

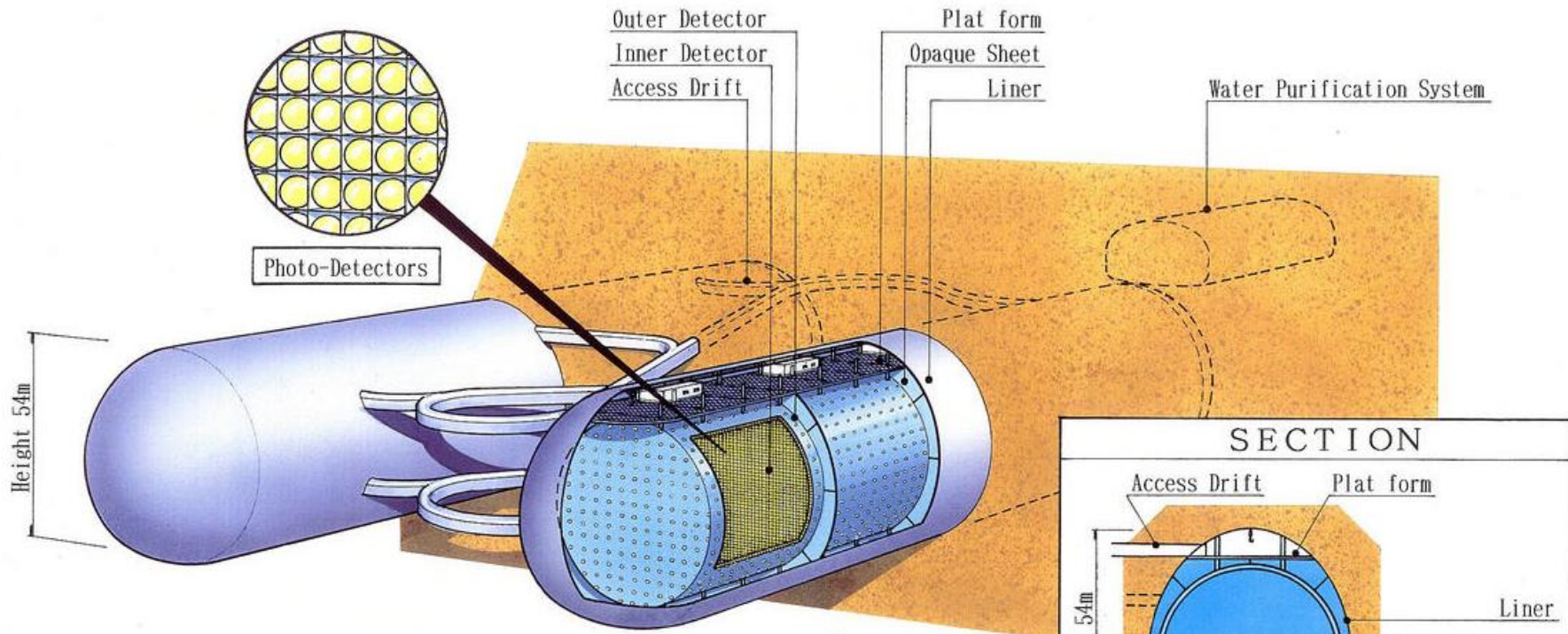
Hyper-Kamiokande

K. Okumura (ICRR Univ. of Tokyo)
NNN07 @ Hamamatsu, Oct 4. 2007

Outline

- ▶ Introduction
- ▶ Physics of HK
- ▶ Prospect for HK electronics
- ▶ Summary

Introduction



- total ~1 Mton (2 x 0.5 Mton) water Cherenkov detector at Kamioka site
- fiducial mass 0.54 Mton in total
- Number of 20-inch PMTs : ~100,000 for 1 PMT / m² (20% photo coverage)
~200,000 for 2 PMT / m² (40% photo coverage)
- Overburden : 600~700 m of rock (2.7 g/cm³ density, 1600~1900 m.w.e.)

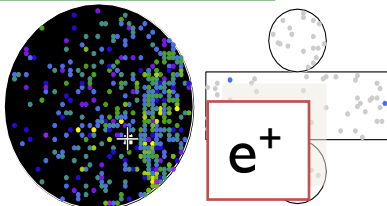
Physics of HK

- ▶ Nucleon decay
 - ▶ $p \rightarrow e^+ \pi^0$, $p \rightarrow \nu K^+$, other decay mode
- ▶ Long baseline oscillation experiments
 - ▶ θ_{13} , CP violation, mass hierarchy
- ▶ Atmospheric neutrinos
- ▶ Super nova neutrinos
- ▶ Relic super nova neutrinos

$p \rightarrow e\pi^0$ Monte Carlo

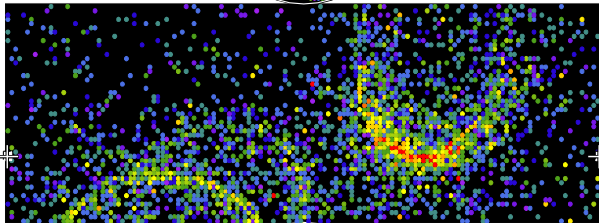
Super-Kamiokande

Run 999999 Event 294
102-11-06:00:06:35
Inner: 3849 hits, 8189 pE
Outer: 4 hits, 2 pE (in-time)
Trigger ID: 0x03
D well: 946.1 cm
PC mass = 909.0 MeV/c²

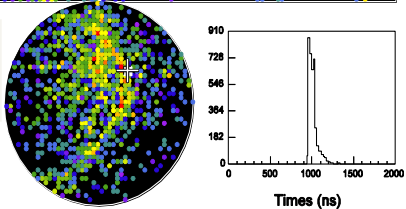


Charge (pe)

• >15.0
• 13.1-15.0
• 11.4-13.1
• 9.8-11.4
• 8.2-9.8
• 6.9-8.2
• 5.6-6.9
• 4.5-5.6
• 3.5-4.5
• 2.6-3.5
• 1.9-2.6
• 1.2-1.9
• 0.8-1.2
• 0.4-0.8
• 0.1-0.4
• <0.1



$\pi^0 \rightarrow \gamma\gamma$



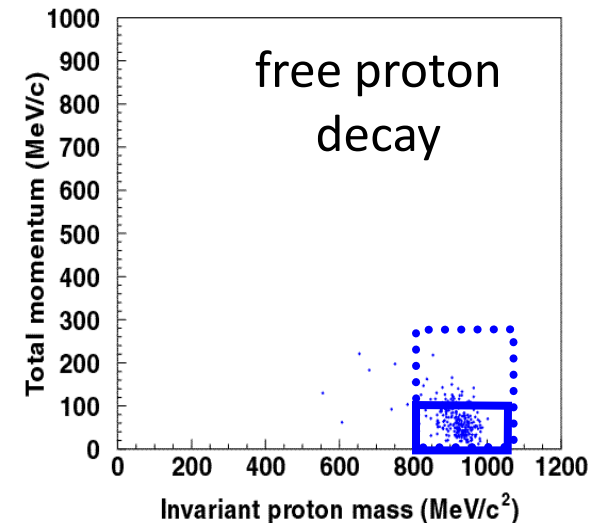
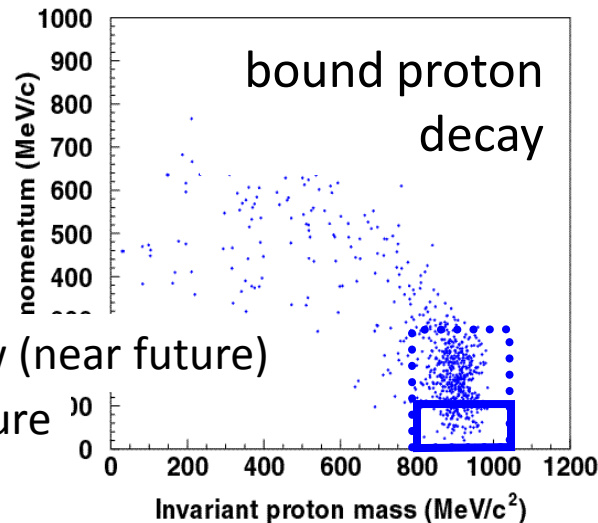
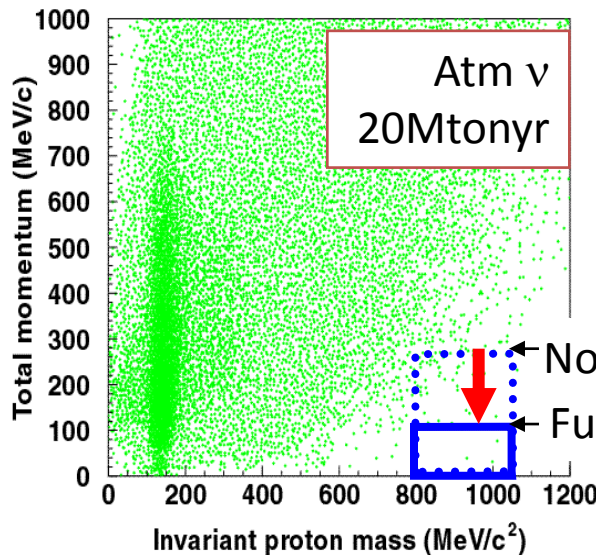
on

