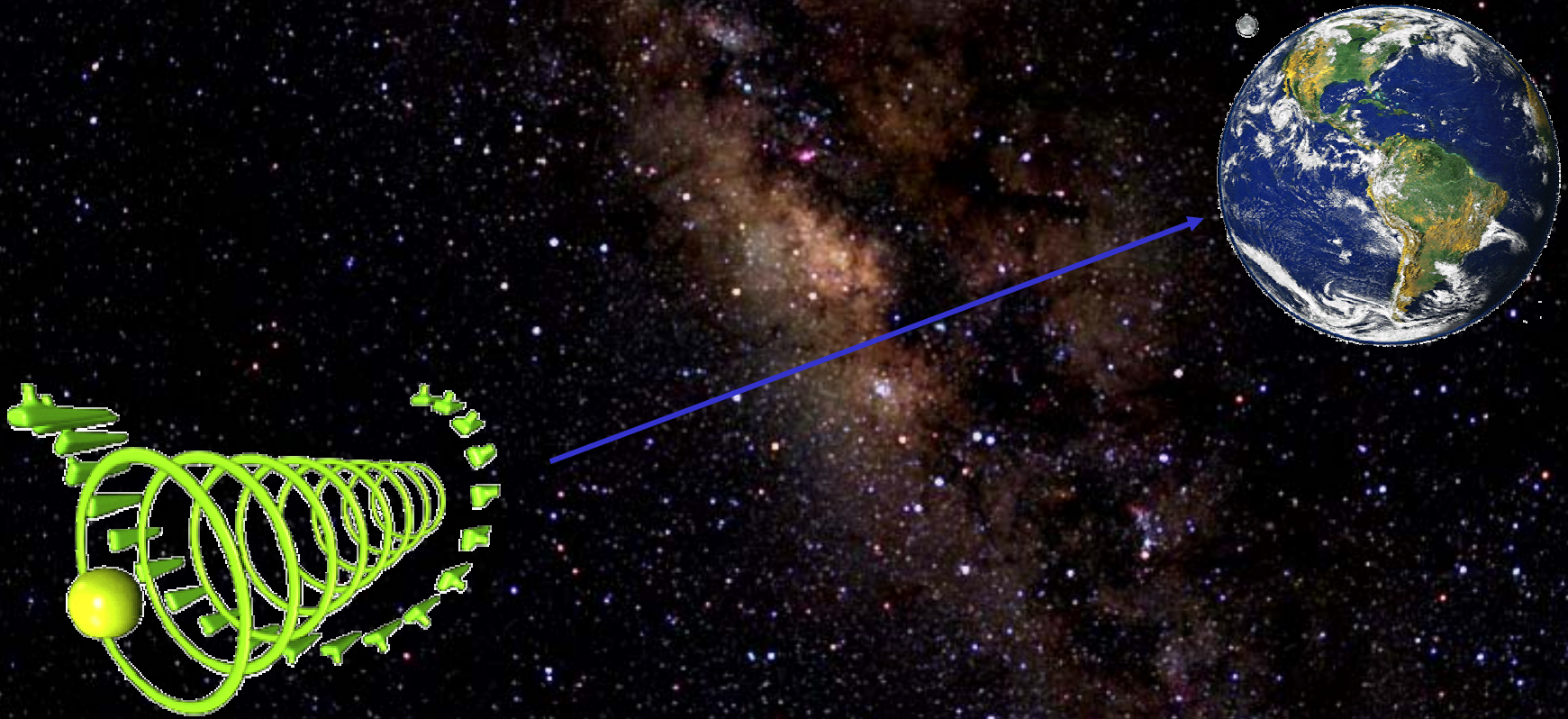
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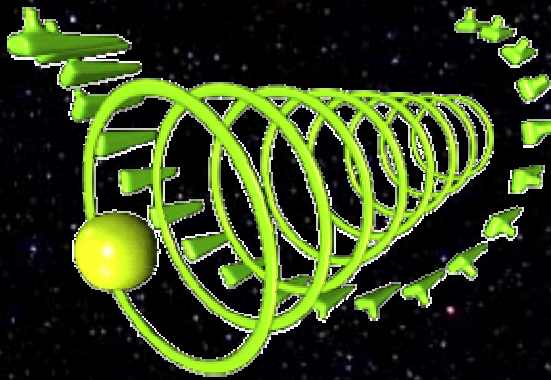


Why is it so difficult  
to spot the origins of Cosmic Rays ?



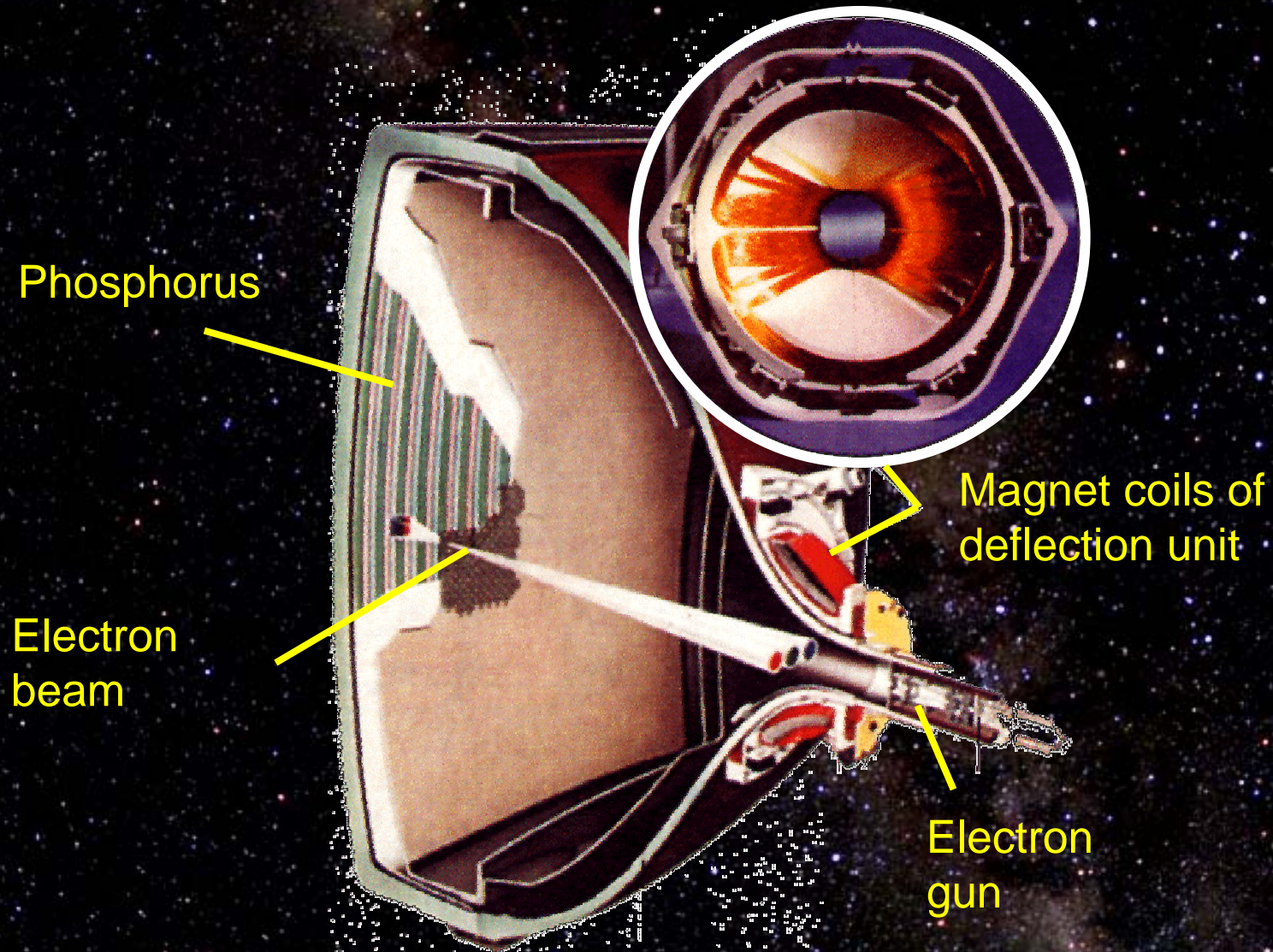


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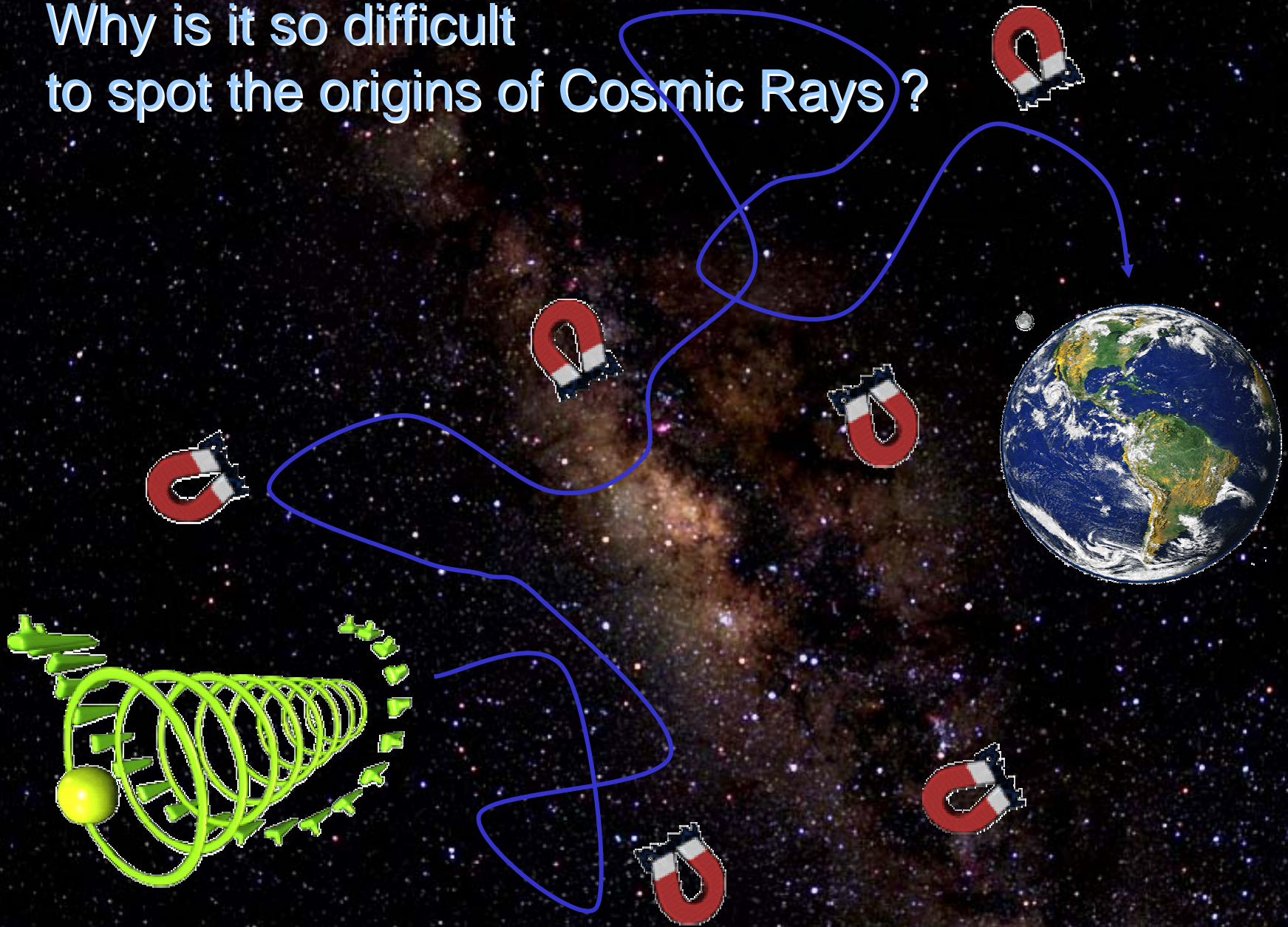


# Magnetic fields bent particle tracks !





Why is it so difficult  
to spot the origins of Cosmic Rays?





Think!

What objects could be cosmic accelerators ?  
Are there other ways to find them ?





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**The suspects:**

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The evidence:

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Astrophysics with Cherenkov  
telescopes

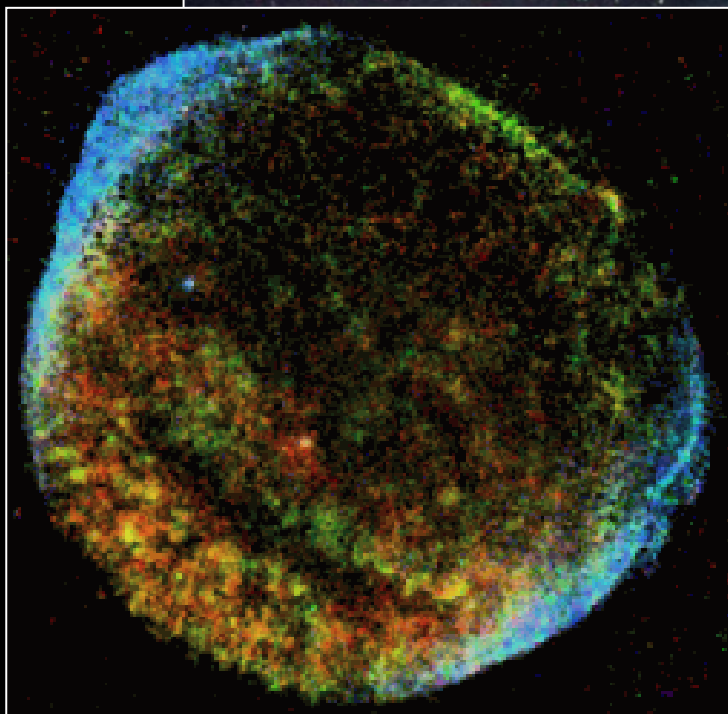
The judgement



# Supernova 1006

T. Tezel

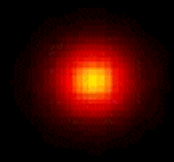
50 light years  
0.5° diameter



以後七月以前、客星入羽林中  
1006 一條院寛弘三年四月二  
日、癸酉、夜以降騎官中有大客星、如熒惑、光明動耀、  
連夜正見南方、或云、騎陣將軍星變本體增光歟  
1054 後冷  
泉院天喜二年四月中旬以後丑時、客星出嵯參度、見東  
方、孝天關星、大如歲星、二條院永萬二年四月廿二日、

from Meigetsuki

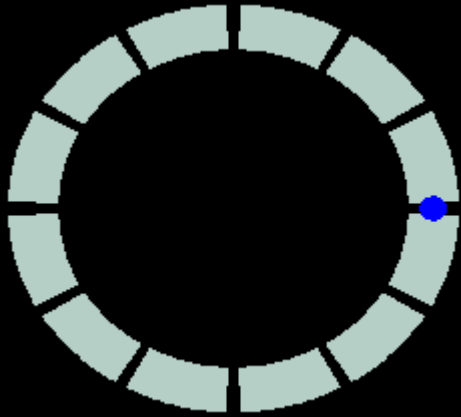
# Galactic accelerators: Supernovae ?





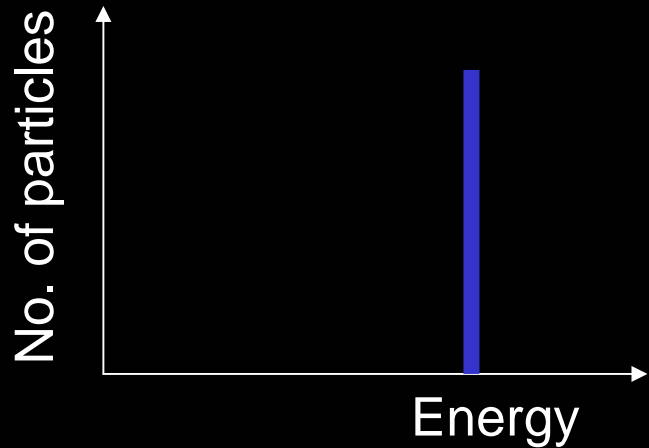
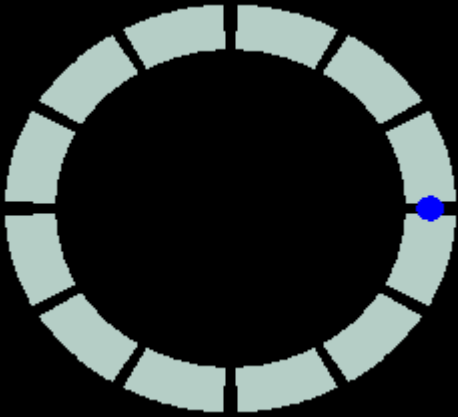
# How might such cosmic accelerators work?