Search for $p \rightarrow e^+\pi^0$

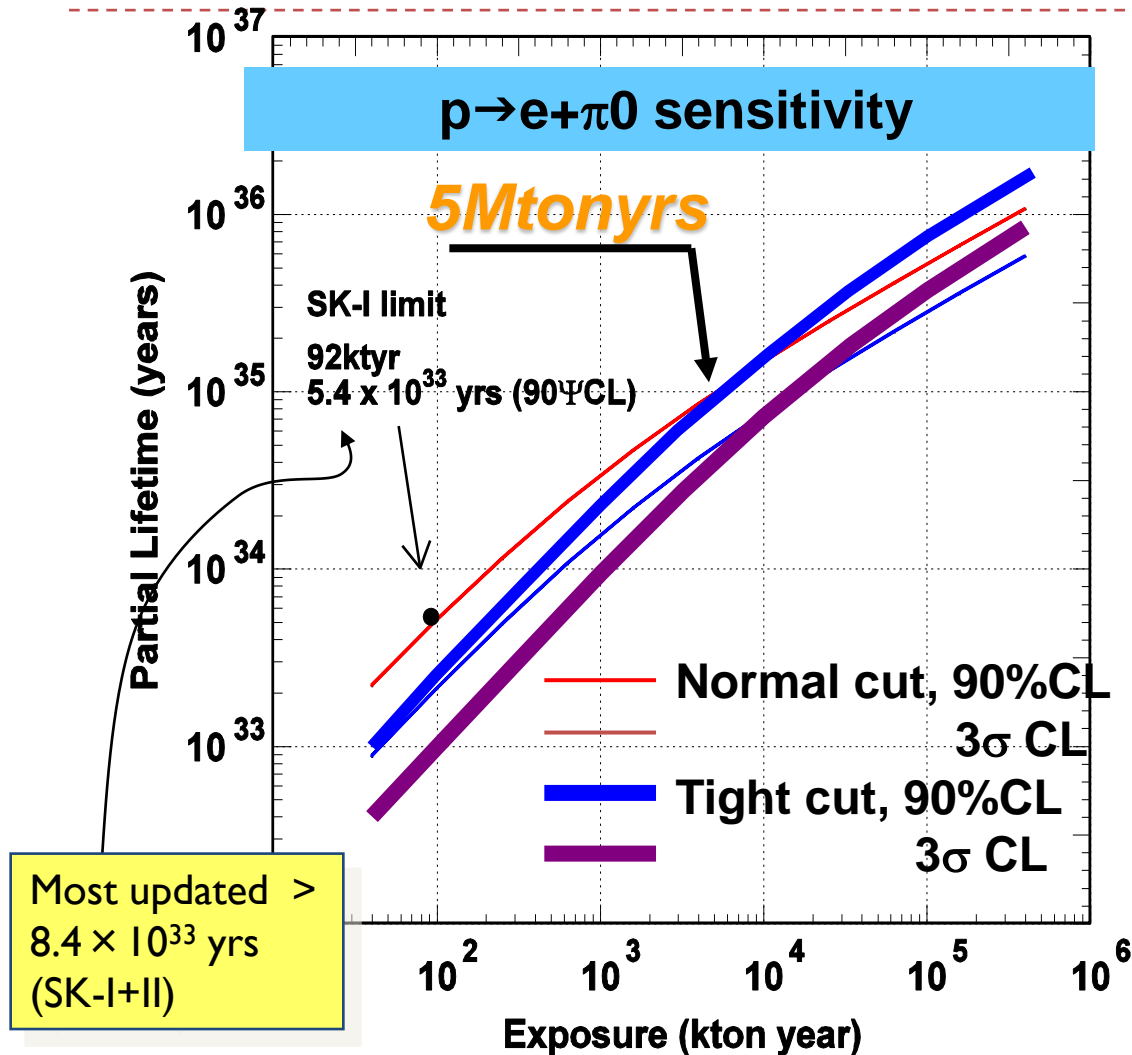
M.Shiozawa NNN05

- $P_{\text{tot}} < 250 \text{ MeV/c}$,
BG 2.2ev/Mtyr, eff=40%
- ↓
- $P_{\text{tot}} < 100 \text{ MeV/c}$,
BG 0.15ev/Mtyr, eff=17%

Main target is free proton decays for the tight cut.

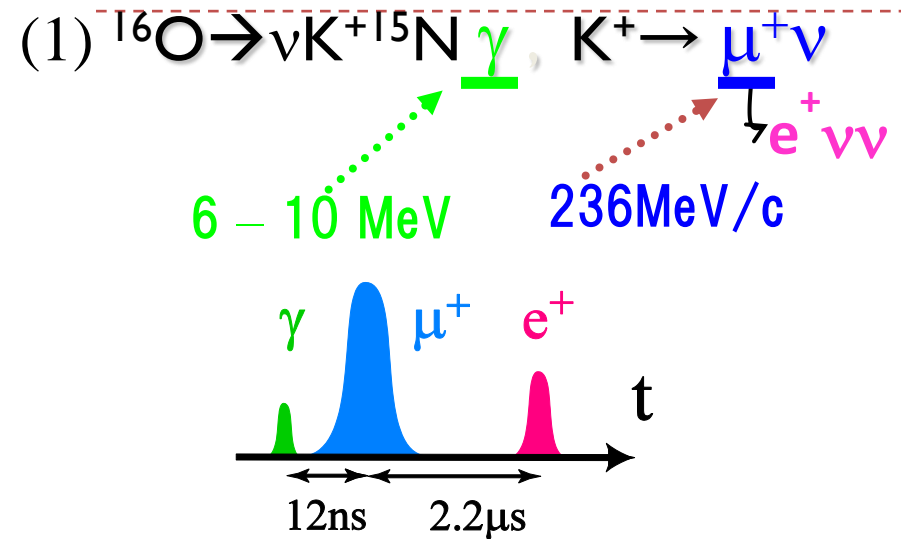


Lifetime sensitivity for $p \rightarrow e^+ \pi^0$



5Mtonyrs \rightarrow
 $\sim 10^{35}$ years@90%CL
 $\sim 4 \times 10^{34}$ years@3 σ CL

$$p \rightarrow \nu K^+$$



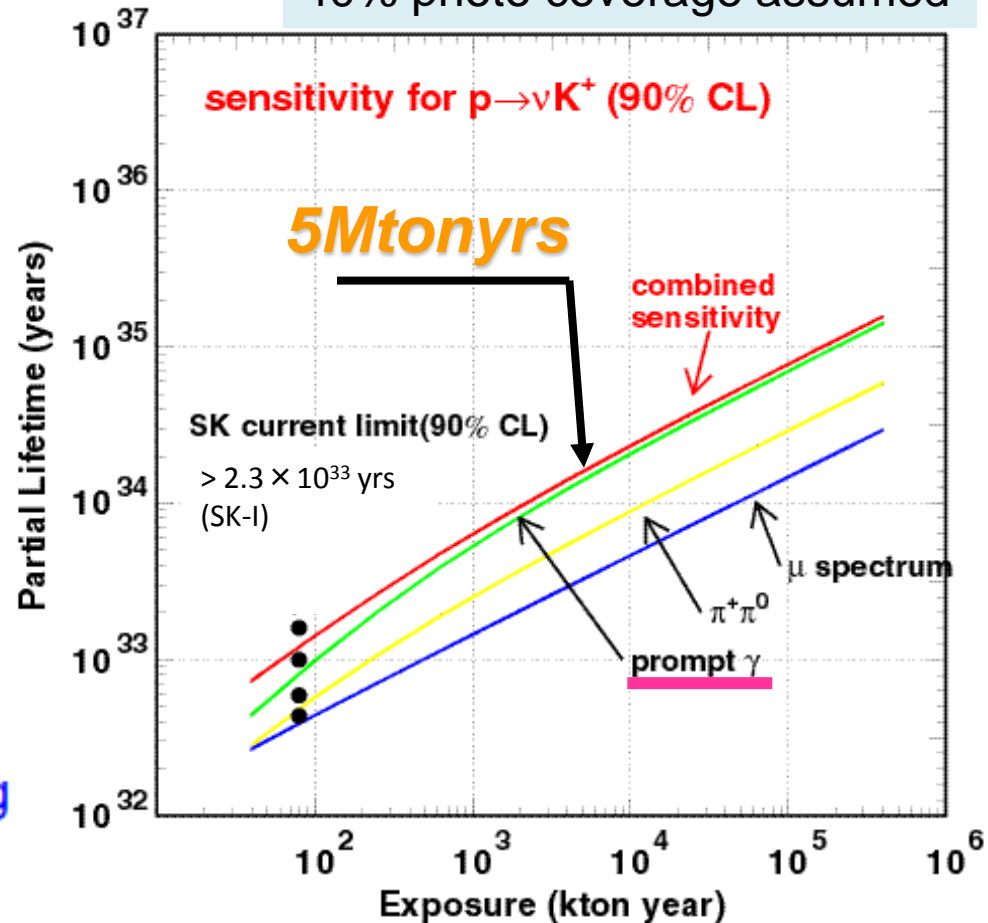
- 1 μ and 1 decay electron
- $215 < P_\mu < 260 \text{ MeV}/c$
- no proton
- maximize NHIT $_\gamma$ in the 12 ns sliding time window, $7 \leq \text{NHIT}_\gamma \leq 60$