Man-made accelerators



# How might such cosmic accelerators work?

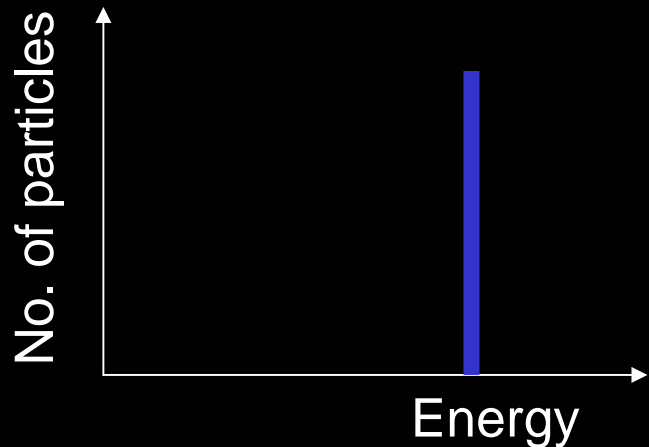
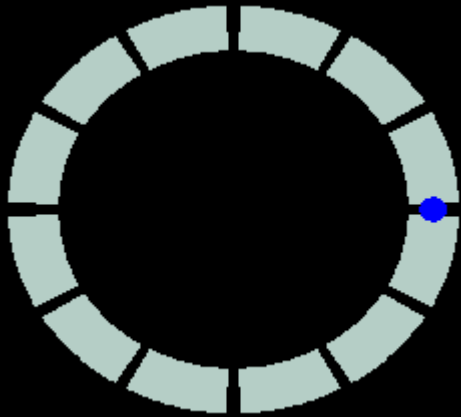
Man-made accelerators



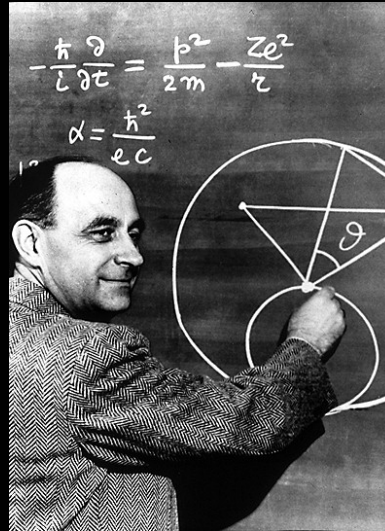


# How might such cosmic accelerators work?

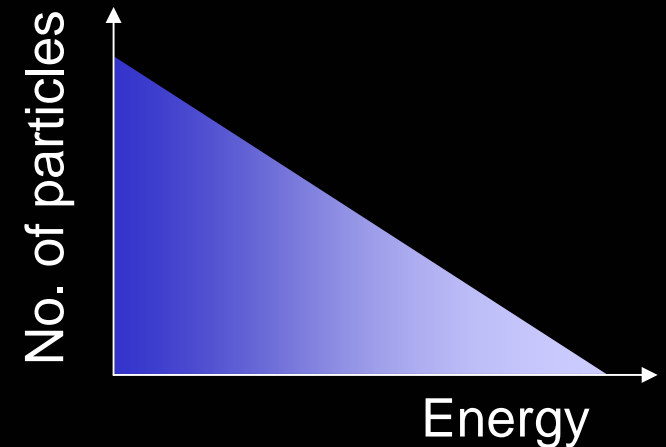
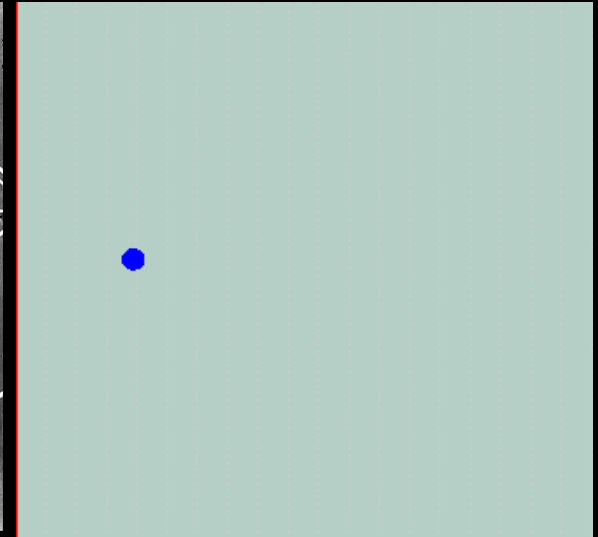
Man-made accelerators



Nature's accelerators



Enrico  
Fermi



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telescopes

The judgement

*but is this really true  
can we observe it ?*





# Galactic accelerators: Supernovae ?

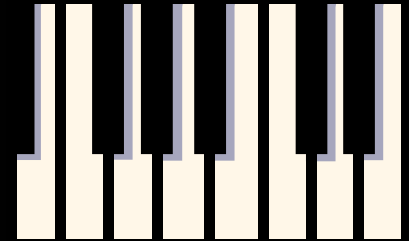
- Image the source using neutral particles
- Astronomy with gammas and neutrinos



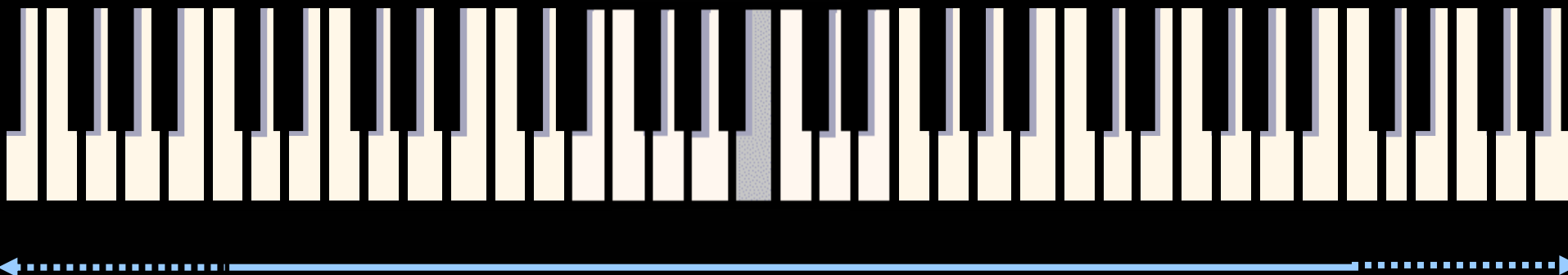
# The spectrum of electromagnetic radiation: from radio waves to gamma rays



Visible light encompasses  
one octave in wavelength



... but the spectrum of electromagnetic radiation  
from space extends over more than 70 octaves ...



... which means that Nature is playing a 15 m long piano ...

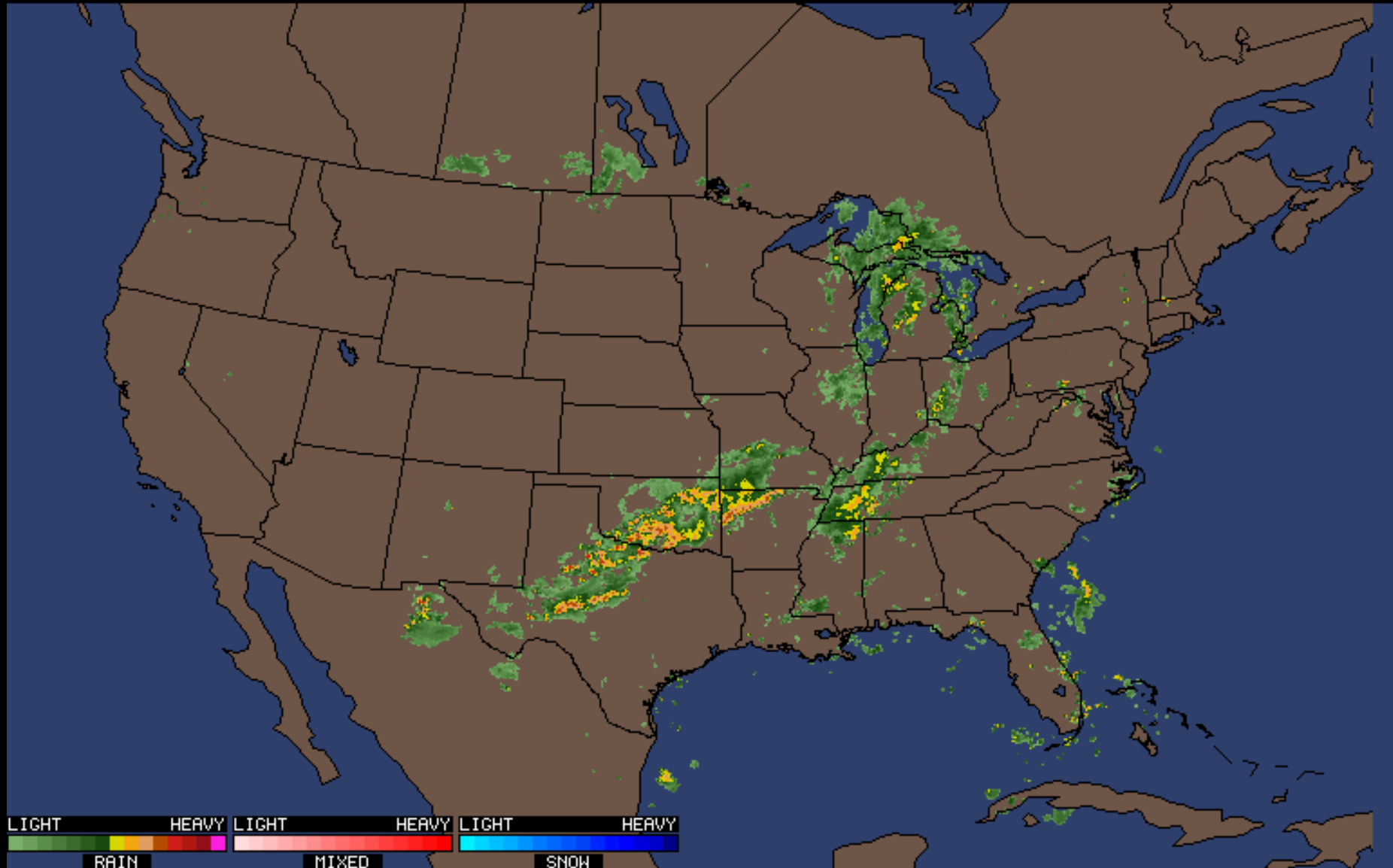


# Infrared images



# Radio waves

04:00 14-AUG-2002 GMT ©Copyright WSI Corporation <http://www.wsi.com>





X-ray  
images





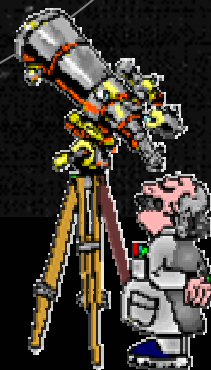
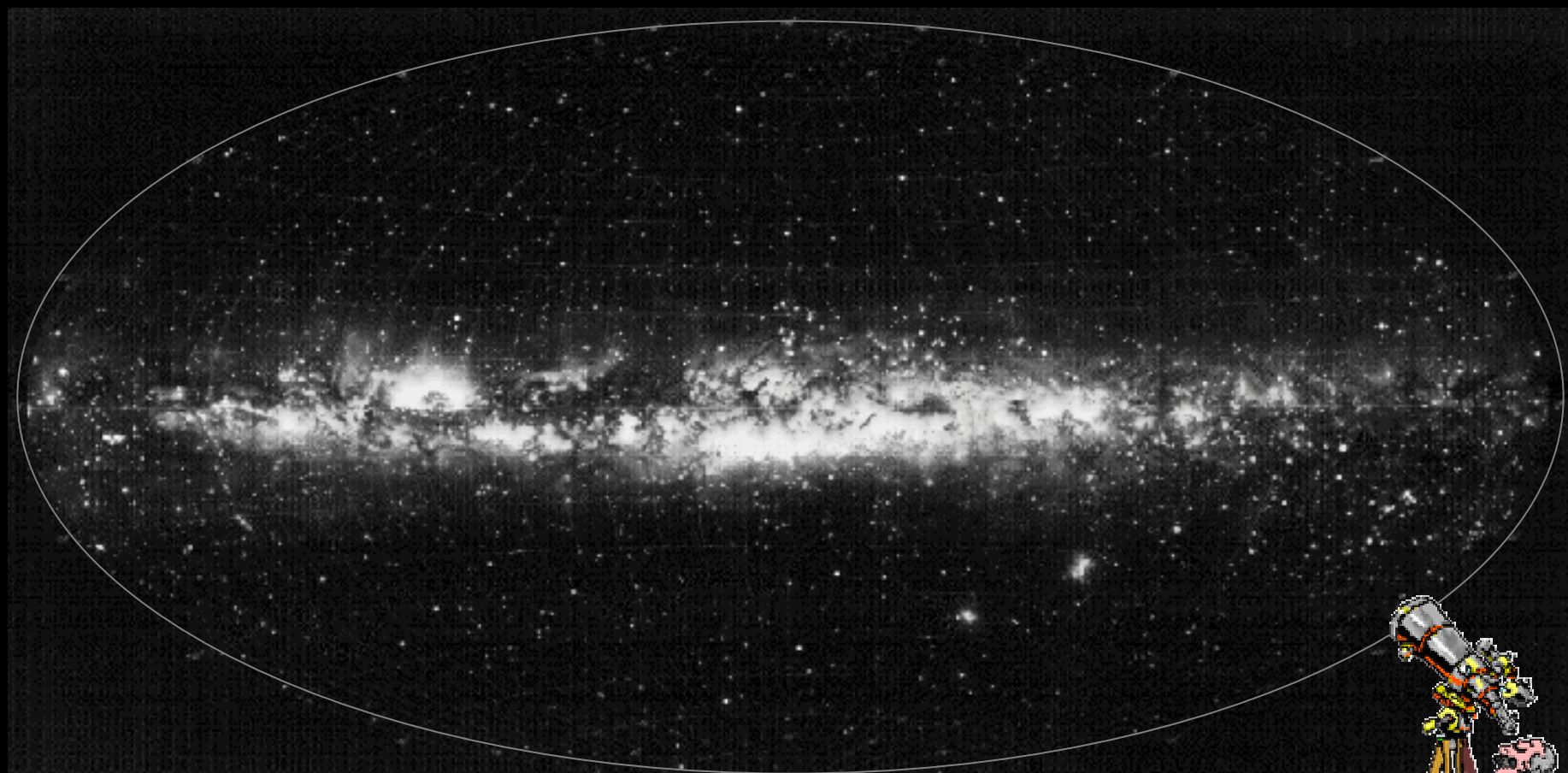
Radio

Infrared

Visible Light  
(eV)