(Also, searches for

(2) $p \rightarrow \nu K^+ (K^+ \rightarrow \pi^+ \pi^0)$

(3) $P \rightarrow \nu K^+ (K^+ \rightarrow \mu^+ \nu, \text{ without } \gamma)$

40% photo coverage assumed

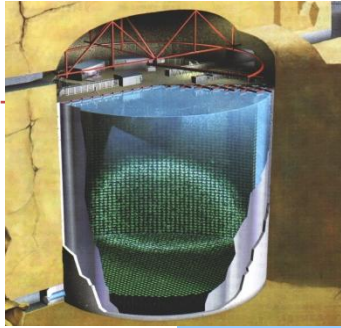


5Mtonyrs \rightarrow

$\sim 2 \times 10^{34} \text{ years @ 90\% CL}$

$\sim 0.8 \times 10^{34} \text{ years @ } 3\sigma \text{ CL}$

T2K



Super-K: 22.5 kt



0.75MW proton beam

Kamioka

$\sim 1\text{GeV}$ ν beam

**J-Parc
(Tokai)**

Super Kamiokande 295km JAERI (Tokai)

Hyper-K: 0.54Mton

4MW proton beam

1st phase

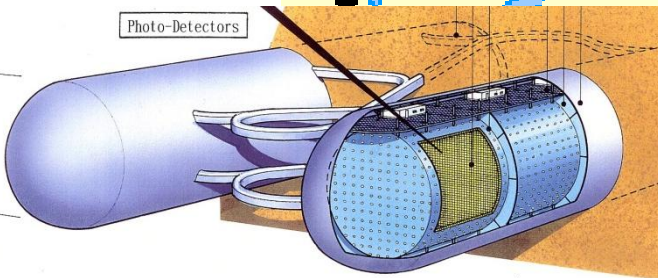
$\nu_\mu \rightarrow \nu_\mu$ precise measurement ($\Delta m_{23}^2, \theta_{23}$)

$\nu_\mu \rightarrow \nu_e$ appearance (θ_{13})

2nd phase (T2K-II)

CP violation

mass hierarchy

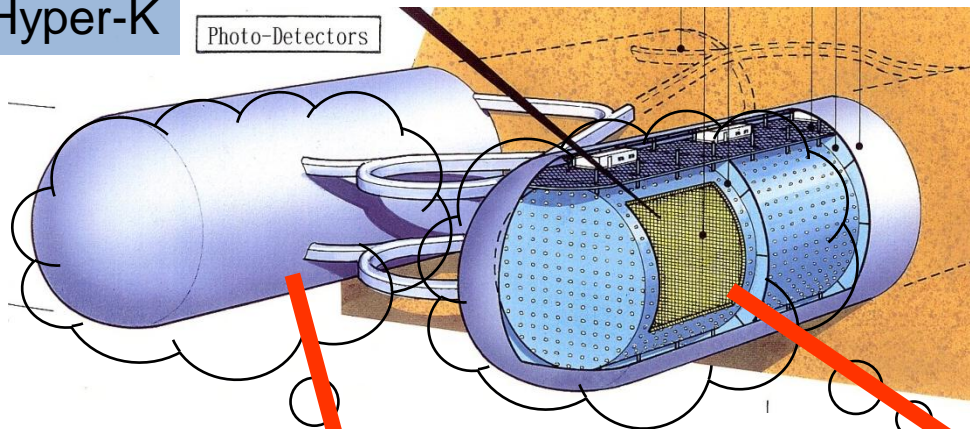


T2KK

Recently T2KK extension is extensively discussed

Hyper-K

Photo-Detectors



two identical detector at different baseline length

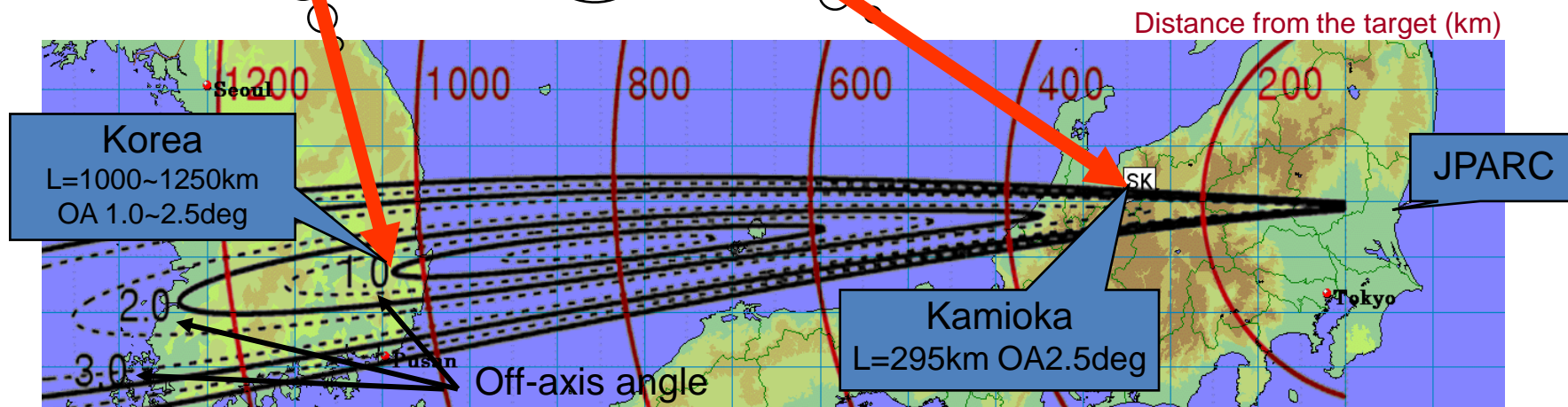
$L = 295\text{km} \rightarrow 1000\sim 1250\text{km}$

CP phase effect enhanced by L

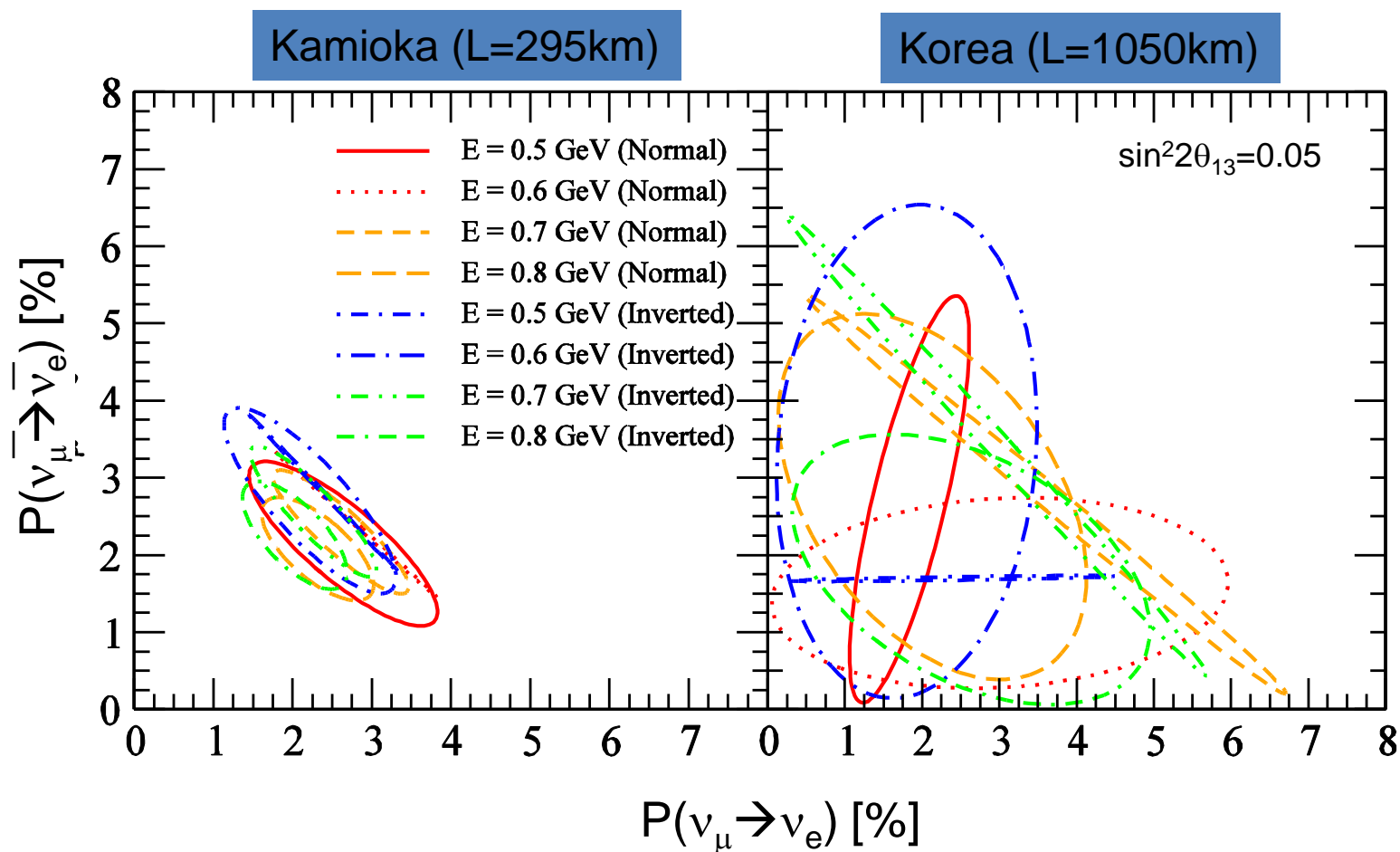
event rate decrease to $\sim 1/10$

same detector \rightarrow

reduce systematic errors



Oscillation Probabilities at Kamioka and Korea

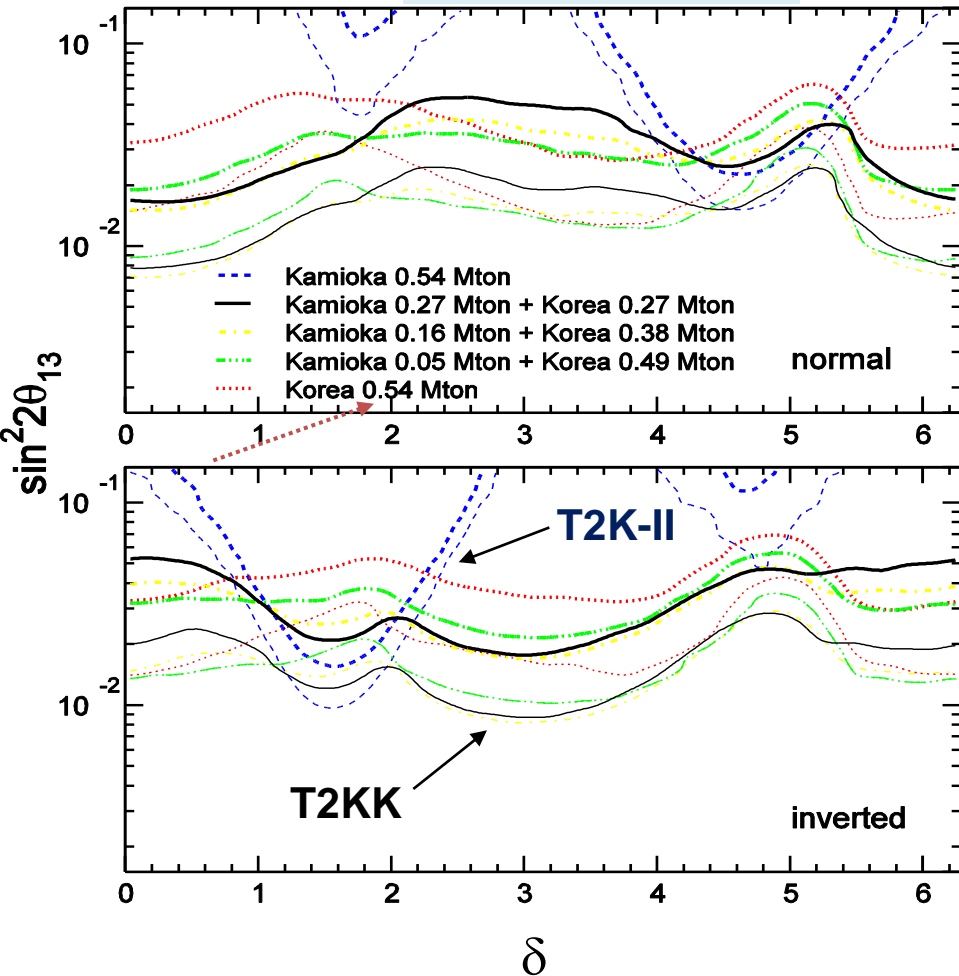


large energy dependence of bi-probabilities
sensitive to mass hierarchy

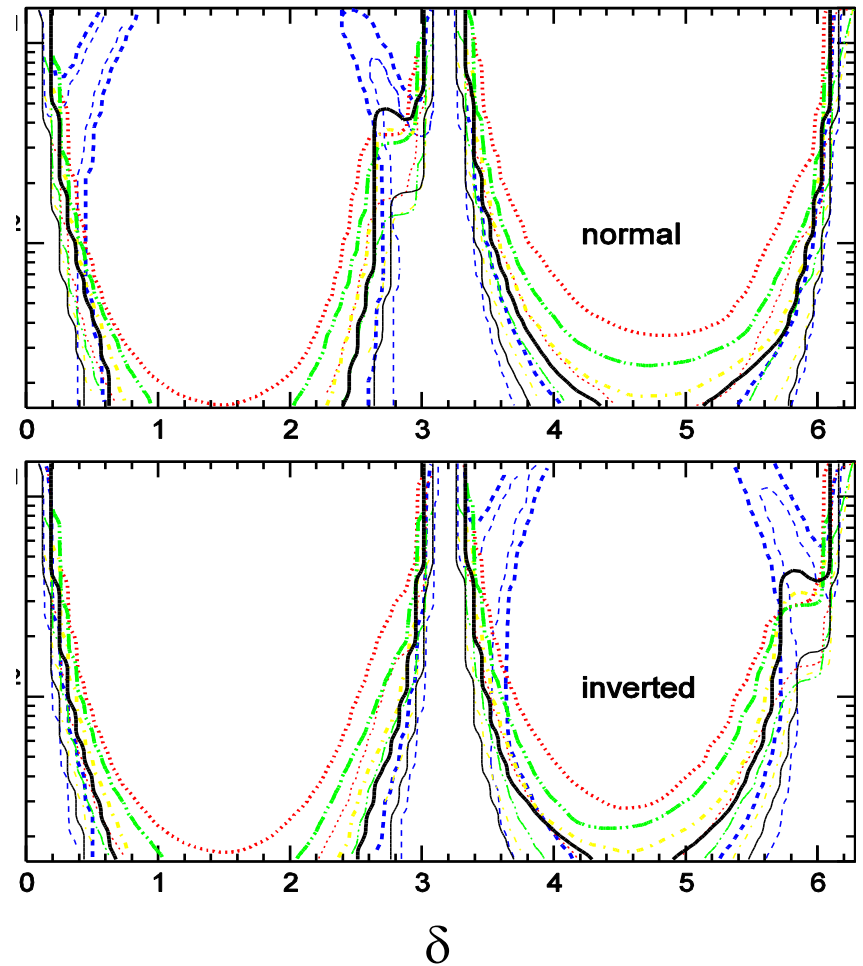
T2K-II and T2KK sensitivities

4yr ν + 4yr $\bar{\nu}$
 4MW J-PARC
 OA2.5° L=1050km

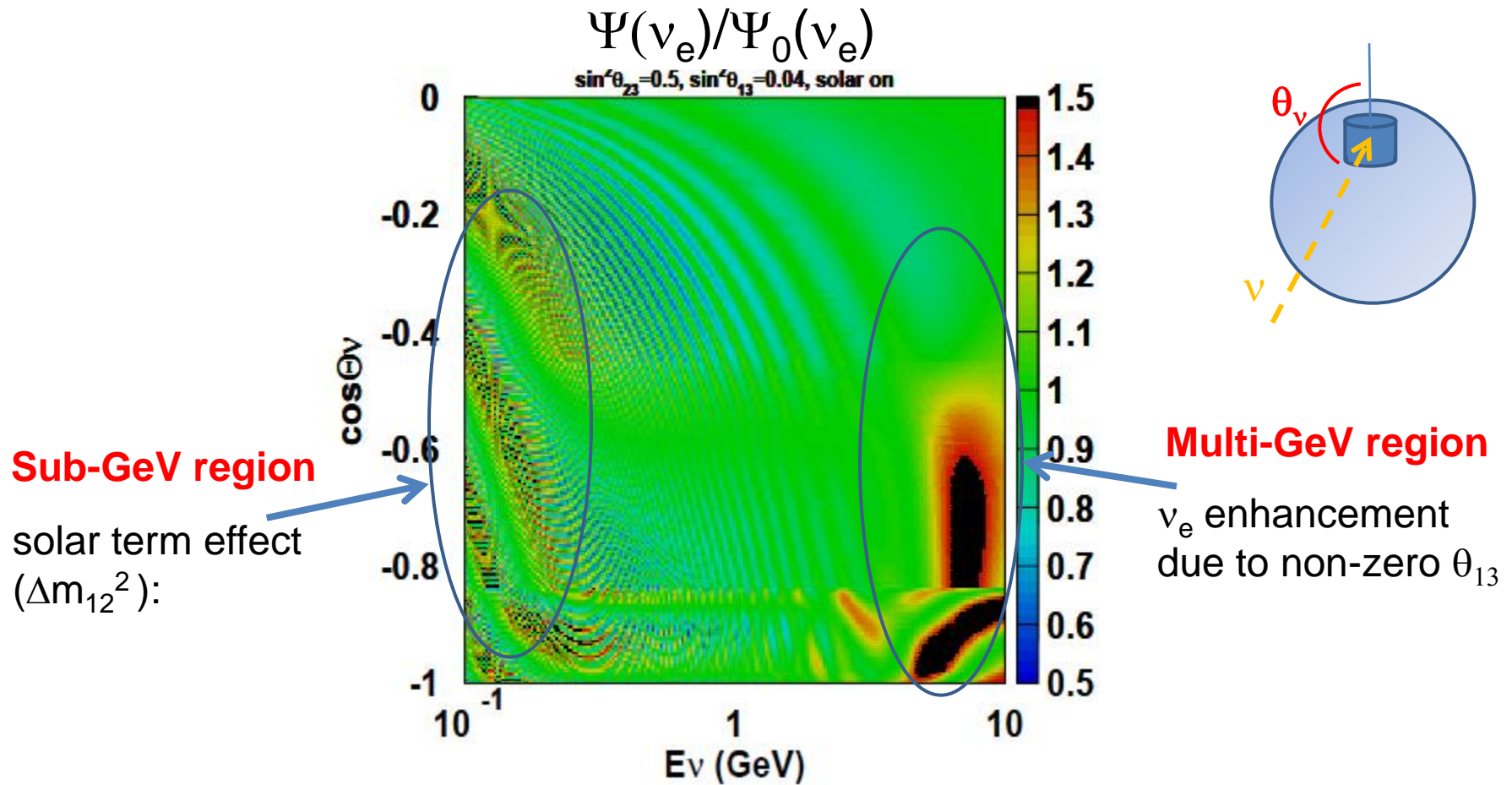
Mass Hierarchy



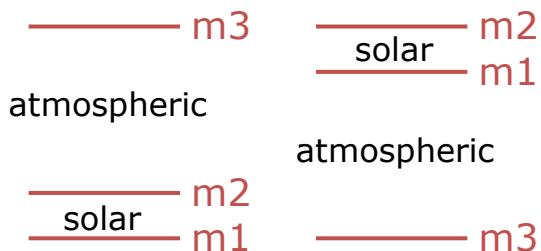
CP Violation



Matter effect of atm. ν in Earth