X-rays

Gamma rays







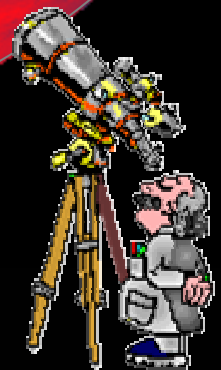
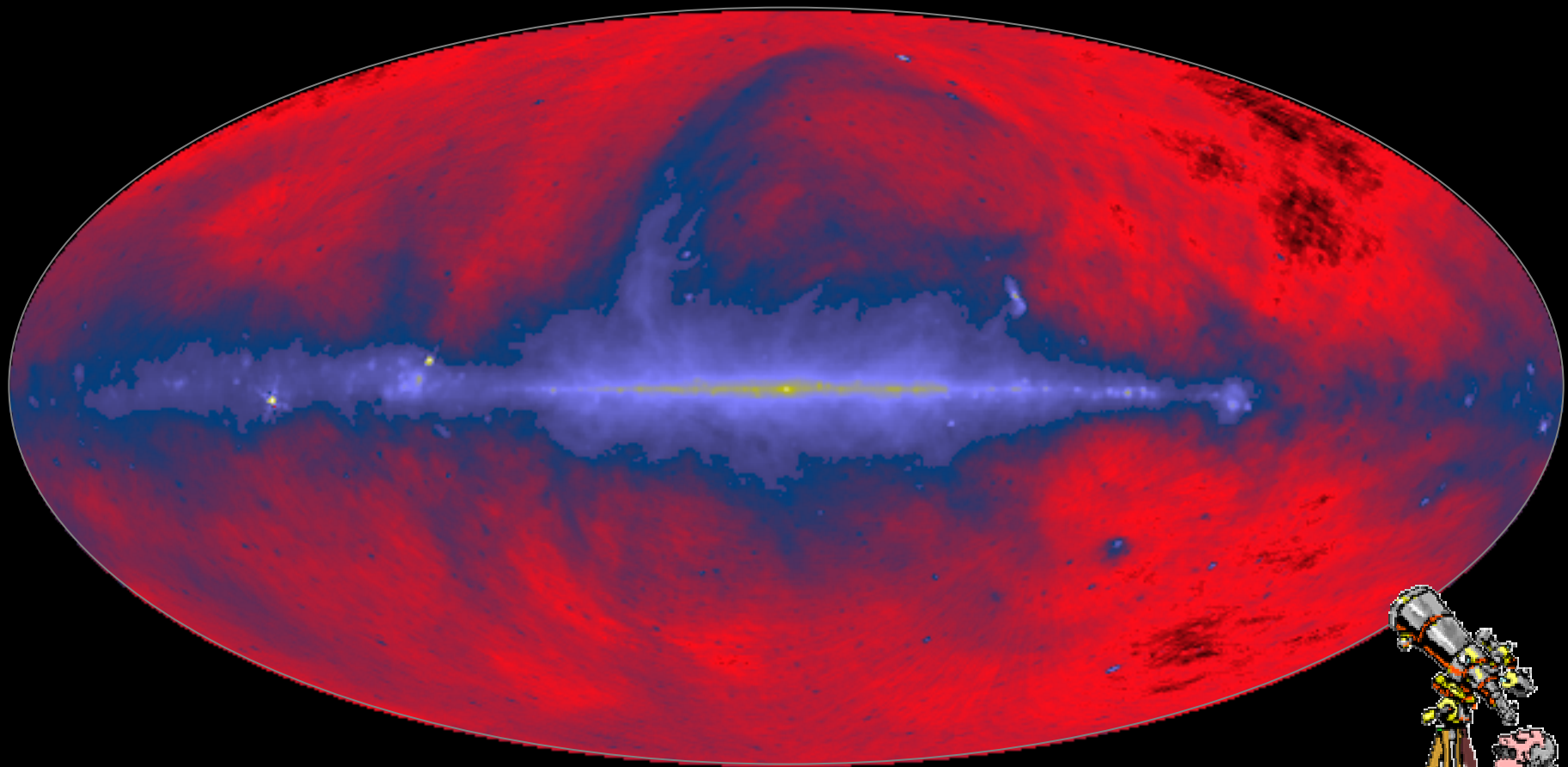
Radio  
( $10^{-6}$  eV)

Infrared

Visible Light

X-rays

Gamma rays





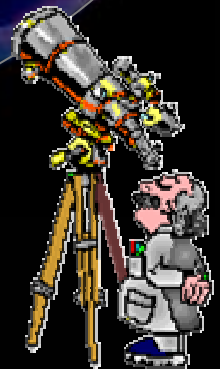
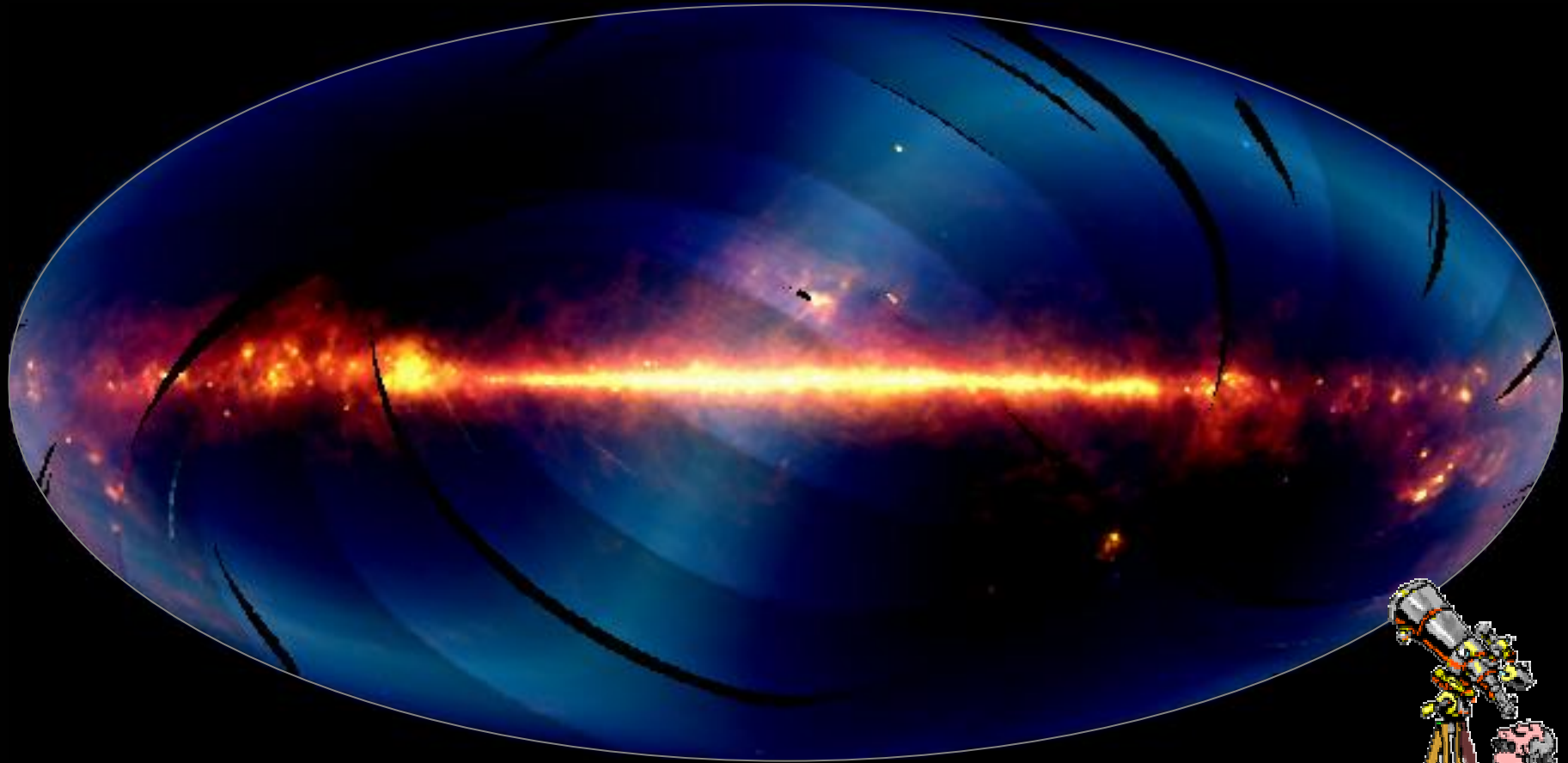
Radio

Infrared  
( $10^{-2}$  eV)

Visible Light

X-rays

Gamma rays







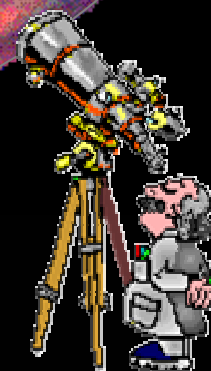
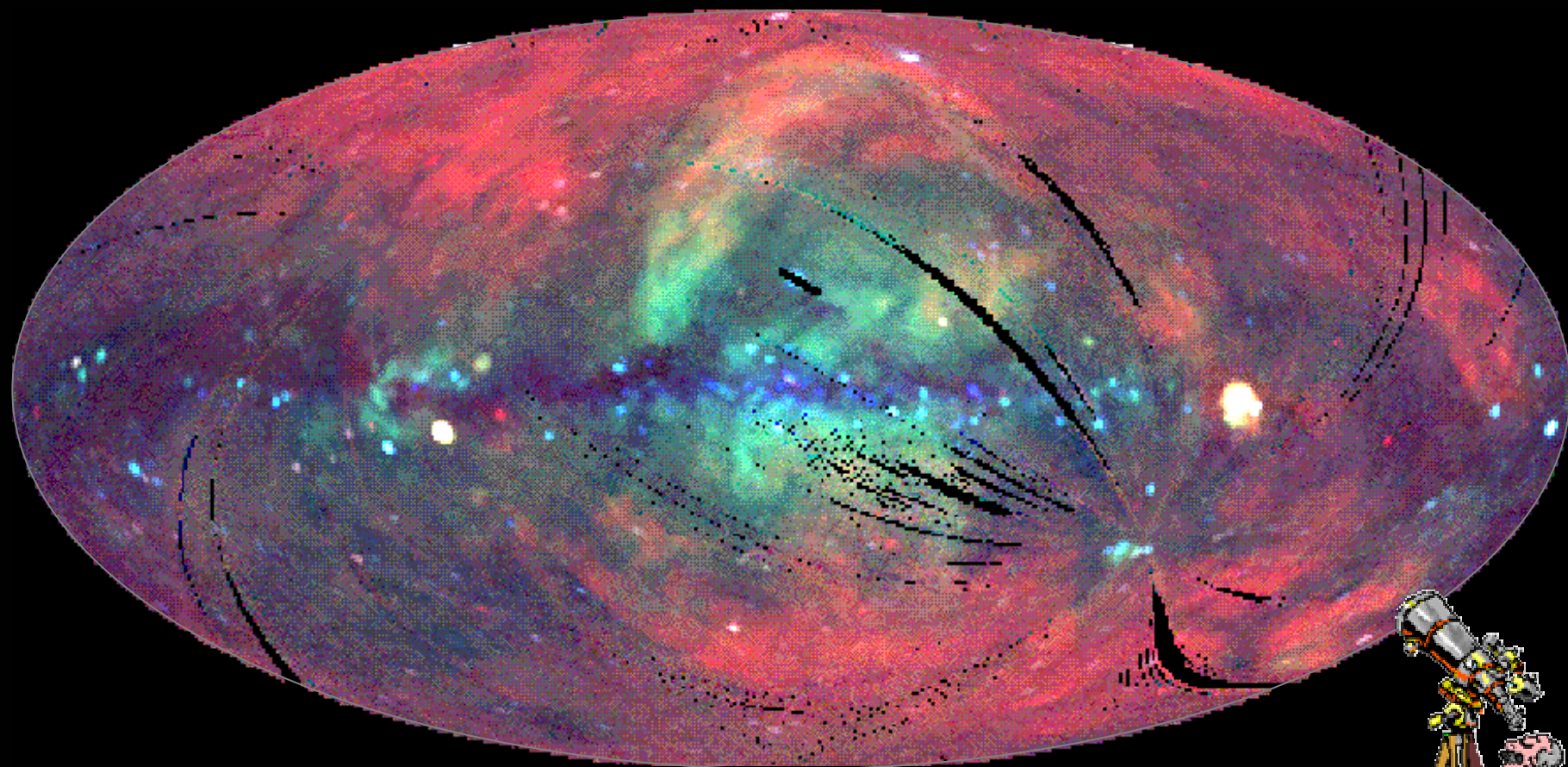
Radio

Infrared

Visible Light

X-rays  
( $10^3$  eV)

Gamma rays





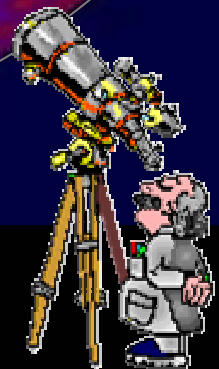
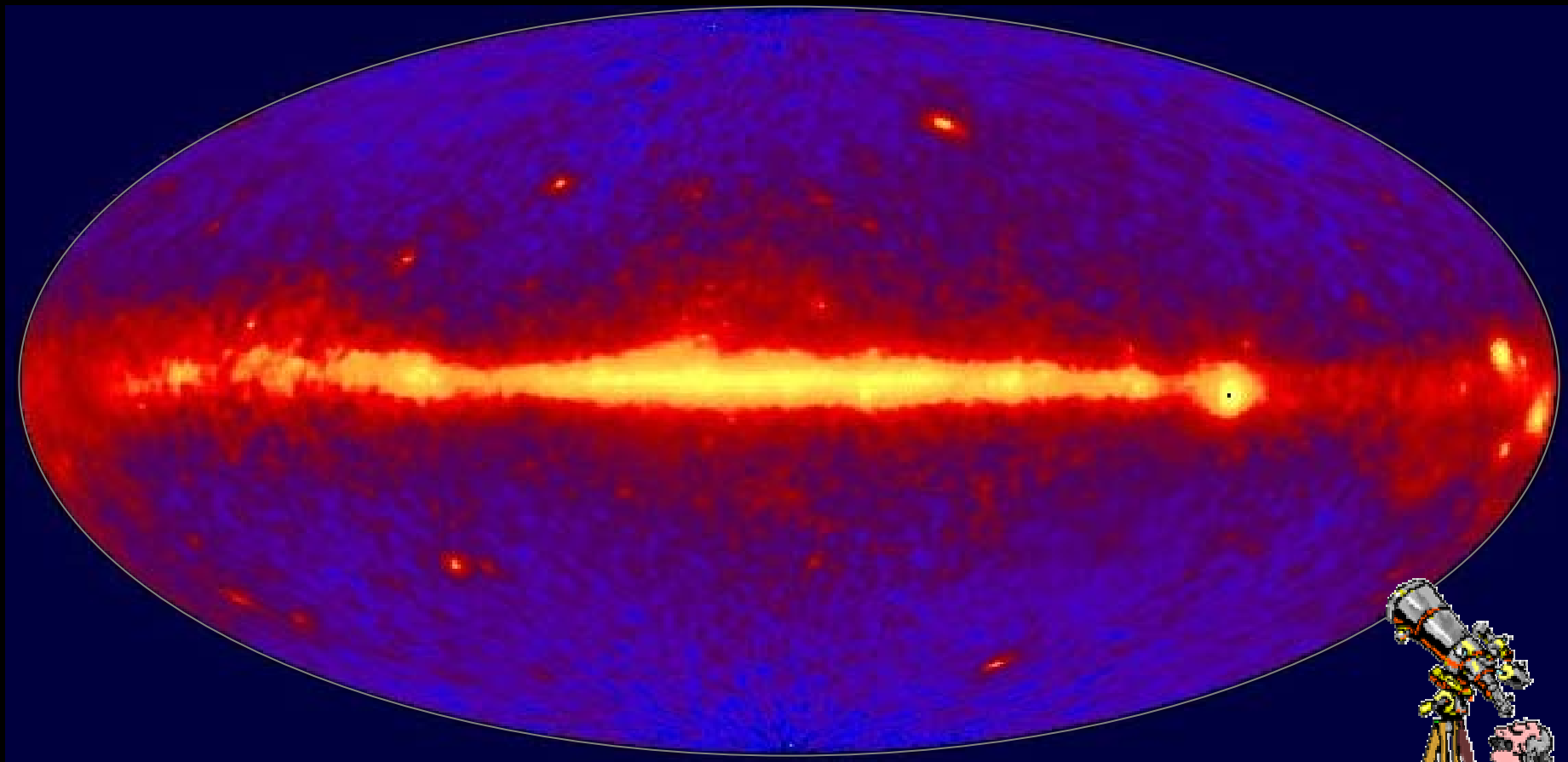
Radio

Infrared

Visible Light

X-rays

Gamma rays  
( $10^9$  eV)





Gamma  
(TeV =  $10^{12}$  eV)



TeV Gamma rays  
from the Crab Nebula  
Whipple Cherenkov Telescope  
1989

Gamma  
ray

Particle  
cascade

# Detection of TeV Gamma Rays

using Cherenkov  
telescopes

~ 10 km

Cherenkov light

1°

~ 120 m

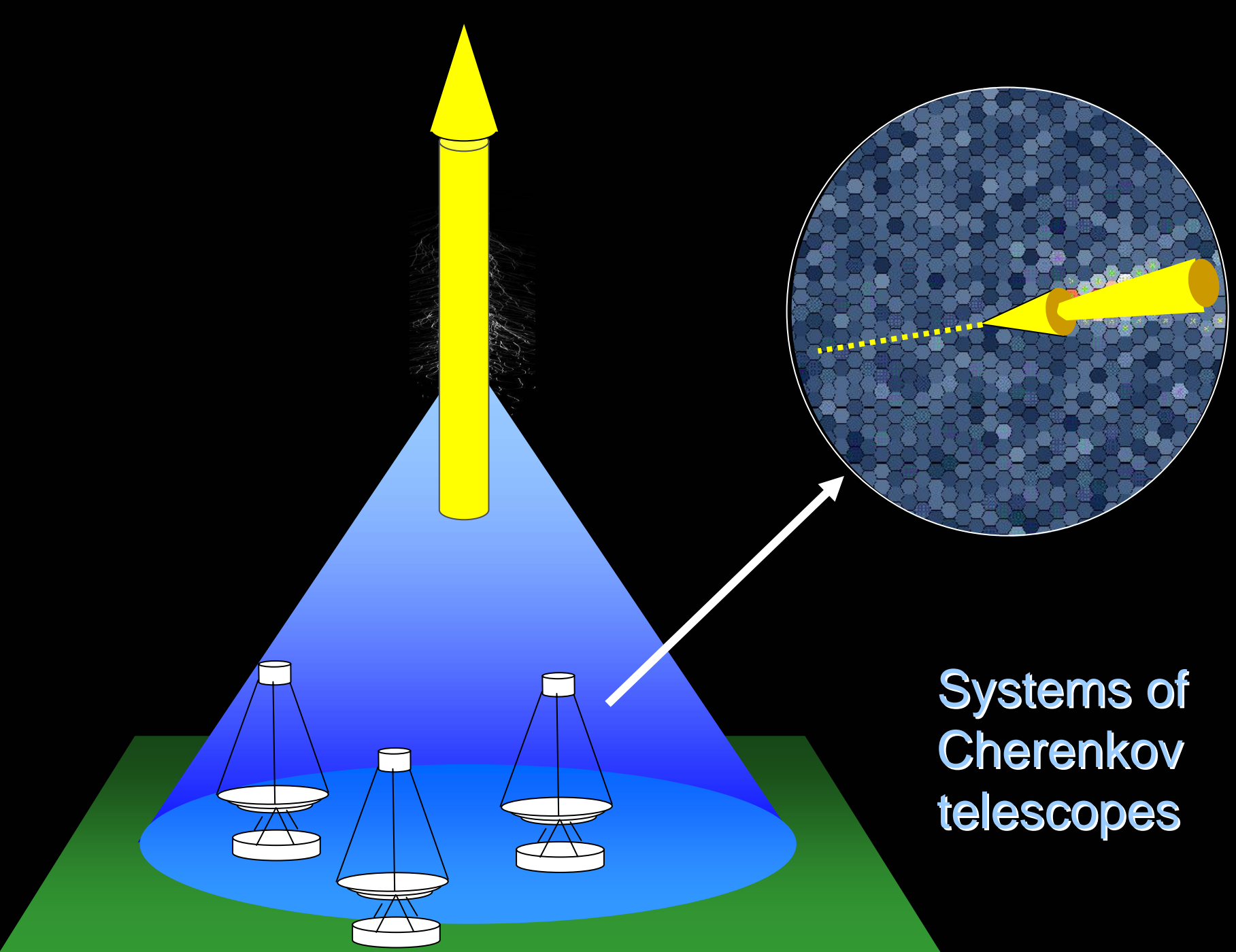




# Cherenkov light

is the equivalent of the sonic boom





Systems of  
Cherenkov  
telescopes



Air showers look  
a bit like meteors

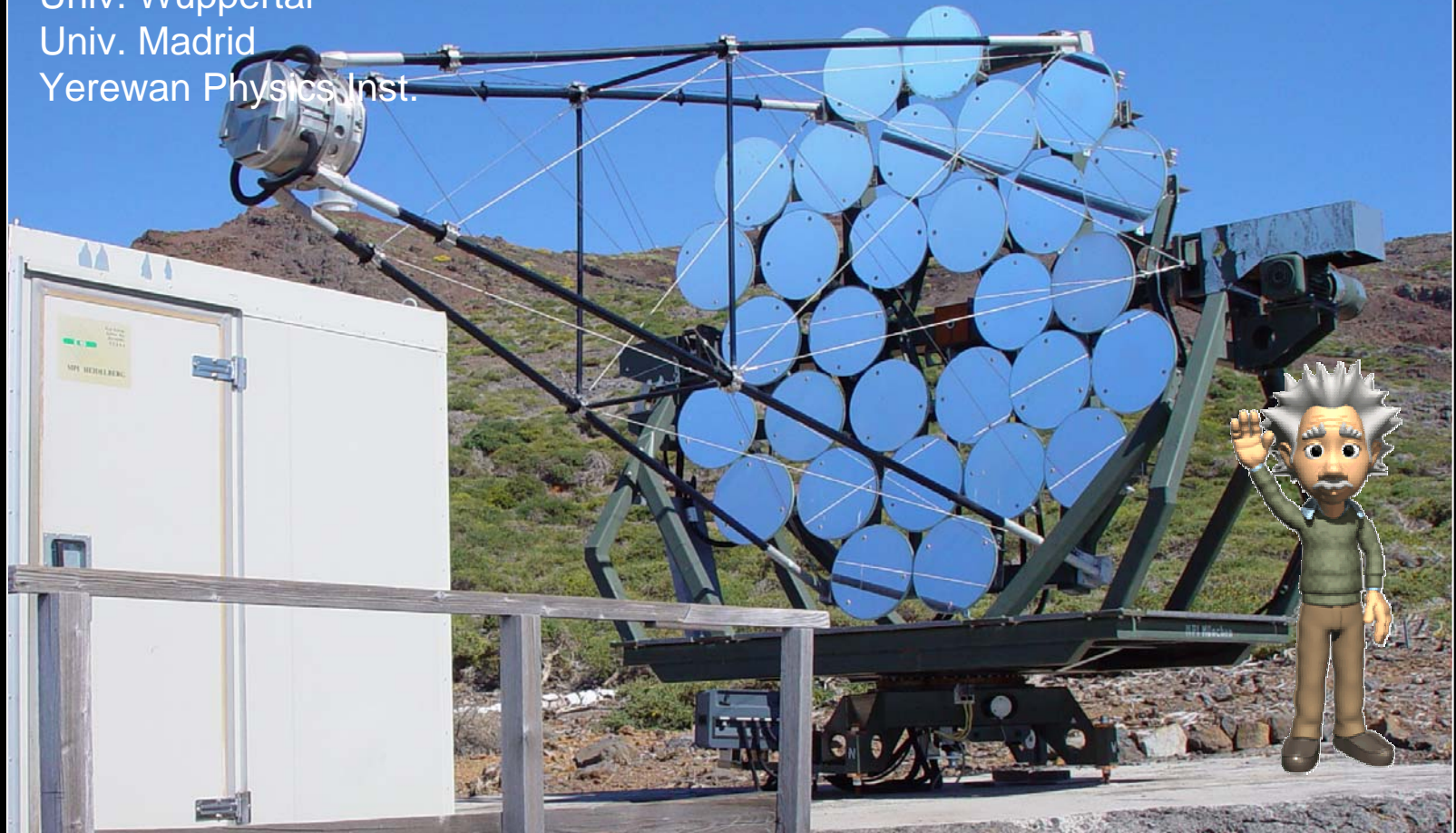
METEORS



(from Sky & Telescope)

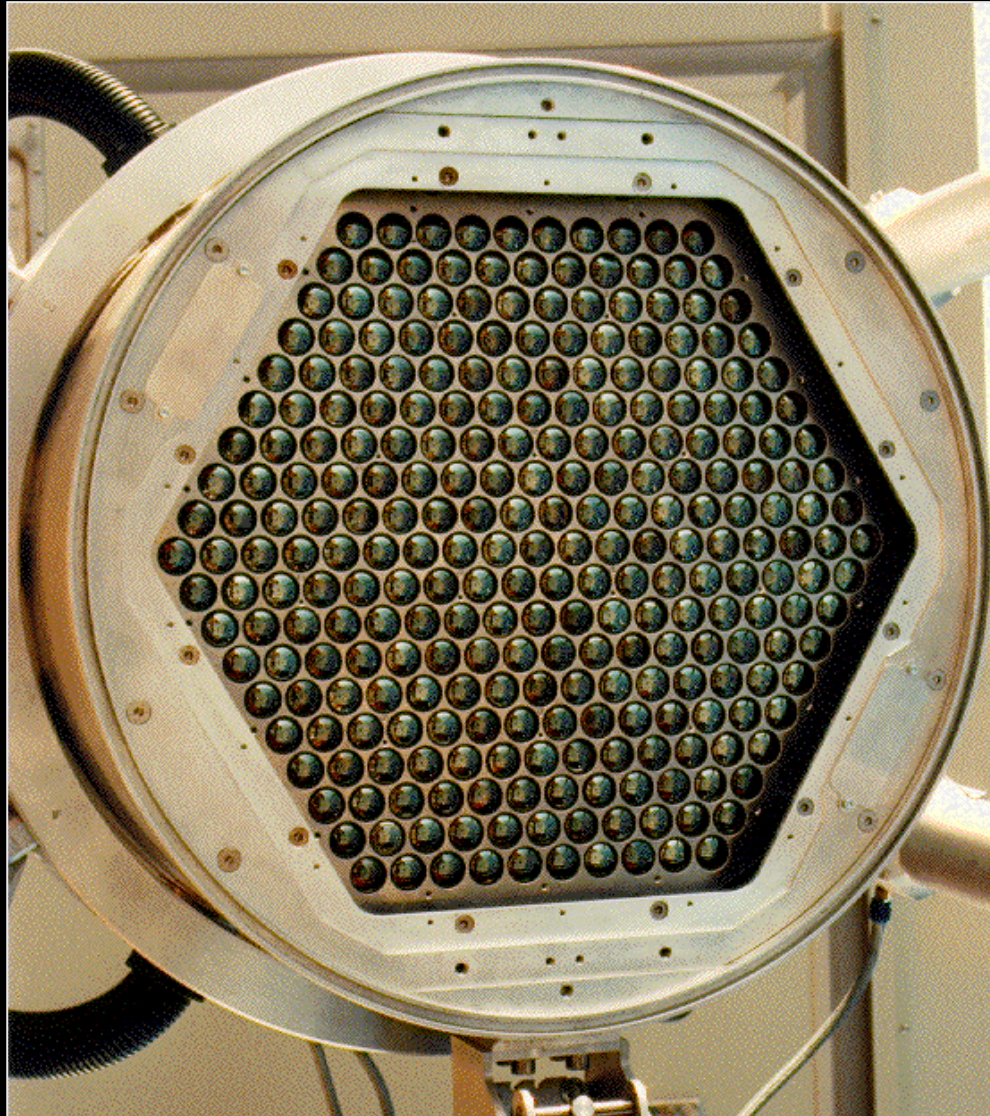
# The HEGRA Teleskopes on La Palma

MPI für Kernphysik  
MPI für Physik, München  
Univ. Hamburg  
Univ. Kiel  
Univ. Wuppertal  
Univ. Madrid  
Yerewan Physics Inst.

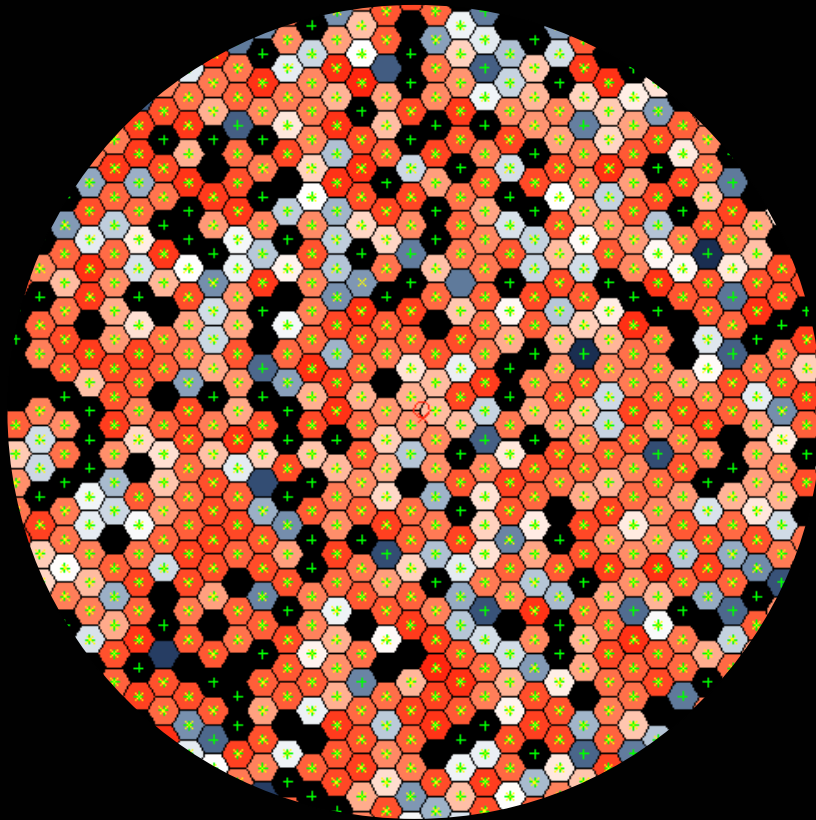




# The HEGRA Cameras



# Key feature of camera: speed !



1/10000  
(100  $\mu$ s)

1/100000  
(10  $\mu$ s)

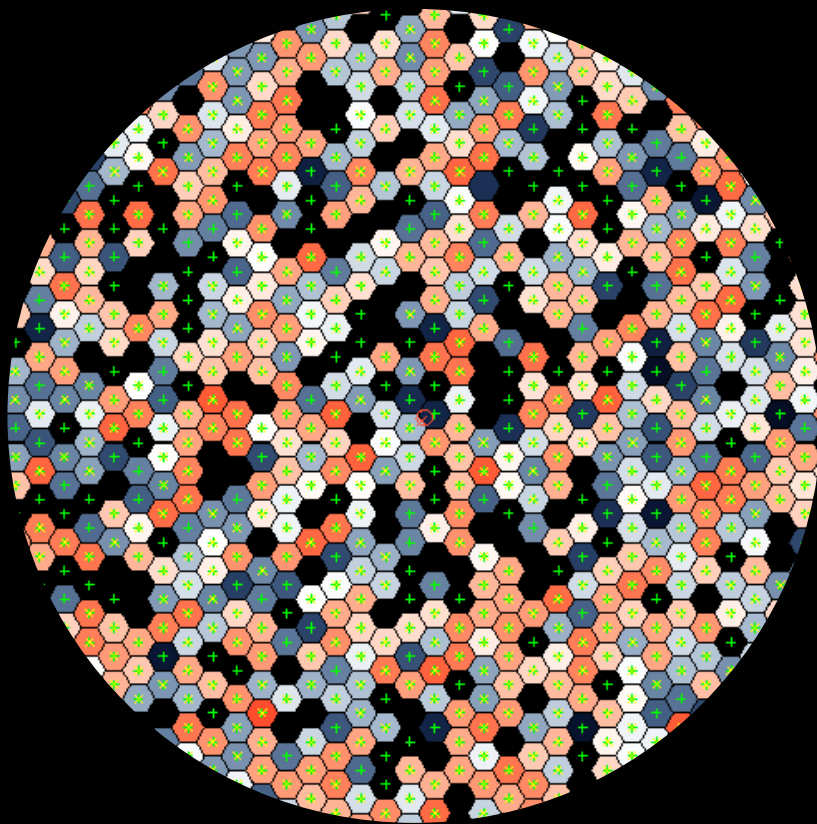
1/1000000  
(1  $\mu$ s)