Probing sign of Δm^2_{23} with atm. ν



$$\Delta m^2_{23} = 0.0025 \text{ eV}^2$$

$$s^2_{23} = 0.5$$

$$s^2_{13} \sim 0.04$$

$$4.5 \text{ Mton} \cdot \text{yr}$$

Normal Hierarchy
($\Delta m^2 > 0$)

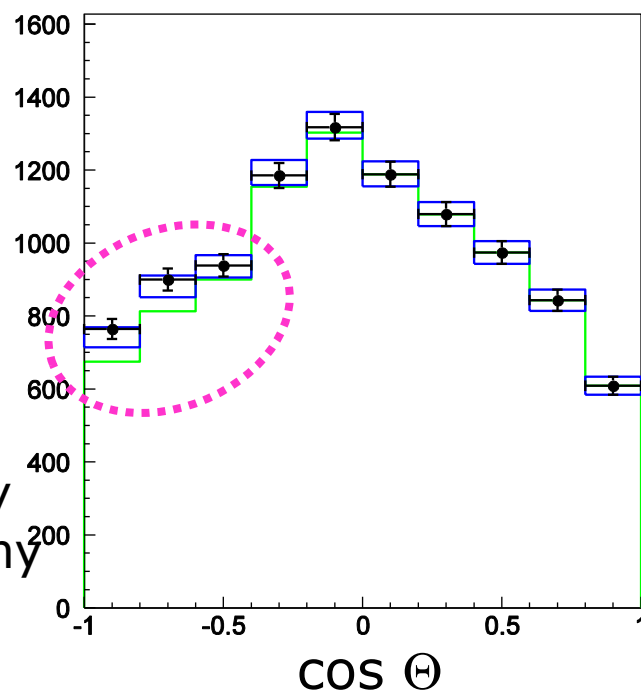
Inverted Hierarchy
($\Delta m^2 < 0$)

resonance only for ν_e

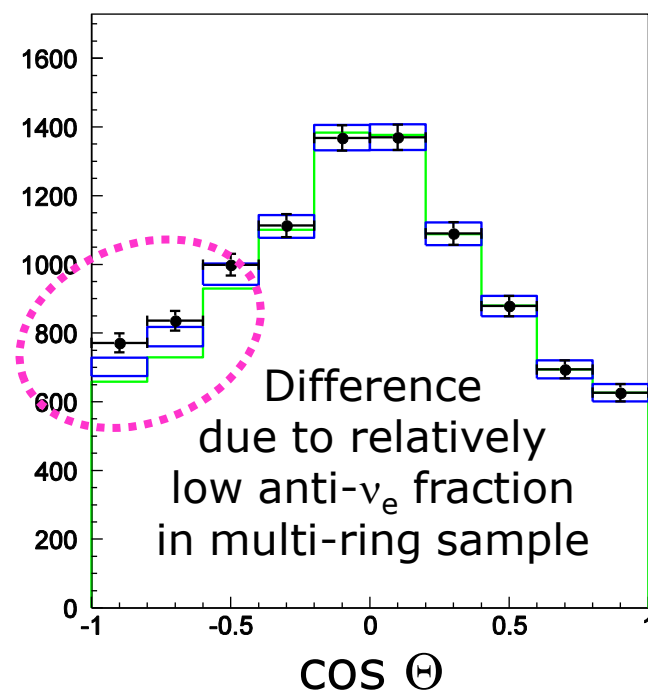
resonance only for $\bar{\nu}_e$

—+— Normal hierarchy
—|— Inverted hierarchy
—|— Null oscillation

single-ring e-like



multi-ring e-like



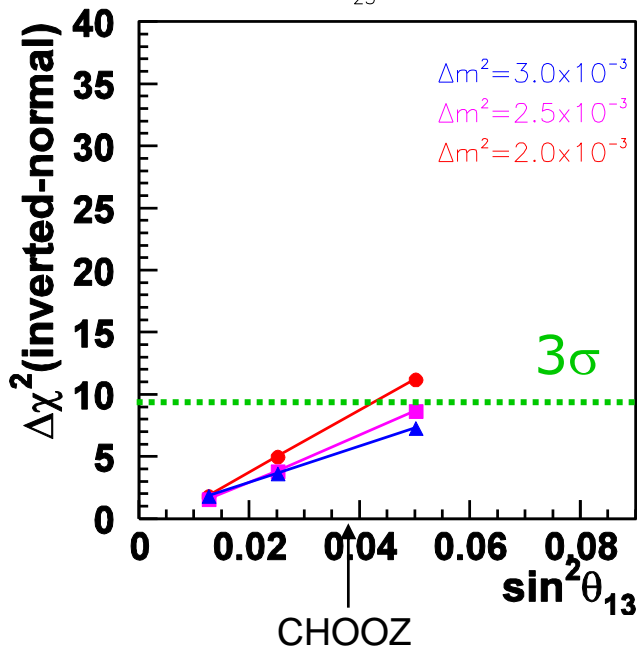
2.5~5 GeV/c

χ^2 difference (inverted – normal)