1/10000000  
(100 ns)

1/100000000  
(10 ns)







$1/10000$   
(100  $\mu\text{s}$ )

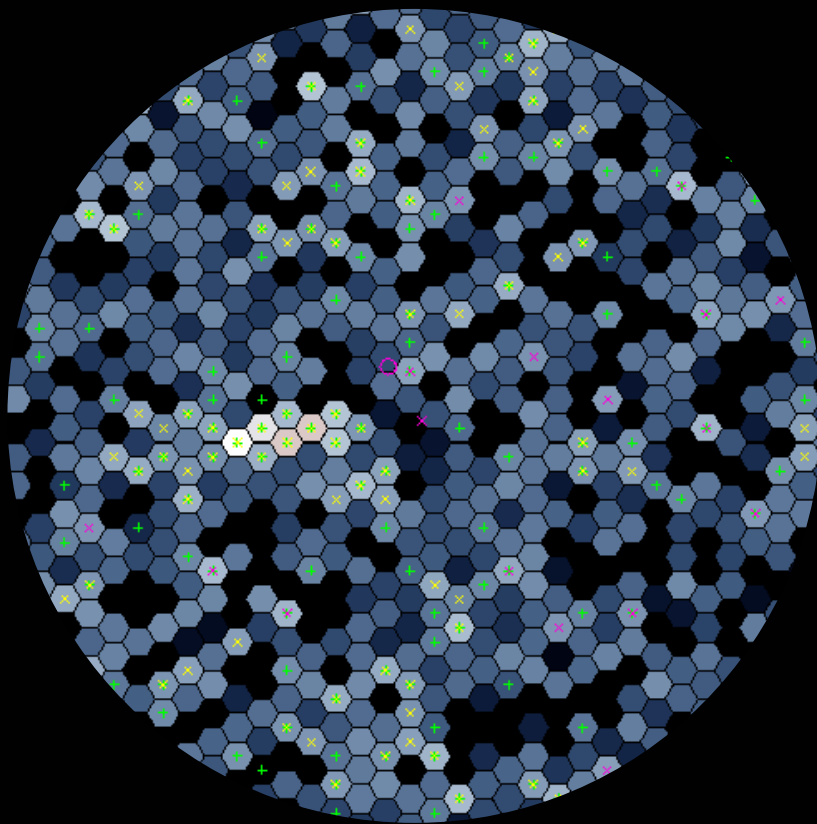
$1/100000$   
(10  $\mu\text{s}$ )

$1/1000000$   
(1  $\mu\text{s}$ )

$1/10000000$   
(100 ns)

$1/100000000$   
(10 ns)





**1/10000**  
**(100 μs)**

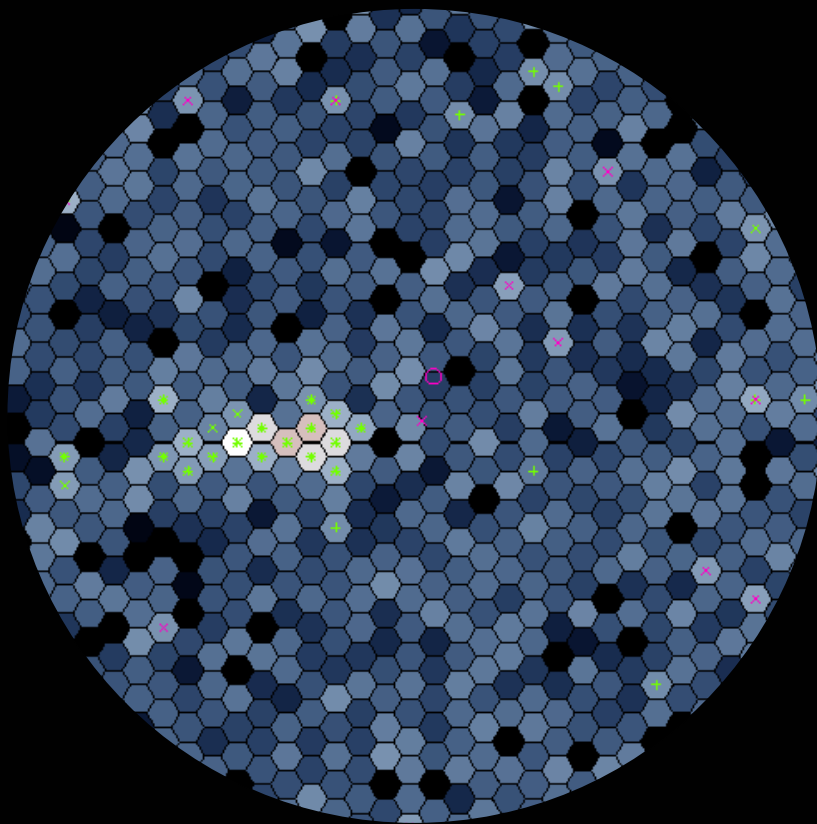
**1/100000**  
**(10 μs)**

**1/1000000**  
**(1 μs)**

**1/10000000**  
**(100 ns)**

**1/100000000**  
**(10 ns)**





$1/10000$   
(100  $\mu\text{s}$ )

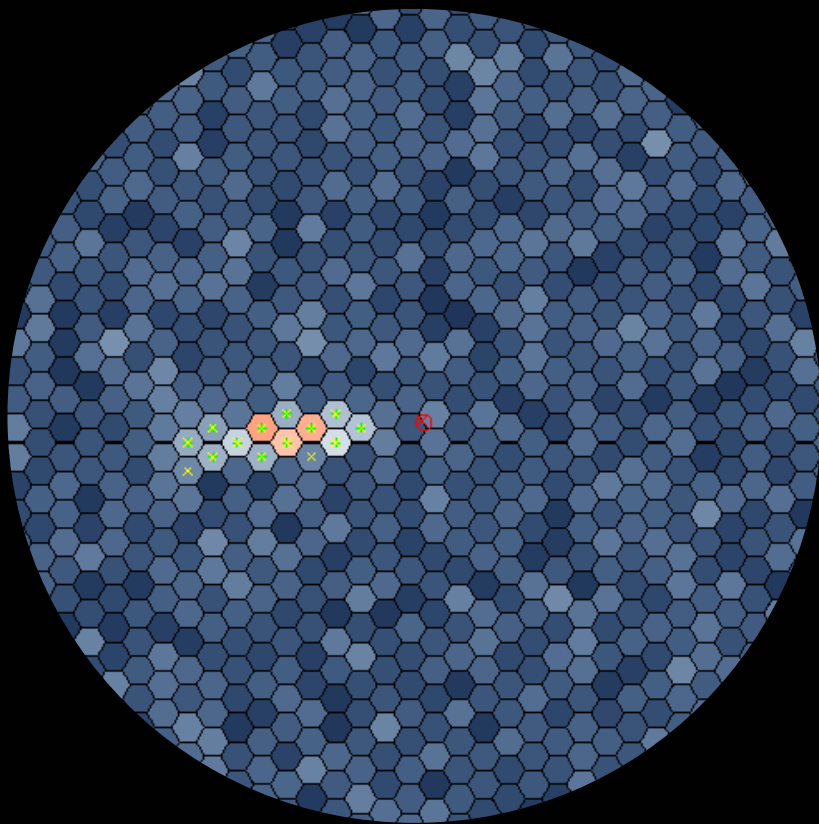
$1/100000$   
(10  $\mu\text{s}$ )

$1/1000000$   
(1  $\mu\text{s}$ )

$1/10000000$   
(100 ns)

$1/100000000$   
(10 ns)





**1/10000  
(100  $\mu$ s)**

**1/100000  
(10  $\mu$ s)**

**1/1000000  
(1  $\mu$ s)**

**1/10000000  
(100 ns)**

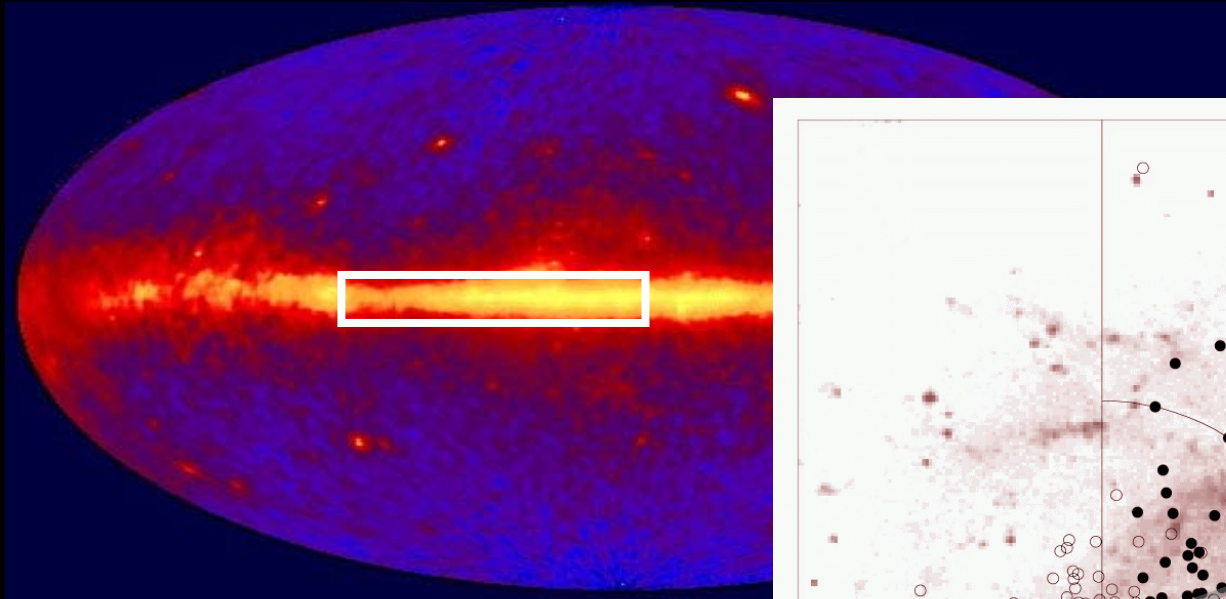
**1/100000000  
(10 ns)**



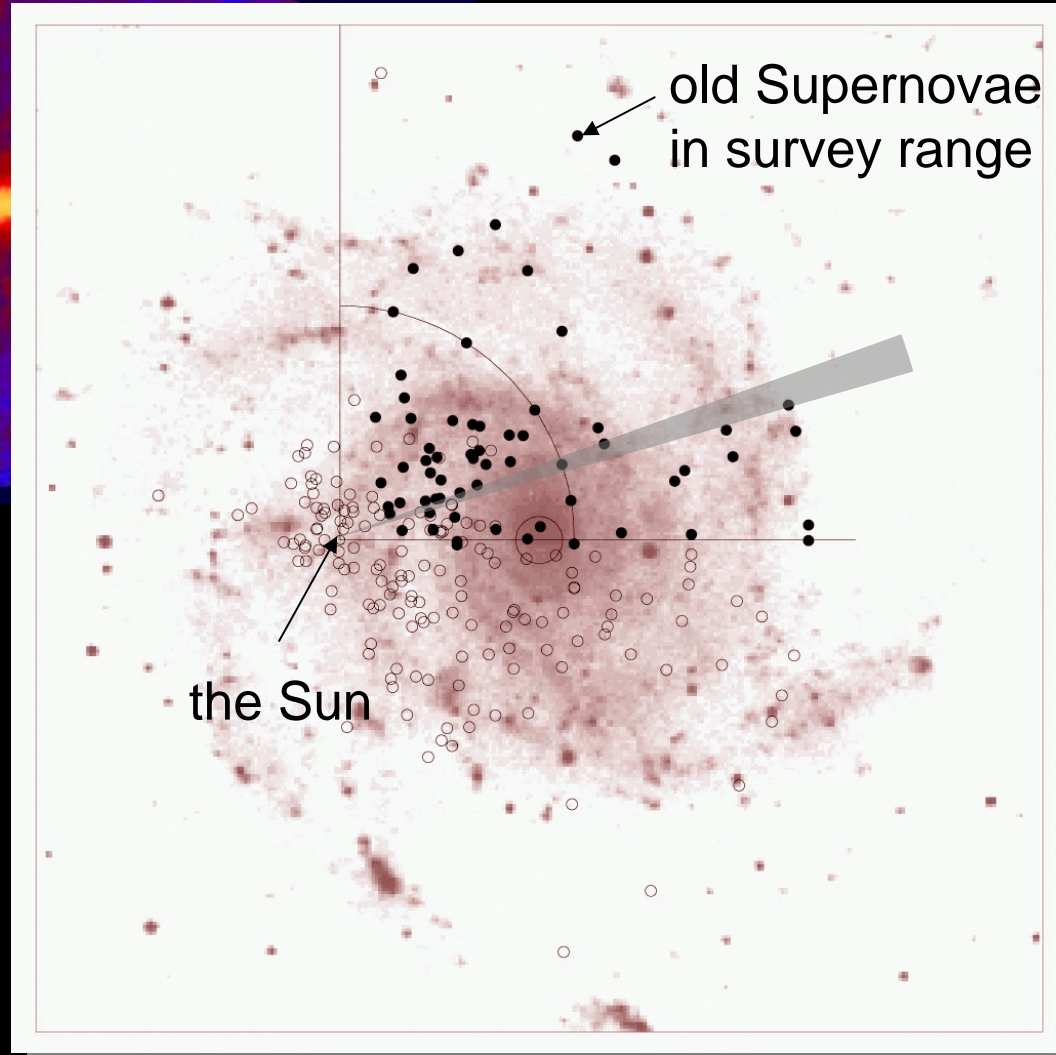


# Very high energy gamma rays from Supernovae ?

## The HEGRA Galactic plane survey



- 60 Supernovae investigated
- No gamma signal found
- Sensitivity should have been (just) sufficient