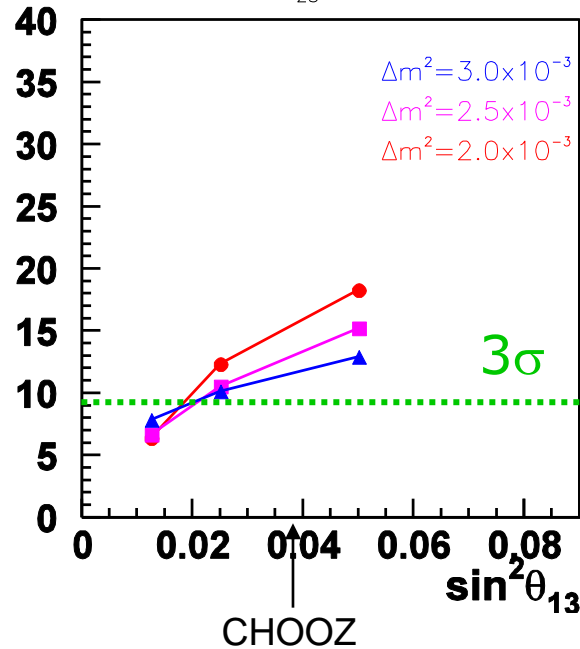
1.8 Mton·yr

True : normal mass hierarchy

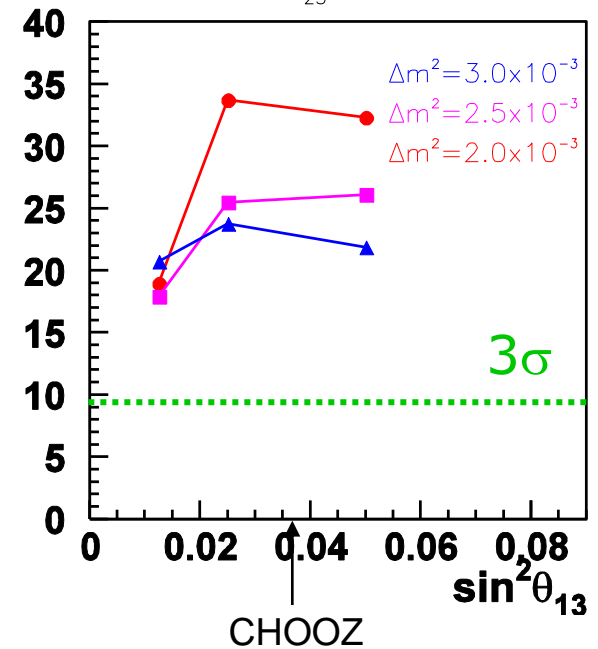
$\sin^2\vartheta_{23}=0.35$



$\sin^2\vartheta_{23}=0.50$



$\sin^2\vartheta_{23}=0.65$



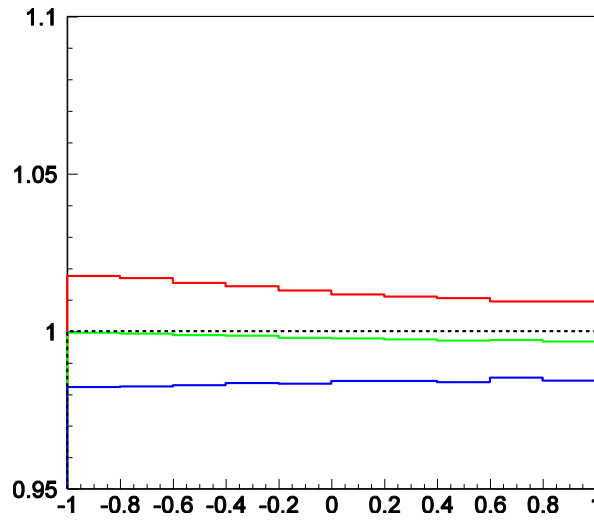
θ_{23} octant

$$\sin^2 2\theta_{23} = 0.96 \longrightarrow \sin^2 \theta_{23} = 0.4 \text{ or } 0.6$$

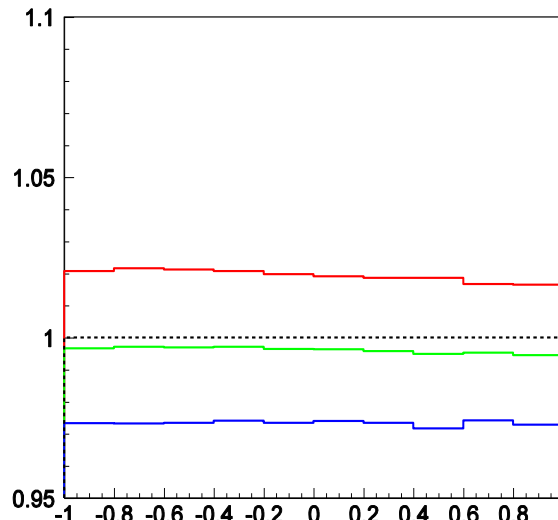
$$\begin{aligned} \Delta m_{12}^2 &= 8.3 \times 10^{-5} \text{ eV}^2 \\ \Delta m_{23}^2 &= 2.5 \times 10^{-3} \text{ eV}^2 \\ \sin^2 2\theta_{12} &= 0.82 \\ \sin^2 \theta_{13} &= 0 \end{aligned}$$

sub-GeV e-like

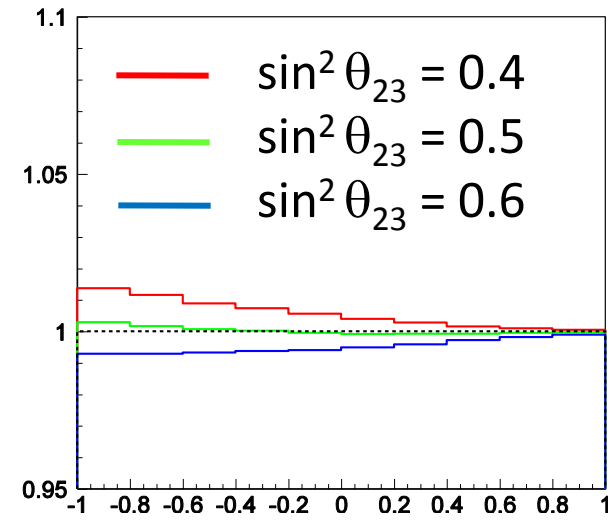
(P_e : 100 ~ 1330 MeV)



(P_e : 100 ~ 400 MeV)



(P_e : 400 ~ 1330 MeV)



$\cos \theta_{\text{zenith}}$

(some opposite effect for μ -like events.)



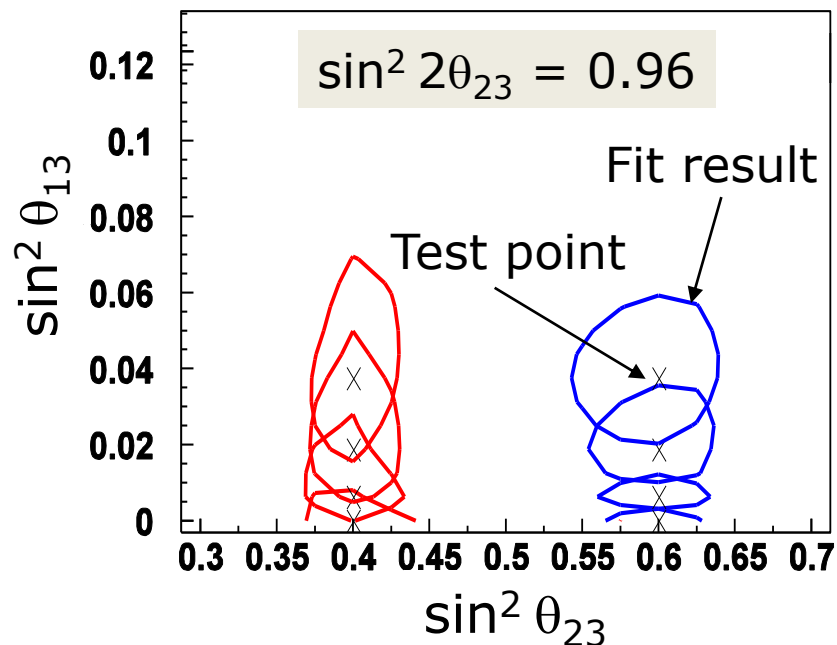
μ/e ratio @low energy is useful to discriminate $\theta_{23} > \pi/4$ and $< \pi/4$.

Discrimination between $\theta_{23} > \pi/4$ and $< \pi/4$ with the (12) and (13) terms

$$\begin{aligned}\sin^2 \theta_{23} &= 0.40 \sim 0.60 \\ \sin^2 \theta_{13} &= 0.00 \sim 0.04 \\ \delta_{\text{cp}} &= 45^\circ\end{aligned}$$

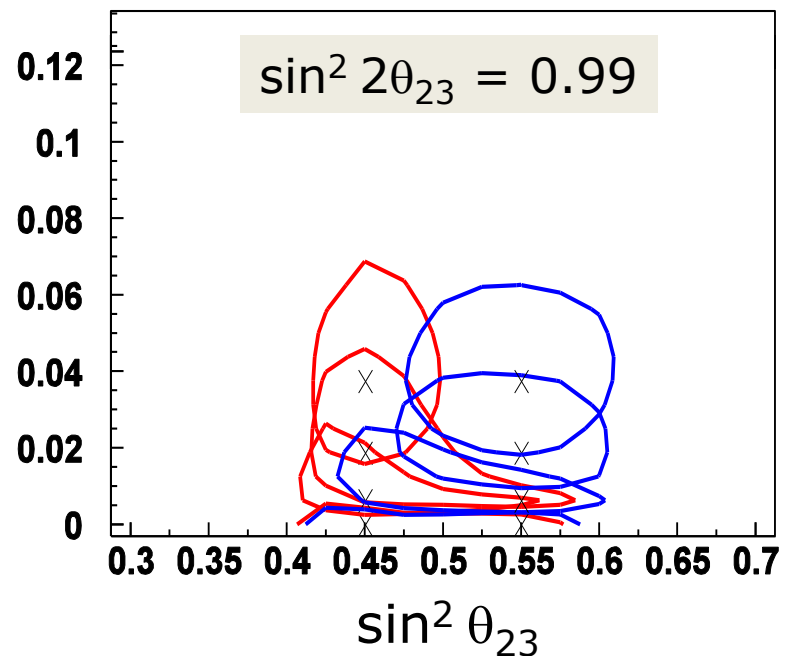
1.8 Mton·yr

90% CL



Discrimination between $\theta_{23} > \pi/4$ and $< \pi/4$ is possible for all θ_{13} .

90% CL



Discrimination between $\theta_{23} > \pi/4$ and $< \pi/4$ is marginally possible only for $\theta_{13} > 0.04$.

Physics summary

▶ Proton decay

- ▶ $p \rightarrow e\pi^0$: $\sim 1 \times 10^{35}$ yr @ 5Mton (90% C.L.)
- ▶ $p \rightarrow \nu K^+$: $\sim 2 \times 10^{34}$ yr @ 5Mton (90% C.L.)