# The CANGAROO telescopes (Japan & Australia)

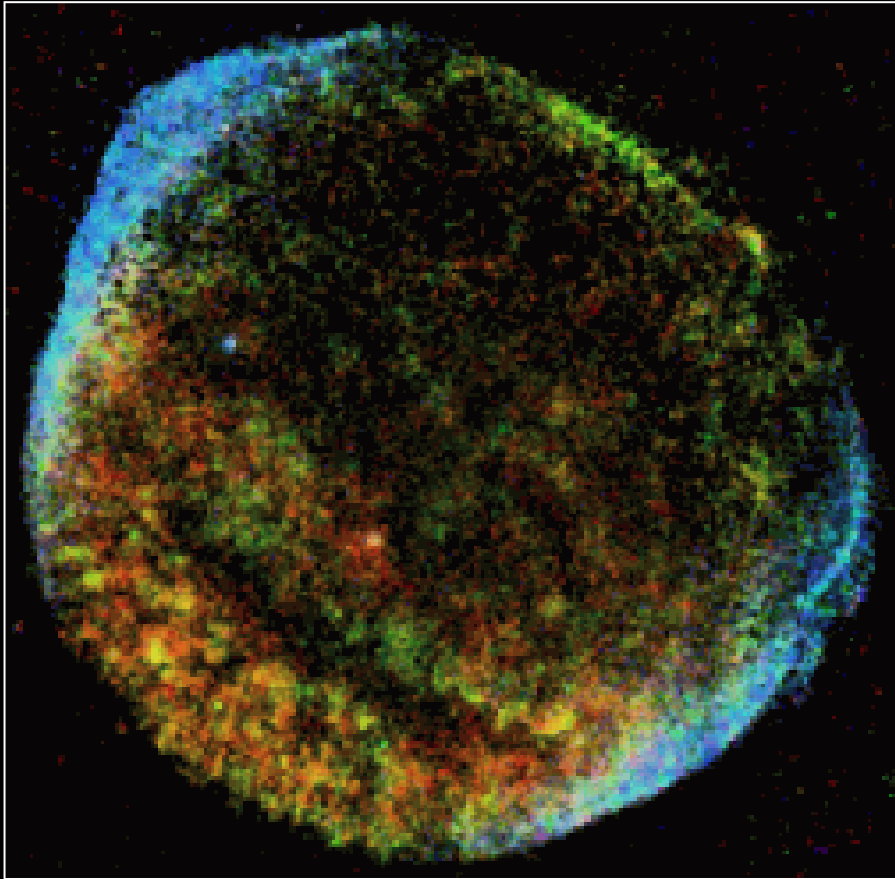


For a long time,  
the most powerful  
instrument watching  
southern skies

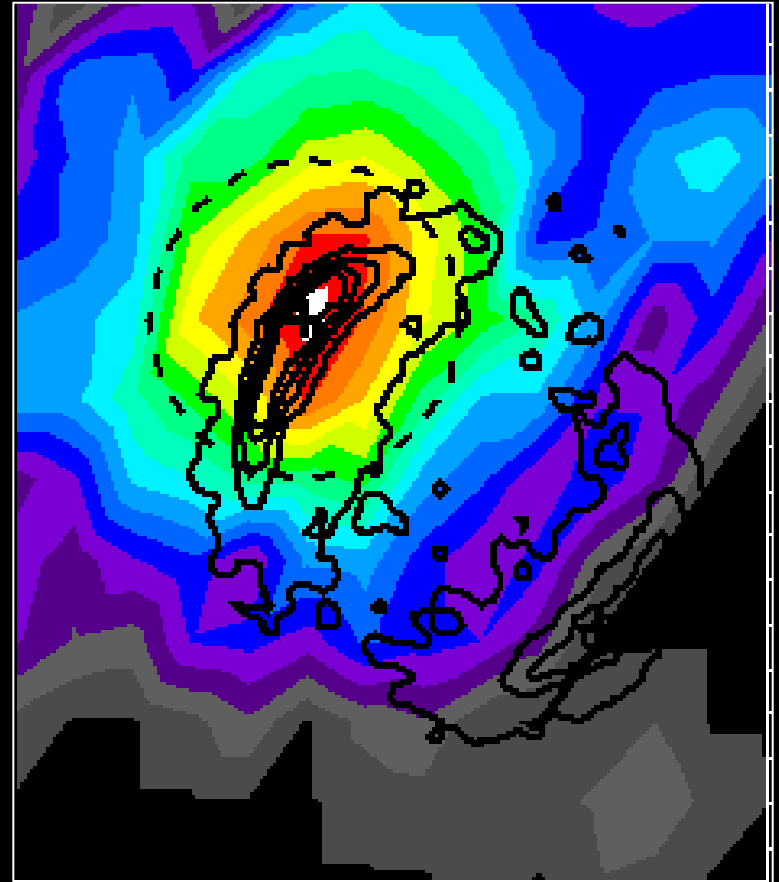




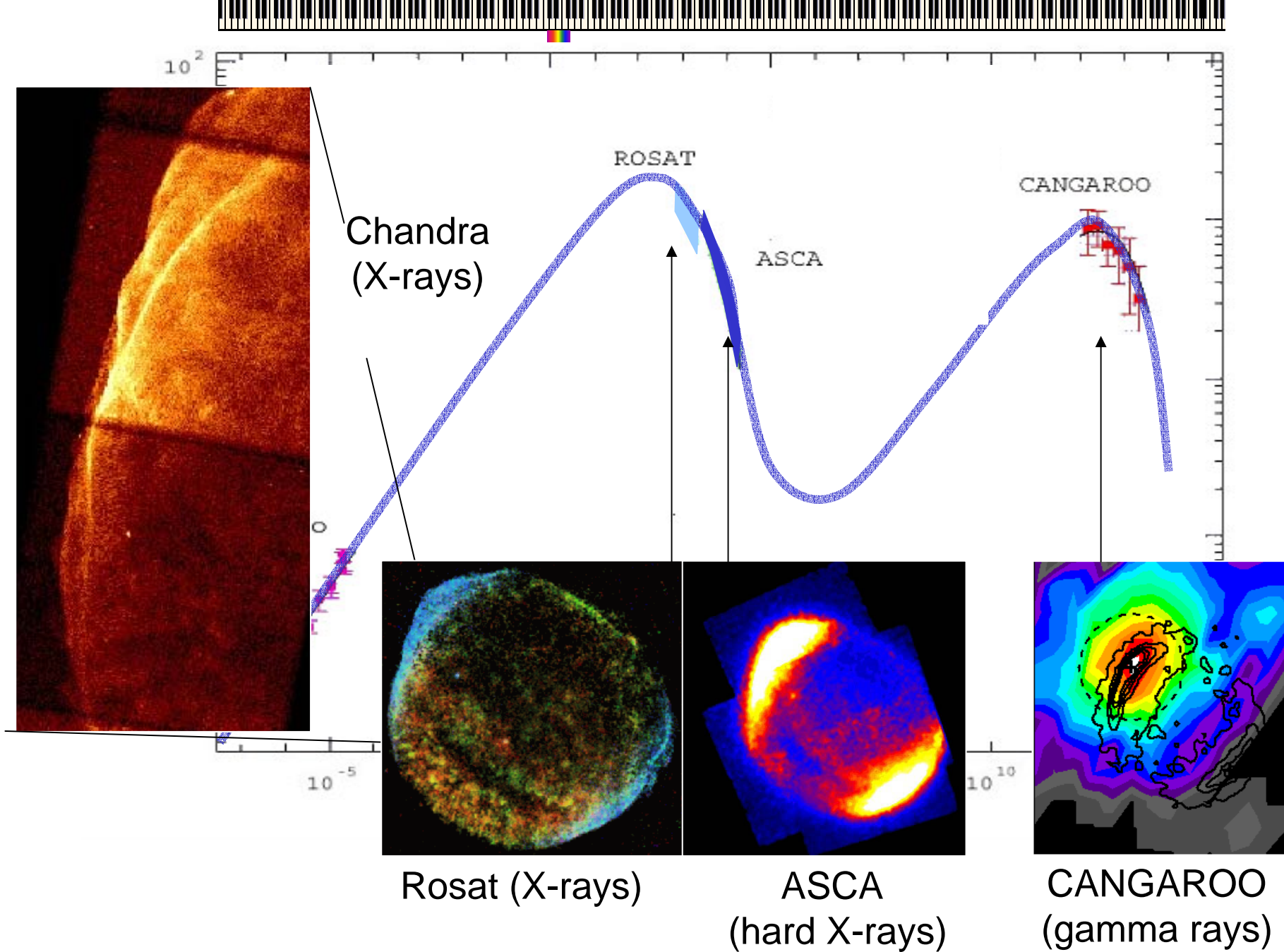
# Supernova 1006



X-ray image  
(ROSAT)



VHE gamma-ray image  
(CANGAROO 1997)







Energy-  
flux

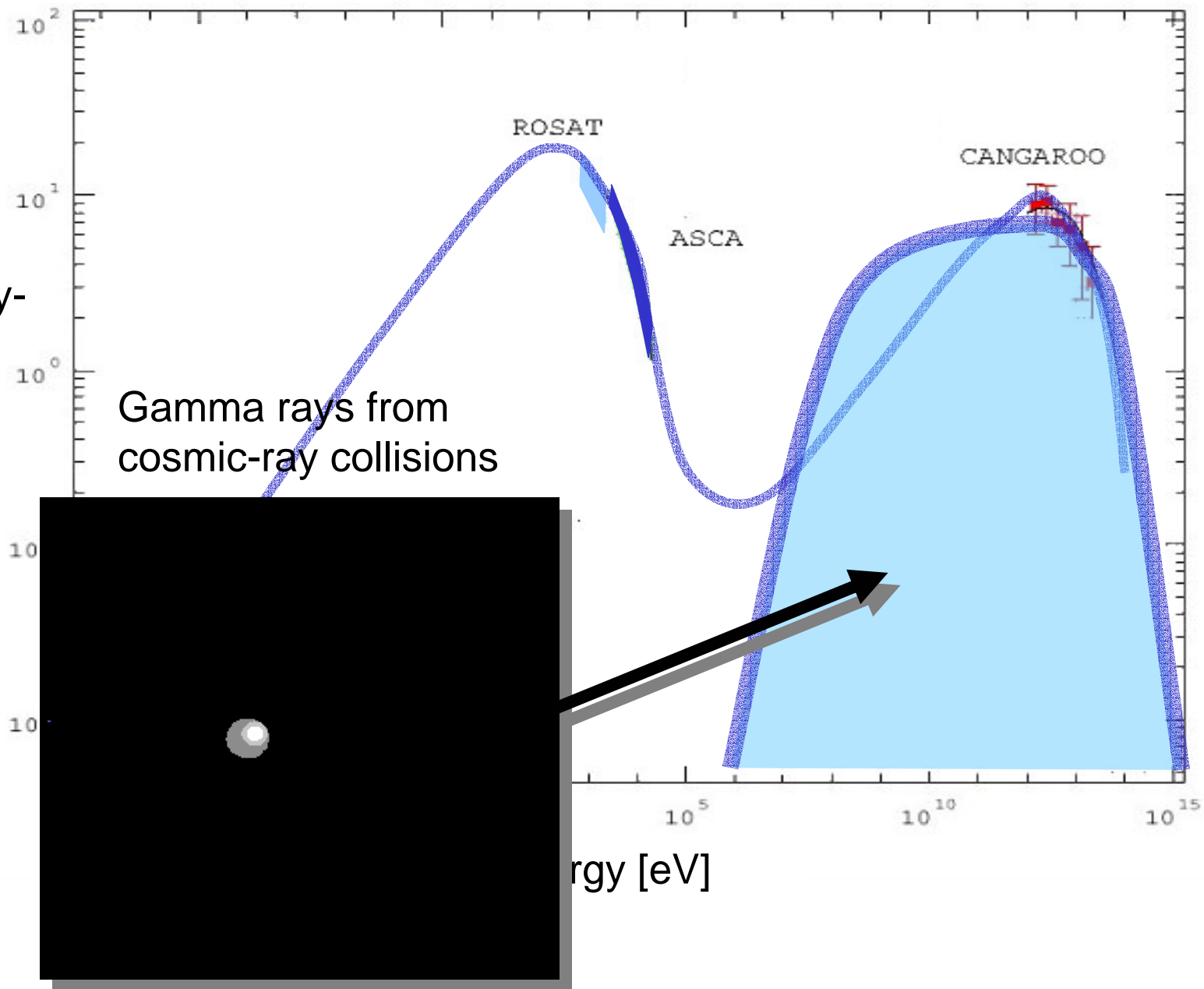
Gamma rays from  
cosmic-ray collisions

ROSAT

ASCA

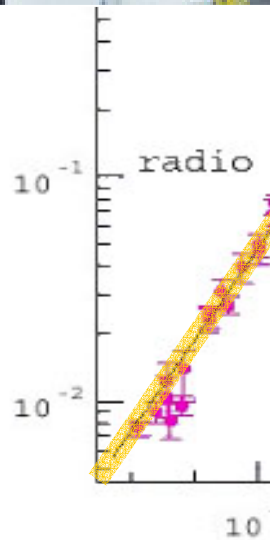
CANGAROO

Energy [eV]

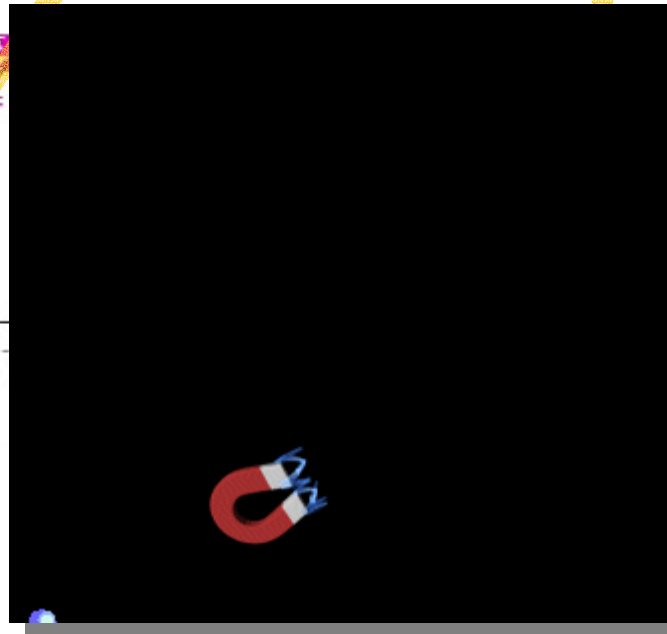




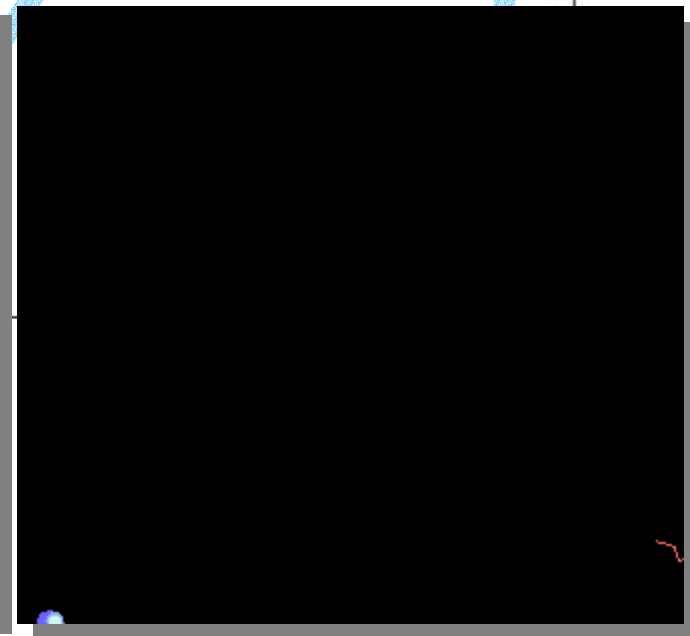
This supernova is clearly a cosmic accelerator!  
But it seems to accelerate electrons (rather than nuclei) !  
*Have we found the wrong accelerator ???*