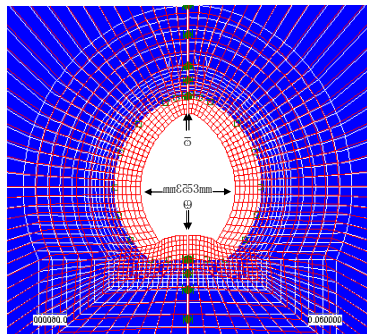
▶ Accelerator oscillation experiments

- ▶ T2K 2nd phase will be able to probe CP violation, and mass hierarchy partially
- ▶ T2KK will extend mass hierarchy sensitivities significantly for any CP phase

▶ Atmospheric neutrino

- ▶ mass hierarchy and θ_{23} octant can be probed due to matter effect in Earth. it will be complementary measurement with long baseline experiment

R&D status of HK



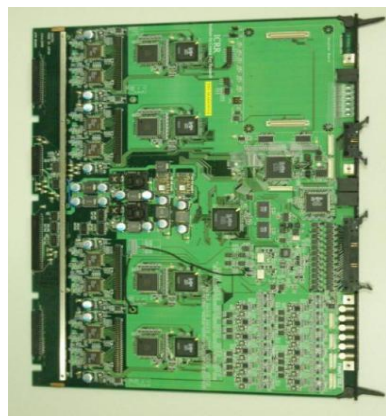
Kamioka Site study and detector structural

→ N. Wakabayashi



Photo sensor (HPD)

→ T. Tanaka



Recently we are developing new electronics for Super-K which are also considering future Mton detector use. I will discuss about the prospect.

Development of new electronics for Super-K

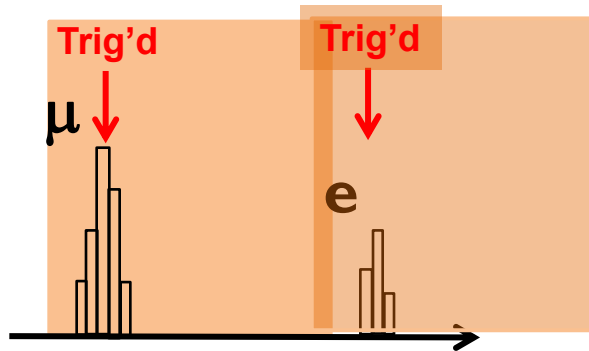
→ S. Nakayama

Requirement for Mton detector DAQ

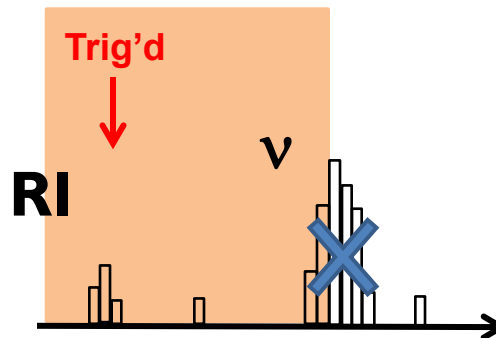
- ▶ Proton decay and neutrinos are rare events
 - ▶ can happen anytime
 - ▶ need to make trigger by event itself
 - ▶ need to minimize dead time
- ▶ Backgrounds
 - ▶ Cosmic ray muon background (10Hz for SK, ~kHz for HK)
 - ▶ Radioactive background events
 - ▶ (may) need to reduce data size
- ▶ Taking successive events is important
 - ▶ $p \rightarrow \nu K^+$, decay electron

Difficulty in self trigger system

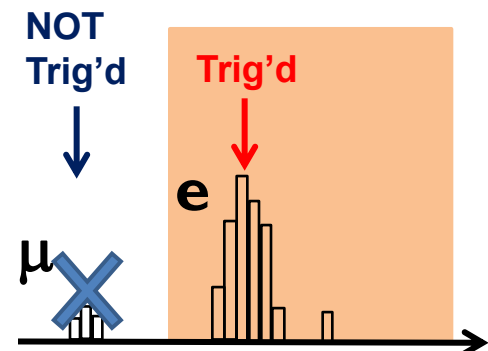
Decay electron



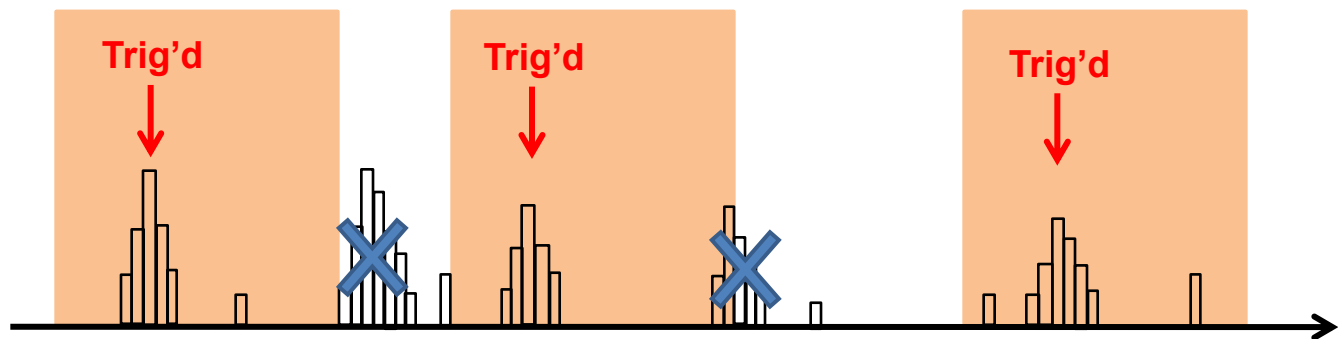
Disturbed ν trigger by background hits



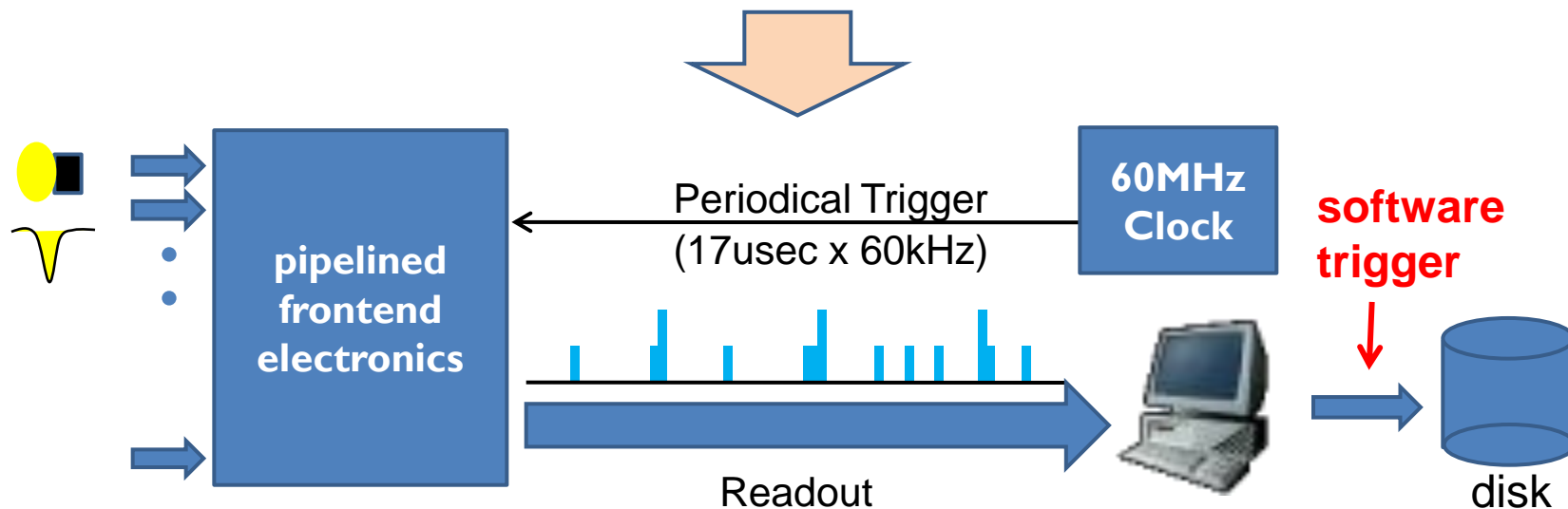
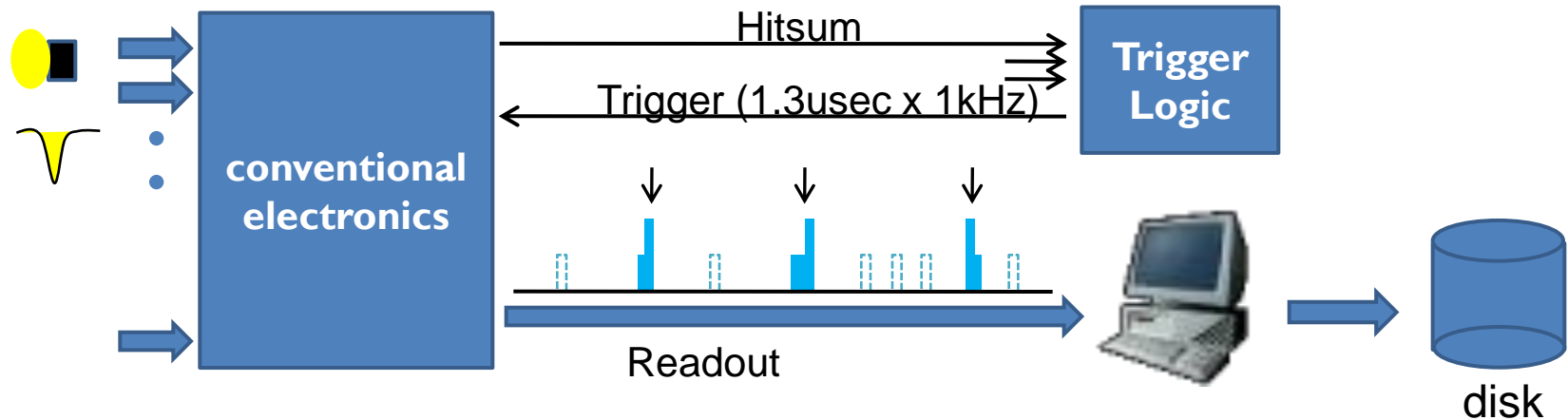
atm $\nu\mu$ BG for relic SN ν



Extreme high rate SN burst !!