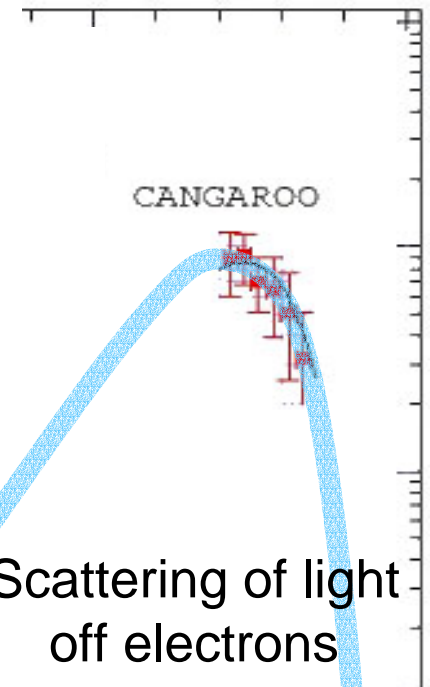
Synchrotron radiation  
by electrons



Scattering of light  
off electrons

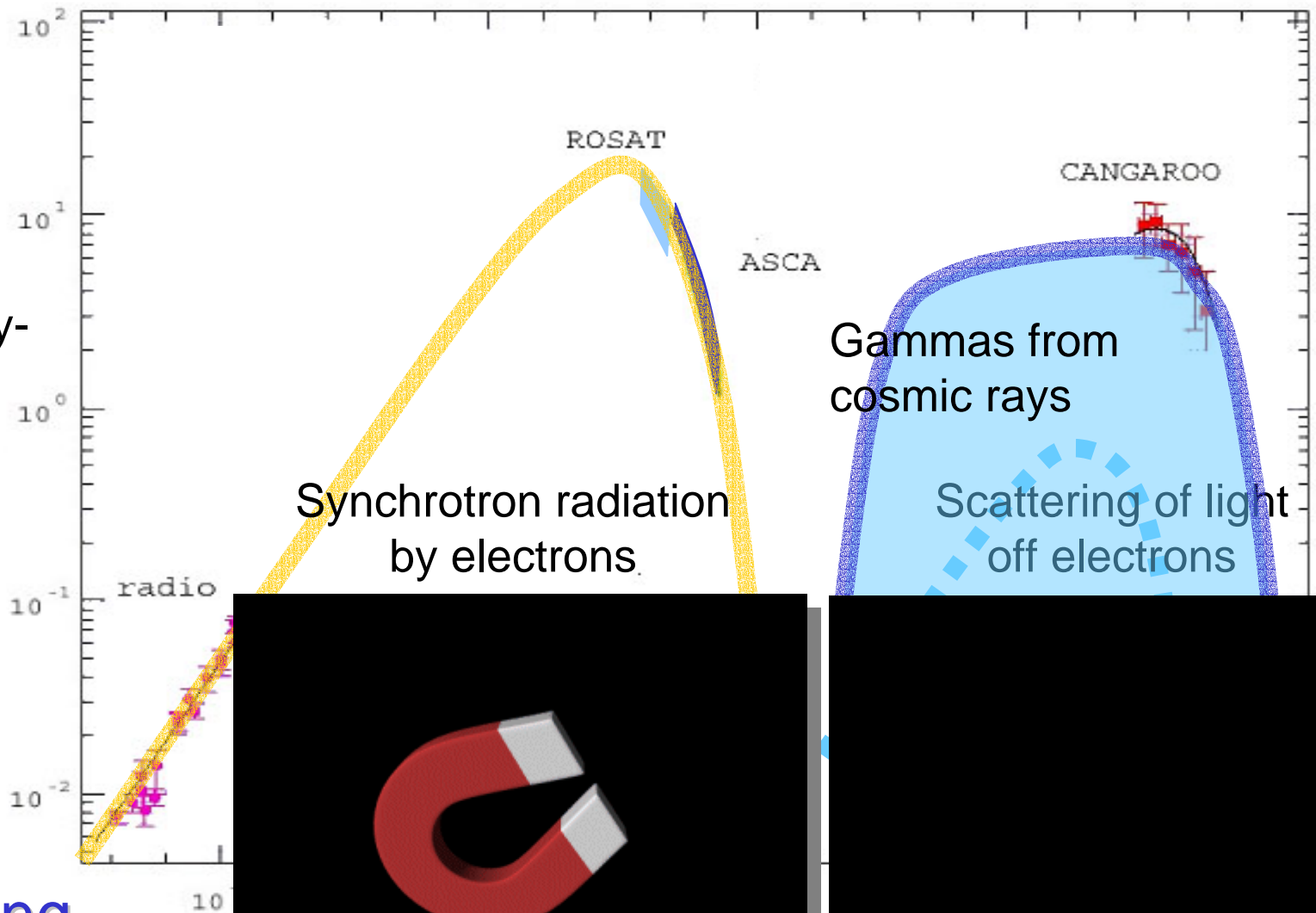


CANGAROO





Energy-  
flux



Strong  
magnetic  
field



# 1999: Supernova Cassiopeia A detected with HEGRA (232 h exposure)



Again, a double-humped spectrum ...  
Equal amounts of energy in X-rays and high energy gamma rays ...  
but ... this object should have a very high magnetic field!

Type II SN  
(Animation: G. Pühlhofer)



# The acceleration of cosmic-ray protons in the supernova remnant RX J1713.7–3946

R. Enomoto<sup>\*</sup>, T. Tanimori<sup>†</sup>, T. Naito<sup>‡</sup>, T. Yoshida<sup>§</sup>, S. Yanagita<sup>§</sup>, M. Mori<sup>\*</sup>, P. G. Edwards<sup>||</sup>, A. Asahara<sup>†</sup>, G. V. Bicknell<sup>¶</sup>, S. Gunji<sup>#</sup>, S. Hara<sup>☆</sup>, T. Hara<sup>‡</sup>, S. Hayashi<sup>\*\*</sup>, C. Itoh<sup>§</sup>, S. Kabuki<sup>\*</sup>, F. Kajino<sup>\*\*</sup>, H. Katagiri<sup>\*</sup>, J. Kataoka<sup>†</sup>, A. Kawachi<sup>\*</sup>, T. Kifune<sup>††</sup>, H. Kubo<sup>†</sup>, J. Kushida<sup>☆</sup>, S. Maeda<sup>\*\*</sup>, A. Maeshiro<sup>\*\*</sup>, Y. Matsubara<sup>‡‡</sup>, Y. Mizumoto<sup>§§</sup>, M. Moriya<sup>☆</sup>, H. Muraishi<sup>|||</sup>, Y. Muraki<sup>‡‡</sup>, T. Nakase<sup>¶¶</sup>, K. Nishijima<sup>¶¶</sup>, M. Ohishi<sup>\*</sup>, K. Okumura<sup>\*</sup>, J. R. Patterson<sup>##</sup>, K. Sakurazawa<sup>☆</sup>, R. Suzuki<sup>\*</sup>, D. L. Swaby<sup>##</sup>, K. Takano<sup>☆</sup>, T. Takano<sup>#</sup>, F. Tokanai<sup>#</sup>, K. Tsuchiya<sup>\*</sup>, H. Tsunoo<sup>\*</sup>, K. Uruma<sup>¶¶</sup>, A. Watanabe<sup>#</sup> & T. Yoshikoshi<sup>☆☆</sup>

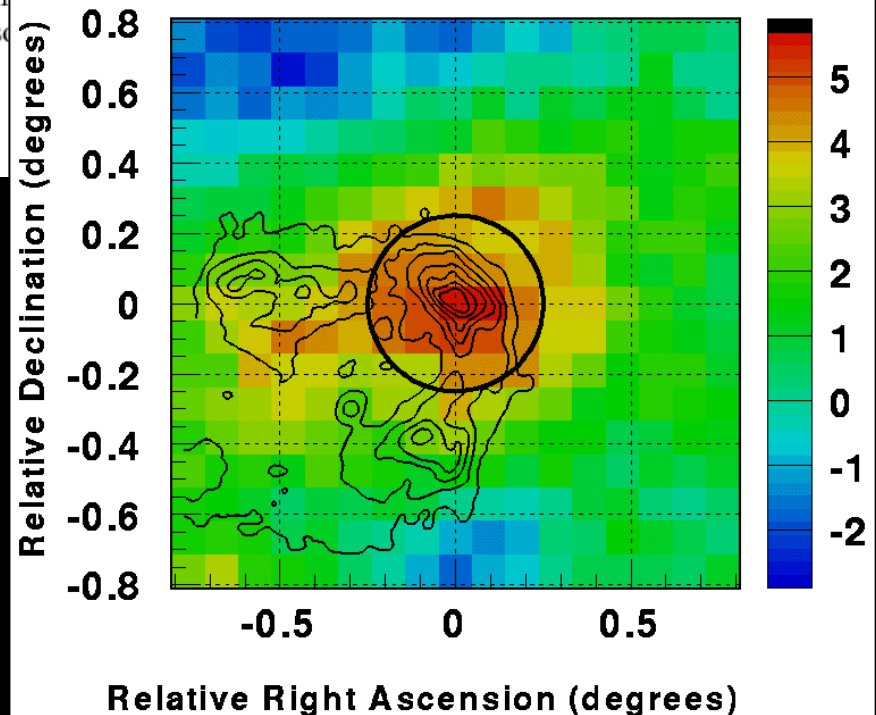
<sup>\*</sup> Institute for Cosmic Ray Research, University of Tokyo, Kashiwa, Chiba 277-8582, Japan

RX J1713.7-3946  
CANGAROO 2000/02

molecular cloud complex<sup>10</sup>.

The CANGAROO air Cerenkov telescope, which is intended to detect very high energy  $\gamma$ -rays, is located near Woomera, South Australia. The 3.8-m telescope<sup>9</sup>, which operated from 1994 to 1998, was replaced in 2000 by a 10-m reflector with a 552-pixel camera of  $0.115^\circ$  square photomultiplier pixels. Observations of RX J1713.7–3946 were carried out 23–26 July and 19–27 August 2000, and 20 May–26 June 2001 with the 10-m telescope. After selecting data taken at high elevation angles ( $>60^\circ$ ) in good weather conditions, a total of 2,332 min on-source and 1,789 min off-source data remained for further analysis.

The differential fluxes of  $\gamma$ -rays from RX J1713.7–3946 are plotted in Fig. 1. The number of excess events was determined from the plots of image orientation angle (see ref. 11) for on- and off-



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the origin of Cosmic Rays

The suspects:

Black holes & exploding stars

The evidence:

Blue flashes from

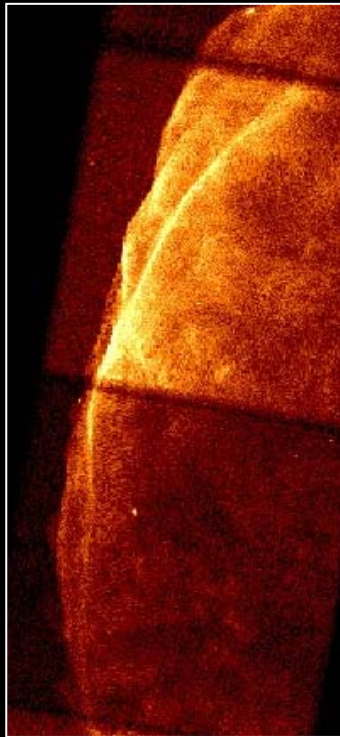
Astrophysical  
telescope



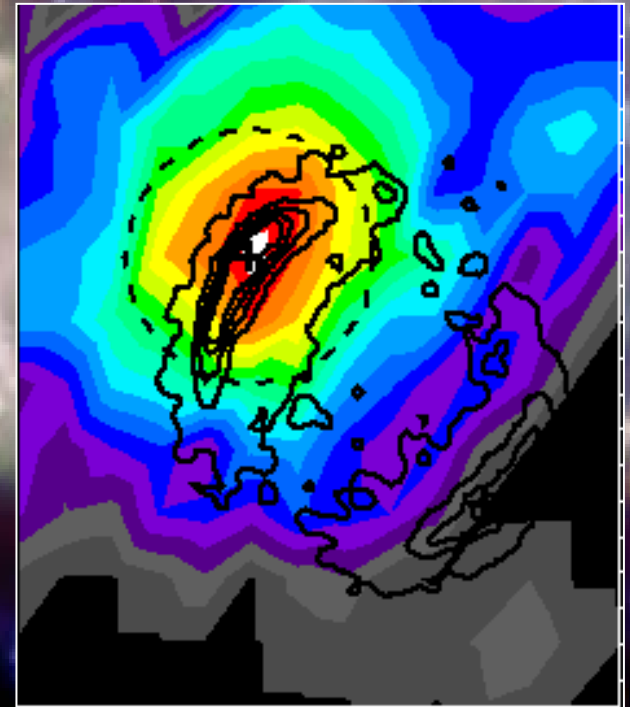
**The judgement**



No question,  
Supernovae are cosmic accelerators !  
But are they THE sources of cosmic rays ?



- Compare in detail the distribution of X-rays and gamma rays
- Understand the magnetic fields
- Find more objects



**Built better instruments**



# CANGAROO III



# The H.E.S.S. Project

## High Energy Stereoscopic System









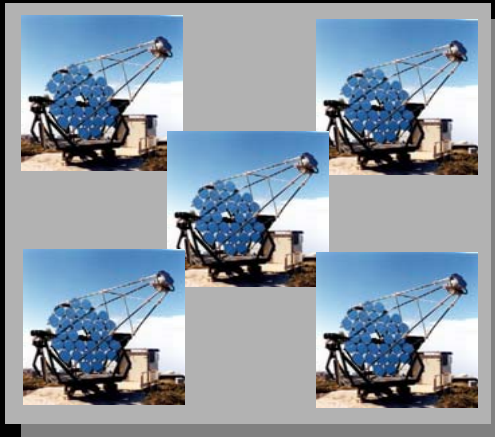




# Progress



First detection of Crab Nebula  
Whipple 1989: 50 h Observation time



HEGRA 1997:  
15 min

The new instruments:  
30 sec



$$\begin{aligned}
 D_{\mu\nu} = & \frac{\Omega^2}{B_0^2} (1 - \mu^2) \Re \sum_{n=-\infty}^{\infty} \int_{-\infty}^{\infty} d^3 k \Big|_0 \\
 & \times \left\{ \frac{c^2}{v^2} (1 - \mu^2) J_n^2(W) R_{\parallel\parallel}(k, \xi) + \frac{1}{2} J_{n+1}^2(W) \right. \\
 & \times \left( P_{\parallel\parallel}(k, \xi) + \mu^2 \frac{c^2}{v^2} R_{\parallel\parallel}(k, \xi) + i\mu \frac{c}{v} [Q_{\parallel\parallel}(k, \xi) - T_{\parallel\parallel}(k, \xi)] \right) \\
 & + \frac{1}{2} J_{n-1}^2(W) \left( P_{\perp\perp}(k, \xi) + \mu^2 \frac{c^2}{v^2} R_{\perp\perp}(k, \xi) \right. \\
 & + i\mu \frac{c}{v} [T_{\perp\perp}(k, \xi) - Q_{\perp\perp}(k, \xi)] \Big) - \frac{1}{2} J_n^2(W) J_{n+1}(W) [e^{2i\psi} \\
 & \times (P_{\perp\perp}(k, \xi) - \mu^2 \frac{c^2}{v^2} R_{\perp\perp}(k, \xi) + e^{-2i\psi} (P_{\perp\perp}(k, \xi) - \mu^2 \frac{c^2}{v^2} R_{\perp\perp}(k, \xi) \\
 & - i\mu \frac{c}{v} [T_{\perp\perp}(k, \xi) + Q_{\perp\perp}(k, \xi)]) \Big] + \frac{ic\sqrt{1 - \mu^2}}{\sqrt{2}v} J_n(W) \\
 & \times [J_{n+1}(W) (e^{i\psi} T_{\parallel\parallel}(k, \xi) - e^{-i\psi} Q_{\parallel\parallel}(k, \xi) \\
 & - (e^{i\psi} T_{\perp\perp}(k, \xi) - e^{-i\psi} Q_{\perp\perp}(k, \xi))] \Big)
 \end{aligned}$$

**E=MC<sup>2</sup>**









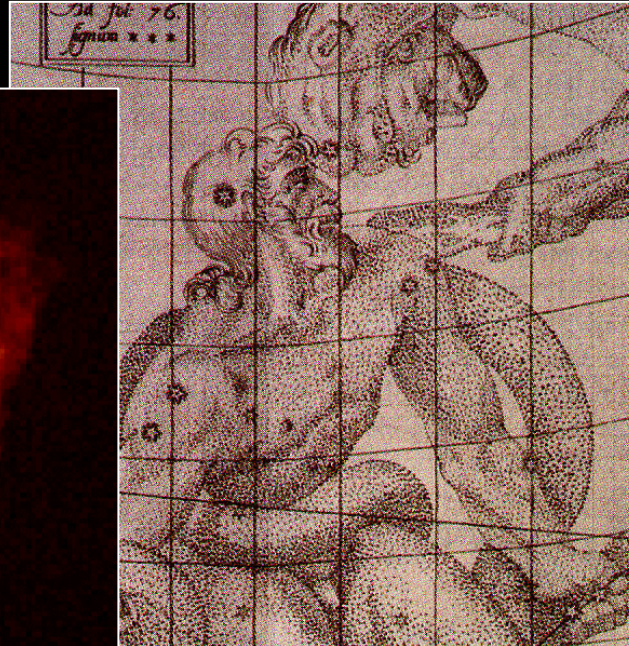
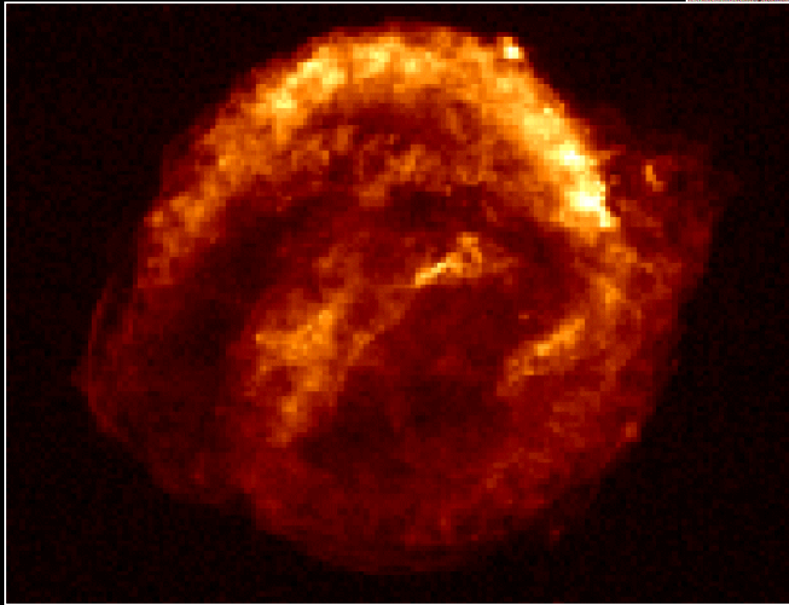


# International Cosmic Ray Conference



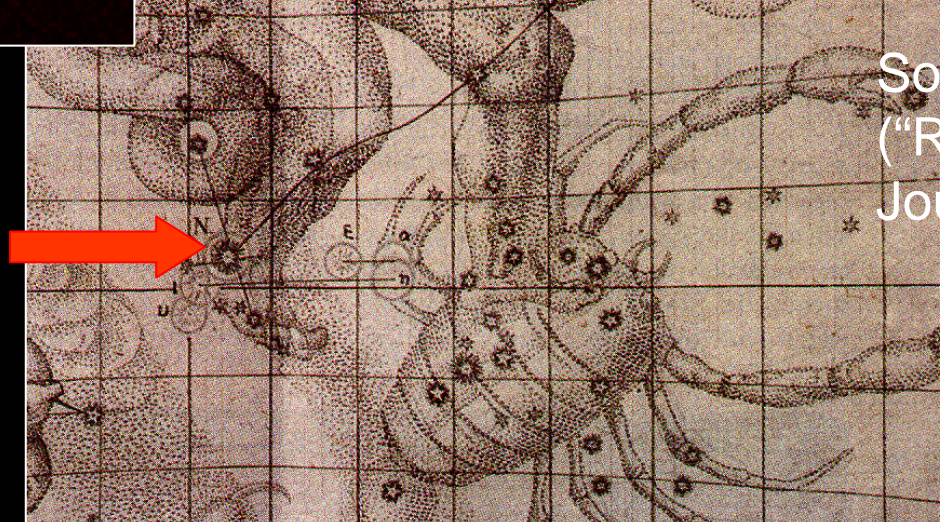


# 1604: Kepler's Supernova



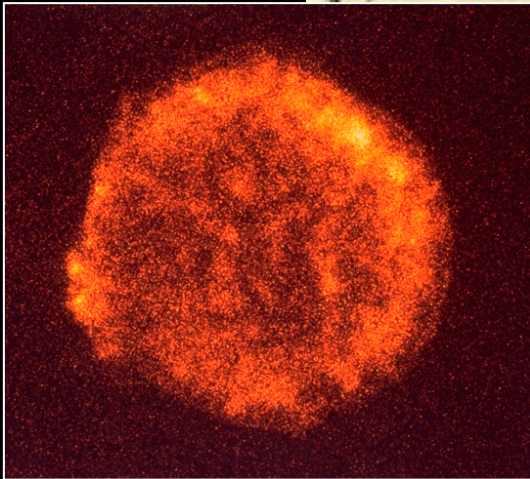
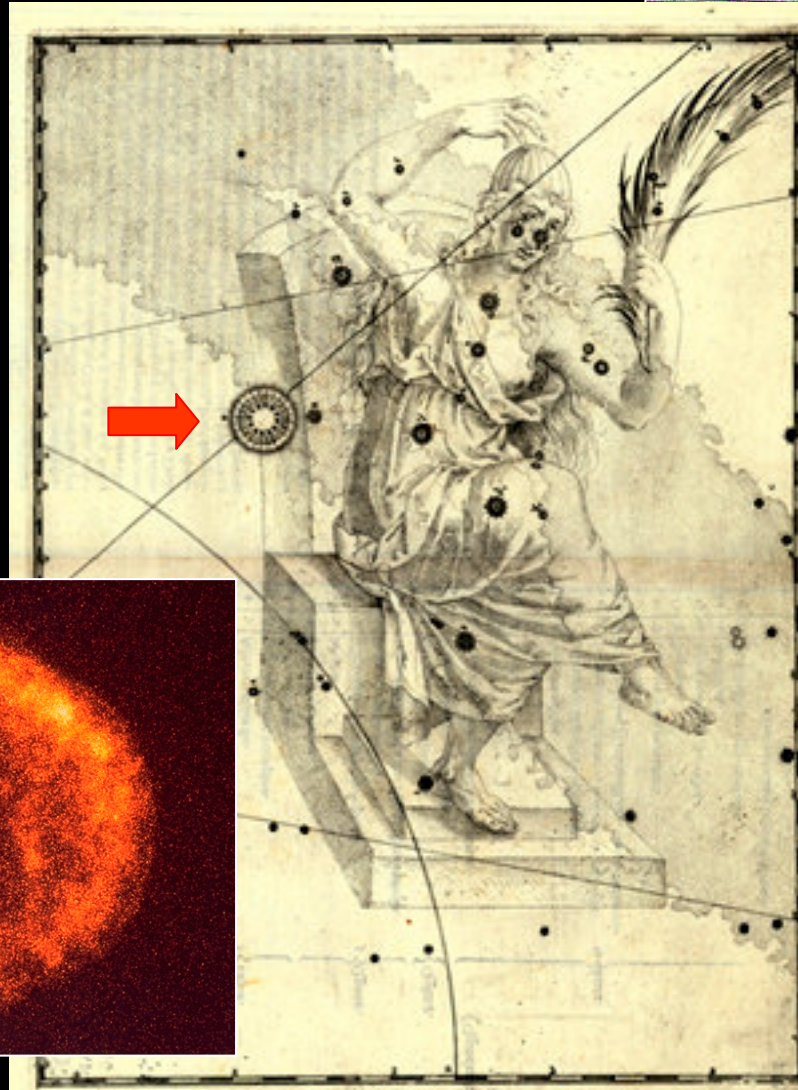
○未時。上柳別殿受針。○壬申。○卯時辰時沉霧夜一更客星見於  
大江星上。在尾宿十一度去極一百九度大如歲星色黃赤動搖。○朝。  
王世子問。安。○憲府。啟曰。兩司劍蕪春秋並令仕進于。貴顯  
莖諫體面與庶官自別以郎廳供仕之際必有虧損拘碍之弊臺諫盡  
帶之負請勿進參高陽郡守權悅濫率成婚子弟多有貽弊之事請命  
罷職麟山僉使朴命壽至率京妻二人侵虐軍卒日以貿易皮物為事  
貪虐汎濫之狀不一而足請。命罷職。○大浦僉使李雲以本道鄉軍  
性且恃妄不合巨鎮邊將請。命通差。○各曰依啓。○癸酉。○巳時午  
時日暈夜一更客星見於大江星上。在尾宿十一度去極一百九度大  
如歲星色黃赤動搖五更月暈。○朝。王世子問。安。○午時。上柳  
別殿受針。○甲戌。○辰時太白見於巳地夜一更客星見於大江星上  
在尾宿十一度去極一百九度大於歲星色黃赤動搖。○上不豫。○朝  
王世子問。安。○午時。上柳別殿受針。○乙亥。○朝。王世子問。安  
○天朝遊擊董正誼入來。上命宰臣申欽迎慰于門外又遣汪書奎  
楊問安于所館處遊擊接見後引出第二門外送之云。○命原任大

Sonjo Sillok  
("Royal Korean  
Journal")



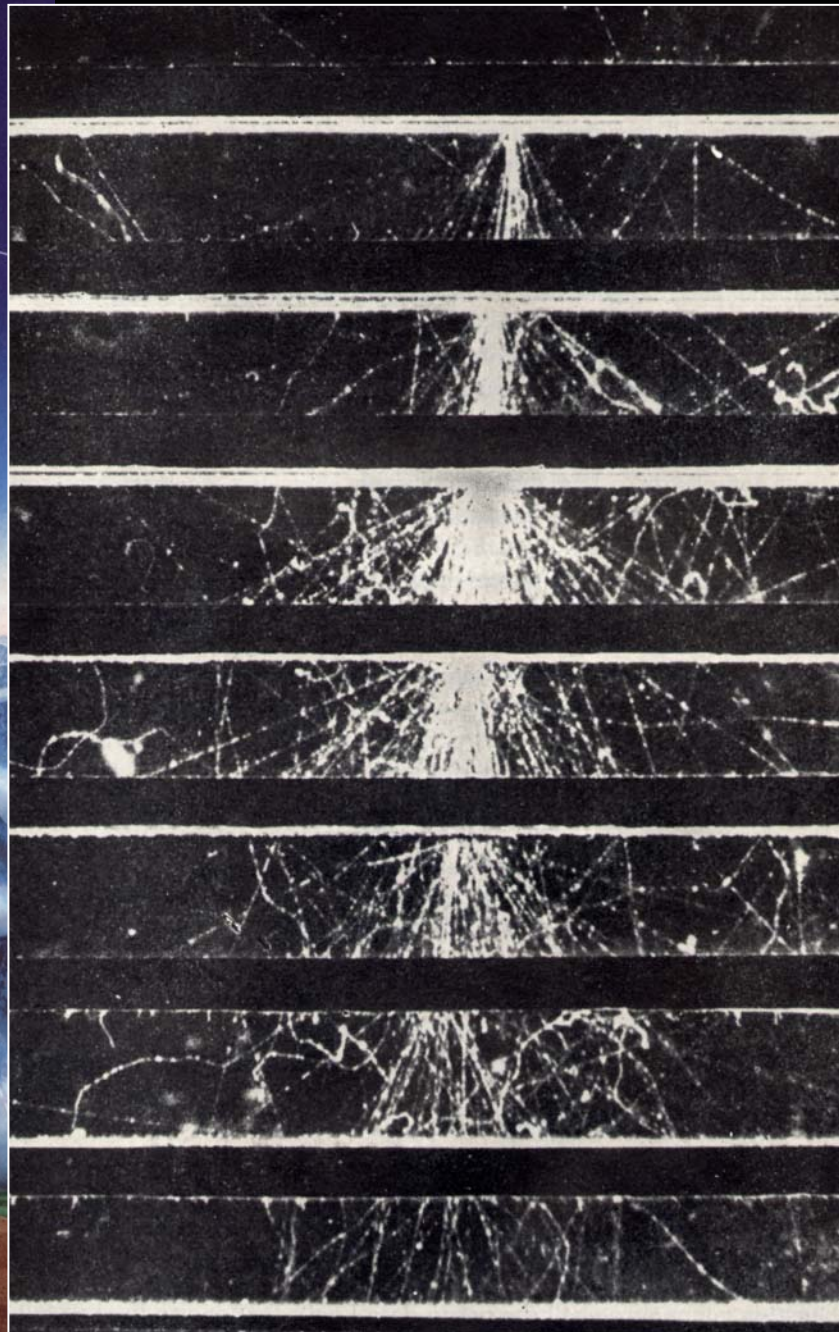


# 1572: Tycho's supernova



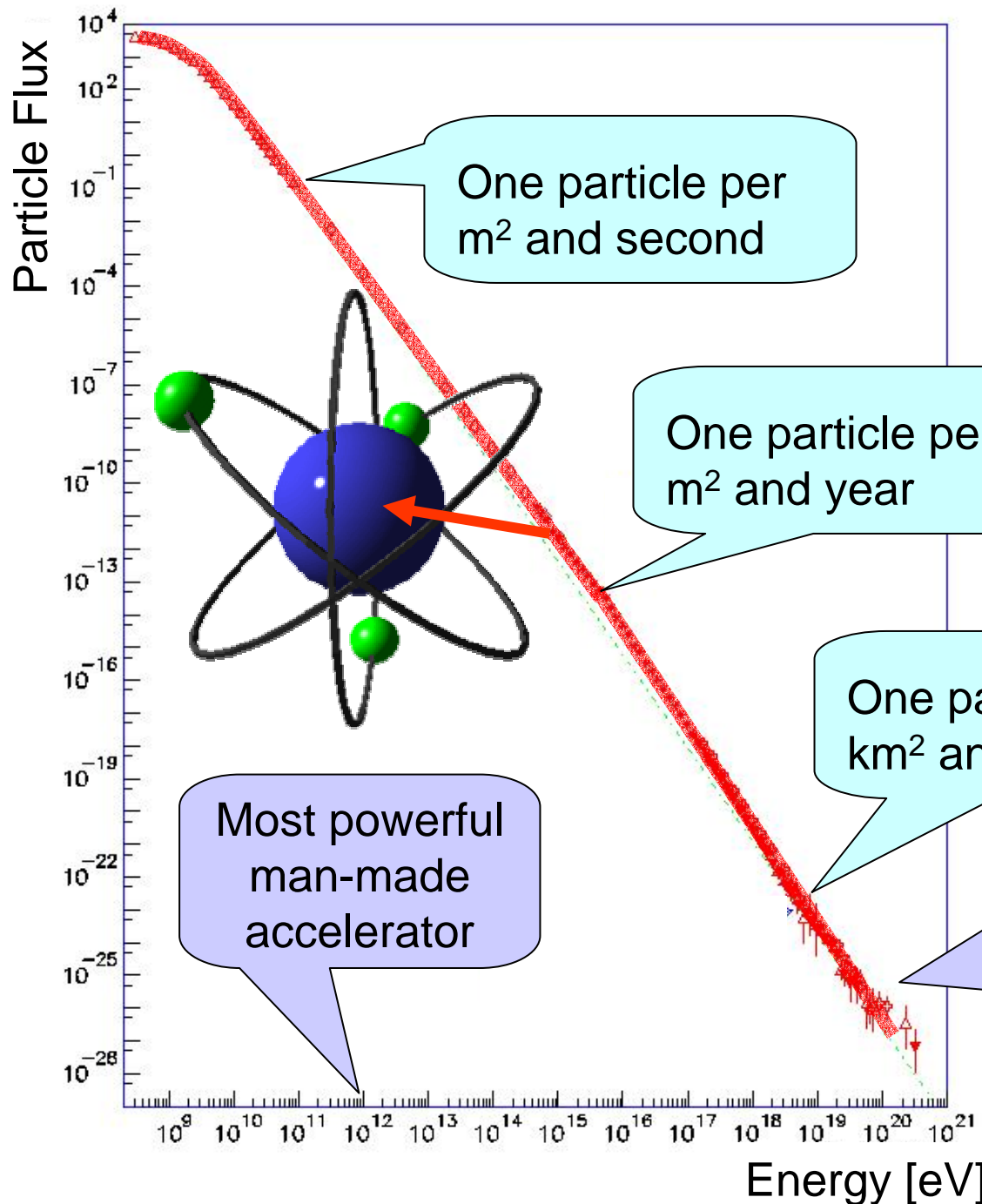


# Air showers





# The flux of Cosmic Rays



■ Energy scales

■ Flux scales

■ Composition:  
mostly atomic  
nuclei

Kinetic energy of 1 gram  
of these particles  
provides world energy  
supply for 1000 years !