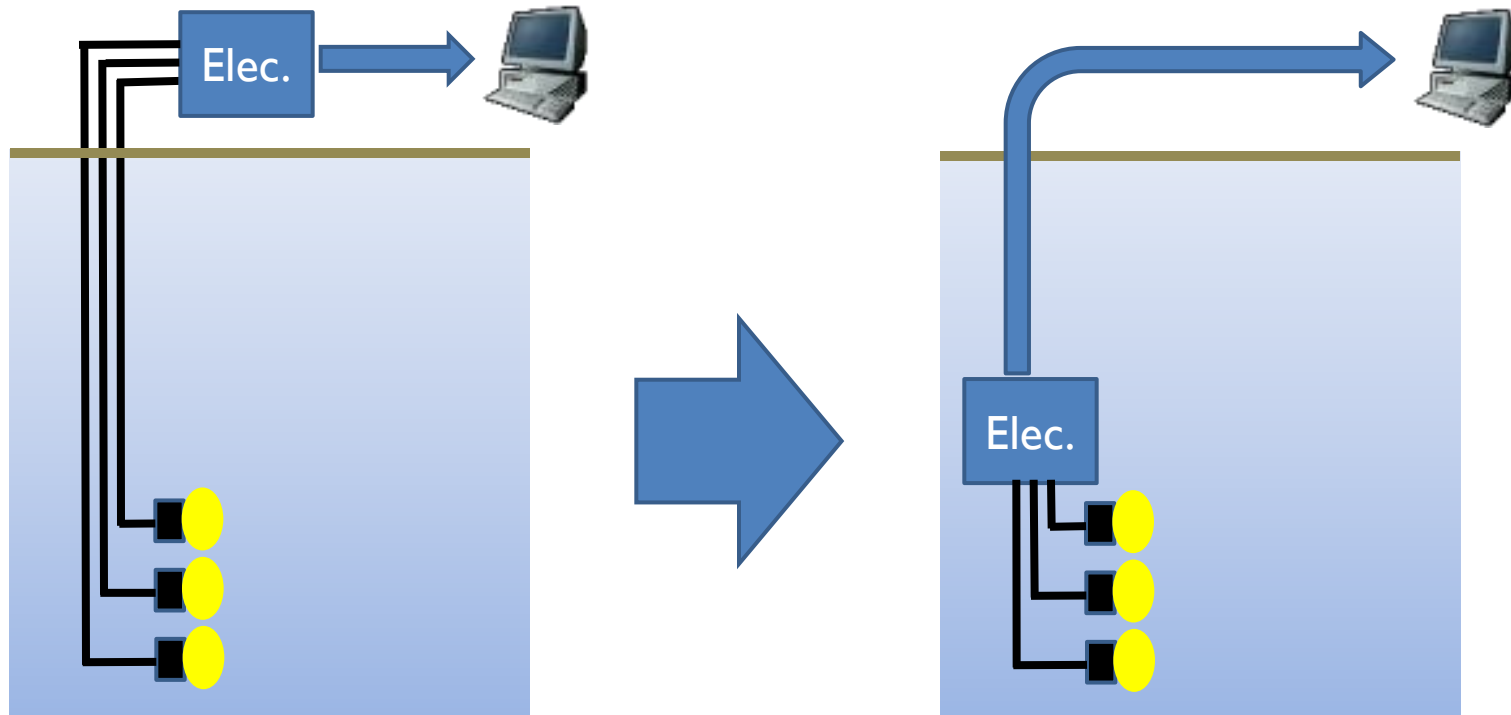
We propose “record-every-hit”



Advantage of this system

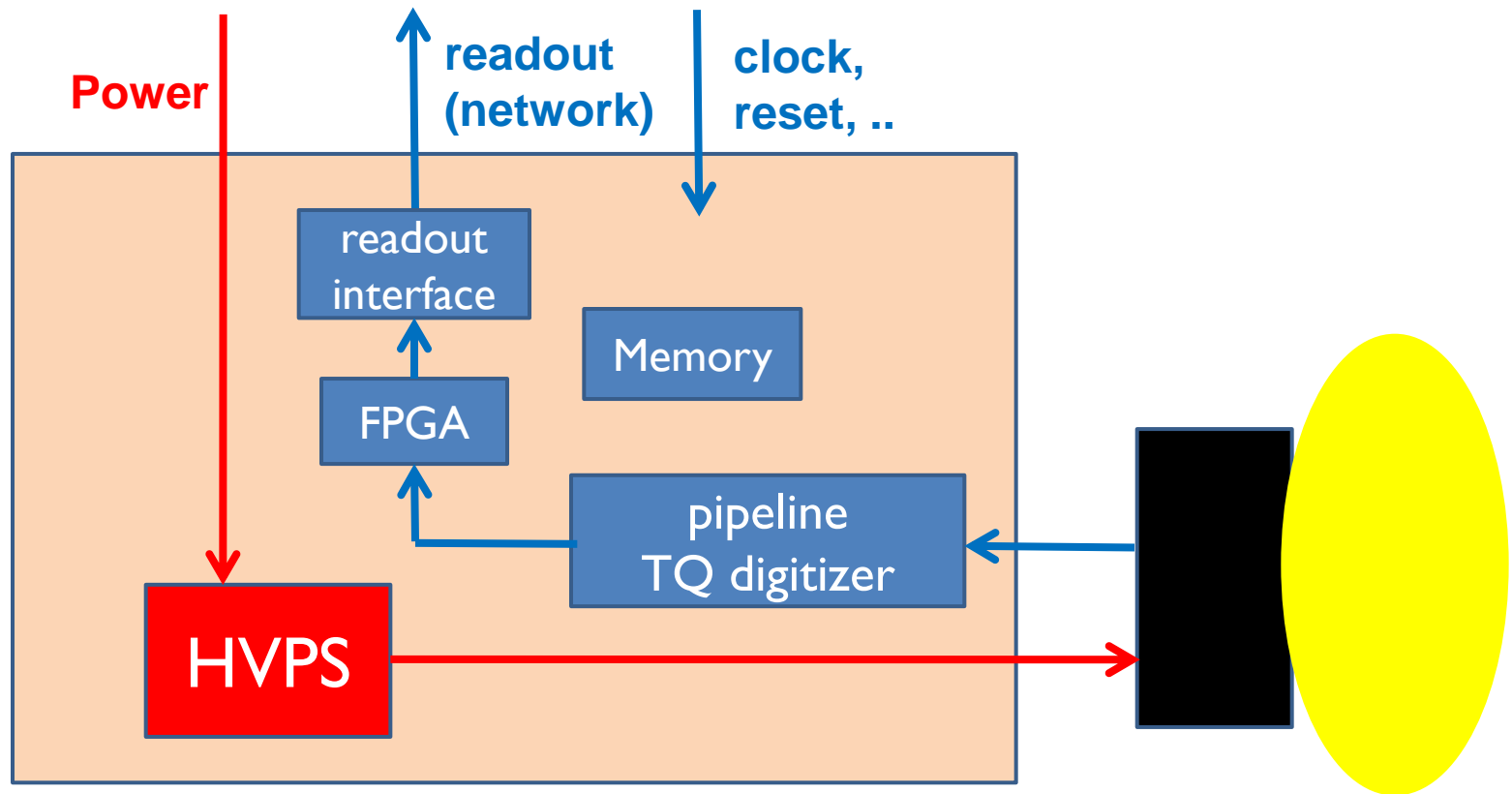
- ▶ hardware trigger is no more necessary
 - ▶ reduce hardware resources for assembling HW trigger
- ▶ dead time is zero !
- ▶ Intelligent and flexible software trigger is available
 - ▶ reduce RI background by online event fitter information
 - ▶ reduce cosmic-ray muons by online
 - ▶ gate window is changeable
 - ▶ wide timing window for proton decay, atm. ν events ($\sim 30\mu\text{sec}$)
 - ▶ narrow timing window for solar neutrino
 - ▶ super nova burst events

Electronics at Photo sensor side



Placement of electronics at photo sensor has the following merits :

- better timing resolution
- reduce hardware resources and cost (cable, connector, etc.)
- easier installation, no cooling inside water (?)



Summary

- ▶ Hyper-K is very attractive detector and it has a potential for probing a lot of physics
 - ▶ Proton decay
 - ▶ T2K-II and T2KK
 - ▶ Atmospheric neutrinos
- ▶ Prospect of electronics system for Hyper-K
 - ▶ Pipelined electronics and `record-every-hit` system will make simple and efficient
 - ▶ Electronics at photo sensor side will make much simpler and improve detector performance